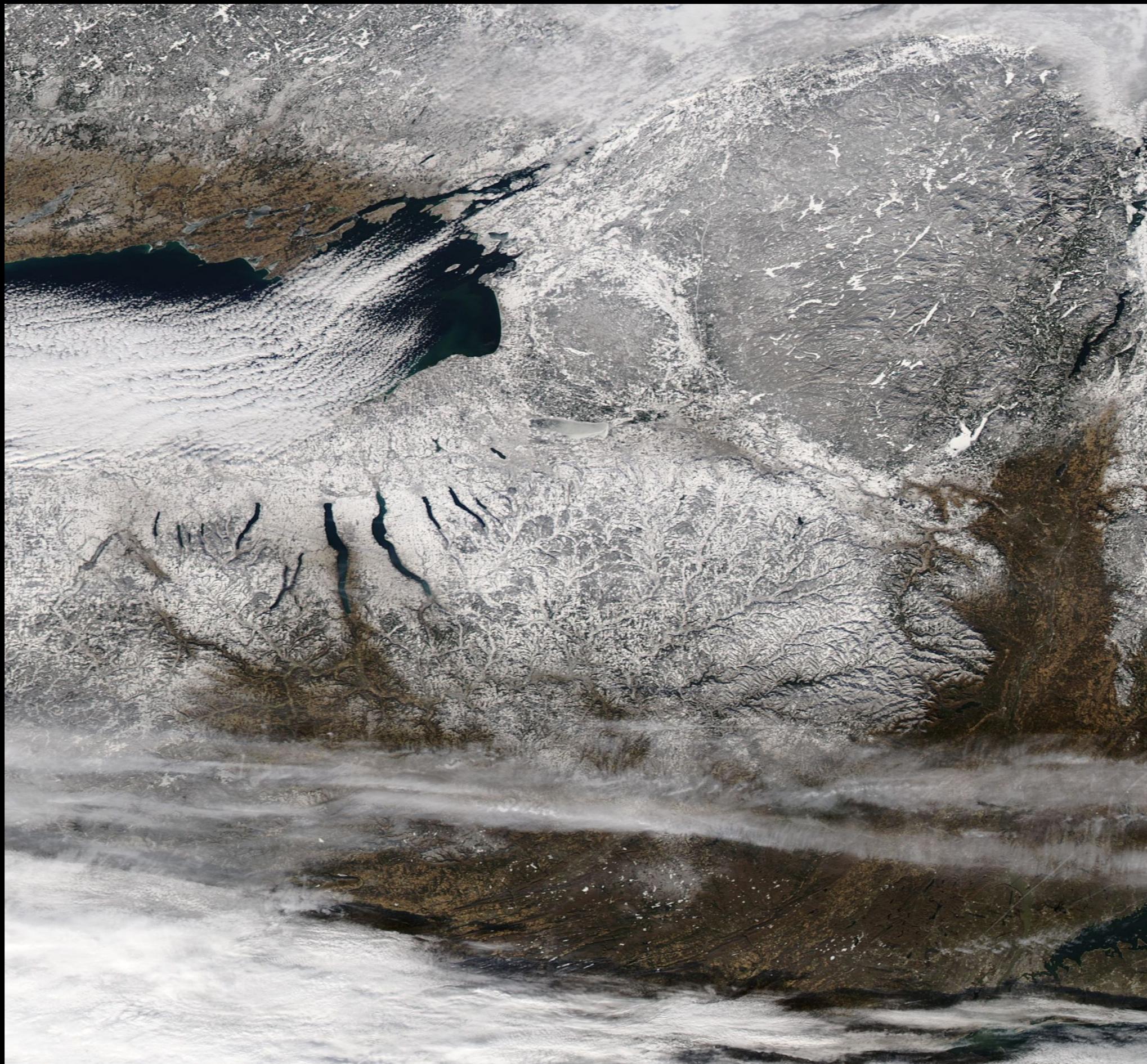


Basic Physical Principles in Remote Sensing

Xi Yang

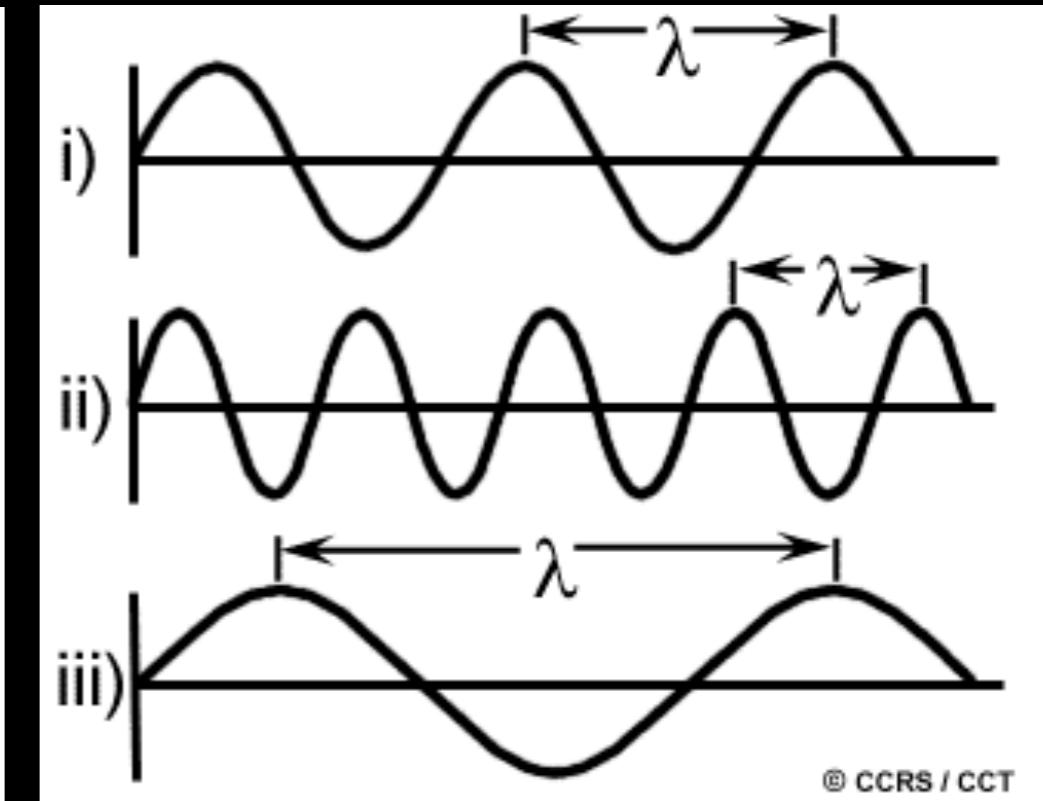
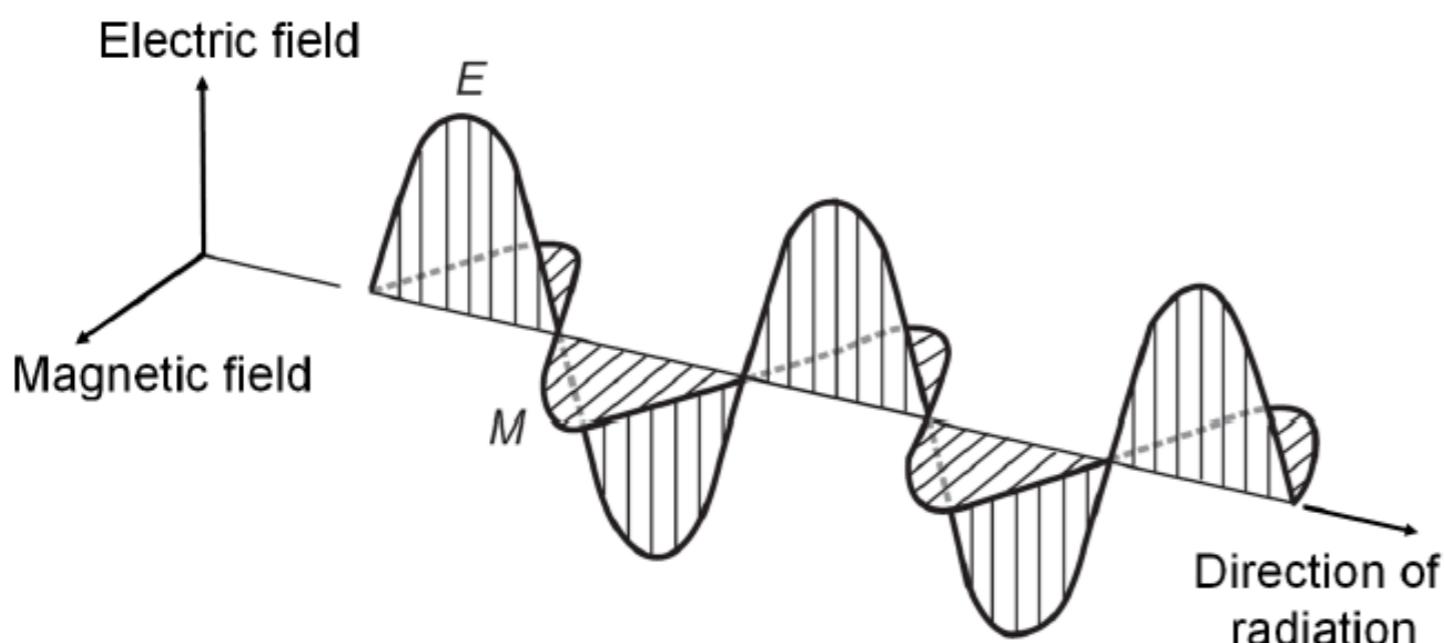
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Outline of Week 2

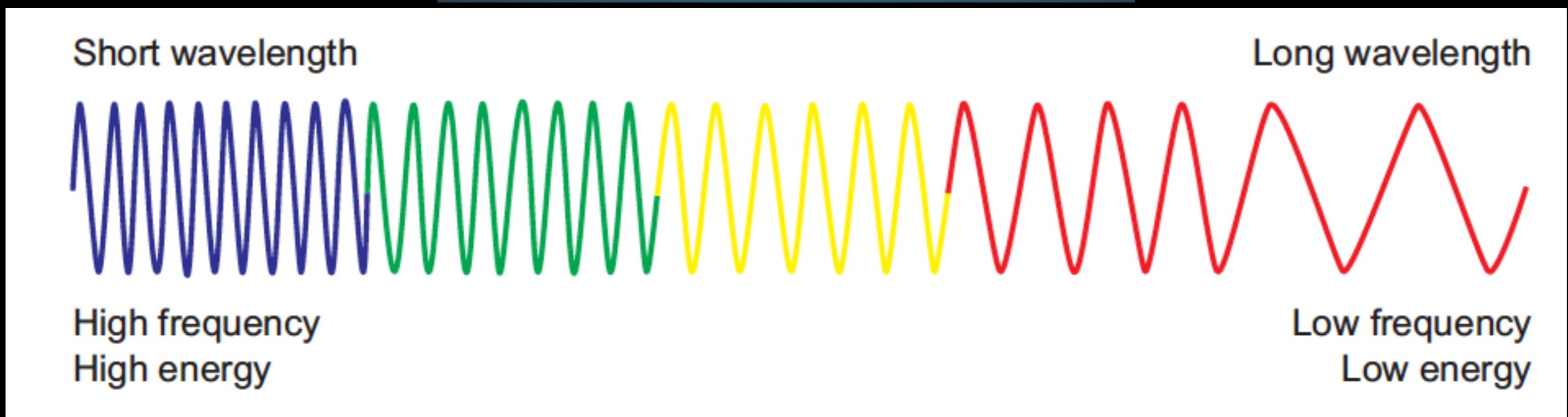
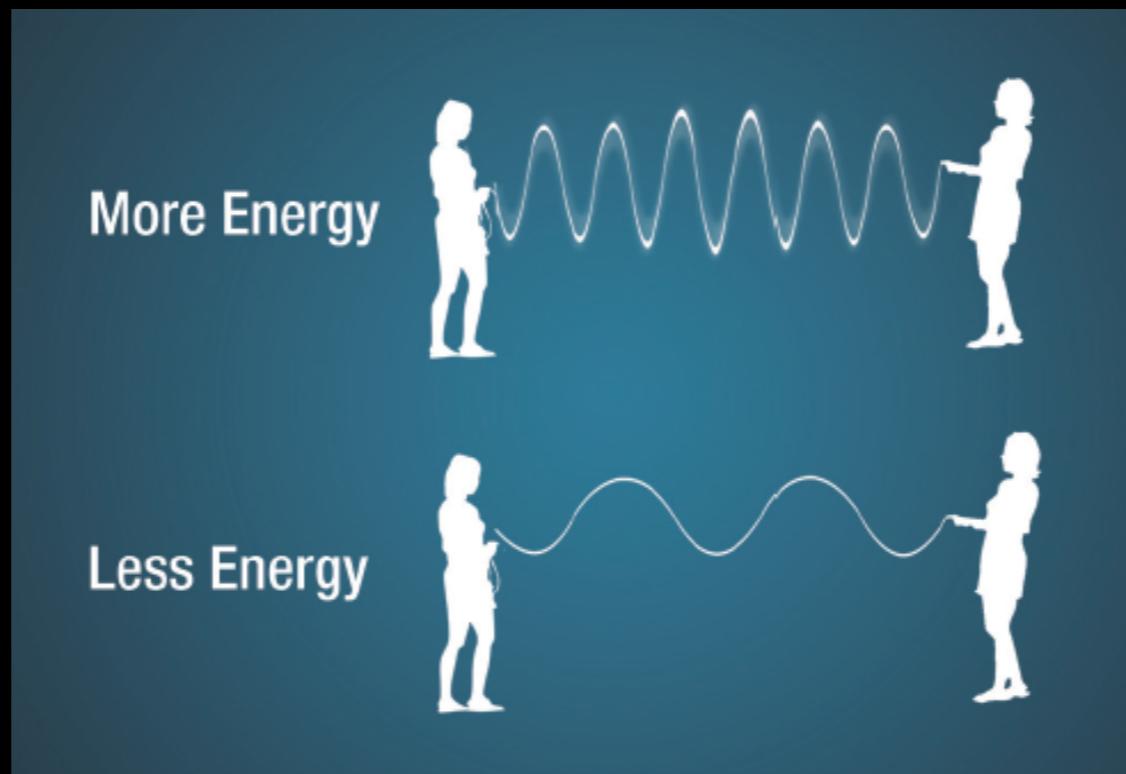
- Electromagnetic radiation (EMR)
- Key concepts in remote sensing
- Laws of EMR
- Solar radiation and surface reflectance

Electromagnetic wave



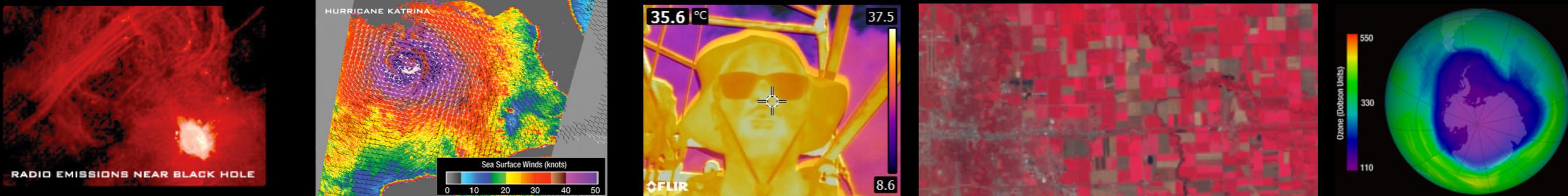
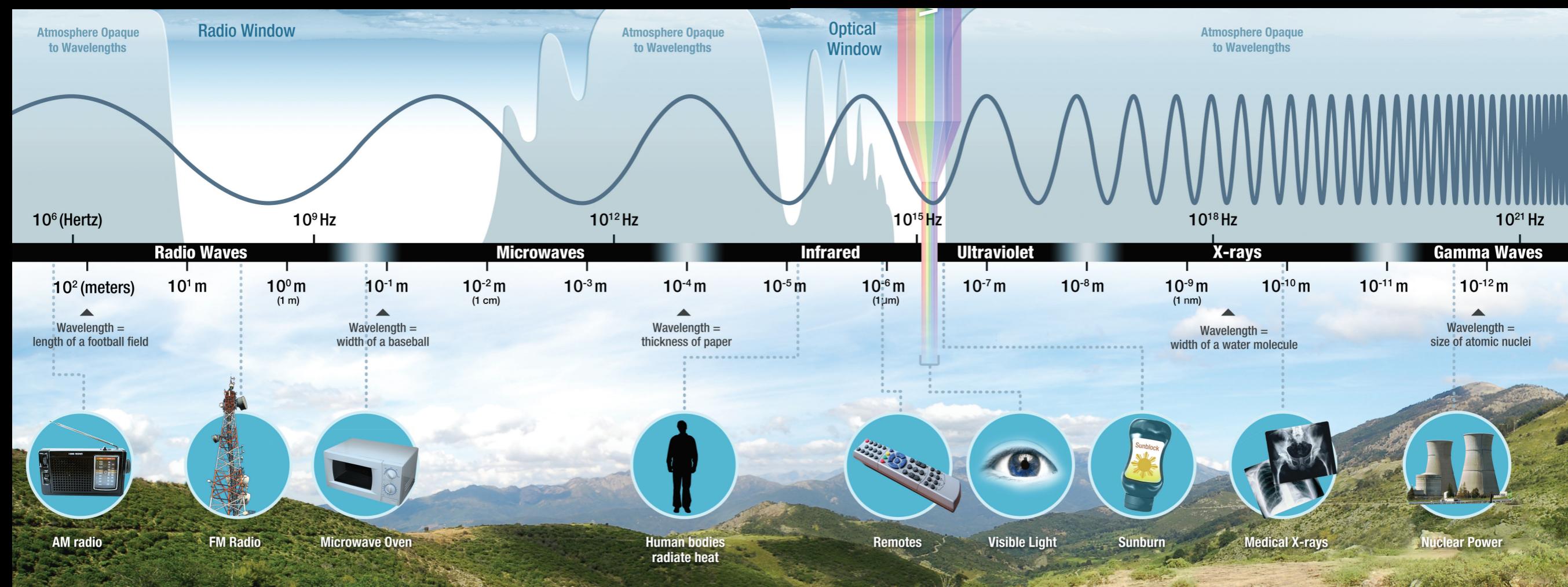
- Wavelength (λ): the distance between successive crests of a wave
- Frequency (v) = c/λ .
- Wavenumber (l/λ), unit: cm^{-1} .

Energy of electromagnetic wave

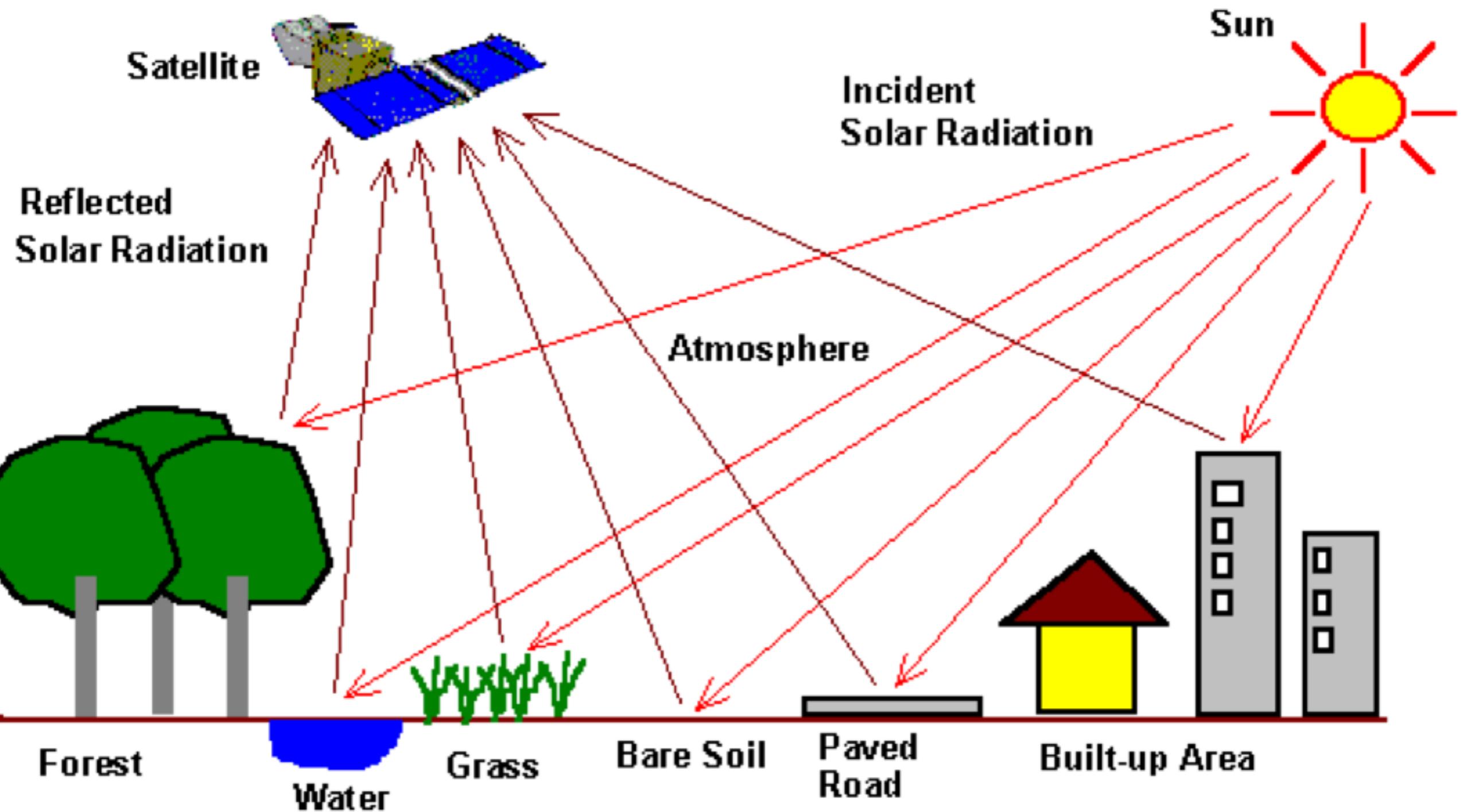


- $Q = hc/\lambda = hv$. Q is in joule, h is Planck's constant ($6.626 \times 10^{-34} \text{ J s}$).

Types of EMR



How does EM radiation reach the satellite?

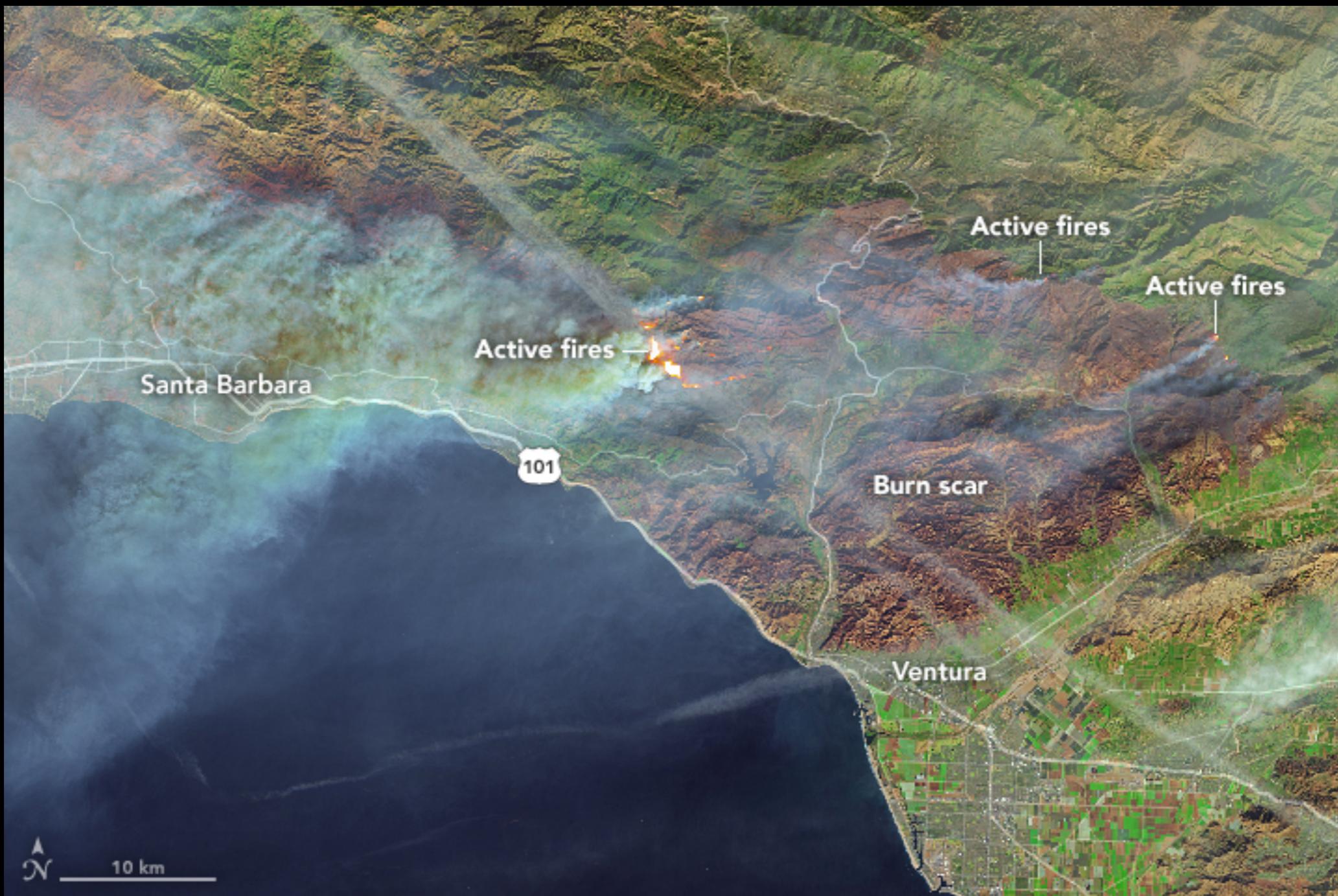


Near infrared and optical remote sensing



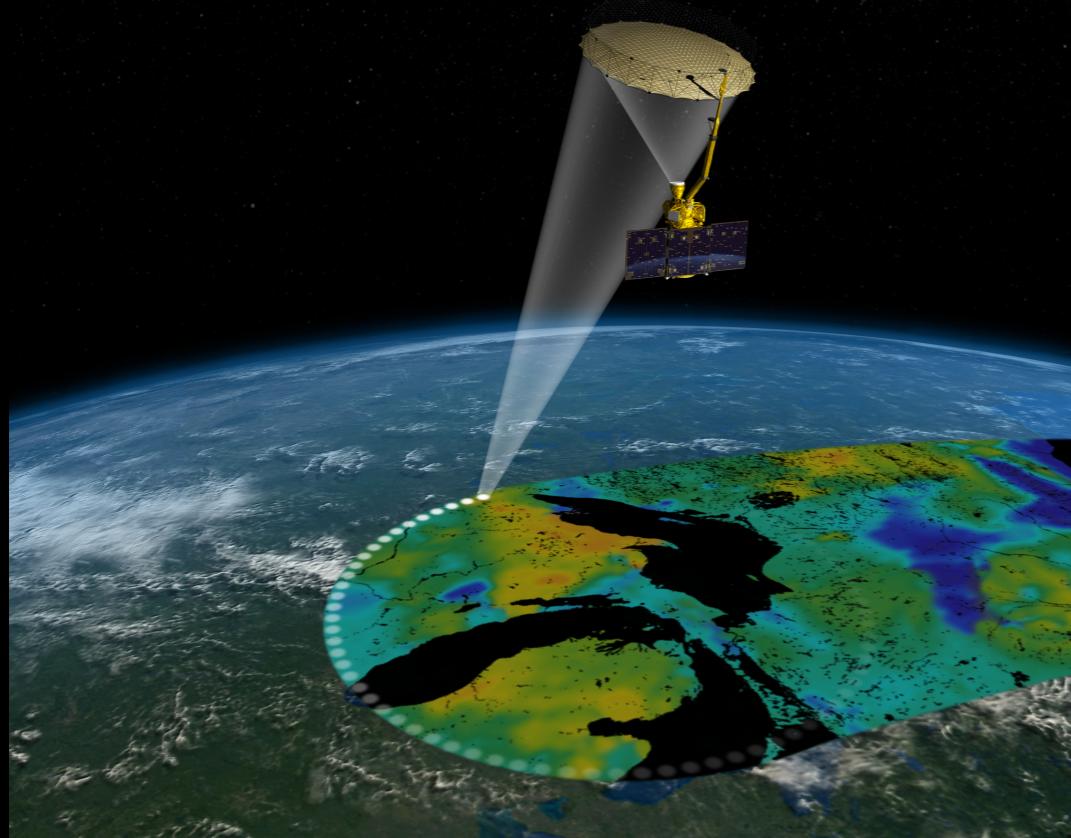
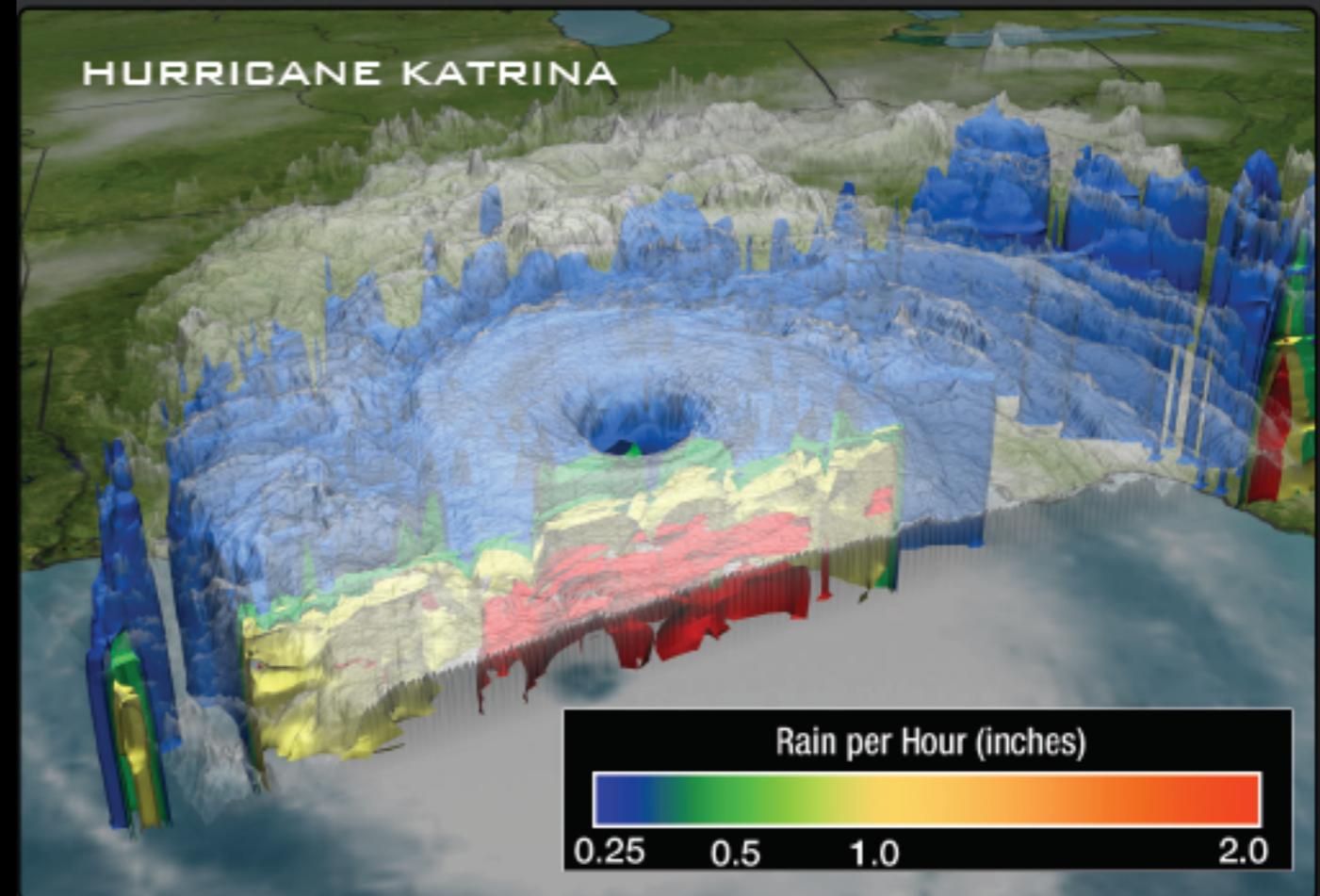
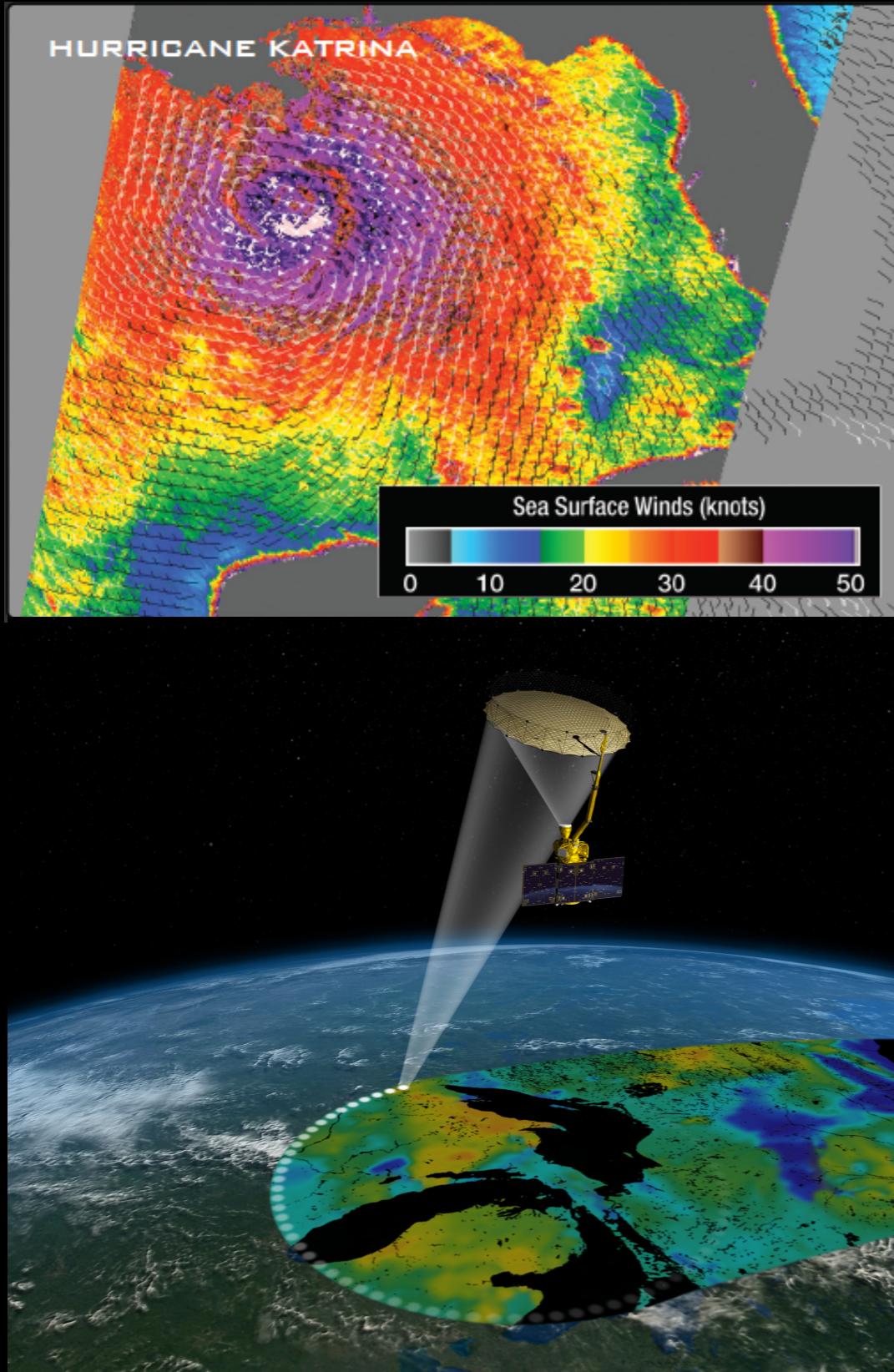
Vegetation; Urbanization

Thermal remote sensing



It is based on observations of visible and shortwave infrared light from OLI and of surface heat from Landsat's Thermal Infrared Sensor (TIRS). Active fires appear orange; the burn scar is brown. Unburned vegetation is green; developed areas are gray.

Microwave remote sensing



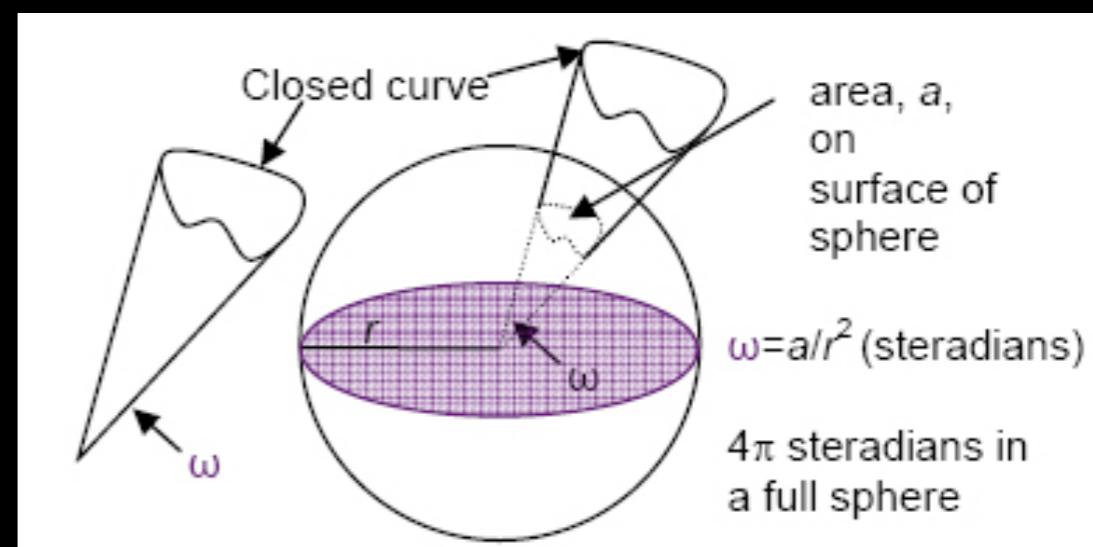
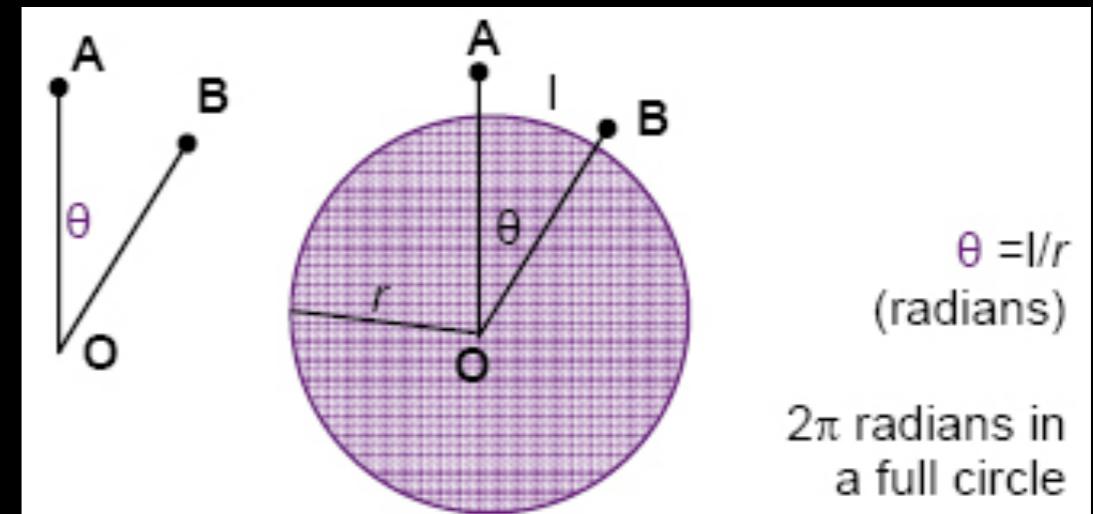
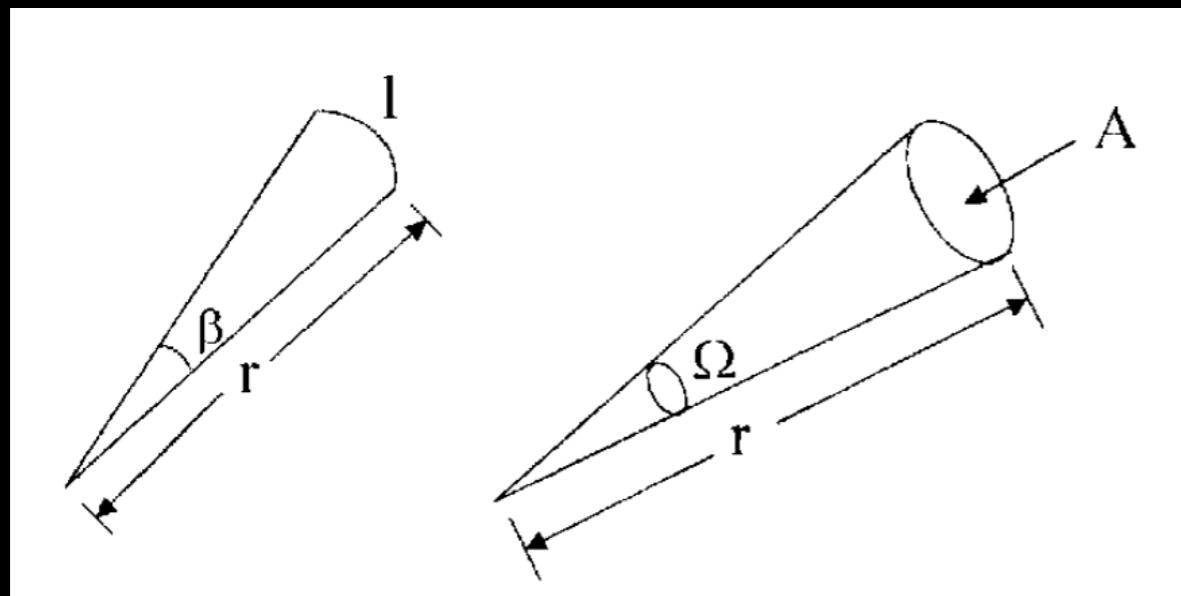
GPS; Sea surface winds;
rainfall; soil moisture.

Basic concepts

- Radiant Energy (Q , unit: joule): the total energy radiated toward a surface.
- Radiant Flux ($\Phi = \delta Q / \delta t$, unit: joule/s=watt): the energy per unit of time.
 - Photon flux: no. of photons per unit time.
- Irradiance ($E = \delta \Phi / \delta A$, unit: w/m²): the radiant flux arriving on a surface per unit area [flashlight vs. the Sun]
- Exitance or Emittance (M , w/m²): the radiant flux *leaving* a surface

Solid angle

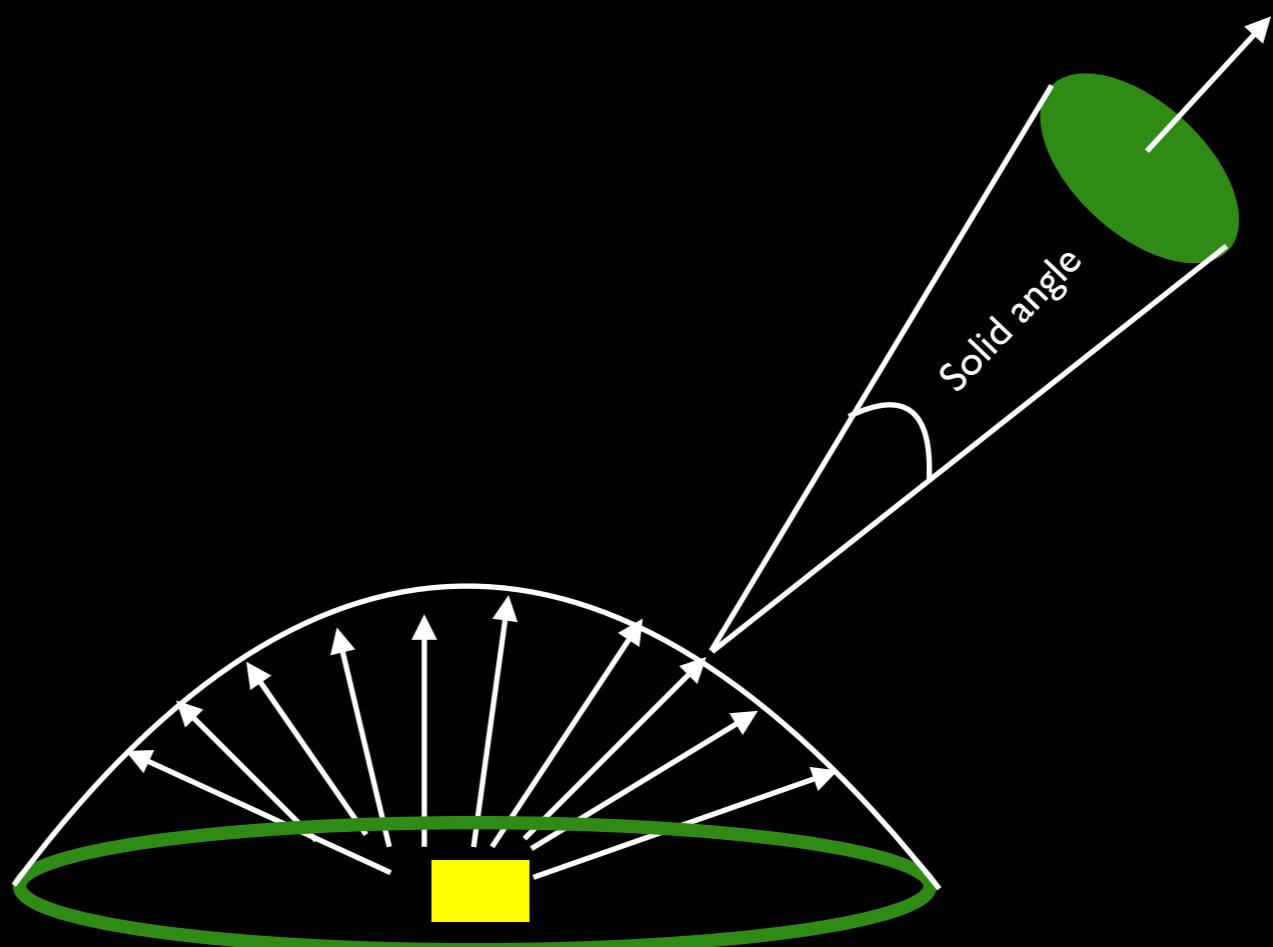
- Solid angle (Ω , unit: steradian): a solid angle is, analogous to an angle in the 2D world, but in a 3D world.



What is the solid angle of a sphere?

Radiance

- Radiance (L , unit: $\text{W m}^{-2} \text{ sr}^{-1}$): the total energy exiting in a certain direction per unit area per solid angle of measurement.



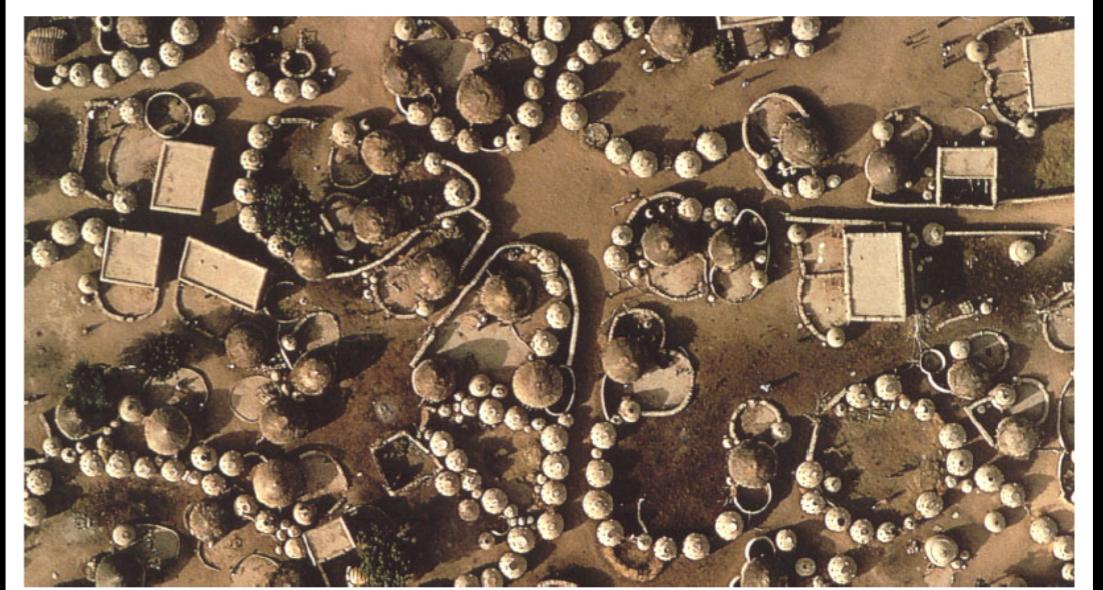
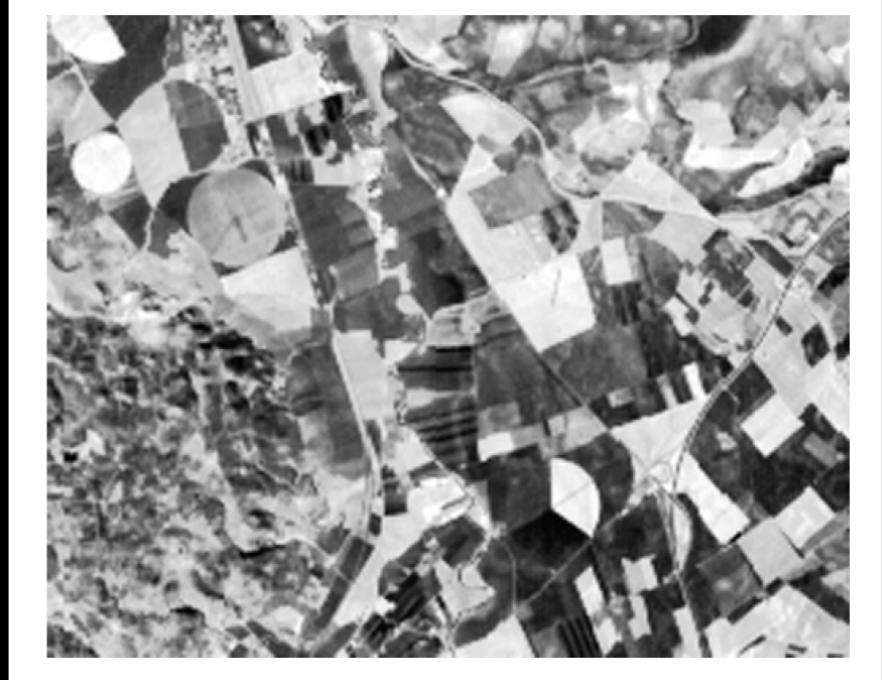
Spectral irradiance/radiance

- Spectral Irradiance ($E = \delta\Phi/\delta A/\delta\lambda$, unit: $\text{W/m}^2/\text{nm}$): the radiant flux arriving on a surface per unit area per unit wavelength.
- Spectral Radiance (L , unit: $\text{W m}^{-2} \text{ sr}^{-1} \text{ nm}^{-1}$): the total energy exiting in a certain direction per unit area per solid angle of measurement per unit wavelength.

$$1 \text{ m} = 10^3 \text{ mm} = 10^6 \mu\text{m (micron)} = 10^9 \text{ nm}$$

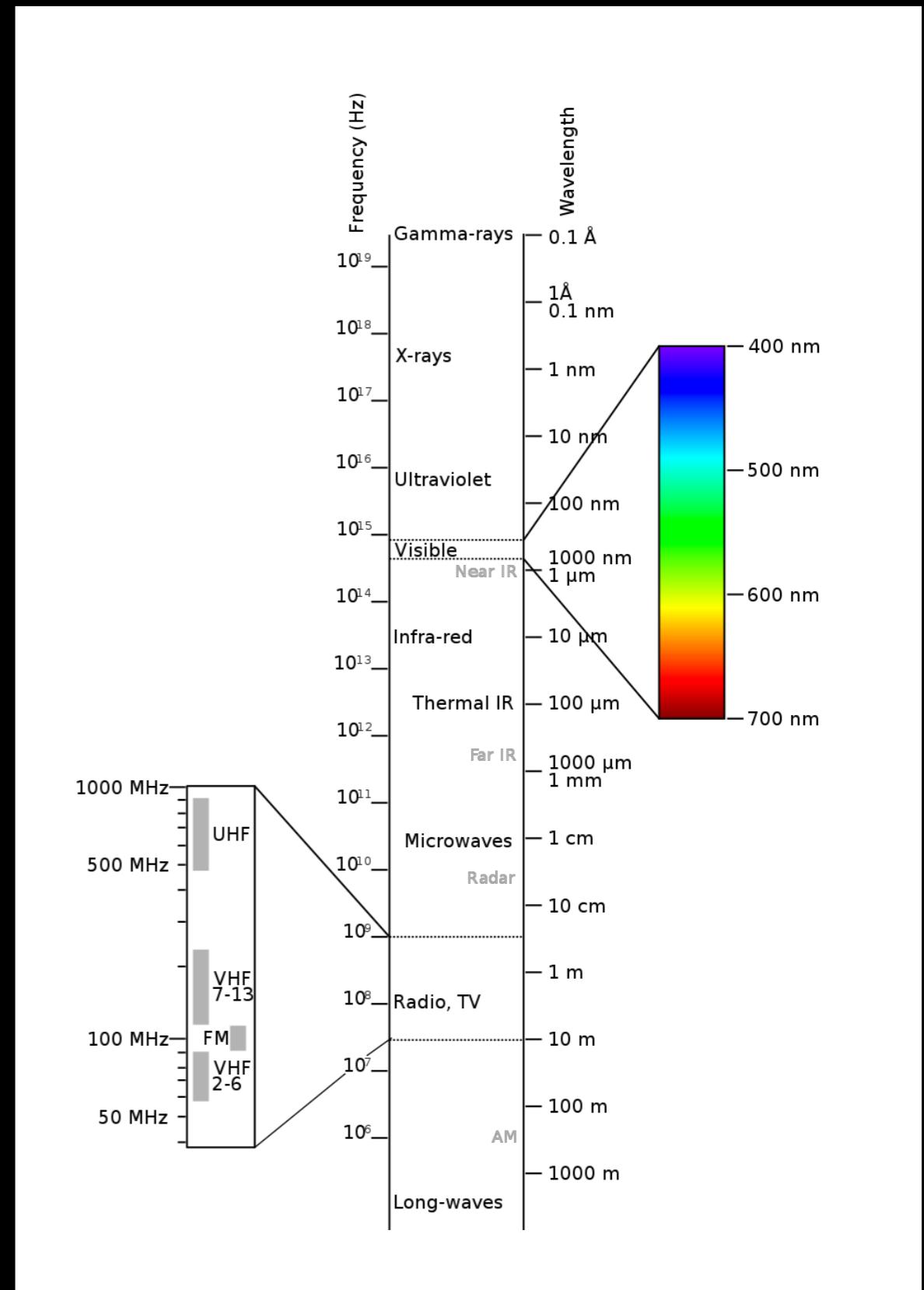
Visual interpretation of remote sensing data

- Brightness
- Color
- Texture
- Shape and size
- Shadows
- Period of acquisition



EMR wavelengths

- **Visible** (400 nm-700nm).
Red: 620-700 nm; Green: 500-578 nm; Blue: 446-500 nm.
- **Infrared** (0.7 um-100um).
NIR: 0.7 - 1.4 um; SWIR: 1.4-3 um; TIR: 3-15um. Far IR: 15-1000um
- **Microwave** (1 mm - 1 m).

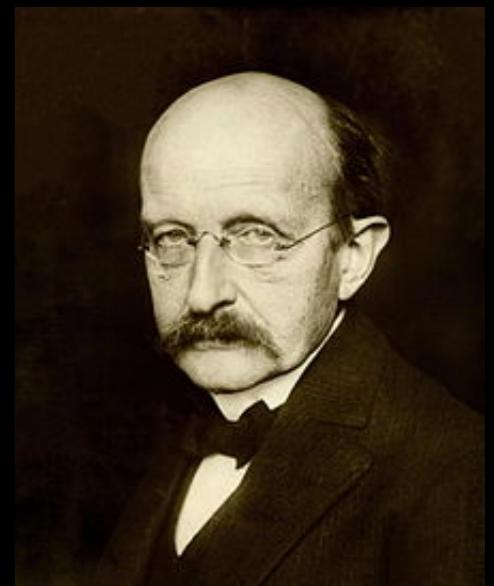


Electromagnetic radiation laws (I)

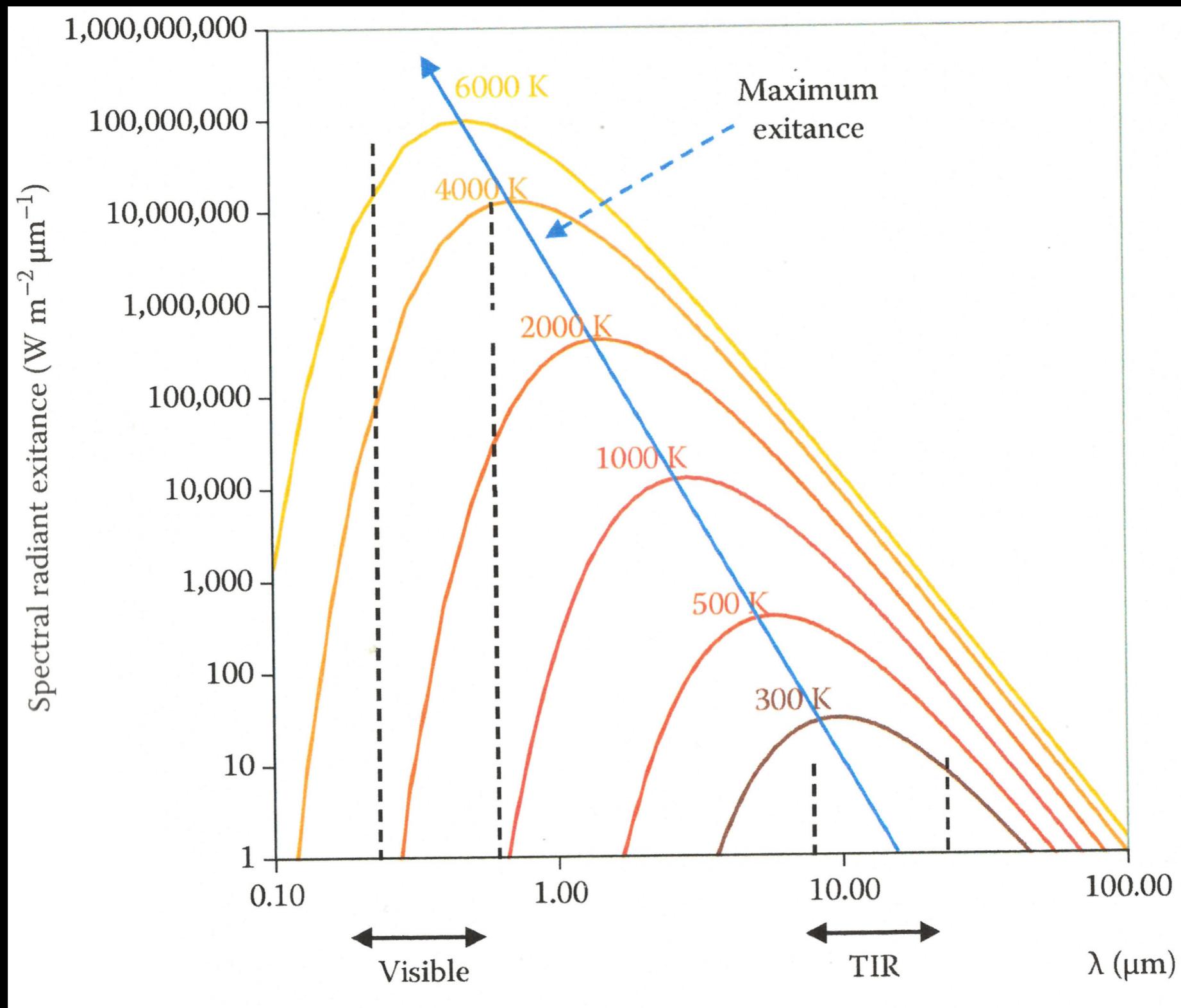
- Planck's law of radiation: the spectral distribution of the emittance of a blackbody is a function of its temperature and wavelength.

$$M(\lambda) = \frac{c_1}{\lambda^5 (e^{(c_2/\lambda T)} - 1)} = f(\lambda, T)$$

$$c_1 = 3.741 \times 10^8 \text{ W m}^{-2} \text{ um}^{-1};$$
$$c_2 = 1.438 \times 10^4 \text{ um K}$$



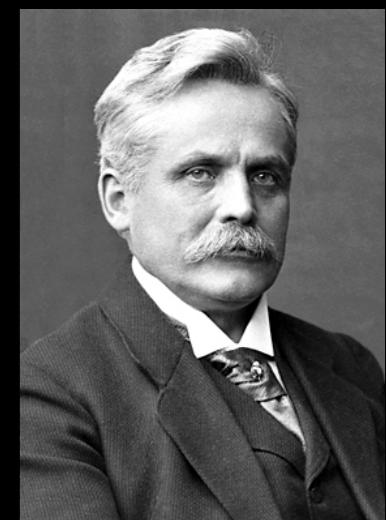
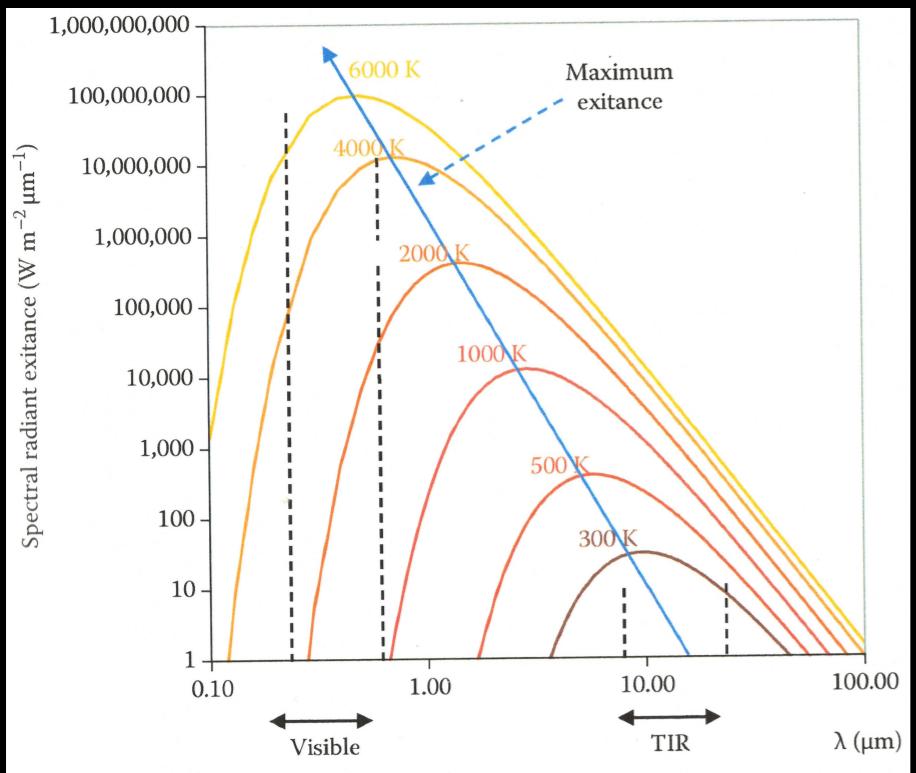
Blackbody radiation curves



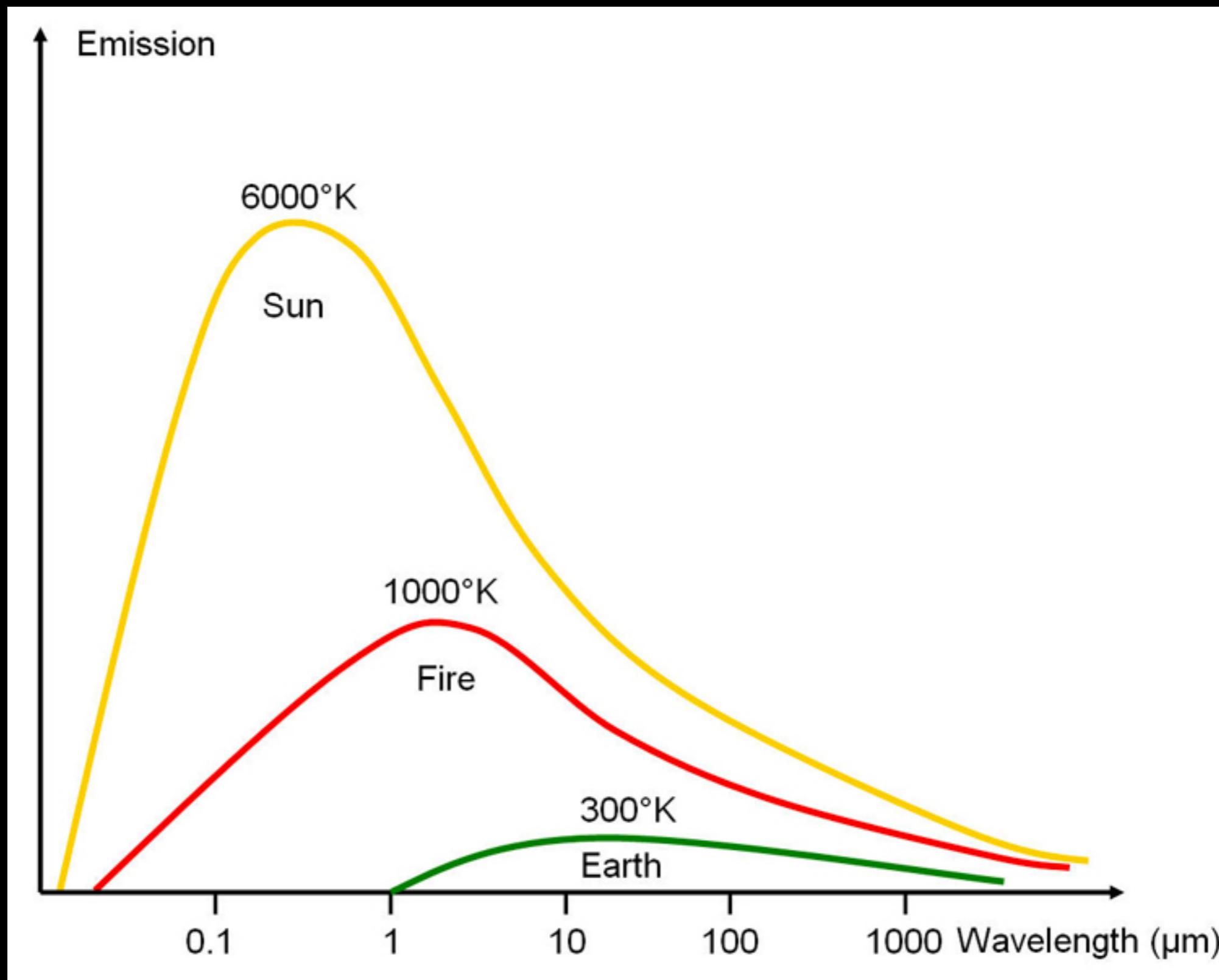
Electromagnetic radiation laws (II)

- Wien's Displacement Law:

$$\lambda_{\max} = \frac{2898}{T}$$



A remote sensing application



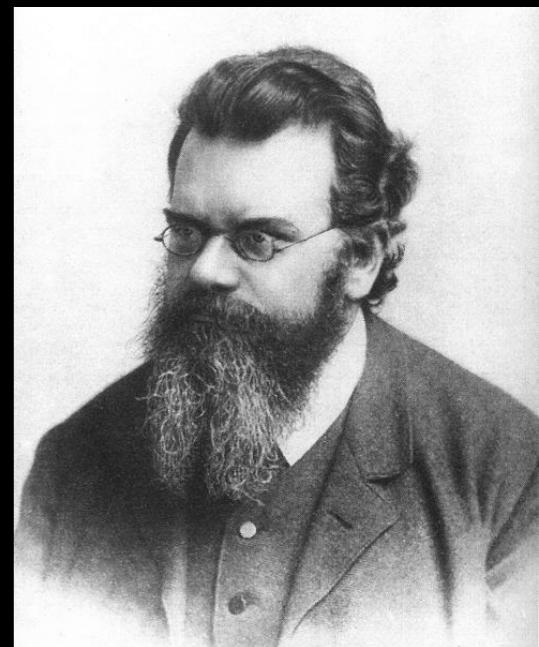
Electromagnetic radiation laws (III)

- Stefan-Boltzmann law:

$$M = \sigma T^4$$



$$\sigma = 5.6697 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$$



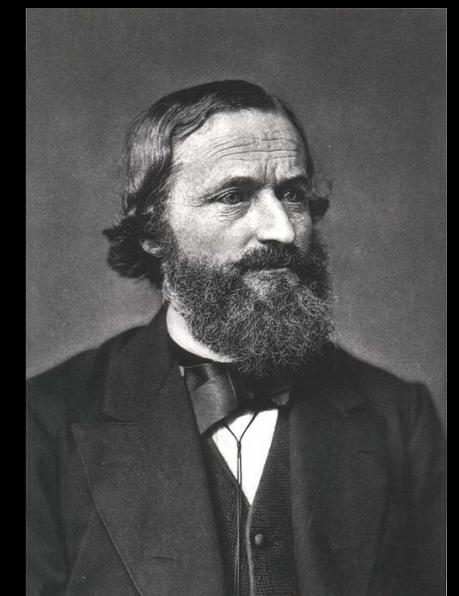
Electromagnetic radiation laws (IV)

- Kirchhoff's law:

$$M = \varepsilon \sigma T^4$$

$$\varepsilon=0.1$$

$$\sigma=5.6697 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$$



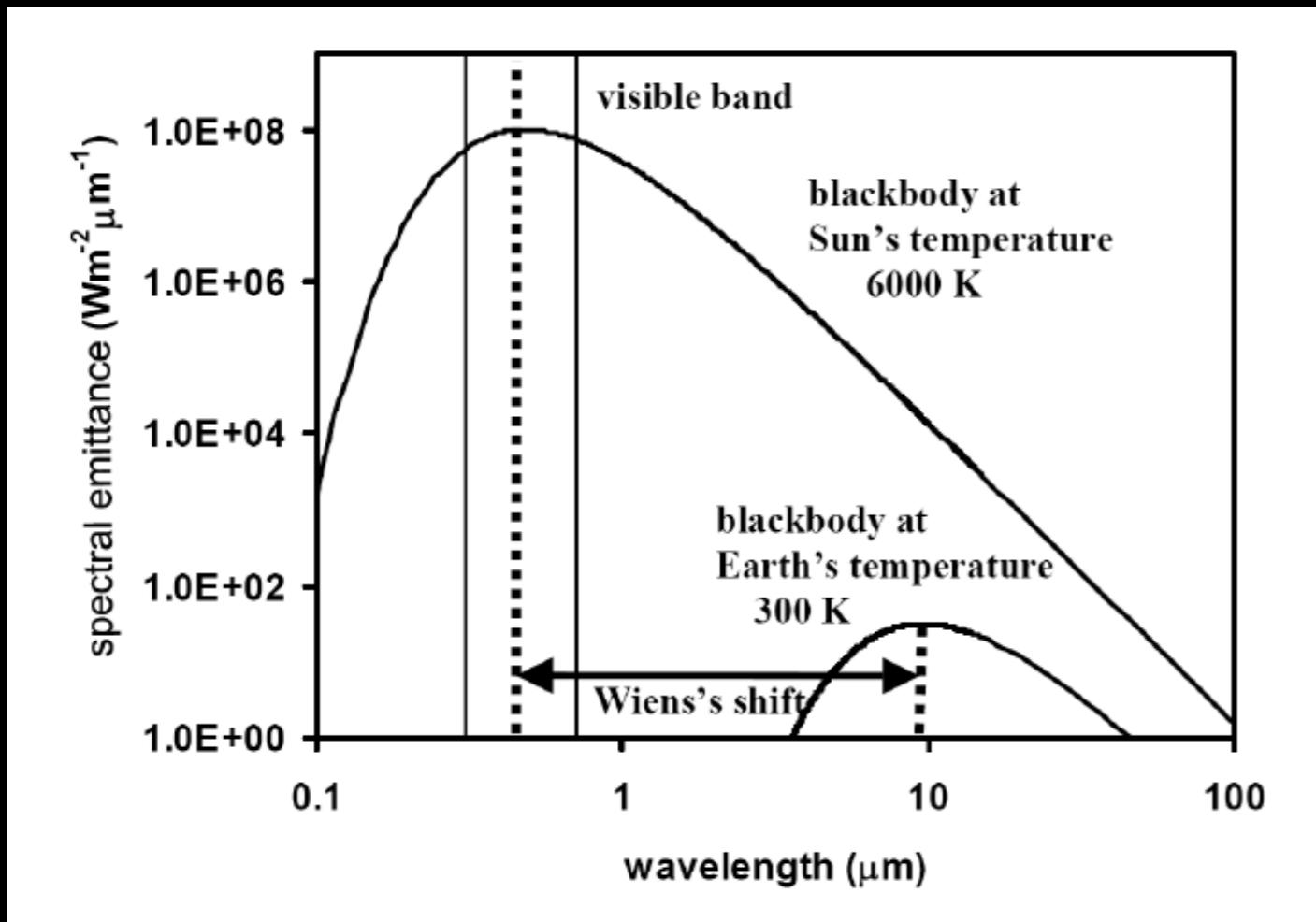
Electromagnetic radiation laws summary

- Four EMR laws tell us:

The magnitude and spectral distribution of energy depends on 1) the temperature; 2) the emissivity.

Knowing the temperature of the object, we can determine the most suitable wavelength for detecting objects of interest.

Sun and Earth can be treated as a blackbody

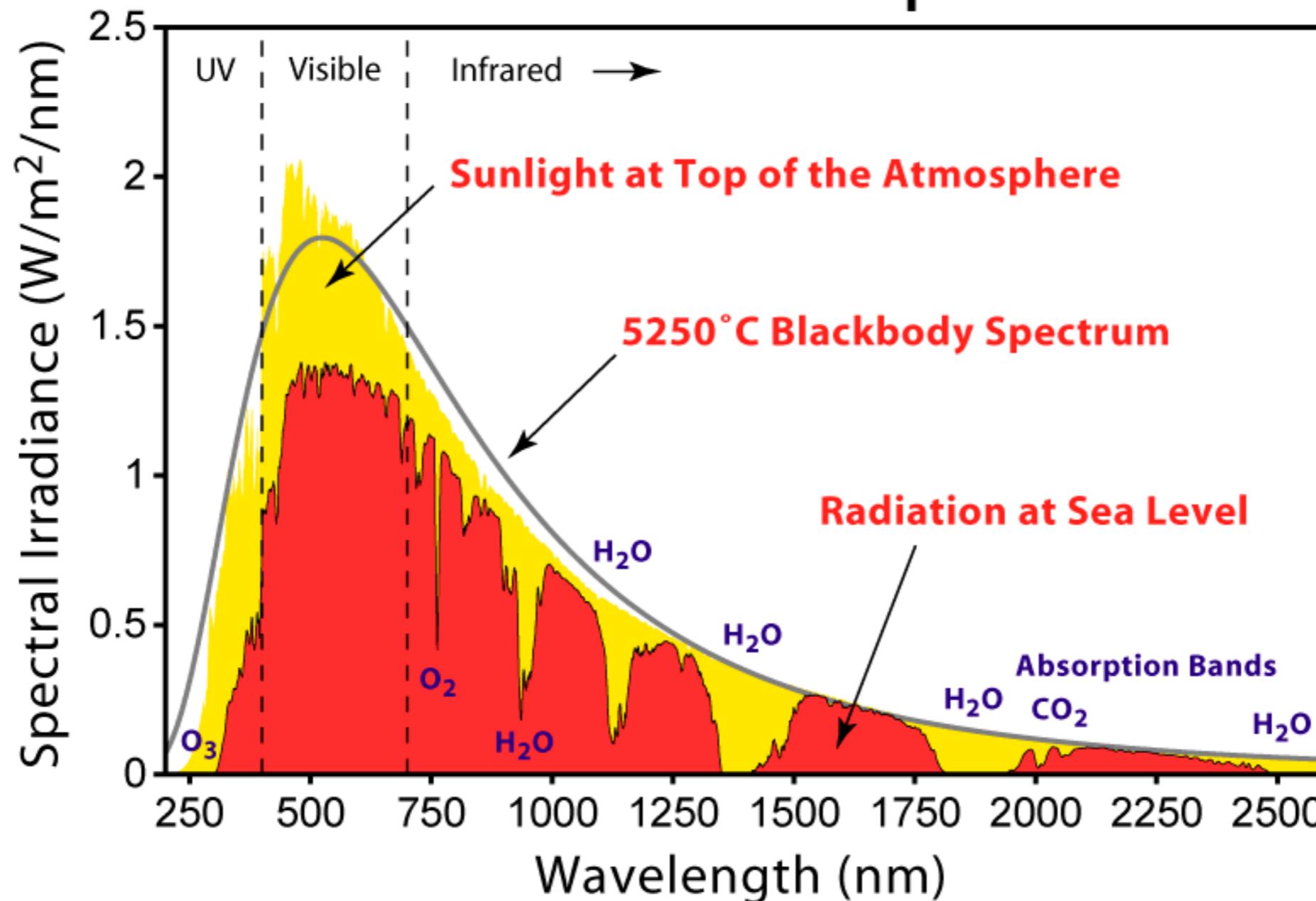


And the solar radiation (irradiance) at the top-of-atmosphere (E) is:

$$E = M \frac{R^2}{D^2}$$

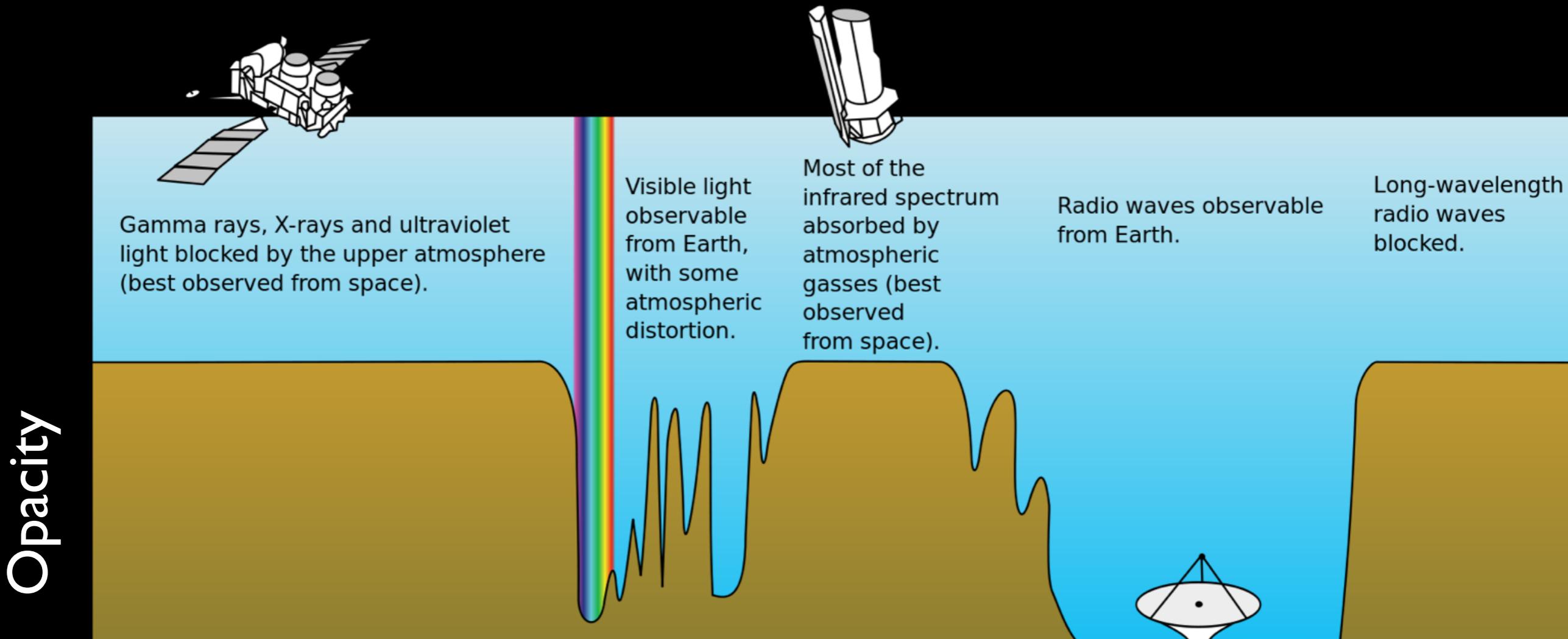
M is the exitance at the solar surface
R is radius of the sun
D is the sun-earth distance

Solar Radiation Spectrum

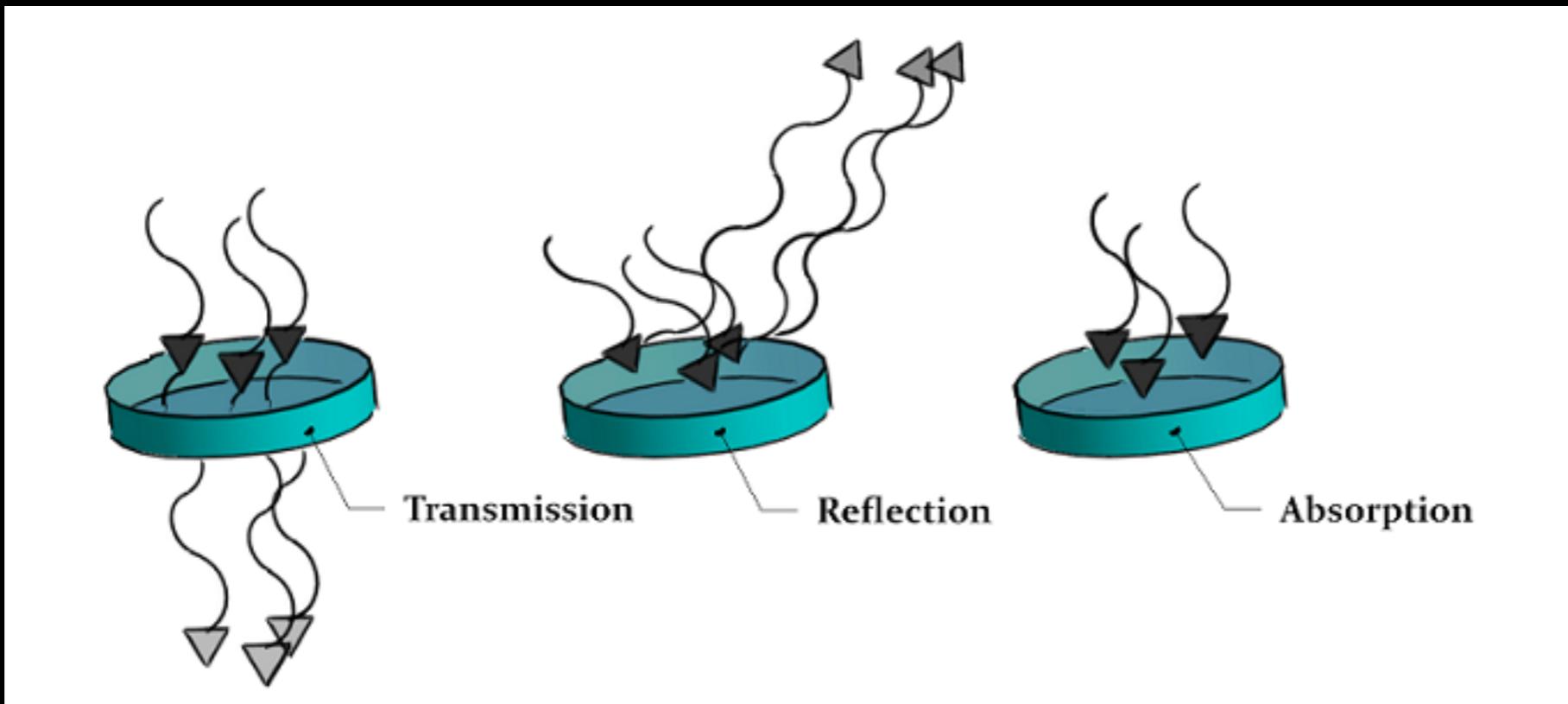


- ~ 40% of the solar energy is in visible wavelength (400 - 700nm)

Atmospheric window



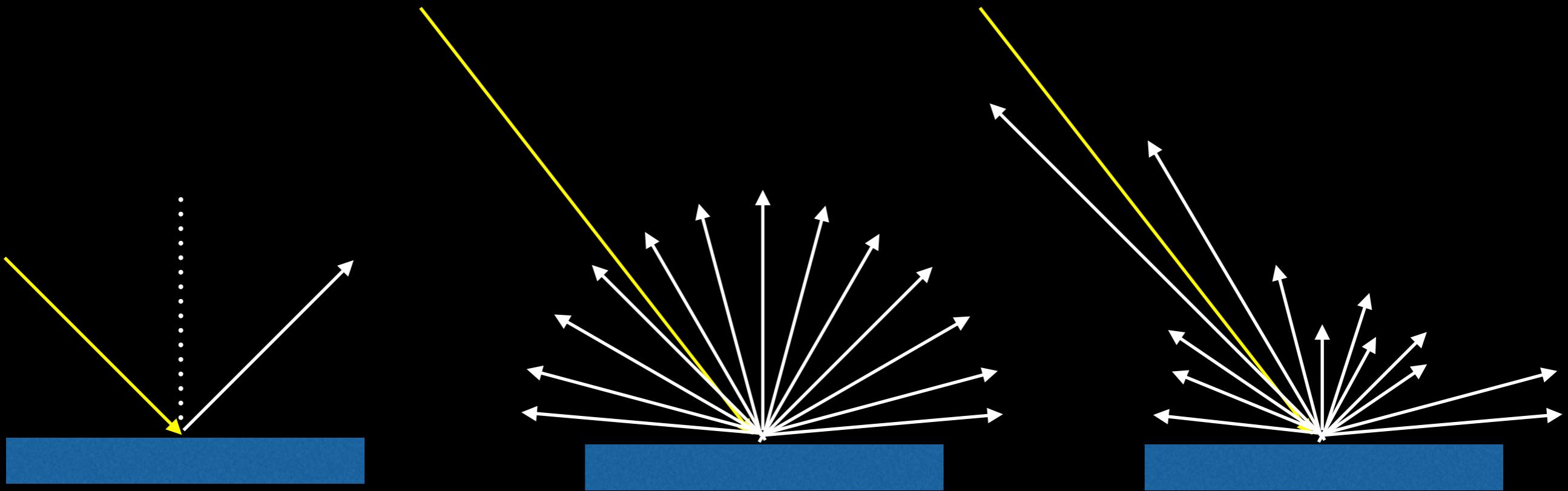
Reflectance



Incoming radiation = Transmission + Reflection + Absorption

$$I = Transmittance + Reflectance + Absorbance$$

Three types of surfaces



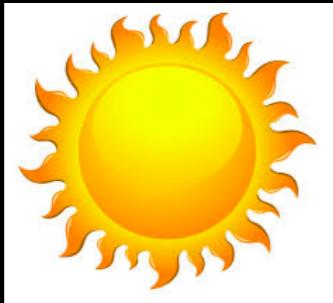
Specular

Lambertian
(isotropic)

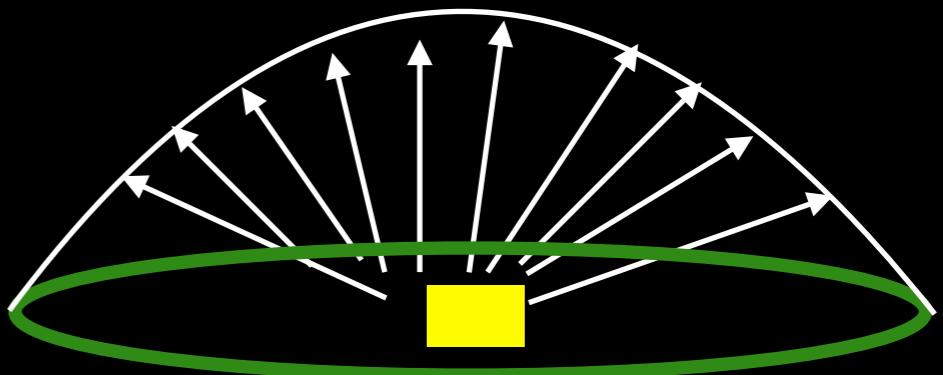
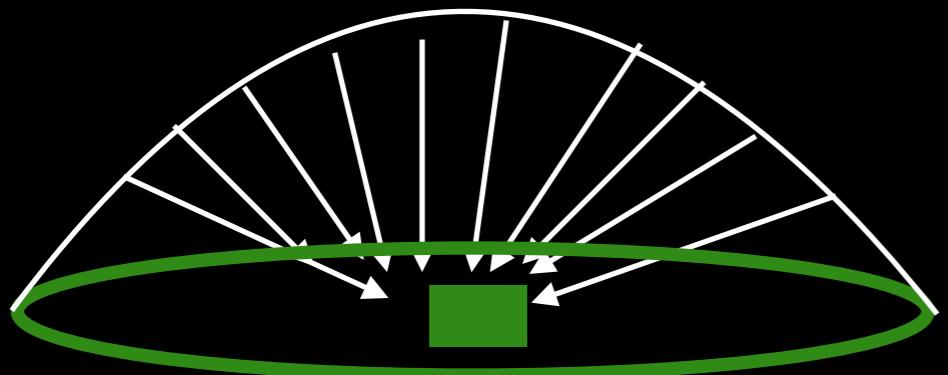
Anisotropic

For a Lambertian surface, if we know the radiance in one direction, L , then the exitance $M = \pi L$.

Albedo (α)

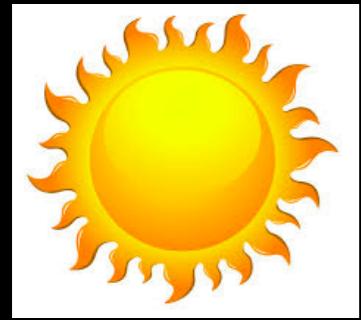


Ocean: 0.06
Conifer: ~0.10
Deciduous: 0.15
Grass: 0.25
Ice: 0.5-0.7
Fresh snow: 0.8

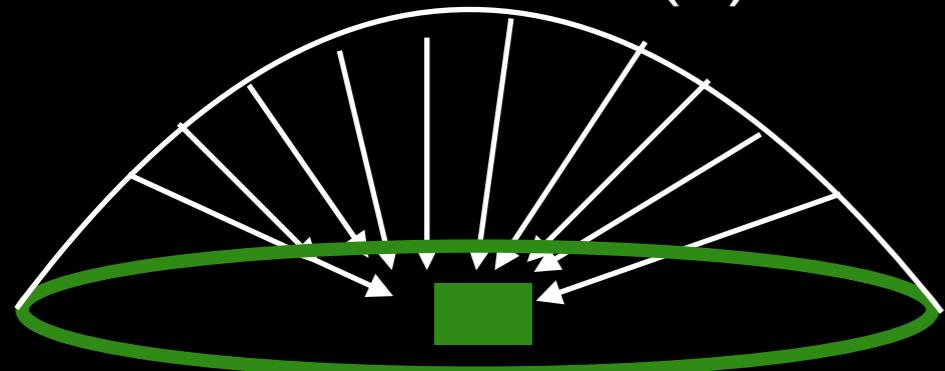


$$\alpha = \text{exitance/irradiance} = M/E$$

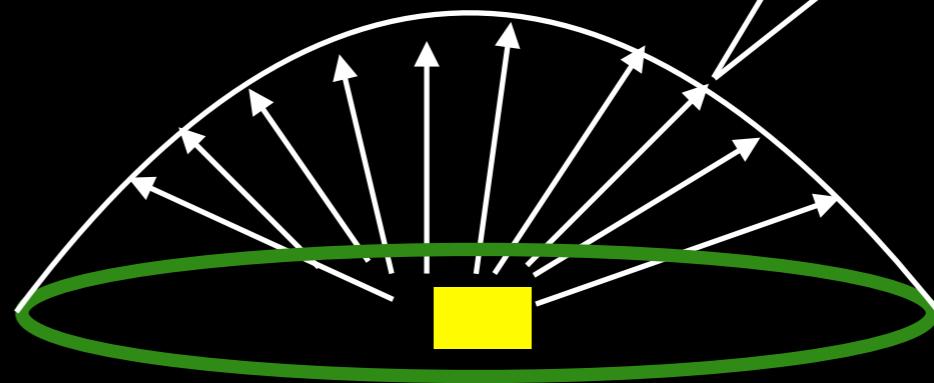
Reflectance in remote sensing (ρ_λ)



Irradiance (E)

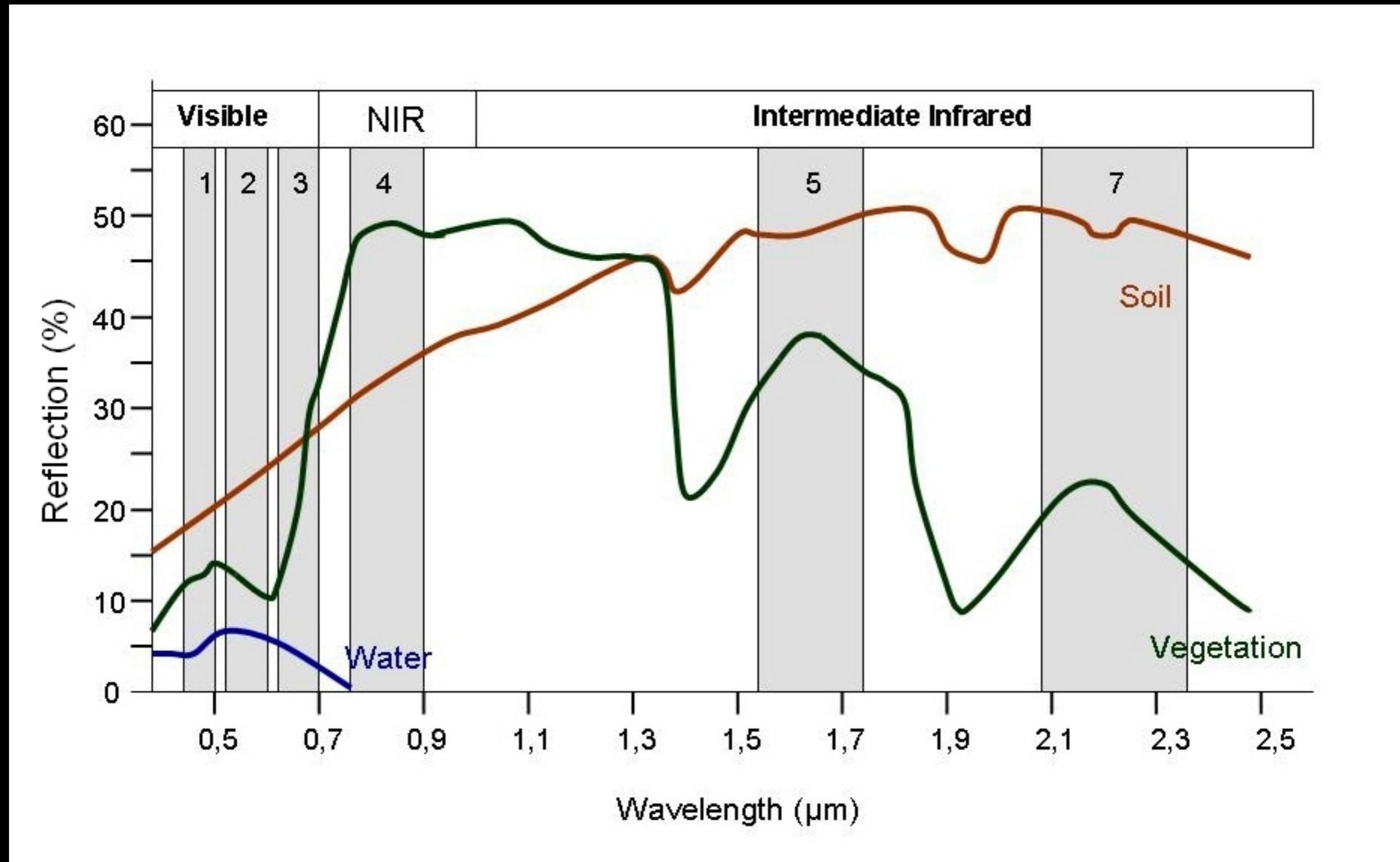


Radiance (L)

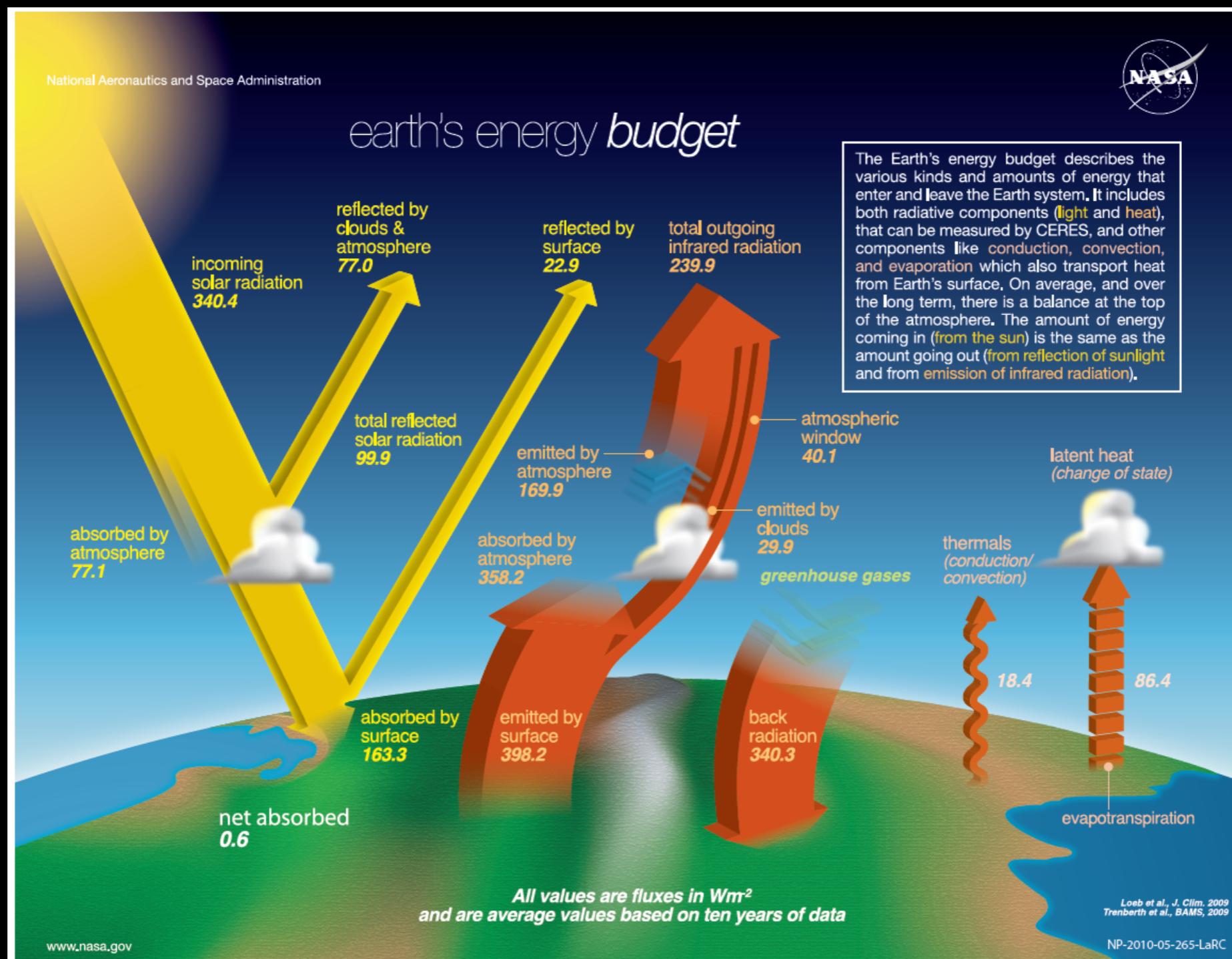


$$\rho_\lambda = \pi L / E$$

Reflectance of typical objects



Earth's energy balance



<https://www.youtube.com/watch?v=O0B8Yi7AZvQ>

Recap

- Wavelength, frequency, and energy.
- Regions of EM wave.
- Solid angle. Irradiance. Radiance.
- Data of the Week.
- Laws of EMR.
- Solar radiation. Reflectance. Albedo