

### EXERCISE #1

$$\textcircled{a} \quad 1+z = \frac{\lambda_{\text{obs}}}{\lambda_{\text{rest}}}, \quad v_{\text{rot}} = c \cdot \frac{\Delta\lambda}{\lambda}$$

• Lyman  $\alpha$  rest wavelength =  $1216 \text{ \AA} = \lambda_{\text{rest}}$

• From spectrum:  $\lambda_{\text{obs}} = 1214 \text{ \AA}$

$$z = \frac{\lambda_{\text{obs}}}{\lambda_{\text{rest}}} - 1 = \left( \frac{1214}{1216} \right) - 1 = -0.00164 \Rightarrow \text{M31 is blueshifted}$$

$$v_{\text{rot}} = c \cdot \frac{\Delta\lambda}{\lambda} = c \cdot \left( \frac{\lambda_{\text{obs}} - \lambda_{\text{rest}}}{\lambda_{\text{rest}}} \right) = -490 \frac{\text{km}}{\text{s}}$$

$$\textcircled{b} \quad d = 780 \text{ kpc}, \quad H_0 = 71 \frac{\text{km}}{\text{s Mpc}}$$

$$v_r = H_0 d = 71 \frac{\text{km}}{\text{s Mpc}} \cdot (780 \times 10^{-3} \text{ Mpc}) = 55 \frac{\text{km}}{\text{s}} \quad ??$$

$$\textcircled{c} \quad z = 0.05$$

$$v_r = cz = 2.98 \times 10^5 \frac{\text{km}}{\text{s}} \times 0.05 = 14.9 \times 10^3 \frac{\text{km}}{\text{s}}$$

$$d = \frac{v_r}{H_0} = 14.9 \times 10^3 \left[ \frac{\text{km}}{\text{s}} \right] \times \frac{1}{71} \left[ \frac{\text{Mpc s}}{\text{km}} \right] = 210 \text{ Mpc}$$

### EXERCISE #2

$$\textcircled{a} \quad \gamma + \gamma \rightarrow e^+ + e^-$$

$$E = mc^2 = k_B T = 0.51 \text{ MeV}$$

$$\Rightarrow T_{50} = \frac{mc^2}{k_B} = 5.94 \times 10^9 \text{ K}$$

$\Rightarrow$  at  $T \lesssim 6$  billion kelvin the electron-positron pairs can no longer be produced efficiently  $\Rightarrow$  no density of photons with energies above pair production threshold is too small

$$(b) T[K] \sim 1.5 \times 10^{10} (t[s])^{-1/2}$$

$$\eta = \frac{n_\gamma}{n_b} = 10^9 \Rightarrow 10^9 \text{ photons for each stable baryon}$$

$$T \sim 1.5 \times 10^{10} (1[s])^{-1/2} = 1.5 \times 10^{10} K$$

$$E = k_B T = 1.3 \text{ MeV}$$

$$(c) p + n \rightleftharpoons D + \gamma \quad m_p = m_n = 1.67 \times 10^{-27} \text{ kg}$$

$$\rightarrow m_D = 3.34 \times 10^{-27} \text{ kg} = 1875.61 \text{ MeV}/c^2$$

• Energy needed to disintegrate D:  $E > 2.23 \text{ MeV}$

$$\begin{aligned} \Delta m &= Z m_p + (A - Z) m_n - m_{nuc} \Rightarrow A = 2 \text{ \& } Z = 1 \text{ for D} \\ &= m_p + m_n - m_{nuc} \end{aligned}$$

Deuteron binding energy:

$$\bullet E_B = \Delta m c^2 = (m_p + m_n - m_{nuc}) c^2 = \underline{\underline{2.23 \text{ MeV}}}$$

$\Rightarrow$  Energy required to photo-disintegrate D must be greater than  $2.23 \text{ MeV} \Rightarrow E_{\gamma} > 2.23 \text{ MeV}$

$$\bullet E = \hbar \omega \rightarrow \omega = \frac{E}{\hbar} = \frac{2.23 \times 10^6 \text{ eV}}{6.58 \times 10^{-16} \text{ eV}\cdot\text{s}} = 3.39 \times 10^{21} \text{ Hz}$$

• Deuterons are present in such small amount since it is the element with lowest binding energy and thus it is easy to break up