

Time Symmetry

Rainer Spurzem

September 29, 2021

1 “Leap Frog”

Hamiltonian, relative two-body motion without perturbations

$$H = \frac{\mathbf{p}^2}{2\mu} - \frac{GMm}{|\mathbf{r}|} \quad (1)$$

with canonical conjugate variables $\mathbf{p} = \mu\mathbf{v}$, and \mathbf{r} ; H is separable:

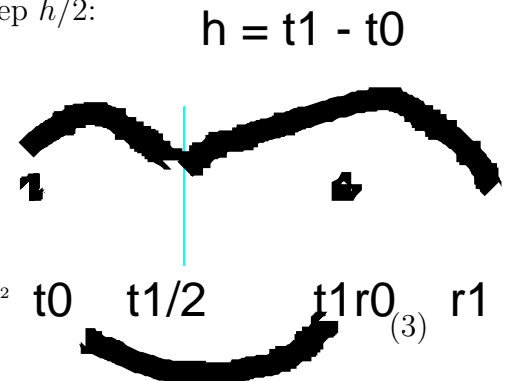
$$H = H_1(\mathbf{p}) + H_2(\mathbf{r}) = \frac{\mathbf{p}^2}{2\mu} + U(\mathbf{r}) \quad (2)$$

In the “Leap Frog” we start with an initial time step $h/2$:

$$r_{1/2} = r_0 + \frac{1}{2}hp_0$$

Then taking turns:

$$\begin{aligned} p_1 &= p_0 - h \left. \frac{dU}{dr} \right|_{r=r_{1/2}} \\ r_{3/2} &= r_{1/2} + hp_1 \end{aligned}$$



This solution is **time symmetric** (for more than two particles: only if all particles have the same time step). It means that the numerical solution is the solution of an approximate Hamiltonian $H = H_0 + hH_1 + \dots$

symplectic

2 Hermite

Remember from our previous lectures, we first do the predictor step (for all particles in regular step, and for all neighbours in irregular step):

$$\begin{aligned} \mathbf{x}_p(t) &= \frac{1}{6}(t-t_0)^3\dot{\mathbf{a}}_0 + \frac{1}{2}(t-t_0)^2\mathbf{a}_0 + (t-t_0)\mathbf{v}_0 + \mathbf{x}_0 \\ \mathbf{v}_p(t) &= \frac{1}{2}(t-t_0)^2\dot{\mathbf{a}}_0 + (t-t_0)\mathbf{a}_0 + \mathbf{v}_0 \end{aligned} \quad (4)$$

We get then the corrector step for the i particle (in bolded here):

$$\begin{aligned}\mathbf{x}_c(t) &= \mathbf{x}_p(t) + \frac{1}{24}(t-t_0)^4 \mathbf{a}^{(2)} + \frac{1}{20}(t-t_0)^5 \mathbf{a}^{(3)} \\ \mathbf{v}_c(t) &= \mathbf{v}_p(t) + \frac{1}{6}(t-t_0)^3 \mathbf{a}^{(2)} + \frac{1}{24}(t-t_0)^4 \mathbf{a}^{(3)}.\end{aligned}\quad (5)$$

Remember, that $\mathbf{a}^{(2)}$ and $\mathbf{a}^{(3)}$ have been obtained from the special Hermite step (Taylor series for accelerations), while \mathbf{a} and $\dot{\mathbf{a}}$ have been computed directly from Newton's law and its time derivative (sums over all other particles). This method is **not** time symmetric. It means that our numerical solution does not belong to a Hamiltonian function; the best we can say is following Quinlan & Tremaine (1995) that it jumps from one Hamiltonian to another one, never leaving the "possible" space of solutions.

Makino (1997) has proposed to drop the highest order term in \mathbf{x} , and shown that then the method is still 4th order and time-symmetric:

$$\begin{aligned}\mathbf{v}_c(t) &= \mathbf{v}_0(t) + \frac{1}{2}(t-t_0)(\mathbf{a}_1 + \mathbf{a}_0) - \frac{1}{12}(t-t_0)^2(\dot{\mathbf{a}}_1 - \dot{\mathbf{a}}_0) \\ \mathbf{x}_c(t) &= \mathbf{x}_0(t) + \frac{1}{2}(t-t_0)(\mathbf{v}_c + \mathbf{v}_0) - \frac{1}{12}(t-t_0)^2(\mathbf{a}_1 - \mathbf{a}_0).\end{aligned}\quad (6)$$

You can use the Hermite interpolation formulae (previous lecture manuscript) and show the equivalence. This method can be also used to iterate to higher precision, by using \mathbf{x}_c and \mathbf{v}_c in a second step to re-compute \mathbf{a}_1 and $\dot{\mathbf{a}}_1$, until convergence is reached. In that way we avoid the spurious trick to use low-order predicted coordinates.

Eq. (5) and Eq. (6) are equivalent, use Hermite step

3 Regularisation as canonical transformation

We used in classical KS (Kustaanheimo & Stiefel) regularisation:

$$g(\mathbf{r}, \mathbf{p}, t) ds = \mathbf{r} ds = d\mathbf{t} \quad \mathbf{r} = \mathbf{u}^{**2} \quad (7)$$

and the Poincaré-Transformation of the Hamiltonian

$$\Gamma = g(H - E) \quad (8)$$

leads to the following canonical equations (the last equation originates from taking $p_0 = E$ as canonically conjugate to the time t):

$$\begin{aligned}\mathbf{p}' &= -\frac{\partial \Gamma}{\partial \mathbf{q}} \\ \mathbf{q}' &= \frac{\partial \Gamma}{\partial \mathbf{p}} \\ t' &= \frac{\partial \Gamma}{\partial p_0} = g\end{aligned}\quad \mathbf{r}' \quad (9)$$

with the derivative $'$ with respect to s . For the isolated two-body problem we get the new Poincaré-Hamiltonian

$$\Gamma = u^2 \left(\frac{\mathbf{p}^2}{8u^2} - \frac{GM}{u^2} - E \right) \quad (10)$$

so for $E < 0$:

$$\Gamma = \frac{1}{8}\mathbf{p}^2 + |E|\mathbf{u}^2 = GM \quad (11)$$

Note that in case of 3D particle coordinate space we have to use the 4D quaternion space for the vectors \mathbf{u} and \mathbf{p} .

4 Algorithmic Regularisation

This new idea of Mikkola (2000) and also Preto & Tremaine (1999) we start again with the isolated two-body Hamiltonian:

$$H = H_1(\mathbf{p}) + H_2(\mathbf{r}) = T(\mathbf{p}) + U(\mathbf{r}) \quad (12)$$

Now a new regularising time transformation is used:

$$|U|^{-1}ds = dt \quad (13)$$

and the Poincaré transformed Hamiltonian is

$$\Gamma = |U|^{-1}(T(\mathbf{p}) + U(\mathbf{r}) - E) \quad (14)$$

It is not separable, so it looks bad, but

$$\Lambda = \log(1 + \Gamma) = \log(T(\mathbf{p}) - E) - \log|U(\mathbf{r})| \quad (15)$$

is separable!! And one can show that now new canonical equations are valid:

logarithmic Hamiltonian

$$\begin{aligned} \mathbf{p}' &= -\frac{\partial \Lambda}{\partial \mathbf{r}} = -\frac{1}{1 + \Gamma} \cdot \frac{\partial \Gamma}{\partial \mathbf{r}} \\ \mathbf{r}' &= \frac{\partial \Lambda}{\partial \mathbf{p}} = \frac{1}{1 + \Gamma} \cdot \frac{\partial \Gamma}{\partial \mathbf{p}} \\ t' &= \frac{\partial \Lambda}{\partial p_0} = \frac{1}{1 + \Gamma} \cdot \frac{\partial \Gamma}{\partial p_0} \end{aligned} \quad (16)$$

$$\mathbf{p}' = d\mathbf{p}/ds \sim (\mathbf{p}_1 - \mathbf{p}_0)/ds$$

AR regularization (2 bodies) AR C

$$\begin{aligned} \mathbf{p}' &= -\frac{1}{|U|} \frac{\partial U}{\partial \mathbf{r}} \\ \mathbf{r}' &= \frac{1}{T - E} \cdot \frac{\mathbf{p}}{m} \\ t' &= \frac{1}{T - E} \end{aligned} \quad (17)$$

These equations can be used to make a new “Leap Frog” - the time-transformed leap frog. Algorithmic regularization is based on this.

5 New Developments - very short

GPU

Rantala, Naab, Springel, et al. 2021: frost: a momentum-conserving CUDA implementation of a hierarchical fourth-order forward symplectic integrator

Rantala et al. 2020: MSTAR - a fast parallelized algorithmically regularized integrator with minimum spanning tree coordinates

Chin & Chen 2005: Symplectic Integrators

Planetary Systems: Mercury, new variants of Hernandez