

# A Novel Framework for Highlight Reflectance Transformation Imaging

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## Abstract

We propose a novel pipeline and related software tools for processing the multi-light image collections (MLICs) acquired in different application contexts to obtain shape and appearance information of captured surfaces, as well as to derive compact relightable representations of them. Our pipeline extends the popular Highlight Reflectance Transformation Imaging (H-RTI) framework, which is widely used in the Cultural Heritage domain. We support, in particular, perspective camera modeling, per-pixel interpolated light direction estimation, as well as light normalization correcting vignetting and uneven non-directional illumination. Furthermore, we propose two novel easy-to-use software tools to simplify all processing steps. The tools, in addition to support easy processing and encoding of pixel data, implement a variety of visualizations, as well as multiple reflectance-model-fitting options. Experimental tests on synthetic and real-world MLICs demonstrate the usefulness of the novel algorithmic framework and the potential benefits of the proposed tools for end-user applications.

*Keywords:* Multi Light Image Collections, Highlight Reflectance Transformation Imaging, Photometric Stereo, Image Enhancement

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## 1. Introduction

2 Multi-light image collections (MLICs) are an effective mean to gather detailed information  
3 on the shape and appearance of objects. They are, thus, widely used in many application contexts.

4 The basic idea of the approach is to visually characterize objects by capturing multiple im-  
5 ages of the surface of interest from a fixed point of view, changing the illumination conditions  
6 at each shot. The acquired data is then processed to extract shape and material information.  
7 While some techniques exist for general variable environmental illumination [1, 2], the most  
8 widespread approach in most application fields, including Cultural Heritage (CH) [3], medical  
9 interventions [4, 5], and underwater data gathering [6], considers a single calibrated camera tak-  
10 ing multiple images of a scene illuminated by a single moving light. A large variety of efficient  
11 computational tools have been devised in this context to extract information from the captured  
12 image stack in order to effectively solve different problems, such as feature detection and en-  
13 hancement, reconstruction of normals and 3D shapes, and creation of relightable images.

14 Photometric Stereo (PS) is probably the most widely known technology based on MLICs.  
15 It exploits priors on reflectance functions to derive local normals and global 3D shapes [7]. Re-  
16 flectance Transformation Imaging (RTI) [8, 9, 10, 11] extends the PS idea by interpolating MLIC  
17 reflectance data with parametric functions (Polynomial Texture Maps, PTM; Hemispherical Har-  
18 monics, HSH; Discrete Modal Decomposition, DMD), which can be used to estimate and display  
19 images relighted from arbitrary angles and incorporating other image enhancements for visually  
20 revealing surface details not detectable from a single view [10, 12, 13]. Other techniques try,  
21 rather, to improve understanding by highlighting specific making surface and material proper-  
22 ties. For instance, Raskar et al. [14] exploited multi-light images to extract depth edges with a  
23 simple heuristics and used the result to create non-photorealistic rendering methods, while Fattal  
24 et al. [15] used them to generate enhanced images emphasizing shape and surface detail.

25 RTI is possibly the most widely applied MLIC technique. This kind of imaging has rapidly  
26 become a widely used solution for the documentation, recording and decoding of Cultural Her-  
27 itage (CH) objects, as it supports an offline analysis of the artifacts, supporting and going beyond  
28 simulated raking light analysis, and allows the estimation of image enhancements emphasizing  
29 details [16, 17]. Furthermore, the reflectance interpolation coefficients derived from MLIC pro-  
30 cessing, or the image features extracted from the image stack, can be used to characterize and  
31 classify materials, as shown in a number of works [18, 19, 20].

32 The widespread use of RTI for visual surface characterization, especially in the CH domain,  
33 is also due to the fact that it can be performed with a low-cost, flexible, and easy to use setup  
34 based on freehand light positioning and highlight-based light direction estimation (H-RTI) [3].  
35 In the H-RTI image-capture technique, the reflection of the light source on one or more reflective  
36 spheres visible in each shot enables the processing software to calculate the light direction for  
37 each image, providing great robustness and flexibility in subject size and location.

38 The classic H-RTI acquisition setup and processing pipeline, however, are based on strong  
39 assumptions on lights (ideally constant in direction and intensity) and camera model (ortho-  
40 graphic), not necessarily matching typical acquisition conditions [3, 21]. In particular, due to the  
41 lack of uniformity in illumination intensity and direction, the results obtained with this simple  
42 setup may vary widely between acquisitions, and may be unsuitable for quantitative analyses,  
43 which include normal estimation, roughness or material segmentation/classification, as well as  
44 monitoring over time. Exploitation of H-RTI data is thus often limited to rough qualitative anal-  
45 ysis of single acquisitions.

46 In this article, we revise the H-RTI approach, presenting a novel practical setup and a set  
47 of tools that relax the aforementioned strong assumptions. Our solution offers a better support  
48 for qualitative analysis of MLICs and enables the addition of quantitative analysis on top of the  
49 classic RTI method. Our main contributions are the following:

- 50 • a novel practical setup and processing pipeline that can cope with the effects of perspective  
51 camera distortion, non-point lights, spatially varying illumination, variable light distance,  
52 as well as camera vignetting. Per-pixel light directions are estimated from highlights on  
53 multiple reflective spheres, taking into account perspective correction and performing di-  
54 rection interpolation, while illumination variations are compensated by an algorithm ex-  
55 ploiting light intensity measured on matte white targets positioned around the object of  
56 interest.
- 57 • An easy to use tool to perform/control all the processing pipeline, not requiring to rely  
58 on external image processing applications and storing reordered pixel information with

59 associated light directions in a dedicated structure that can be effectively used for post-  
60 processing (e.g. Photometric Stereo, RTI, feature detection, and visualization).

- 61 • An easy to use tool to complete the pipeline with enhanced visualizations, as well as with  
62 shape and material information recovery operations.

63 Our novel combination of several flexible auto-calibration techniques into a single framework  
64 aims to provide a significant step towards a practical quantitative and repeatable analysis using  
65 simple and low-cost free-form acquisition strategies.

66 The paper is organized as follows. Sec. 2 provides a quick overview of existing RTI ac-  
67 quisition setups and light calibration approaches, while in Sec. 3 our algorithmic contribution  
68 is described. The tools that implement the proposed approach are presented in Sec. 4. Sec. 5  
69 demonstrates, with experimental tests, the advantages of our improved pipeline, as well as the  
70 potential advantages of its use for practical applications.

71 **2. Related work**

72 Multi-light acquisition, processing, and analysis are broad research subjects, and a full review  
73 is out-of-scope for this article. We concentrate here only on the most-related methods to perform  
74 RTI acquisition and processing. For a wider coverage, we refer the reader to established surveys  
75 in surface reflectance capture [22], multi-light computational frameworks [7], digital modeling  
76 of material appearance [23], and geometric analysis in cultural heritage [24].

77 A wide variety of RTI acquisition setups exist, ranging from low-cost and transportable  
78 kits [21] to different sizes of fixed light domes[21, 25, 26]. Recently, some dome solutions have  
79 been presented that use both visible and invisible light wavelengths [27, 28]. Dome solutions al-  
80 low for pre-calibration of lights, but they are, in general, expensive and not flexible, thus limiting  
81 the potential changes in light numbers, positions and types, and the size of captured surfaces.

82 Our goal is, rather, to improve the classic hand-held light capture, which is low-cost, simple  
83 to implement, and allows for a more flexible choice of the number and the positions of the light  
84 sources; these factors are very important, for example, when dealing with non-Lambertian, shiny  
85 materials. Moreover, it is easy to extend the presented pipeline to the multi- or hyper-spectral  
86 domain at a much lower cost than multiple-light setups, which can easily require hundreds of  
87 illuminators or filters.

88 Free-form hand-held setups are widely used in the Cultural Heritage domain as powerful  
89 Computational Photography tools by many end users, especially to create relightable images and  
90 for detail enhancement. This large diffusion is mainly due to publicly available packages such as  
91 RTIBuilder and RTIViewer [21], which employ the H-RTI capture setup and use manual annota-  
92 tion of reflective spheres and highlight-based light direction estimation. However, these tools rely  
93 on limiting assumptions about lighting and camera, i.e., uniform and far point light (collimated  
94 rays) and orthographic camera with an ideal lens. Since the computation of surface attributes  
95 leads to significant errors and provides variable results for each acquisition, the applications of  
96 this method for geometrical reconstruction, material acquisition, and quantitative analysis are  
97 limited. Conversely, we want to adopt here more realistic lighting and camera models, taking  
98 into account optical effects such as vignetting, non-uniform light emission, and light attenuation  
99 with distance.

100 The calibration of real illumination is a well-known topic in Computer Vision, and, specif-  
101 ically, in the Photometric Stereo (PS) field [7]. While some methods try to implicitly consider

real light sources within the particular PS framework [29, 30], others are more focused on the explicit calibration of various lighting properties. Some methods make assumptions on light form factor, e.g., near point light [31] or linear light source [32], and try to exploit the illuminated scene to extract the light position and direction. For instance, Ahmad et al. [31] exploit diffused maxima regions in the framed model, and derive from them the light directions. Others perform calibration by sampling light on calibration targets of known shape and albedo (e.g., reflective spheres or diffuse targets). Corsini et al. [33] use high-dynamic range images of two reflective balls to acquire the spatially-varying illumination of a real-world scene, and it focuses more on the environment light effect rather than of the computation of a per-pixel light direction and intensity. Ackermann et al. [34] present a study and validation through error statistics of both a forward and backward geometric point light source calibration by using sets of different numbers of reflective spheres. Although it proposes a very simple and robust way to compute light direction, it considers a point light model without taking into account non-uniform light intensity. Other methods strongly rely on a specific, fixed light form factor (e.g., LED light [35, 36], and model the intensity with the corresponding fall-off due to both distance and angle to the light principal axis. Xie et al. [36] also consider vignetting effects. Unfortunately, those methods are not applicable to the case of a general variable illumination due to non-ideal lamps or lenses. Some works thus try to cope with non-uniform intensity without imposing an analytical light model [37, 38]. Similarly to us, they use a flat reference object with known albedo to sample an arbitrary lighting vector field and to calibrate it using a flat-fielding approach. They don't use polynomial interpolation, but they exploit measured spatially-varying intensities to compensate the input images, and to convert the problem into a standard collimated case. Differently to our work, they require different acquisitions for the calibration step and the actual capture; this is possible only with a fixed light configuration, but it is not applicable to a more general free-form, hand-held multi-light acquisition. In our approach, we use multiple spheres to estimate a light direction field, and use measures on a planar white target to estimate the intensity of each light ray, infilling missing data with a low-degree interpolation, thus reconstructing an approximation of the entire light field illuminating the scene.

The pipeline presented here was preliminarily proposed in our previous conference papers [39, 20]. The pipeline improves the classical highlight-based RTI capture framework by estimating per-pixel interpolated light direction and creating intensity-corrected images simulating constant illumination on a reference plane. We here provide a more thorough exposition, but also significant new material, including the support for a non-orthographic camera model, a new organization of data that facilitates processing and display, the presentation of easy-to-use software interfaces to perform all the processing steps and novel experiments to demonstrate the advantages of the proposed methods. Finally, we have attempted to further clarify the steps in our methods to facilitate their implementation and to make the transfer between abstract concepts and actual code as straightforward as possible.

### 3. Improved Highlight RTI pipeline

Our complete acquisition and processing pipeline is shown in Fig. 1. We acquire and take as input a Multi-Light image collection. Light information may be in principle known for each image if coming from a calibrated system (light dome). If lights are not known and calibrated, as in hand-held light acquisition, the classical solution is to assume uniform intensity and direction and use a reflective sphere for estimating light direction from highlight position (H-RTI).

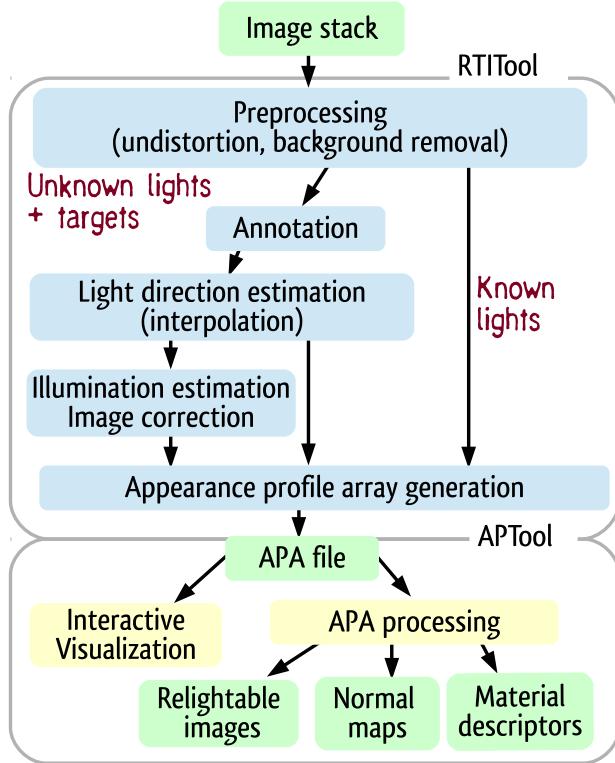


Figure 1: The proposed MLIC processing pipeline.

146 Our first contribution is a more complete setup (Fig. 2(a)) to characterize lights directly from  
 147 images, improving H-RTI. This setup includes, in addition to several (typically four) reflective  
 148 spheres, a matte white planar frame around the object being captured. The multiple spheres  
 149 are used to derive a more accurate per pixel interpolated direction, while the frame is used to  
 150 estimate a correction for light non-uniformity and vignetting, as described in Sec. 3.4. Several  
 151 instantiations of this concept are possible. In particular, if an object is captured in a typical  
 152 laboratory setup, the white frame can be replaced by a Lambertian surface covering the plane  
 153 supporting the object. Moreover, in outdoor acquisitions of large objects, spheres at the corner  
 154 of the visual fields and multiple co-planar Lambertian targets on the acquisition reference plane  
 155 could be placed, as well, and used for the subsequent calibration procedures. In order to simplify  
 156 generic on-site acquisitions, we realized a modular frame building set, which combines 3D-  
 157 printed supports for spheres with aluminum bars of different lengths covered by approximately  
 158 Lambertian coating (Fig. 2(b)). This allows the creation of rigid frames that can be placed in  
 159 horizontal, vertical or arbitrary orientations. The current version holds 5cm wide spheres, but we  
 160 plan to realize sets of different sizes.

161 Before the acquisition, we assume that we have already (and once) calibrated the internal  
 162 characteristics of the camera, in order to obtain the radiometric response function and the  
 163 lens parameters. The capture process outputs an image stack, which is preprocessed and semi-  
 164 automatically annotated with custom software (see Sec. 4) to find the position of the spheres and



Figure 2: **Capture setup.** (a) The acquisition setup consists in a stationary DSLR digital camera, a hand-held light source, some calibration targets (a Lambertian white background/frame and four black reflective spheres), and the object being acquired positioned on a planar support perpendicular to the camera axis. (b) Camera view of the scene.

165 rectangular regions of the white frame (Fig. 3). From the positions of highlights, the incident  
 166 light direction is estimated at the highlight location and interpolated across the whole image.  
 167 Then, illumination intensity is corrected at each pixel location. This is done by multiplying the  
 168 local value by the factor that would make the locally interpolated white frame intensity match a  
 169 reference value multiplied by the cosine of the local light direction elevation angle.

170 After that, each pixel is associated with a calibrated reflectance profile (appearance profile),  
 171 coupled with calibrated light parameters. Those are used to provide the user with an interactive  
 172 data visualization, and to perform various processing operations on reflectance data. For instance,  
 173 as in typical RTI settings, we fit reflectance data to a low-dimensional analytic representation, in  
 174 order to extract a small set of coefficients that can compactly describe the image stack at each  
 175 pixel. Then, we use this information to relight the object, to compute geometric attributes (e.g.,  
 176 normal maps or 3D surface reconstruction), or to extract meaningful appearance features and  
 177 descriptors for material classification and recognition.

178 All the procedures can be controlled by two software tools that will be described in detail  
 179 in Sec. 4: one dedicated to the preprocessing and reorganization of pixel data (RTITool), one to  
 180 reflectance data fitting, normals estimation, visualization and analysis (APTool).

181 In the rest of this section, we provide details on the major pipeline components: preprocessing  
 182 to prepare data for further elaboration (Sec. 3.1), perspective light direction estimation from  
 183 highlight on a single sphere (Sec. 3.2), reconstruction of lper-pixel light direction by interpola-  
 184 tion of results on multiple detected highlights (Sec. 3.3), light intensity correction by exploiting  
 185 interpolated directions and measures on a matte planar target (Sec. 3.4), storage of the calibrated  
 186 per-pixel information in a 3D appearance profile array (Sec. 3.5), and, finally, basic processing  
 187 of appearance profile data to recover shape and reflectance parameters (Sec. 3.6).

### 188 3.1. Preprocessing

189 Image preprocessing consists mainly in the removal of ambient light and undistortion. These  
 190 two transformations are applied to all the images in the collection before they are fed to the light

191 direction estimation step. The ambient light is captured by acquiring an extra image of the scene  
 192 with the handheld light source turned off. The undistortion is performed according to the intrinsic  
 193 camera parameters estimated a priori with a standard calibration procedure. Auxiliary to the  
 194 annotation of the black reflective spheres and the white Lambertian frame, stand the maximum  
 195 image and respectively, the minimum image estimation. The maximum of the image collection  
 196 discards the shadow around the spheres, hence improving the visual acuity, while the minimum  
 197 image maps the projected shadows areas that should be avoided when selecting consistent highly  
 198 reflective regions (Fig. 3).

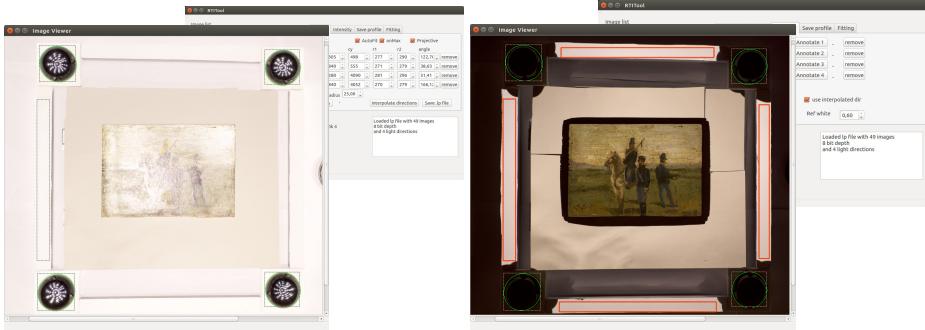


Figure 3: Snapshots of the RTITool interface during the annotation of reflective spheres (left) over the maximum image estimated from the MLIC stack, and the annotation of the Lambertian frame performed on the minimum image estimates from the stack (right).

### 199 3.2. Perspective Light direction estimation

200 In the general case of free-form RTI acquisition without known lights, we compute the  
 201 highlight-based light direction by releasing the orthographic projection hypothesis used in  
 202 previous classic solutions [3] and implemented in the well-known RTIBuilder package. This allows  
 203 the computation of light direction when the reflective spheres are at the margin of the image and  
 204 appear relevantly distorted (elliptical) in the image.

205 In the current algorithm and implementation, we assume known intrinsic parameters of the  
 206 camera: optical center,  $\vec{o}_x, \vec{o}_y$ , pixel size,  $\vec{s}_x, \vec{s}_y$ , and focal length  $f$ . They are loaded from files  
 207 in the software tool. However, if we have a scene with multiple reflective spheres, we could, in  
 208 principle, exploit them also to calibrate the camera including distortion parameters [40]. We plan  
 209 to include this feature in future version of the package.

210 Once we have identified the projected sphere outline, that is an ellipse, we can easily locate  
 211 the extrema of the major axes, with known coordinates in camera frame  $\vec{p} = ((p_x - o_x)s_x, (p_y -$   
 212  $o_y)s_y, f)$  and  $\vec{q}$  (Fig. 4). Note that the knowledge of the pixel size  $\vec{s}_x, \vec{s}_y$  is not necessary. We can  
 213 only add knowledge of aspect ratio  $s$  to the focal length expressed in pixels. From  $\vec{p}$  and  $\vec{q}$ , we  
 214 can easily compute the direction of the vectors  $\vec{a}, \vec{b}$  pointing to the corresponding tangent points  
 215 on the sphere  $\vec{P}, \vec{Q}$

$$\vec{a} = (\vec{p} - \vec{O}) / \|\vec{p} - \vec{O}\|$$

$$\vec{b} = (\vec{q} - \vec{O}) / \|\vec{q} - \vec{O}\|$$

216 217 This also allows us to estimate the unit vector  $\vec{w}$  pointing to the center of the sphere  $\vec{C}$ :

$$\vec{w} = (\vec{a} + \vec{b}) / \|\vec{a} + \vec{b}\|$$

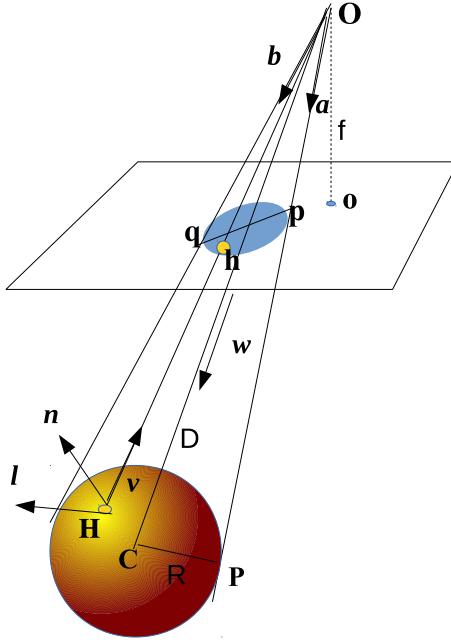


Figure 4: Light direction can be easily computed also assuming releasing the orthographic projection constraint.

218     The triangle  $\vec{O}\vec{C}\vec{P}$  has known angles  $\arcsin(\vec{d} \cdot \vec{w})$ ,  $\arccos(\vec{d} \cdot \vec{w})$ ,  $\pi/2$  and a known side  $R$ . We  
219     can, thus, estimate the distance  $D$  of the sphere center from the camera center, and the coordinates  
220     of the sphere center in camera coordinates  $C = D\vec{w}$ .

221     Since we have multiple spheres on a plane, we can then estimate the plane orientation/position  
222     from the estimated centers with a simple fit.

223     Once we estimate the position of the projected highlight  $\vec{h}$ , we can solve for the 3D highlight  
224     position and the light direction estimation, by computing the view unit vector

$$\vec{v} = (\vec{d} - \vec{h}) / \|\vec{d} - \vec{h}\|$$

225     and the equation of the line from the origin to the highlight

$$\vec{X} = -t\vec{v}$$

226     Solving the equation system that combines this equation and the sphere equation

$$(X_1 - C_1)^2 + (X_2 - C_2)^2 + (X_3 - C_3)^2 = R^2$$

227     we can find two intersections. The one closest to the origin is the highlight position in 3D  $\vec{H}$ .  
228     The unit normal in the point is then

$$\vec{n} = (\vec{H} - \vec{C}) / \|\vec{H} - \vec{C}\|$$

229     and reflecting  $\vec{v}$  with respect to  $\vec{n}$  we can estimate the light direction  $\vec{l}$ .

230     In our tool, we implemented this algorithm coupled with a simple ellipse detector based on  
231     local image binarization and ellipse fitting obtained with OpenCV implementation of Fitzgib-  
232     bon's method[41].

233 **3.3. Multiple spheres setup and light direction interpolation**

234 By putting a sphere at the margin of the image, we reduce the odds that it casts shadow on the  
 235 object. The perspective model allows us to do this even in non-ideal conditions and wide field of  
 236 views.

237 For most light sources, in addition, the assumptions of a parallel and uniform beam across the  
 238 entire scene is also far from being fulfilled, and errors introduced in this case are not negligible,  
 239 as shown in the experimental section. We, therefore, strive to obtain, when possible, a better  
 240 per-pixel light direction estimation by using multiple (typically four) spheres placed close to  
 241 image corners, estimating directions for the various highlight positions, and linearly interpolating  
 242 estimated light direction across the image.

243 If this configuration is chosen, rather than just estimate and store a light direction for each  
 244 image, we estimate for each image the coefficients of a linear interpolation of the directions that  
 245 are later used to recover per pixel light direction values. Coefficients are saved in our special-  
 246 ized data structure (appearance profile array, APA) and used to support a better estimation of  
 247 reflectance parameters.

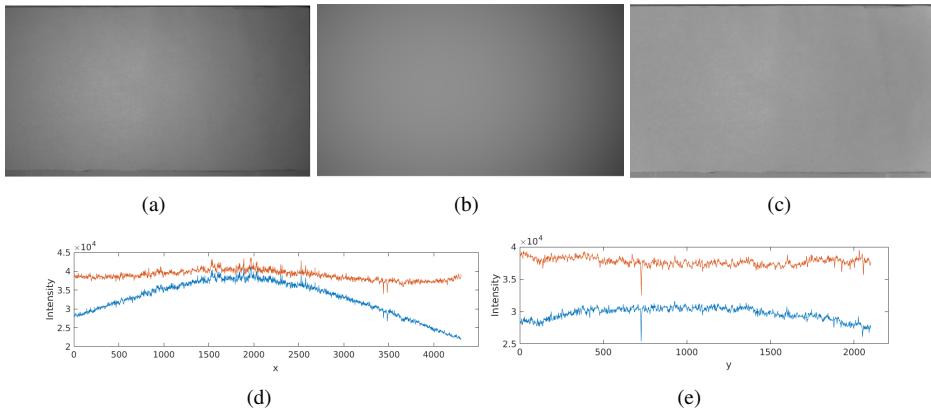


Figure 5: Intensity correction procedure of image with only white background: from the annotated planar frame (not visible in the cropped image (a), a polynomial estimate of the illumination in the whole image is performed (b), and used to estimate a corrected image (c). Intensity profiles along the central line and column in the original and corrected images are compared in (d),(e).

248 **3.4. Light intensity correction**

249 The non-uniformity of the beam intensity can be reasonably corrected with a solution that  
 250 can be applied in many practical acquisition scenarios. The idea, here, is to place a planar frame  
 251 around the object of interest, with an approximately Lambertian coating. By detecting the region  
 252 in the images where the target is illuminated, excluding the parts that can be shadowed in some  
 253 images, we can use the measured pixel values on the target to calibrate the pixel values on the  
 254 acquired objects, in order to simulate an acquisition made with a truly constant light intensity,  
 255 at least on the plane of the frame. Ideally, for a Lambertian surface, the brightness of the region  
 256 should be constant (if the light direction is constant). In practice, we measure a non-negligible  
 257 non-uniformity using common lights and cameras, due to non-uniformity of the light beam, as  
 258 well as to vignetting effects of the lenses.

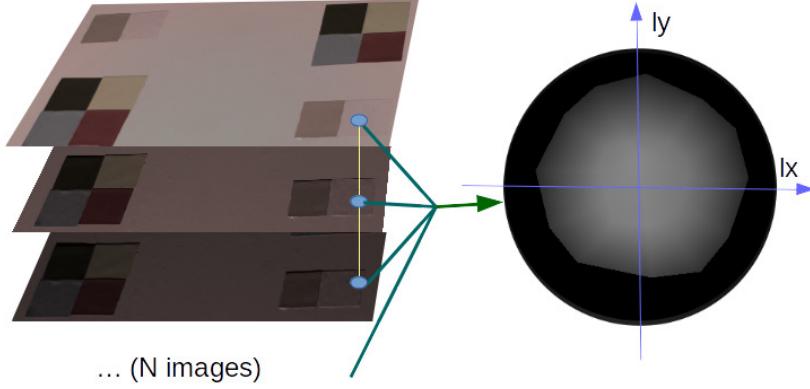


Figure 6: Appearance profile visualization using Delaunay Triangulation based interpolation of the intensity values at the pixel location in the  $l_x, l_y$  plane.

By fitting a polynomial function of the light direction over the brightness estimated in the frame regions, we can compute a correction factor for the estimated light, making the image illuminated as if the light intensity has a standard reference value, and the local light direction estimated at each pixel location (we obtain this by weighting the expected Lambertian reflectance of paper by the actual local cosine factor). This light normalization can correct different non-uniformity causes. Of course it assumes that the light intensity is not changing with depth in the region of interest. Since beam variations are expected to be smooth, we use a quadratic interpolation of the reflectance to extend the reference illumination to the entire plane of interest starting from the reference values on the target. It should be noted that even if the current software fits a quadratic model, a more complex function will be investigated in the future. Fig. 5 shows the effect of the correction procedure in an image with only a planar diffusive surface (spheres and calibration frames were outside the cropped region of interest). The procedure successfully flattens the intensity profiles due to spotlight shape and vignetting.

### 3.5. Appearance profile array files storage

In order to simplify data processing steps, we store the data stack after in a reorganized array structure, where all the per-pixel information is represented sequentially to allow model fitting or pixel processing without the necessity of loading all the data in memory or to allocate large array in processing software. The file structure used (appearance profile array, APA) is composed of a header and data section. The header describes the encoding choices (8 or 16 bits, RGB or chromaticity+luminance, constant or interpolated light directions) and the light direction information (vector elements or interpolation coefficients). The data section stores pixel values in a 3D array. Fig. 6 shows the information encoded in appearance profile: all the brightness information of a pixel location is stored together and can be represented in  $l_x, l_y$  coordinates and interpolated for a better display. The shape of the resulting function is characteristic of both shape and material properties. We tested both Delaunay Triangulation based interpolation and Radial basis functions to obtain visualizations of the local appearance map. Using these interpolation algorithms, relighted images can thus also be directly displayed without the need for simplified parametric representations of the local reflectance as a function of light direction.

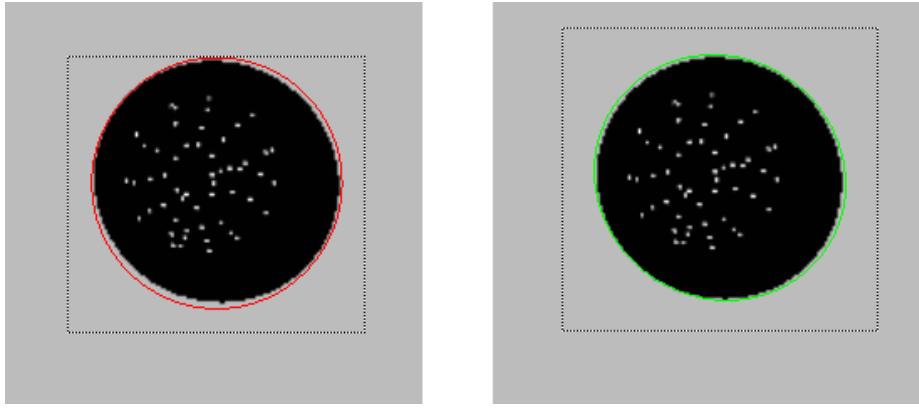


Figure 7: Annotation of reflective spheres on a synthetic dataset. With the orthographic assumption, automatic segmentation is not accurate and annotated circles cannot, in any case, match exactly the real object contours (left). With the perspective mode the segmentation is more accurate, and this results in a quite higher accuracy of the estimated light direction, as shown in Sec.5.1

**287 3.6. MLIC data basic processing: Photometric stereo, PTM/HSH fitting**

**288** Apart from creating relighted images with interpolation in the light direction space, the MLIC  
**289** stacks encoded as (intensity-corrected) appearance profiles can be processed with the standard  
**290** algorithms used to recover shape and reflectance parameters. Given for each pixel the light  
**291** direction  $(l_x(i), l_y(i), l_z(i))$  and the (corrected) reflectance  $L(i)$  known for  $N$  light directions  $\vec{l}(i)$ ,  
**292** basic Photometric Stereo estimates albedo  $a$  and normals  $\vec{n}$  assuming the Lambertian model and  
**293** solving the overconstrained system

$$L(i) = a [n_x, n_y, n_z] [l_x(i), l_y(i), l_z(i)]^T \quad i = 1 \dots N \quad (1)$$

**294** PTM fitting approximates the reflectance function with a polynomial function, also using a  
**295** least squares solution to find coefficients. The classical form [10] is

$$L(i) = [a, b, c, d, e, f] [l_x^2(i), l_y^2(i), l_x(i)l_y(i), l_x(i), l_y(i), 1]^T \quad (2)$$

**296** but different polynomial function have been proposed, as well as different fitting functions,  
**297** such as Hemispherical Harmonics or Discrete Modal Decomposition [12, 13]. Implementing  
**298** different function fitting is quite simple, and their ability to represent the real reflective behavior  
**299** of the material depends clearly on the kind of material analyzed.

**300** Furthermore, it must be considered that non-local effects, such as interreflections and pro-  
**301** jected shadows, create local anomalous behaviors of the laws directly linking light angle and  
**302** reflected color. To cope with these effects, and also to separate diffusive behavior from specular  
**303** highlights, robust fitting methods have been proposed [42, 43], trying to remove outliers from  
**304** the parameters estimation procedure.

**305 4. Simple tools for RTI data processing**

**306** We designed two software tools to process image stacks captured by a camera in RTI settings.

307        The first tool, RTITool, is aimed at performing all the preprocessing steps to transform ac-  
308        quired images to appearance profile array data cropped in the region of interest, prepared so that  
309        they can be used easily to estimate normals, relightable images and feature maps both with our  
310        own tools or other photometric stereo and RTI fitters. RTItool takes as input image and calibra-  
311        tion information, and is able to perform all the calibration steps described in the previous section  
312        to cope with the difficulties of free-form acquisitions.

313        The second tool, APTool, is aimed at processing appearance profile array data using different  
314        algorithms, creating and exporting albedo, normal maps, relightable RTI files (e.g., PTM files),  
315        as well as displaying derived images (multi-light image enhancements, relighted images from  
316        novel illumination directions) on a canvas window. Both tools are still a work in progress, but  
317        current versions, with all the capabilities described in the paper are available at the web site  
318        <http://www.andreagiachetti.it/rtitools>. Code has been developed on a Linux platform, but, as it  
319        has been realized in C++ using Qt and OpenCV libraries, it could be easily ported on a variety  
320        of computing architectures.

321        **4.1. RTITool**

322        This program allows the user to load image sets, both through image lists or through files  
323        with filenames and associated light directions typically used in current RTI tools. Users can then  
324        perform all the processing pipeline with various options, working with both 8-bit and 16-bit depth  
325        images and generating APA files for the entire image size or cropped regions. The interface is  
326        designed to simplify all the annotating tasks. For example, in order to easily annotate reflective  
327        spheres, annotation and automatic fitting algorithms are by default done on the maximum image,  
328        showing the highest luminance pixels over the stack, removing shadows and evidencing the black  
329        object (Fig. 3, left). In the same way, the annotation of the white frame is done by showing the  
330        minimum image, displaying the lowest per-pixel luminance, so as to easily avoid annotating  
331        regions that can be shadowed from some light directions (Fig. 3, right).

332        The annotation of reflective spheres is semi-automatic. The user is asked to draw a rectangle  
333        including each sphere image. The circles (in case of orthographic assumption) or the ellipses (in  
334        case of perspective) are automatically estimated and drawn. Users can also visually refine the  
335        segmentation by interactively changing the curve parameters on the interface. Fig. 7 shows the  
336        inaccuracy of classical circular annotation (left), fixed on the same image by the ellipse fitting.  
337        In both cases, light directions can be estimated and stored. Note that even an apparently small  
338        deviation from the orthographic model, as the one shown in the figure, may result in an increase  
339        of one order of magnitude of error in light direction estimation (see Sec. 5.1).

340        **4.2. APTool**

341        The processing of the raw MLIC stack performed with the RTITool ends by the storage of  
342        the data structure allowing the sequential processing of pixel information (light directions and  
343        associated corrected or non corrected intensity values). This information can be used to estimate  
344        normals and albedo using photometric stereo, creating novel relighted or enhanced images by  
345        interpolating or mixing the different pixel values, fitting reflectance models storing relightable  
346        images like PTM or HSH standard files, and more. We developed, for these purposes, a second  
347        software tool called APTool, which loads preprocessed arrays and allows the generation of nor-  
348        mal and albedo maps derived from PS or the estimation of PTM coefficients. Robust versions of  
349        the fitters are also available. The idea is to include, in the future, different fitting and visualiza-  
350        tion algorithms to the software in order to support different kinds of end-user applications. Apart

351 from fitting models and saving classical RTI files, the tool currently allows direct visualization  
 352 of relighted images given a novel light direction, through direct interpolation of samples based  
 353 on radial basis functions (Fig. 8). By selecting image locations (single points or rectangular  
 354 regions), it is also possible to visualize a 2D intensity map represented in  $l_x, l_y$  space of the local  
 355 appearance profile, obtained by scattered data interpolation of the known samples (Fig. 8(b)).  
 356 We have experimented with a Delaunay triangulation based and a Radial Basis Function interpo-  
 357 lation of the samples, and provide an RBF implementation in the delivered tool.

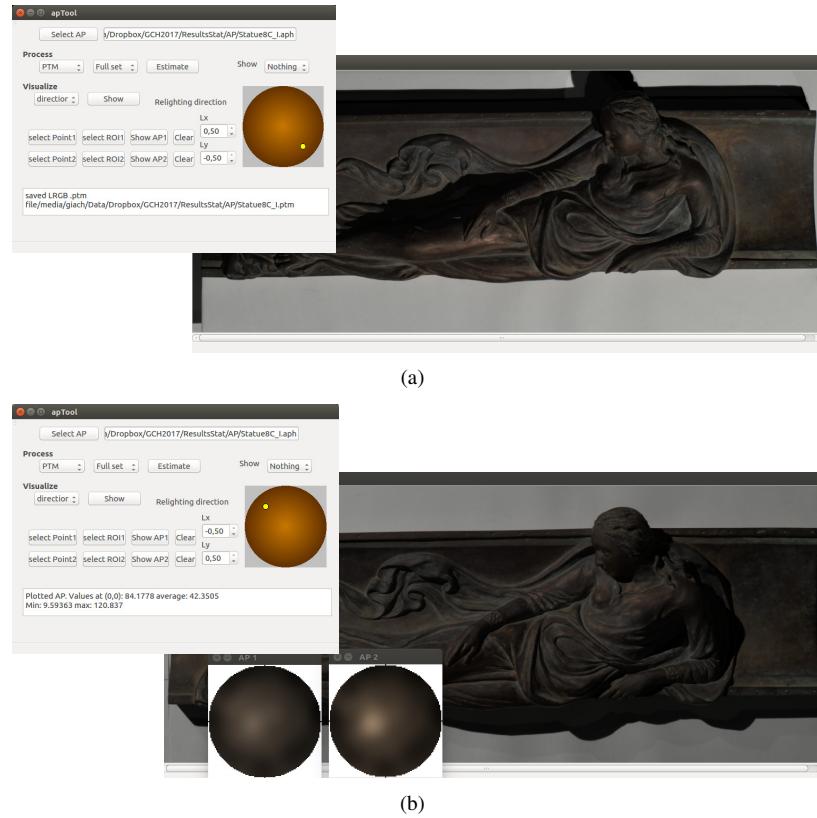


Figure 8: APTool used to show directly a Radial Basis function interpolation simulating direct relighting from ( $l_x = 0.5, l_y = -0.5$ ) (a) and from ( $l_x = -0.5, l_y = 0.5$ ) (b). In the second case, we also selected and compared AP profiles in selected image points.

## 358 5. Experimental results

359 In order to demonstrate the usability of our pipeline and the effects of new algorithms, we  
 360 performed a series of experiments covering different kinds of MLIC capture and processing.  
 361 Our tests include both synthetic datasets and real-world ones. The real-world experiments were  
 362 performed using a DSLR Nikon D810 camera with an architecture based on a CMOS sensor with  
 363 removed IR cut-off filter. The size of the sensor is 36x24mm and the spatial resolution of the full  
 364 format image area is 36MP. To the digital camera a full frame AF-S FX Nikkor 50mm f/1.8G lens

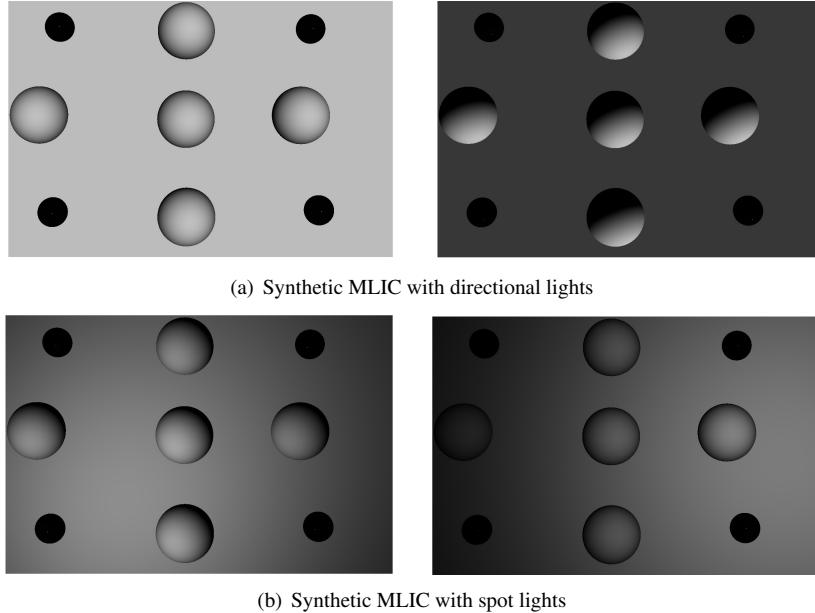


Figure 9: (a) Two images of a synthetic dataset simulating a white plane with some large bumps and 4 reflective spheres, acquired by a fixed camera under different parallel and constant illumination. (b) Two images of a synthetic dataset with the same geometry, but illuminated by simulated spot lights.

365 was attached. As in this paper we present results using visible light, an IDAS-UIBAR III optical  
 366 filter was used to gather only the signal from the visible range of the electromagnetic spectrum.  
 367 The sensor of the digital camera was checked for linearity, by taking images covering a wide  
 368 range of exposures, from very low to very high and then plotting the brightness as a function  
 369 of exposure. The camera was geometrically calibrated by computing the intrinsic parameters,  
 370 two radial and two tangential distortion coefficients with the GML Camera Calibration Toolbox.  
 371 However, any other calibration tool can be used for this purpose.

372 *5.1. Accuracy of light direction estimation*

373 In order to evaluate the errors in light direction estimation when the orthographic camera  
 374 model is not perfectly followed, we created a synthetic RTI dataset by rendering a scene with 4  
 375 reflective spheres near image corners, placed on top of a white Lambertian surface not exactly  
 376 perpendicular to the camera axis, and with some spherical bumps, illuminated with perfectly  
 377 parallel rays along 50 known directions (Fig. 9 a) or with the same number of simulated spot  
 378 lights (Fig. 9 b). Using RTItool, we annotated the elliptic sphere profiles and estimated the light  
 379 directions at each sphere position as described in Sec. 3.2. We compared the results with those  
 380 obtained with our tool in the orthographic approximation, by annotating the circle circumscribed  
 381 to the ellipse. We also compared our results with those obtained similarly with the widely used  
 382 RTIBuilder package [21]. Note that the circular annotation cannot be precise, as the sphere mask  
 383 is actually elliptic, as shown in Fig. 7, and this happens in most real images.

384 A comparison of the errors obtained (difference between the average of the four sphere es-  
 385 timations and ground truth) reveals that, despite the limited eccentricity of the ellipses, with the

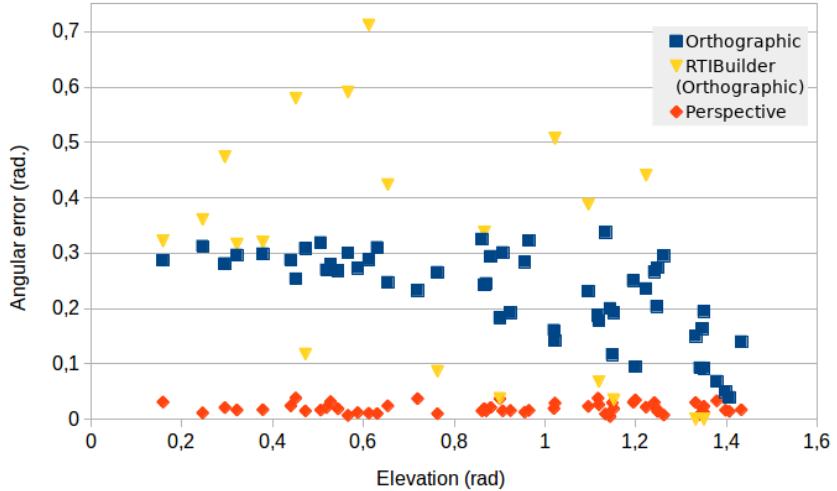


Figure 10: Angular errors due to wrong orthogonal camera assumption can be quite relevant. The use of a perspective model makes the values quite accurate.

386 perspective model we have the error reduced by an order of magnitude. The average errors for  
 387 the 50 directions of Fig. 9 a are, in fact, 0.02 radians for the perspective estimation, 0.21 for the  
 388 orthographic estimation done with RTITool, and 0.23 for the orthographic estimation done with  
 389 RTIBuilder. Fig. 10 shows errors for each of the 50 single light directions sorted by elevation.

390 The local accuracy of light direction estimation is then improved by estimating local values  
 391 by interpolating the values of the sphere at the corners. In order to show the amount of error  
 392 reduction, we conducted two experiments. First, we created another synthetic dataset similar to  
 393 the previous one, but where the images are illuminated with simulated spot lights approximately  
 394 pointed towards the center of the target. Light direction and intensity for each pixel are thus not  
 395 uniform, as in most typical real-world scenarios. In this case, the per pixel average error in the  
 396 constant estimation (average of the values of the four spheres) is significantly higher than the  
 397 error value coming from interpolation. Fig. 11 shows the average errors for the single images  
 398 plotted versus elevation angle of the spotlight orientation. With interpolation, the error, averaged  
 399 on all pixels of all images, is reduced from 0.17 to 0.05 degrees.

400 We also performed experiments on a real acquisition of a calibration target. In this case, we  
 401 captured images of a flat plane perpendicular to the camera axis, putting four reflective spheres at  
 402 the corners of the image area and a fifth in the center. In the set of images captured, the average  
 403 difference between the direction measured in the top left corner and the one measured in the  
 404 image center was 0.146 radians. Fig. 12 shows that, for small elevation angles, the error is higher  
 405 due to the larger effect of quantization error in highlight localization. If we estimate a per-pixel  
 406 linear interpolation of light direction, we reduce the average error to 0.121 radians.

#### 407 5.2. Accuracy of normal reconstruction

408 The improved per-pixel light direction estimation and the procedure to correct illumination  
 409 increase the quality of the MLIC-based reconstructions, as demonstrated, for instance, by our  
 410 tests with Photometric Stereo and normal estimation. Using the same simulated dataset with

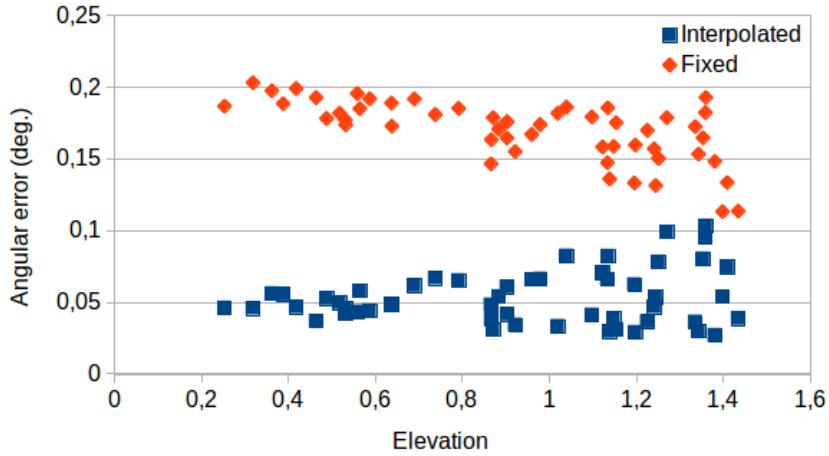


Figure 11: Angular errors of light direction estimated on a sphere with respect to real one in the simulated "spotlight" dataset of Fig. 9 b, plotted against elevation of the principal ground truth spotlight direction. Interpolation strongly increase the average accuracy of the per-pixel estimation.

	Angular err (rad)	Std. Dev
Single Directions, no light correction	0,482	0,271
Interpolated directions, no light correction	0,417	0,249
Single Directions, light correction	0,262	0,188
Interpolated directions, light correction	0,252	0,183

Table 1: Angular errors reduction in per-pixel normal estimation with Photometric Stereo on the synthetic "spotlight" dataset of Fig. 9 b using interpolated light direction estimation and intensity correction.

411 spotlight illumination, we estimated surface normals (and albedo) by solving the classical least-  
 412 squares problem under the assumption of Lambertian surfaces. We then compared per-pixel re-  
 413 constructed normals with the ground truth values. Table 5.1 shows that the average angular error  
 414 is strongly reduced both by the light direction interpolation and the light intensity correction.

415 The effects of light correction can be also appreciated when reconstructing normals of chal-  
 416 lenging real-world objects from MLICs using Photometric Stereo. To show this, we acquired  
 417 images of a set of coins placed over a flat background. We used our reconstruction pipeline and  
 418 the RTITool to recover the appearance profile arrays and then used our APTTool to reconstruct  
 419 normal maps.

420 The set is composed by a bronze Roman coin (*quadrans*) dated 9 B.C. and damaged by  
 421 scratches, and two 10 cent Italian coins. One exemplar, dated 1931 is made of copper and is  
 422 severely degraded, while the second exemplar, dated 1939, is made of a special alloy with nickel,  
 423 called Bronzital, which has been used to improve corrosion resistance.

424 Normal maps obtained with Photometric Stereo have been compared with an (approximate)  
 425 reference solution derived from a high resolution 3D reconstruction of the same coins made  
 426 with an optical microprofilometer based on conoscopic holography [44]. This device is able  
 427 to capture reliable profilometric measurements down to the scale of micron on different kinds  
 428 of materials, reflective or diffusive. Our microprofilometer is based on an Optimet conoscopic

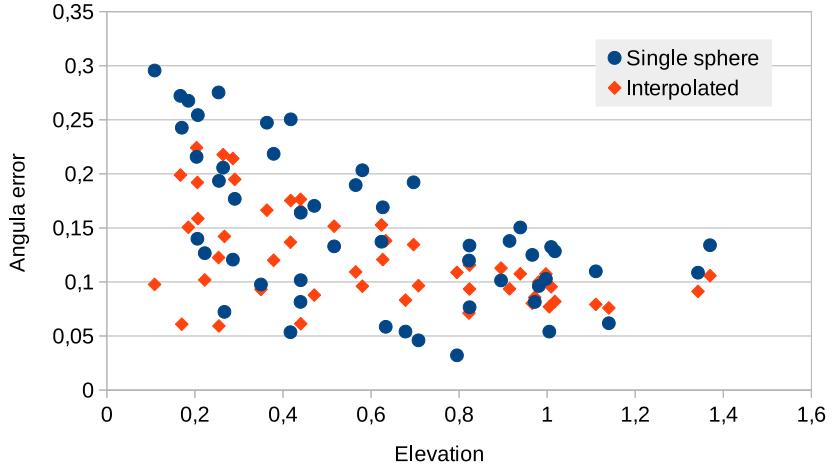


Figure 12: Angular difference of the direction estimated on a sphere near the corner with respect to the one estimate on a sphere near the image center (blue dots), plotted as a function of elevation for a complete MLIC scan (50 photos). Replacing the corner estimate with the linear interpolation of the four corner values in the central position, we can get reduced errors (red squares).

429 probe mounted on sub-micrometric linear stages in order to scan a region up to  $30 \times 30 \text{ cm}^2$  in one  
 430 session. Reference coin models have been reconstructed with a transversal resolution (XY grid)  
 431 of 50 microns.

432 Depth maps derived from these models were finally registered with the estimated RTI normal  
 433 maps using a similarity transform optimized to match the correspondence of manually selected  
 434 points (12 landmarks). This initial registration was then refined by locally optimizing mutual  
 435 information in image space.

436 Fig. 13 shows the three coins and the related differences between RTI-based normal maps and  
 437 the reference normal maps estimated from microprofilometric data, both in case of non-corrected  
 438 image brightness, and with the light correction procedure described in Sec. 3.4. It is evident  
 439 that light correction sensibly improves the reconstruction quality, as quantitatively reported in  
 440 Table 5.2. The light correction procedure reduces the median errors, on average, by 27%.

	Median angular distance (rad.)	
	Non-corrected	Corrected
Bronzital 10c	0.117	0.079
Copper 10c	0.068	0.053
Quadrans	0.171	0.108

Table 2: Median angular distances of the RTI estimated normals from the reference microprofilometer normals. The calibration procedure reduces the errors on average of 27 percent.

### 441 5.3. Recovery of reflectance properties of materials

442 Our intensity correction methods are also important to better recover material properties. To  
 443 demonstrate this fact, we placed a matte paper target with different albedo regions in different

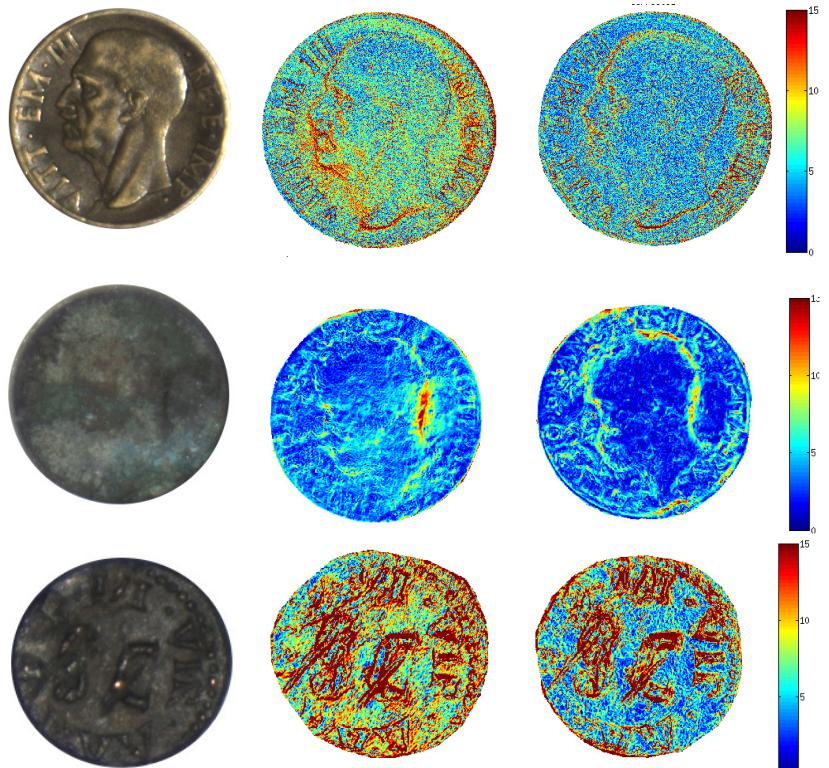


Figure 13: **Angular errors.** Color-coded angular errors (degree) of RTI estimated normals wrt ground truth from microprofilometry. Left: non-calibrated results. Right: results with light calibration.

444 positions on a flat planar background (Fig. 14).



Figure 14: Three images of a MLIC capture of a planar surface with flat paper targets with different albedo regions.

445 If we visualize the interpolated appearance profiles estimated on a pixel in a selected region,  
446 in this case with flat perpendicular surface and approximately Lambertian behavior, we should  
447 see a function that, represented in  $l_x, l_y$  coordinates, should present a regular and symmetric  
448 function. Fig. 15 shows that plots of interpolated appearance profiles on non-corrected images are  
449 not symmetric and different in different regions of the same material if images are not corrected  
450 with our procedure. Conversely, light correction results in profiles similar to those expected and  
451 similar in different parts of the image where the material is the same.

452 This effect can be quantitatively measured by evaluating the average albedo of the patches of

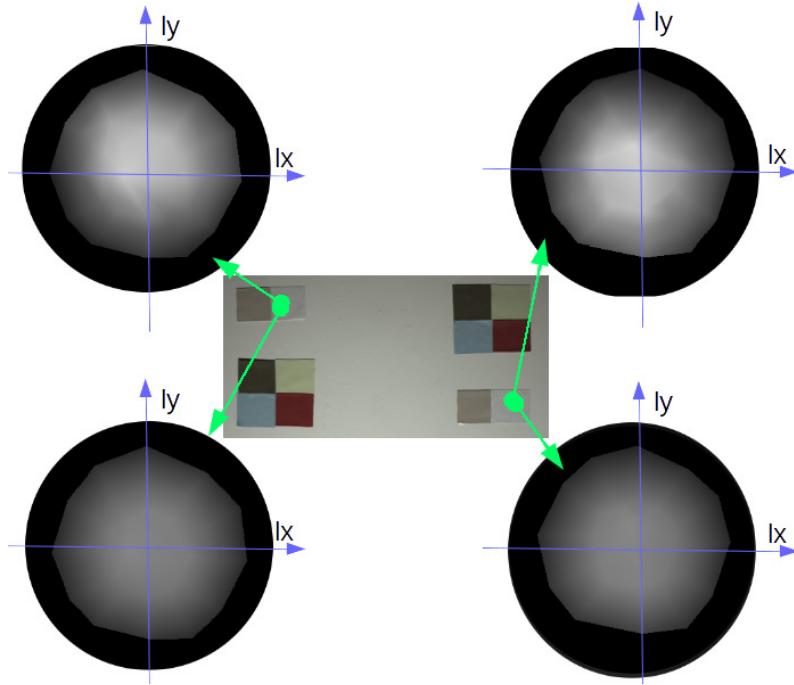


Figure 15: Appearance profiles corresponding to the same material on flat patches are similar and regularly shaped when computed on light-corrected images (bottom), while are irregular when lighting is not uniformed (top)

453 the same paper types put in two different positions in the scene of Fig. 14. Without corrections,  
 454 the albedo of paper patches of the same type placed in different image position differs up to 7%.  
 455 The difference is strongly reduced with the light direction and intensity correction procedure, as  
 456 shown in Table 5.3.

457 By matching the reflectance of the Lambertian frame to a reference value, we can also esti-  
 458 mate the consistency of albedo measurements among different image captures. Table 5.3 shows  
 459 that the measurements obtained in a second acquisition are largely different (often more than  
 460 30% of the value), even if a similar protocol and the same light source has been used. It is ac-  
 461 tually sufficient to change the distance of the source to have different results. However, the use  
 462 of the correction procedure results in similar albedo values (difference lower than 2.5% of the  
 463 value)

464 Another important effect of the light correction procedure is the repeatability of reflectance  
 465 parameters estimation in different MLIC captures without light calibration. Table 5.3 shows  
 466 the albedo of the same paper types of the previous experiment estimated on a different MLIC  
 467 capture of a plane with paper targets glued on it. Without light correction, the light intensity is  
 468 quite different, even if we tried to use a similar configuration. Clearly a small difference in light  
 469 distance results in different illumination and estimated albedo. The light correction procedure,  
 470 by contrast, makes the estimated parameters similar.

471 Processed RTI data is often used to segment different materials not easily recognized in color  
 472 images [18]. Such kind of results can be improved by our light correction procedures. To show  
 473 this, we have performed two RTI acquisitions of a polished silver sample partially covered by

		Albedo Pos. 1	std	Albedo Pos.2	std	diff/mean
Paper1	corrected	0,557	0,003	0,559	0,004	0,39%
	non-corrected	0,726	0,009	0,708	0,009	2,53%
Paper2	corrected	0,406	0,003	0,411	0,003	1,38%
	non-corrected	0,515	0,007	0,539	0,018	4,54%
Paper3	corrected	0,391	0,003	0,391	0,005	0,17%
	non-corrected	0,501	0,009	0,514	0,007	2,73%
Paper4	corrected	0,157	0,004	0,156	0,003	0,55%
	non-corrected	0,206	0,005	0,203	0,006	1,30%
Paper5	corrected	0,551	0,003	0,545	0,005	1,08%
	non-corrected	0,734	0,008	0,685	0,010	6,96%
Paper6	corrected	0,138	0,006	0,135	0,003	2,10%
	non-corrected	0,184	0,004	0,173	0,007	6,18%

Table 3: Albedo measured on planar patches of the same material can be quite different in different image regions if estimated with classic photometric stereo on non-corrected multi-light image stacks. Our brightness and light-direction correction procedures clearly result in more consistent values.

474 a coating, see Fig. 16(a), and applied unsupervised classification to segment regions with or  
 475 without coating. For each RTI sample, we compute a 7-dimensional descriptor of a 30 pixels  
 476 neighborhood. The descriptor is the average albedo value, to account for material color, plus the  
 477 6 standard deviations of the standard RTI polynomial coefficients, to account for the roughness  
 478 of the sample surface.

479 Unsupervised classification is achieved by performing two-class k-means clustering. We  
 480 measured the similarity of the classification outcomes obtained from the two different acquisi-  
 481 tions, without and with light calibration, see Fig. 16. The only difference between the two ac-  
 482 quisitions of the same sample is the different lighting pattern caused by the free-form approach.  
 483 The coated area is in red, while the uncoated area is in green. In the absence of light calibration,  
 484 the clustering outcome is unstable, as it has only a 20% overlap, while, by performing light cal-  
 485 bration, we improve it up to 99.5% of pixels that have been assigned to the same class, showing  
 486 that with our approach free-form RTI can be used for surface characterization.

#### 487 5.4. Visual analysis of RTI enhancements

488 The typical use of MLIC data done in the Cultural Heritage domain consists in estimating  
 489 relightable images and analyzing them to improve the visualization of object details. To sim-  
 490 ulate this application, we created a mock-up of a complex structure with fine relief details by  
 491 imprinting a leaf on modeling paste, then acquiring the photos with our pipeline supporting light  
 492 correction. We exported both corrected and non corrected appearance profiles with RTITool and  
 493 estimated and exported PTM files with APTool. The files have been analyzed with RTIViewer  
 494 to visualize interesting detail [21]. Fig. 17 shows a detail of a relighted image with the specular  
 495 enhancement proposed in [10]. The result on top right is obtained from the non-corrected data,  
 496 while the one on the bottom right is obtained with the corrected pipeline. In the corrected images,  
 497 it is possible to appreciate a better enhancement of detail and a clearer visualization of nervatures  
 498 and scratches, hardly visible on the uncorrected image.

499 This effect is even more visible in the example of Fig. 18, where specular enhancements ob-  
 500 tained from PTM fitting of the non-corrected and corrected appearance profiles derived from the

		Albedo MLIC 1	Albedo MLIC 2	diff/mean
Paper1	corrected	0,558	0,548	1,81%
	non-corrected	0,717	0,518	32,25%
Paper2	corrected	0,409	0,404	1,18%
	non-corrected	0,527	0,393	29,23%
Paper3	corrected	0,391	0,385	1,60%
	non-corrected	0,508	0,381	28,45%
Paper4	corrected	0,156	0,156	0,23%
	non-corrected	0,205	0,155	27,64%
Paper5	corrected	0,548	0,536	2,22%
	non-corrected	0,710	0,510	32,77%
Paper6	corrected	0,136	0,139	2,06%
	non-corrected	0,178	0,135	27,84%

Table 4: Albedo values measured on a different acquisition of the same material patches of Table 3. The correction procedure results in similar values for similar materials.

501 acquisition shown in Fig. 2(b) are compared. The correction leads to a much better visualization  
 502 of brush strokes.

503 Looking at PTM-based relighting, it is interesting to note that even if light calibration ensure  
 504 a better quality of enhancements due to the improved normals, the removal of specular compo-  
 505 nents results in loss of possibly relevant information about the imaged object. This can be seen  
 506 comparing relighted PTMs with corresponding relighted APA visualized with our tool. Fig.19  
 507 shows this on the painting detail. PTM-based relighting represents similarly regions where the  
 508 surface has different specular behavior and the perception of depth is reduced by the absence of  
 509 specular effects, visible in RBF interpolation.

510 We plan, therefore, to investigate on possible improvements of interactive direct visualiza-  
 511 tion of APA information and on the development of novel enhancement methods that can be  
 512 directly implemented in the APTool to allow a better visual analysis of the information hidden  
 513 in RTI stacks. PTM or HSH encodings are useful as they allow compact storage of relightable  
 514 images, but, imposing a drastically simplified reflectance model discarding relevant information,  
 515 they may result in information loss that may create serious problems to the subsequent surface  
 516 analysis. Our plan is to use smart compression techniques to obtain a compact representation of  
 517 the full APA information allowing an easier handling and more efficient direct visualization.

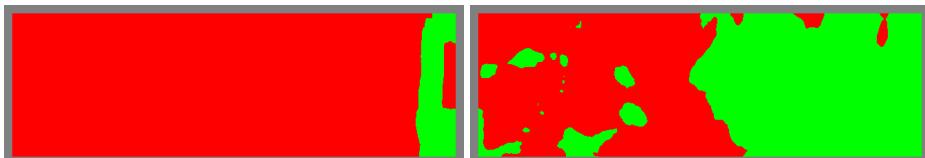
## 518 6. Discussion

519 Highlight RTI is quite popular, especially in the Cultural Heritage domain, to the point that it  
 520 may be considered one of the most successful computational photography techniques in that do-  
 521 main. It can be realized with a simple camera, a simple light source and one (or more) reflective  
 522 spheres. However, the framework commonly used for this task has some limitations, and this can  
 523 result in a low degree of repeatability of measurements, as well as in a poor quality of extracted  
 524 information, leading, in some cases, to the impossibility of effectively using the technique.

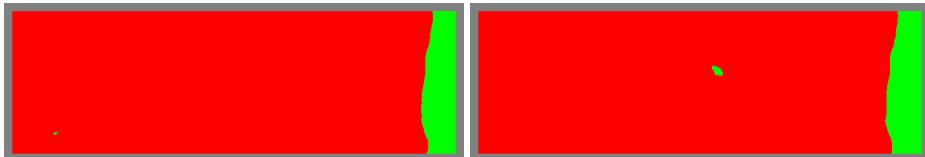
525 In this article, we have shown that, with slight modifications of the standard acquisition  
 526 setup, it is possible to significantly improve the quality of the fusion of a multi-light image



(a) Partially coated silver sample



(b) Clustering results without light correction



(c) Clustering results with light correction

**Figure 16: Unsupervised classification.** In (b),(c), two class k-means clustering (left/right) applied to two different acquisitions of the polished sample in (a) are represented. The coated area is in red while the uncoated area is in green. Without calibration we have a)20% of classification similarity, while we obtain a value of 99.5% by using the calibrated images. This shows the drastically increased level of repeatability of the proposed pipeline with respect to classic free-form RTI.

527 collection, achieving a better reconstruction of shape and material properties of the scene, as  
 528 well as an improved quality of relightable images. Our approach realizes a sort of integration  
 529 of the classic H-RTI technique, usually based on uncalibrated lights and qualitative analysis,  
 530 with the Material Capture and Photometric Stereo approaches targeted at accurate shape (and  
 531 reflectance) reconstruction, but usually requiring very high-density acquisitions and/or light and  
 532 camera calibration.

533 As with all practical setups, the proposed approach has also some limitations. First of all, the  
 534 necessity of placing more targets near the object, and the fact that we assume that the object to be  
 535 imaged is mostly planar. The latter assumption is, however, typically true in H-RTI applications,  
 536 and can be resolved with the same iterative techniques applied in PS settings. Moreover, in our  
 537 current implementation, using our custom designed frame with the four spheres and the coated  
 538 aluminum bars, the size of the object to be captured is limited to a range from about 50x50  
 539 cm to 1mx1m. For larger sizes, the placement of co-planar Lambertian targets to estimate the  
 540 correction may be difficult in on-site acquisitions. We are investigating, however, different light  
 541 correction methods that may take into account depth variations of the illumination. We are also  
 542 investigating improved interpolation methods tuned for standard spot lights.

543 Our current work focuses on the finalization and testing of our processing tools, that will be  
 544 freely available for the scientific community.

545 We are also investigating novel techniques for shape and material reconstruction, as well for

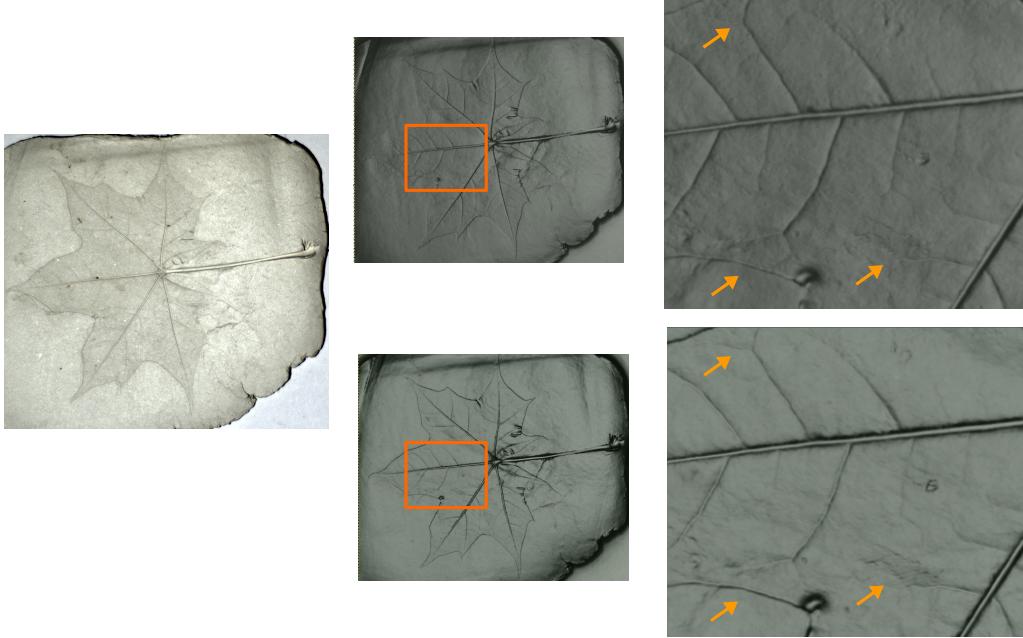


Figure 17: Detail of relighting with specular enhancements of a captured mock up representing a leaf with small imprinted details. Using RTIViewer with the same parameters, the result obtained with the PTM files estimated using corrected images (bottom) allows a better perception of small details.

546 feature detection from MLIC. A challenging problem is, for example, the development of robust  
 547 fitting techniques able to recover material reflectance information independently from shape.  
 548 Apart from the difficulty in modeling reflectance, releasing hypotheses of Lambertian behavior,  
 549 it is also necessary to consider that pixel information is not always depending only on local shape  
 550 and reflectance, but also to global effects like inter-reflections and projected shadows. The use  
 551 of classic outlier removal procedures, proposed in previous works [43], may be problematic due  
 552 to the relatively low number of samples and more specific heuristics for outlier rejection may be  
 553 more effective.

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(a) Non corrected

(b) Corrected

Figure 18: Relighting with  $l_x = 0, l_y = 0$  and specular enhancement of PTMs estimated on the non-corrected appearance profile array (a) and corrected appearance profile array (b) coming from the acquisition of Fig. 2. The second one shows much better the relief of the brush strokes and the painting style.



(a)

(b)

Figure 19: Painting details relighted from the same direction ( $l_x = 0, l_y = 0$ ) using (a) interpolation based on PTM coefficients (b) Radial Basis Functions interpolation on APA data. Material differences and relief details are not visible on the PTM visualization while are correctly perceived in the direct relighting.

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