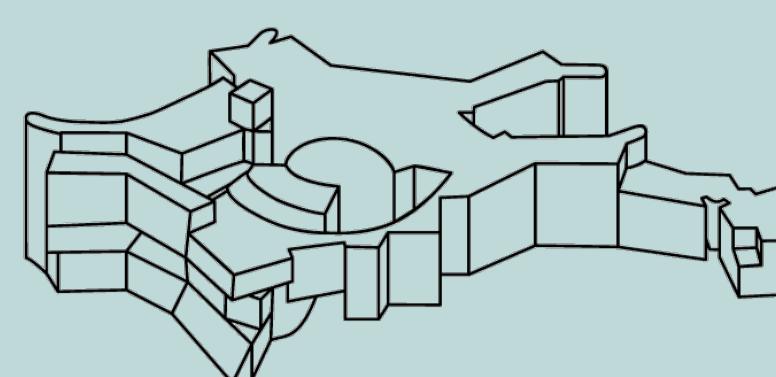


ONe white dwarf mergers with AREPO



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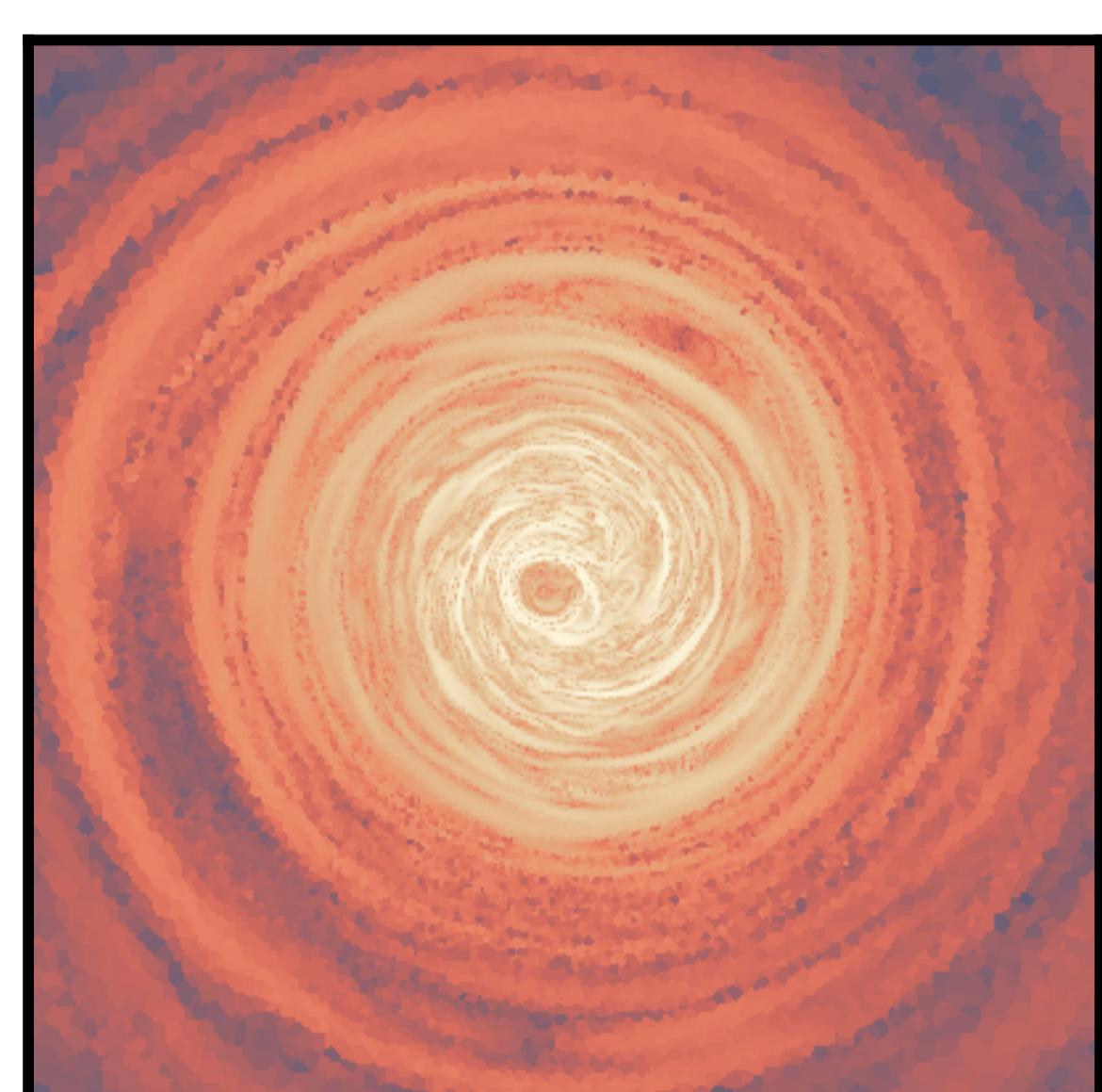
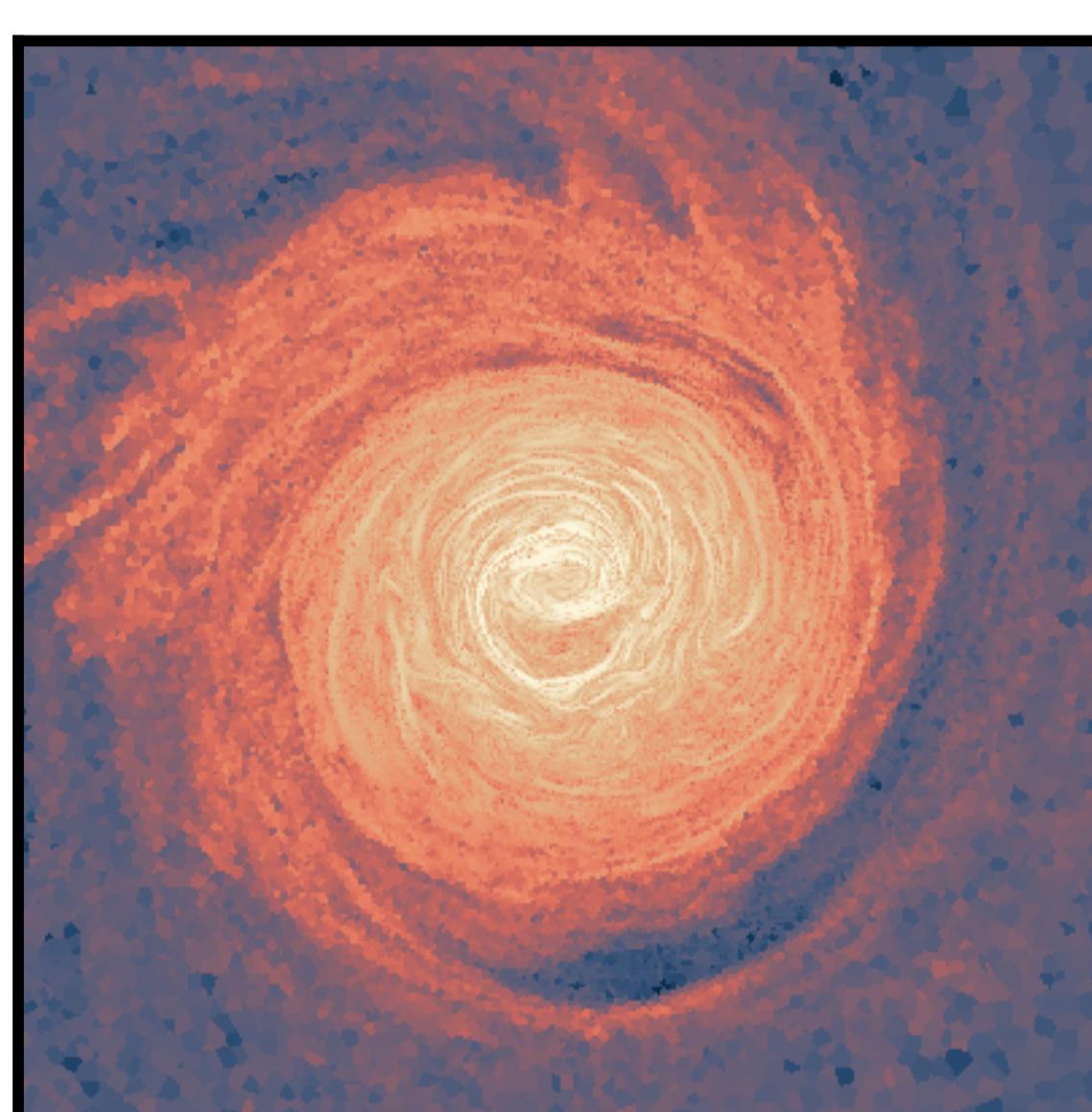
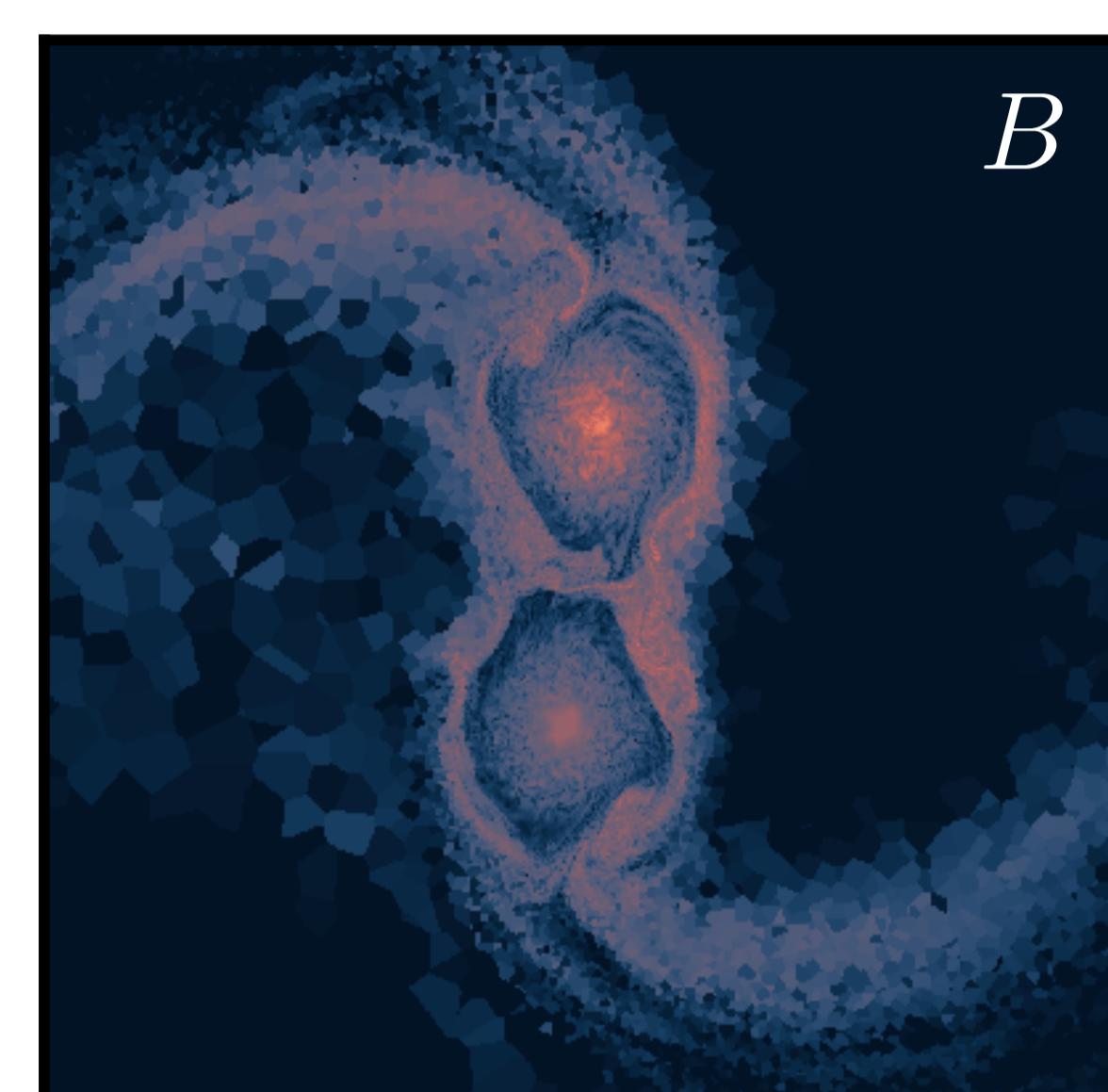
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WORK IN
PROGRESS
Stay
tuned!

B [G]

$10^3 \ 10^4 \ 10^5 \ 10^6 \ 10^7 \ 10^8 \ 10^9 \ 10^{10}$



Aims

We will study a **double ONe WD merger** from immediately before the coalescence up to a few hundred rotations of the product. We are particularly interested in modelling the magnetic field amplification and evolution. This aspect is fundamental, as the product is initially supported by rapid rotation. Understanding the timescale over which the magnetic field removes angular momentum is crucial, as this determines when rotational support is lost and collapse ensues.

Methods

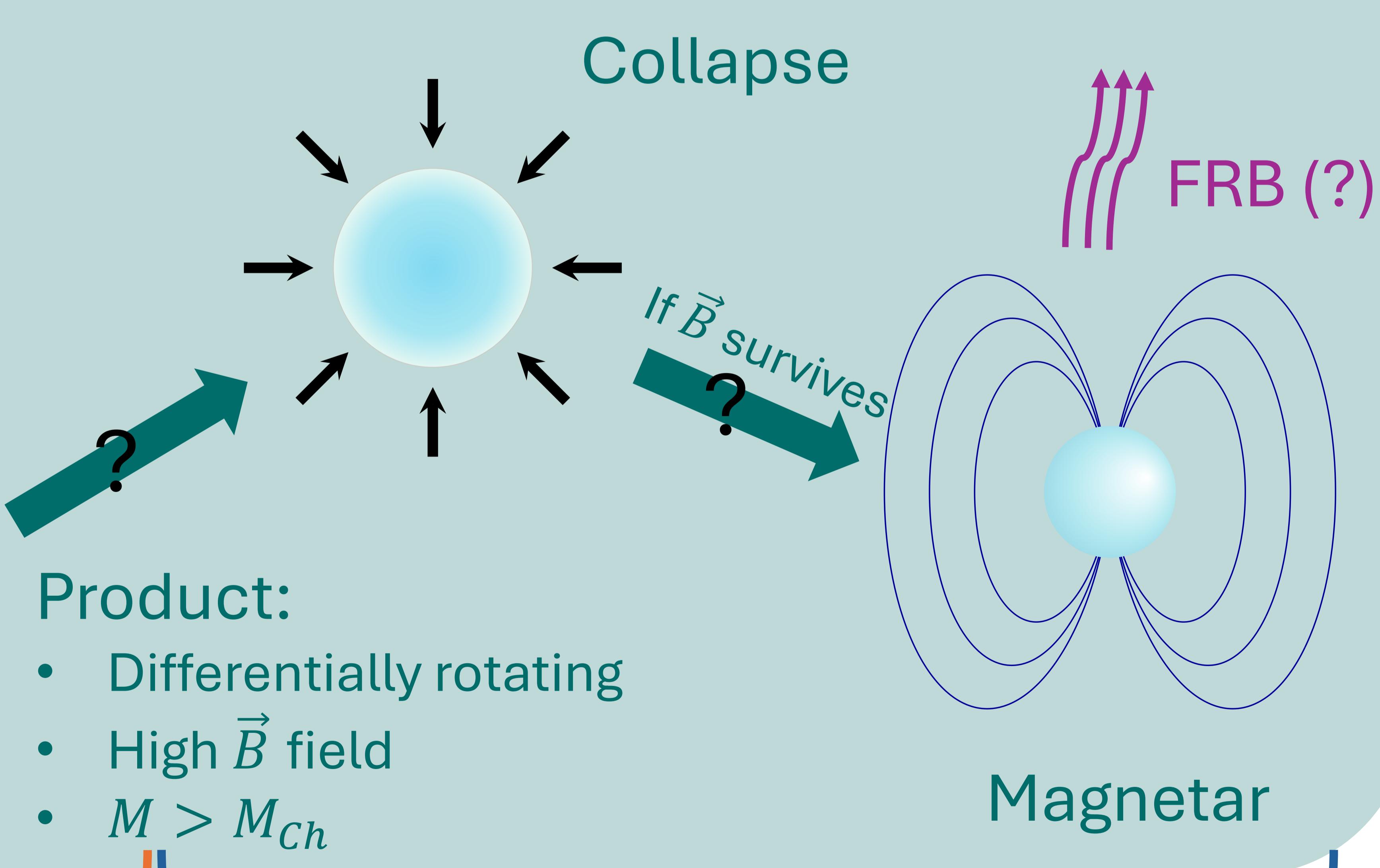
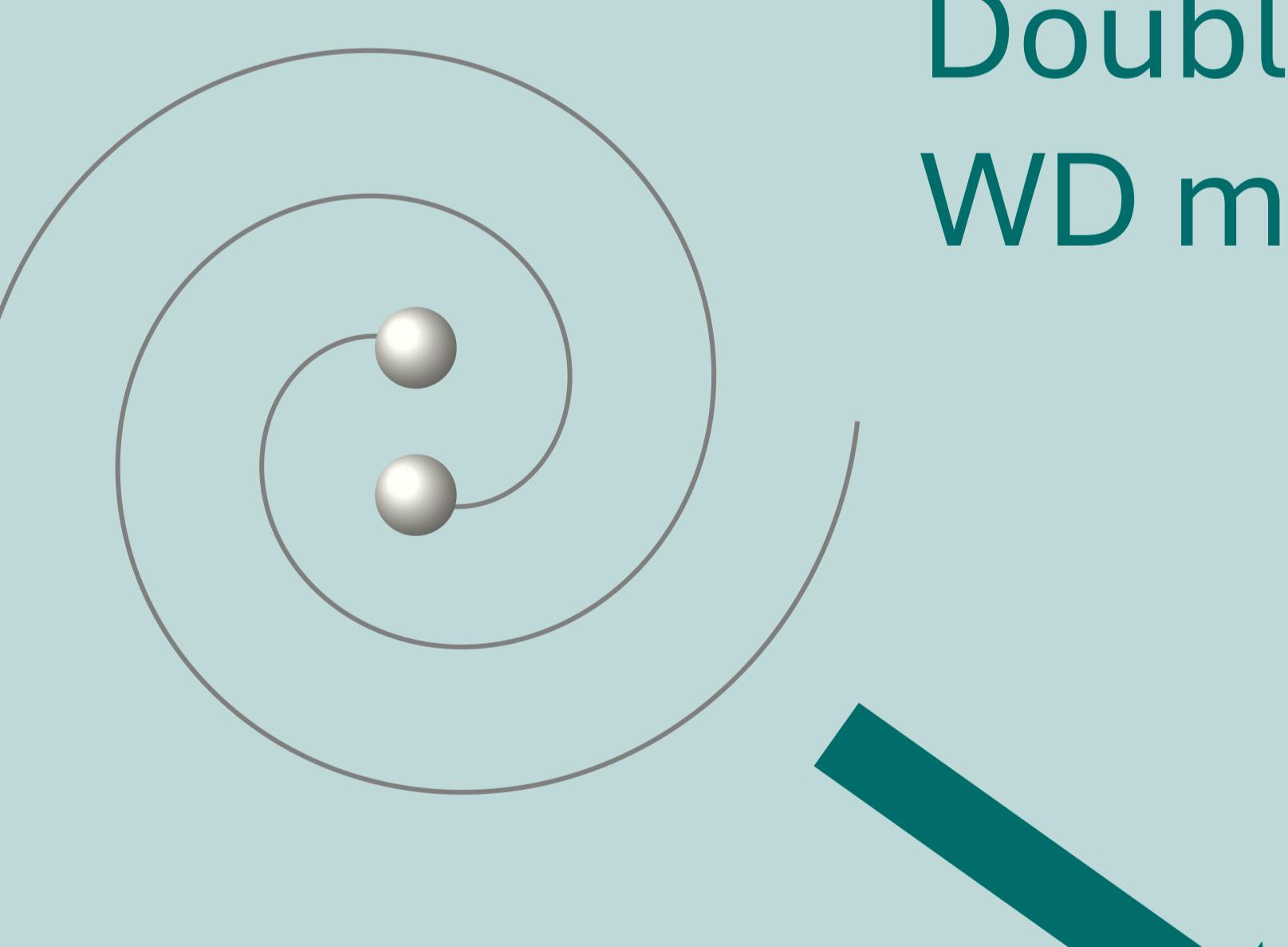
We use the 1D stellar evolution code **MESA**¹³ to generate a realistic starting model for the ONe WDs. Then, we will use the moving-mesh code **AREPO**^{14,15} to simulate the merger and the very first stages of the product following the approach of [8]. AREPO solves the ideal MHD equations in 3D using a second-order finite volume scheme¹⁵. This allows us not only to provide the first 3D hydrodynamical simulation of such a merger, but also allows us to follow the magnetic field amplification and evolution throughout the event. In addition, the moving-mesh scheme is particularly good to simulate chaotic systems such as a merger.

Limitations

Likely limitations will mainly be linked to the limited amount of time for which we will be able to simulate the event, due to the high computational time required by 3D MHD simulations and the build-up of numerical errors at longer times. In addition, the limited time and the complex physics involved will prevent us to simulate the collapse itself.

Project and (possible) implications

Double ONe WD merger



MY PROJECT

Product:

- Differentially rotating
- High \vec{B} field
- $M > M_{Ch}$

Possible Implications

References

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