Constrained Dynamics

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Generalized Inertia



- If reference point O^i is the centre of mass lacktriangle
 - $\qquad \qquad \mathbf{F}_i^i = m^i \begin{bmatrix} \ddot{R}_x^i & \ddot{R}_y^i \end{bmatrix}^\mathrm{T} \text{; } M^i = J^i \ddot{\theta}^i$
 - ▶ Inertia force and moment
- ullet Mass moment of inertia J^i

 - ullet $ar{\mathbf{u}}^i$ position vector of any point on i from O^i
- The virtual work done by the inertia forces
- We can replace these by an equipollent system of forces at any other point

Equipollent Forces



- Remember that the original and equipollent systems do the same virtual work
- \bullet Let us choose a point P^i for deriving the equipollent system ${}^{\bullet}$ RefCoord
- ullet With respect to the reference point O^i
 - $\qquad \qquad \mathbf{r}_P^i = \mathbf{R}^i + \mathbf{A}^i \bar{\mathbf{u}}_P^i \quad \text{and} \quad \delta \mathbf{r}_P^i = \delta \mathbf{R}^i + \mathbf{A}_\theta^i \bar{\mathbf{u}}_P^i \delta \theta^i$
 - lacktriangle One can rewrite second equation as $\delta {f R}^i = \delta {f r}_P^i {f A}_{ heta}^i {ar {f u}}_P^i \delta heta^i$
- Then the virtual work of inertia forces becomes
 - $\qquad \delta W_i^i = \mathbf{F}^{i^{\mathrm{T}}} \delta \mathbf{r}_P^i + (M^i \mathbf{F}^{i^{\mathrm{T}}} \mathbf{A}_{\theta}^i \bar{\mathbf{u}}_P^i) \delta \theta^i$
 - $lackbox{ One can show that } \mathbf{F}^{i^{\mathrm{T}}} \mathbf{A}_{ heta}^i ar{\mathbf{u}}_P^i = (\mathbf{u}_P^i imes \mathbf{F}^i) \cdot \hat{\mathbf{k}}$
 - $\mathbf{u}_P^i = \mathbf{A}^i \bar{\mathbf{u}}_P^i$

Parallel Axis Theorem



- The parallel axis theorem can be obtained as a special case
- ullet Let us assume that P^i is a fixed point
 - lacktriangle This implies that $\delta {f r}_P^i = 0$
- ullet Hence $\delta W_i^i = (M^i \mathbf{F}^{i^{\mathrm{T}}} \mathbf{A}_{ heta}^i \bar{\mathbf{u}}_P^i) \delta heta^i$
- $\bullet \text{ As } P^i \text{ is fixed } \mathbf{F}^i = m^i \mathbf{\ddot{R}}^i = m^i \Big\{ -\ddot{\theta}^i \mathbf{A}^i_{\theta} \ddot{\mathbf{u}}^i_P + (\dot{\theta}^i)^2 \mathbf{A}^i \ddot{\mathbf{u}}^i_P \Big\}$
 - lacksquare Hence $-\mathbf{F}^{i^{\mathrm{T}}}\mathbf{A}_{ heta}^{i}\bar{\mathbf{u}}_{P}^{i}=m^{i}(l_{P}^{i})^{2}\ddot{\theta}^{i}$
- Recall that $M^i = J^i \ddot{\theta}^i$ which means
 - - Mass moment of inertia about an arbitrary point P^i on the rigid body is equal to mass moment of inertia about centre of mass plus product of mass and square of the distance between P^i and centre of mass

Mass Matrix



- ullet The kinetic energy of the body T^i
 - $T^i = \frac{1}{2} \int_{V^i} \rho^i \dot{\mathbf{r}}^{i^{\mathrm{T}}} \dot{\mathbf{r}}^i \ dV^i$
 - f r Now we can write $\dot{f r}^i = \left[f I \quad {f A}^i_ heta ar{f u}^i_P
 ight] \left\{ egin{matrix} \dot{f R}^i \ \dot{ heta}^i \end{array}
 ight\}$
- Using this the kinetic energy becomes

$$T^{i} = \frac{1}{2} \int_{V^{i}} \rho^{i} \begin{bmatrix} \dot{\mathbf{R}}^{i^{T}} & \dot{\theta}^{i^{T}} \end{bmatrix} \begin{Bmatrix} \mathbf{I}^{T} \\ \dot{\mathbf{u}}^{i^{T}} \mathbf{A}_{\theta}^{i^{T}} \end{Bmatrix} \begin{bmatrix} \mathbf{I} & \mathbf{A}_{\theta}^{i} \dot{\mathbf{u}}^{i} \end{bmatrix} \begin{Bmatrix} \dot{\mathbf{R}}^{i} \\ \dot{\theta}^{i} \end{Bmatrix} dV^{i}$$

$$= \frac{1}{2} \begin{bmatrix} \dot{\mathbf{R}}^{i^{T}} & \dot{\theta}^{i^{T}} \end{bmatrix} \begin{Bmatrix} \int_{V^{i}} \rho^{i} \begin{bmatrix} \mathbf{I} & \mathbf{A}_{\theta}^{i} \dot{\mathbf{u}}^{i} \\ \dot{\mathbf{u}}^{i^{T}} \mathbf{A}_{\theta}^{i^{T}} & \dot{\mathbf{u}}^{i^{T}} \dot{\mathbf{u}}^{i} \end{bmatrix} dV^{i} \end{Bmatrix} \begin{Bmatrix} \dot{\mathbf{R}}^{i} \\ \dot{\theta}^{i} \end{Bmatrix}$$

$$= \frac{1}{2} \dot{\mathbf{q}}^{i^{T}} \mathbf{M}^{i} \dot{\mathbf{q}}^{i}$$

Mass Matrix ..2



The mass matrix elements

$$\mathbf{M}^{i} = \begin{bmatrix} \mathbf{m}_{RR}^{i} & \mathbf{m}_{R\theta}^{i} \\ \mathbf{m}_{R\theta}^{i^{\mathrm{T}}} & m_{\theta\theta}^{i} \end{bmatrix}$$
$$\mathbf{m}_{RR}^{i} = \int_{V^{i}} \rho^{i} \mathbf{I} \, dv^{i} = m^{i} \mathbf{I}$$
$$\mathbf{m}_{R\theta}^{i} = \mathbf{A}_{\theta}^{i} \int_{V^{i}} \rho^{i} \bar{\mathbf{u}}^{i} \, dV^{i}$$
$$m_{\theta\theta}^{i} = \int_{V^{i}} \rho^{i} \bar{\mathbf{u}}^{i^{\mathrm{T}}} \bar{\mathbf{u}}^{i} \, dV^{i}$$

- Note that $m_{\theta\theta}$ is a scalar and represents the mass moment of inertia
- \bullet The $\mathbf{m}_{R\theta}^i$ matrix represents inertia coupling between translation and rotation of body
 - $\mathbf{m}_{R heta}^{\imath}=\mathbf{0}$ if body co-ordinate system origin is at centre of mass

Cenrtifugal Inertia Force



The generalized inertia forces can be obtained as

$$\mathbf{Q}_i^i = \frac{d}{dt} \left(\frac{\partial T^i}{\partial \dot{\mathbf{q}}^i} \right)^{\mathrm{T}} - \frac{\partial T^i}{\partial \mathbf{q}^i}^{\mathrm{T}}$$

 \bullet Another way will be to use the virtual work δW_i^i

• From this we can write

$$\qquad \qquad \dot{\mathbf{r}}^i = \mathbf{L}^i \dot{\mathbf{q}}^i; \ \ddot{\mathbf{r}}^i = \mathbf{L}^i \ddot{\mathbf{q}}^i + \dot{\mathbf{L}}^i \dot{\mathbf{q}}^i \ ; \ \ \dot{\mathbf{L}}^i = \begin{bmatrix} \mathbf{0} & -\dot{\theta}^i \mathbf{A}^i \ddot{\mathbf{u}}^i \end{bmatrix}$$

Substitution leads to

$$\delta W_i^i = \int_{V^i} \rho^i \ddot{\mathbf{q}}^{i^{\mathrm{T}}} \mathbf{L}^{i^{\mathrm{T}}} \mathbf{L}^i \delta \mathbf{q}^i \, dV^i + \int_{V^i} \rho^i \dot{\mathbf{q}}^{i^{\mathrm{T}}} \dot{\mathbf{L}}^{i^{\mathrm{T}}} \mathbf{L}^i \delta \mathbf{q}^i \, dV^i$$

Centrifugal Force



The first integral leads to the mass matrix

$$\mathbf{M}^i = \int_{V^i} \rho^i \mathbf{L}^{i^{\mathrm{T}}} \mathbf{L}^i \, dV^i$$

The second integral is the centrifugal inertia force

$$\qquad \qquad \mathbf{Q}_v^i = -\int_{V^i} \rho^i \mathbf{L}^{i^{\mathrm{T}}} \dot{\mathbf{L}}^i \dot{\mathbf{q}}^i \, dV^i$$

- ullet Then we have $\delta W_i^i = [\mathbf{M}^i\ddot{\mathbf{q}}^i \mathbf{Q}_v^i]^{\mathrm{T}}\delta\mathbf{q}^i$
- It can be shown that

$$\mathbf{Q}_{v}^{i} = (\dot{\theta}^{i})^{2} \mathbf{A}^{i} \begin{Bmatrix} \int_{V^{i}} \rho^{i} \bar{\mathbf{u}}^{i} dV^{i} \\ 0 \end{Bmatrix}$$

• For the body co-ordinate system origin at centre of mass

$$\mathbf{Q}_{v}^{i}=\mathbf{0}$$

Newton-Euler Equations of Motion



Using the principle of virtual work in dynamics

- \bullet We have $\delta W_e^i = \mathbf{Q}_e^{i^T} \delta \mathbf{q}^i$ and $\delta W_c^i = \mathbf{Q}_c^{i^T} \delta \mathbf{q}^i$
- From the previous slide we then have

$$[\mathbf{M}^i \ddot{\mathbf{q}}^i - \mathbf{Q}_v^i - \mathbf{Q}_e^i - \mathbf{Q}_c^i]^{\mathrm{T}} \delta \mathbf{q}^i = 0$$

- Since we have isolated a single body we can write
 - $\mathbf{M}^i \ddot{\mathbf{q}}^i = \mathbf{Q}_v^i + \mathbf{Q}_e^i + \mathbf{Q}_c^i$
- For the origin at centre of mass this becomes

$$\begin{bmatrix} \begin{bmatrix} m^i \mathbf{I} & \mathbf{0} \\ \mathbf{0} & J^i \end{bmatrix} \begin{Bmatrix} \ddot{\mathbf{R}}^i \\ \ddot{\theta}^i \end{Bmatrix} = \begin{Bmatrix} (\mathbf{Q}_e^i)_R \\ (\mathbf{Q}_e^i)_\theta \end{Bmatrix} + \begin{Bmatrix} (\mathbf{Q}_c^i)_R \\ (\mathbf{Q}_c^i)_\theta \end{Bmatrix}$$

Multi-body System



• When we have n_b interconnected bodies

$$\mathbf{M}^{i}\ddot{\mathbf{q}}^{i} = \mathbf{Q}_{e}^{i} + \mathbf{Q}_{c}^{i}; i = 1, 2, \dots, n_{b}$$

- We can combine all these and write it in a form
 - $\mathbf{M}\ddot{\mathbf{q}} = \mathbf{Q}_e + \mathbf{Q}_c$
- The structure is

$$egin{bmatrix} \mathbf{M}^1 & \mathbf{0} & \cdots & \mathbf{0} \ \mathbf{0} & \mathbf{M}^2 & \cdots & \mathbf{0} \ dots & dots & \ddots & dots \ \mathbf{0} & \mathbf{0} & \cdots & \mathbf{M}^{n_b} \end{bmatrix} egin{bmatrix} \ddot{\mathbf{q}}^1 \ \ddot{\mathbf{q}}^2 \ dots \ \ddot{\mathbf{q}}^{n_b} \end{pmatrix} = egin{bmatrix} \mathbf{Q}_e^1 \ \mathbf{Q}_e^2 \ dots \ \mathbf{Q}_e^{n_b} \end{pmatrix} + egin{bmatrix} \mathbf{Q}_c^1 \ \mathbf{Q}_c^2 \ dots \ \mathbf{Q}_c^{n_b} \end{pmatrix}$$

Constraint Force Elimination



ullet For a system of n_b connected rigid bodies

As before for frictionless constraint forces we have

$$\sum_{i=1}^{n_b} \delta W_c^i = 0$$

• This then implies that

$$\qquad \qquad \sum_{i=1}^{n_b} \{\mathbf{M}^i \ddot{\mathbf{q}}^i - \mathbf{Q}_e^i\}^{\mathrm{T}} \delta \mathbf{q}^i = 0$$

• The above equation can be re-written in the form



$$\begin{cases}
\delta \mathbf{q}^1 \\
\delta \mathbf{q}^2 \\
\vdots \\
\delta \mathbf{q}^{n_b}
\end{cases}^{\mathrm{T}} \begin{pmatrix}
\mathbf{M}^1 & \mathbf{0} & \cdots & \mathbf{0} \\
\mathbf{0} & \mathbf{M}^2 & \cdots & \mathbf{0} \\
\vdots & \vdots & \ddots & \vdots \\
\mathbf{0} & \mathbf{0} & \cdots & \mathbf{M}^{n_b}
\end{cases} \begin{pmatrix}
\ddot{\mathbf{q}}^1 \\
\ddot{\mathbf{q}}^2 \\
\vdots \\
\ddot{\mathbf{q}}^{n_b}
\end{pmatrix} - \begin{pmatrix}
\mathbf{Q}_e^1 \\
\mathbf{Q}_e^2 \\
\vdots \\
\mathbf{Q}_e^{n_b}
\end{pmatrix} = 0$$

- ullet Since the ${f q}$ are all not independent one cannot write
 - $\mathbf{M}\ddot{\mathbf{q}} \mathbf{Q}_e = \mathbf{0}$
- ullet We now try to partition ${f q}$ into dependent co-ordinates ${f q}_d$ and independent co-ordinates ${f q}_i$
- The constraint equations are C(q, t) = 0

Partition Co-ordinates



ullet For a virtual displacement $\delta {f q}$ we have

- Now \mathbf{q}_d is of dimension n_c with \mathbf{q}_i of dimension $n-n_c$
- The virtual displacement equation above can be re-written as

$$\qquad \qquad \mathbf{C}_{\mathbf{q}_d} \delta \mathbf{q}_d + \mathbf{C}_{\mathbf{q}_i} \delta \mathbf{q}_i = \mathbf{0} \text{ or } \delta \mathbf{q}_d = -\mathbf{C}_{\mathbf{q}_d}^{-1} \mathbf{C}_{\mathbf{q}_i} \delta \mathbf{q}_i$$

Using this we can now write

$$\qquad \qquad \delta \mathbf{q} = \mathbf{B}_i \delta \mathbf{q}_i \, \text{ with } \, \mathbf{B}_i = \left\{ \begin{matrix} -\mathbf{C}_{\mathbf{q}_d}^{-1} \mathbf{C}_{\mathbf{q}_i} \\ \mathbf{I} \end{matrix} \right\}$$

Embedding Technique



- So the original equation $\delta \mathbf{q}^{\mathrm{T}}\{\mathbf{M}\ddot{\mathbf{q}}-\mathbf{Q}_{e}\}=\mathbf{0}$ can be written as
- Remember that the acceleration form of the constraint equations

$$\mathbf{C}_{\mathbf{q}}\ddot{\mathbf{q}} = \mathbf{Q}_d = -(\mathbf{C}_{\mathbf{q}}\dot{\mathbf{q}})_{\mathbf{q}}\dot{\mathbf{q}} - 2\mathbf{C}_{\mathbf{q}t}\dot{\mathbf{q}} - \mathbf{C}_{tt}$$

Partitioning into dependent and independent variables

From this we can write

$$raket \ddot{\mathbf{q}} = \mathbf{B}_i \ddot{\mathbf{q}}_i + oldsymbol{\gamma} \; ; \;\; oldsymbol{\gamma} = egin{cases} \mathbf{C}_{\mathbf{q}d}^{-1} \mathbf{Q}_d \ \mathbf{0} \end{cases}$$

Final Form



- $ullet \left[\mathbf{B}_i^{\mathrm{T}} \mathbf{M} \mathbf{B}_i \ddot{\mathbf{q}}_i = \mathbf{B}_i^{\mathrm{T}} \mathbf{Q}_e \mathbf{B}_i^{\mathrm{T}} \mathbf{M} oldsymbol{\gamma}
 ight]$
- How does one automatically identify the q_i ?
- ullet The Jacobian matrix $\mathbf{C}_{\mathbf{q}}$ is of size $n_c imes n$
- Gaussian Elimination with partial or full pivoting leads to a form

$$\begin{bmatrix} 1 & C_{12} & \cdots & C_{1n_c} & C_{1(n_c+1)} & \cdots & C_{1n} \\ 0 & 1 & \cdots & C_{2n_c} & C_{2(n_c+1)} & \cdots & C_{2n} \\ \vdots & \vdots & \ddots & \vdots & \ddots & \vdots \\ 0 & 0 & \cdots & 1 & C_{n_c(n_c+1)} & \cdots & C_{n_cn} \end{bmatrix} \begin{bmatrix} \delta u_1 \\ \delta u_2 \\ \vdots \\ \delta u_{n_c} \\ \delta v_1 \\ \vdots \\ \delta v_{n-n_c} \end{bmatrix} = \begin{bmatrix} 0 \\ 0 \\ \vdots \\ 0 \\ 0 \\ \vdots \\ 0 \end{bmatrix}$$

Automatic Determination

- The form of the previous equation is
- ullet The original co-ordinates have been re-ordered in $\delta {f u}$ and $\delta {f v}$
- \bullet Note that U is an upper triangular matrix with diagonal entries of 1
 - Will be non-singular as long as there are no redundant co-ordinates
- ullet The vector $\delta {f u} = \delta {f q}_d$ and $\delta {f v} = \delta {f q}_i$

Implementation: Embedded Formulation



- Involves numerical integration of differential equations only
 - Independent co-ordinates only involved
 - ▶ Strongly coupled and smaller in size
- The dependent velocities and co-ordinates found from kinematic constraint equations

$$\dot{\mathbf{q}}_d = -(\mathbf{C}_{\mathbf{q}d})^{-1} \left\{ \mathbf{C}_t + \mathbf{C}_{\mathbf{q}i} \dot{\mathbf{q}}_i \right\}$$

$$\mathbf{q}_d = -(\mathbf{C}_{\mathbf{q}d})^{-1}\mathbf{C}_{\mathbf{q}i}\mathbf{q}_i$$

Alternate Idea



- Estimate initial conditions for the initial configuration of the multi-body system
 - ▶ Both positions and velocities to be a good approximation
- ullet Construct Jacobian matrix $\mathbf{C}_{\mathbf{q}}$ using initial co-ordinates
 - ▶ LU factorization to identify independent co-ordinates
- Constraint equations are non-linear system of equations in dependent co-ordinates
 - Assuming the initial values of independent co-ordinates known
 - Newton-Raphson iteration used
- Above process done without need for decomposition of $\mathbf{C_q}$ into $\mathbf{C_{q\it d}}$ and $\mathbf{C_{q\it i}}$

Newton-Iteration



• The Newton iteration is written in the following form

$$\quad \left\{ \begin{aligned} \mathbf{C_q} \\ \mathbf{I}_d \end{aligned} \right\} \Delta \mathbf{q} = \left\{ \begin{aligned} -\mathbf{C} \\ \mathbf{0} \end{aligned} \right\}$$

- $ightharpoonup \Delta \mathbf{q}$ are the Newton updates
- ▶ \mathbf{I}_d is a Boolean matrix of size $(N-N_c) \times N$
 - ★ Entry of 1 in each row corresponding to independent co-ordinate
 - ★ All other entries are 0 in that row
- ullet We are basically driving $\Delta \mathbf{q}_i$ to be zero
- ullet Above form a square matrix of size $N \times N$ and sparse
 - Sparse solvers can be used
 - lacktriangle No need to decompose $\mathbf{C}_{\mathbf{q}}$

Next Steps



- From the solution of the non-linear algebraic equations all co-ordinates are obtained
- ullet From this one can obtain all N velocities

$$\qquad \left\{ \begin{aligned} \mathbf{C}_{\mathbf{q}} \\ \mathbf{I}_{d} \end{aligned} \right\} \dot{\mathbf{q}} = \left\{ \begin{aligned} -\mathbf{C}_{t} \\ \dot{\mathbf{q}}_{i} \end{aligned} \right\}$$

- ullet $\dot{\mathbf{q}}_i$ assumed to be known from numerical integration
- If the set of independent co-ordinates change during simulation one needs to simply change the location of the 1s and 0s

Numerical Integration



- Most numerical integration algorithms exist for first order ODEs of the form $\frac{d\mathbf{y}}{dt} = \mathbf{f}(\mathbf{y}, t)$
- We need to re-cast the dynamic equations into state-space form
- Let us look at the simplest one Euler method
- ullet Integrating from $t=t_0$ to $t=t_0+h$ we have

$$\mathbf{y}(t_0 + h) = \mathbf{y}(t_0) + \int_{t_0}^{t_0 + h} \mathbf{f}(\mathbf{y}, t) dt$$

If h is small then one can approximate the above as

$$\mathbf{y}(t_0 + h) = \mathbf{y}(t_0) + h\mathbf{f}(\mathbf{y}_0, t)$$

- Can be generalized to a form
 - $\mathbf{y}(t_n + h) = \mathbf{y}(t_n) + h\mathbf{f}(\mathbf{y}_n, t_n)$

Euler's Method



One can also arrive at this from a Taylor series

$$\mathbf{y}(t_n + h) = \mathbf{y}(t_n) + h \frac{d\mathbf{y}}{dt}|_{t_n} + \frac{h^2}{2!} \frac{d^2\mathbf{y}}{dt^2}|_{t_n} + \frac{h^3}{3!} \frac{d^3\mathbf{y}}{dt^3}|_{t_n} + \cdots$$

This can be re-written in the form below

$$\mathbf{y}(t_n + h) = \mathbf{y}(t_n) + h\mathbf{f}(\mathbf{y}_n, t_n) + \frac{h^2}{2!} \left(\frac{\partial \mathbf{f}}{\partial t} + \frac{\partial \mathbf{f}}{\partial \mathbf{y}} \mathbf{f}\right) + \cdots$$

- If we retain only the first term in the Taylor series we derive the Euler's method
 - ▶ This can also be looked as forward difference formula
 - ▶ The error will be of $\mathcal{O}(h)$ assuming the second derivative is bounded
 - ▶ It is called a first order method

Higher Order Methods



- Higher order methods will include more terms in the Taylor series expansion
- For example the mid-point method gives a second order scheme

$$\mathbf{y}(t_n + h) = \mathbf{y}(t_n) + h\mathbf{f}\{y_n + \frac{1}{2}h\mathbf{f}(\mathbf{y}_n, t_n), t_n + \frac{1}{2}h\}$$

• A Taylor series expansion of the function

$$\mathbf{f}\{y_n + \frac{1}{2}h\mathbf{f}(\mathbf{y}_n, t_n), t_n + \frac{1}{2}h\} = \mathbf{f}(\mathbf{y}_n, t_n) + \frac{1}{2}h\mathbf{f}(\mathbf{y}_n, t_n)\frac{\partial \mathbf{f}}{\partial \mathbf{y}} + \frac{1}{2}h\frac{\partial \mathbf{f}}{\partial t} + \cdots$$

- \bullet Substitution will show the expansion satisfies Taylor series to order h^2
 - Error will be of $\mathcal{O}(h^2)$

Generalized Second Order Schemes



A general form for second order numerical integration schemes

$$\mathbf{y}(t_n + h) = \mathbf{y}(t_n) + h \left[(1 - \frac{1}{2\alpha}) \mathbf{f}(\mathbf{y}_n, t_n) + \frac{1}{2\alpha} \mathbf{f} \left\{ \mathbf{y}_n + \alpha h \mathbf{f}(\mathbf{y}_n, t_n), t_n + \alpha h \right\} \right]$$

- We have just used $\alpha = \frac{1}{2}$ to derive the mid-point method
- ullet For lpha=1 we get the Heun's method

$$\mathbf{y}(t_n + h) = \mathbf{y}(t_n) + \frac{1}{2}h \left[\mathbf{f}(\mathbf{y}_n, t_n) + \mathbf{f}(\mathbf{y}_n + h\mathbf{f}(\mathbf{y}_n, t_n), t_n + h) \right]$$

• A choice of $\alpha = \frac{2}{3}$ gives the Ralston's method

Fourth-Order Runge-Kutta Method



$$\mathbf{y}(t_n + h) = \mathbf{y}(t_n) + \frac{h}{6}(\mathbf{k}_1 + 2\mathbf{k}_2 + 2\mathbf{k}_3 + \mathbf{k}_4)$$

$$\mathbf{k}_1 = \mathbf{f}(\mathbf{y}_n, t_n)$$

$$\mathbf{k}_2 = \mathbf{f}(\mathbf{y}_n + \frac{h}{2}\mathbf{k}_1, t_n + \frac{h}{2})$$

$$\mathbf{k}_3 = \mathbf{f}(\mathbf{y}_n + \frac{h}{2}\mathbf{k}_2, t_n + \frac{h}{2})$$

$$\mathbf{k}_4 = \mathbf{f}(\mathbf{y}_n + h\mathbf{k}_3, t_n + h)$$

- A Taylor series expansion will fit upto $\mathcal{O}(h^4)$
- Error for this method will be $\mathcal{O}(h^4)$
- At each time step four function evaluations of $(\mathbf{f}(\mathbf{y},t))$ needs to be done

Stability



- All the methods discussed so far are explicit schemes
 - System state at next time step computed with information at current and previous time steps
 - ▶ If only current time step involved then called one step schemes
 - Else multi-step schemes
- Implicit schemes need current time step and next time step states
 - Involves iteration for solution
- Then why does one use implicit schemes at all?
 - ▶ They are numerically more stable
- We shall look at a simple example to demonstrate the idea

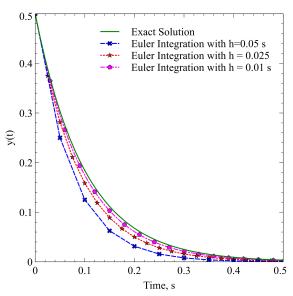
Stability for Euler Scheme



- Let us look at the numerical integration of the following equation
 - $\frac{dy}{dt} = -\alpha y$; $y(0) = y_0$ with exact solution $y(t) = y_0 e^{-\alpha t}$
- Our first-order explicit Euler method is
 - $y_{n+1} = y_n + h f(y_n, t_n) = (1 \alpha h) y_n$
 - From the initial condition we have $y_{n+1} = (1 \alpha h)^n y_0$
- ullet Now we know the solution exponentially decays for positive lpha
- For the numerical solution to behave similarly we need
 - ▶ $|1 \alpha h| < 1$ which leads to h > 0 and $h < \frac{2}{\alpha}$

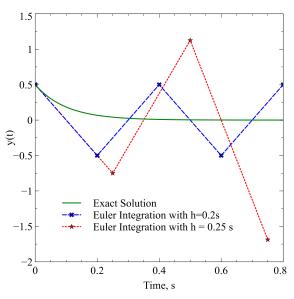
Convergence of Euler Method ($\alpha = 10$)





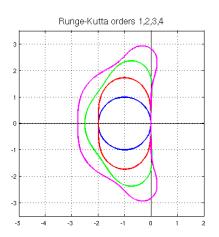
Instability of Euler Method ($\alpha = 10$)





λh plane RK Stability Boundaries





 Courtesy https://www.mathworks.com/matlabcentral/mlcdownloads/downloads/submissions/23972/versions/22/previews/chebfun/examples/ode/html

Implicit Euler Method



- This can usually be solved by
 - ► Fixed-point iteration scheme
 - Newton-Raphson method
- Fixed-point iteration

$$\mathbf{y}_{n+1}^0 = \mathbf{y}_n; \ \mathbf{y}_{n+1}^{k+1} = \mathbf{y}_n + h\mathbf{f}(\mathbf{y}_{n+1}^k, t_{n+1}), k = 0, 1, 2, \cdots$$

- Now let us look at the stability of this method through the same example

Implicit Euler Stability



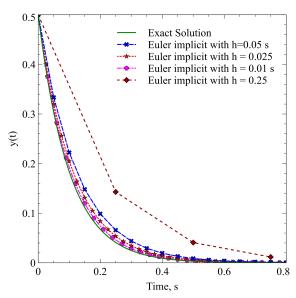
The implicit Euler method for this example is

$$y_{n+1} = y_n - \alpha h y_{n+1}$$
 or $y_{n+1} = \frac{1}{1+\alpha h} y_n = (\frac{1}{1+\alpha h})^n y_0$

- ullet For $\alpha>0$ the exact solution exponentially decays
- The numerical scheme will also decay provided
 - $| \frac{1}{1+\alpha h} | < 1$ which leads to h>0 and $h<-\frac{2}{\alpha}$
 - \blacktriangleright Since $\alpha>0$ the second condition would mean h<0 which is not meaningful
 - ▶ This implies that for any h > 0 solution will decay
- The implicit method is called as A-stable

Implicit Euler Method ($\alpha = 10$)





Explicit Multi-Step Methods



• The Adams method is most popular and of the form

$$\mathbf{y}_{n+1} = \mathbf{y}_n + \sum_{i=1}^s \beta_i \mathbf{f}(\mathbf{y}_{n+1-i}, t_{n+1-i}) \text{ with } \sum_{i=1}^s \beta_i = 1$$

- ightharpoonup s is the number of steps
- The solution can be written as follows

$$\mathbf{y}_{n+1} = \mathbf{y}_n + \int_{t_n}^{t_{n+1}} \mathbf{f}(\mathbf{y}, s) \, ds$$

- In the Adams method f is replaced by a polynomial function
 - ▶ This is generated from knowledge of f at $t_{n+1-s}, t_{n+2-s}, \cdots, t_n$

$$\mathbf{P}_{s}(t) = \sum_{i=0}^{s-1} \mathbf{f}(\mathbf{y}_{n-i}, t_{n-i}) L_{i}(t) ; L_{i}(t) = \prod_{\substack{k=0\\k\neq i}}^{s-1} \left(\frac{t-t_{n-k}}{t_{n-i}-t_{n-k}}\right)$$

Adams-Bashforth Method



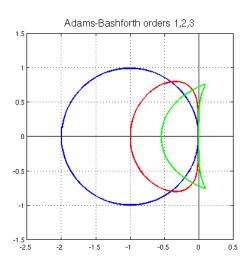
• A 3 step Adams-Bashforth method will be

$$\mathbf{y}_{n+1} = \mathbf{y}_n + \frac{h}{12} \left[23\mathbf{f}(\mathbf{y}_n, t_n) - 16\mathbf{f}(\mathbf{y}_{n-1}, t_{n-1}) + 5\mathbf{f}(\mathbf{y}_{n-2}, t_{n-2}) \right]$$

- ullet A s step Adams-Bashforth method is of order s
- One can see that the multi-step method is not self-starting
 - lackbox One can use a single step method for the first s-1 steps and then switch to the multi-step method
- ullet At each time step only one function evaluation $oldsymbol{(f(y,t))}$ needs to be done

Stability of A-B methods





Courtesy https://www.mathworks.com/matlabcentral/mlcdownloads/downloads/submissions/23972/versions/22/previews/chebfun/examples/ode/html

Adams-Moulton Methods



- The idea of the scheme is exactly the same as the explicit scheme
 - Only difference is that it involves the next step
- For instance the two-step Adams-Moulton method

$$\mathbf{y}_{n+1} = \mathbf{y}_n + \frac{h}{12} [5\mathbf{f}(\mathbf{y}_{n+1}, t_{n+1}) + 8\mathbf{f}(\mathbf{y}_n, t_n) - \mathbf{f}(\mathbf{y}_{n-1}, t_{n-1})]$$

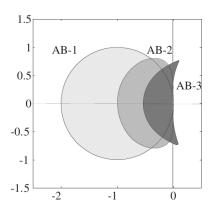
• Similarly a four-step A-M method would be

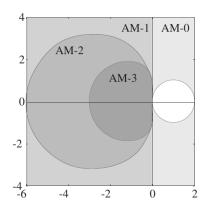
$$\mathbf{y}_{n+1} = \mathbf{y}_n + \frac{h}{720} [251\mathbf{f}(\mathbf{y}_{n+1}, t_{n+1}) + 646\mathbf{f}(\mathbf{y}_n, t_n) - 264\mathbf{f}(\mathbf{y}_{n-1}, t_{n-1}) + 106\mathbf{f}(\mathbf{y}_{n-2}, t_{n-2}) - 19\mathbf{f}(\mathbf{y}_{n-3}, t_{n-3})]$$

- Obviously one has to do a fixed-point iteration or Newton-Raphson solution
- \bullet A n step A-M method is of order n+1 whereas a n step A-B is of order n

A-M Stability







• From: http://www.maths.lth.se/na/courses/FMN081/FMN081-06/lecture23.pdf

Predictor-Corrector Methods



- Designed to reduce the iterative nature of implicit methods
- A predictor step is essentially an explicit method

$$\hat{\mathbf{y}}_{n+1} = \mathbf{y}_n + h\mathbf{f}(\mathbf{y}_n, t_n)$$

This estimate is now used in a corrector step based on an implicit scheme

$$\mathbf{y}_{n+1} = \mathbf{y}_n + \frac{h}{2} [\mathbf{f}(\mathbf{y}_n, t_n) + \mathbf{f}(\hat{\mathbf{y}}_{n+1}, t_{n+1})]$$

• Now one can stop here or do additional corrector steps

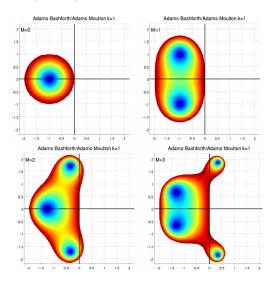
$$\tilde{\mathbf{y}}_{n+1} = \mathbf{y}_n + \frac{h}{2} [\mathbf{f}(\mathbf{y}_n, t_n) + \mathbf{f}(\hat{\mathbf{y}}_{n+1}, t_{n+1})]$$

$$\mathbf{y}_{n+1} = \mathbf{y}_n + \frac{h}{2} [\mathbf{f}(\mathbf{y}_n, t_n) + \mathbf{f}(\tilde{\mathbf{y}}_{n+1}, t_{n+1})]$$

ullet These are called $P(EC)^m$ schemes where P is predictor E is evaluate and C for corrector

Stability of $P(EC)^m$ schemes





From: http://www3.math.tu-berlin.de/Vorlesungen/SS14/NumMath2/exp_StabRegionsPECE.pdf

Variable Time Step Scheme



- How does one determine the step size h?
- The computational effort is proportional to the number of time steps
 - We would want the step size to be as large as possible
 - Not too large that truncation error is big
- ullet We want truncation error to be less than some tolerance ϵ
- ullet Remember we are trying to solve $\dot{y}=f(y,t)$ with $y(t_0)=y_0$
 - Considering single variable for illustration purposes
 - Holds for y

Estimate Error



• For a single step method of order p the approximate solution $\eta(t;h)$ can be written as

$$\eta(t;h) = y(t) + h^p e_p(t) + h^{p+1} e_{p+1}(t) + \cdots$$

• The error estimate is then given by

•
$$e(t;h) = \eta(t;h) - y(t) = h^p e_p(t) + \mathcal{O}(h^{p+1})$$

• If we take a step of size $\frac{h}{2}$ we have

•
$$e(t; \frac{h}{2}) = \eta(t; \frac{h}{2}) - y(t) = (\frac{h}{2})^p e_p(t)$$

- Hence one can now estimate $e_p(t)$ as simply

 - $e_p(t)(\frac{h}{2})^p = \frac{\eta(t;h) \eta(t;\frac{h}{2})}{(2^p 1)}$
- For 4th order Runge-Kutta denominator is 15

Step Size Variation



- ullet Now $e_p(t_0+h)=\dot{e}_P(t_0)h^{p+1}$ assuming $y(t_0)$ is known
- We require $\epsilon = |\dot{e}(t_0)h^{p+1}|$
- Suppose we use a step size H to compute $\eta(t_0+H;H)$ and $\eta(t_0+H;\frac{H}{2})$
- From this we can derive the following

$$\dot{e}_p(t_0) = \frac{2^p}{H^{p+1}(2^p-1)} \left[\eta(t_0 + H; H) - \eta(t_0 + H; \frac{H}{2}) \right]$$

Hence we have

$$\qquad \qquad \qquad \qquad \qquad \qquad \qquad \frac{H}{h} \; = \; \sqrt[p+1]{\frac{2^p}{2^p - 1}} \frac{|\eta(t_0 + H; H) - \eta(t_0 + H; \frac{H}{2})|}{\epsilon}$$

Step Size Control



- If $\frac{H}{h} \gg 2$ error high; replace H by 2h and proceed
- •
- \bullet For error estimate we need to compute $\eta(t_0+H;H)$ and $\eta(t_0+H;\frac{H}{2})$
 - ► This involves 2 additional function evaluations per time step
- ullet Instead one can use two different order methods with same h

$$\hat{y}_{n+1} = y_n + h\Phi_1(y_n, t_n; h); \ \tilde{y}_{n+1} = y_n + h\Phi_2(y_n, t_n; h)$$

$$\Phi_1 = \sum_{k=0}^2 c_k f_k(y, t; h); \ \Phi_2 = \sum_{k=0}^3 \hat{c}_k f_k(y, t; h)$$

$$f_k = f(y + h \sum_{l=0}^{k-1} \beta_{kl} f_l, t + \alpha_k h)$$

• Φ_2 order 3 while Φ_1 is order 2; f_0, f_1 and f_2 from Φ_1 used in Φ_2 with only one additional f_3 computed

Runge-Kutta-Fehlberg Method



\overline{k}	α_k		β_{kl}	c_k	\hat{c}_k	
		l=0	l = 1	l=2		
0	0	0			$\frac{214}{891}$	$\frac{533}{2106}$
1	$\frac{1}{4}$	$\frac{1}{4}$			$\frac{1}{33}$	0
2	$\frac{27}{40}$	$-\frac{189}{800}$	$\frac{729}{800}$		$\frac{650}{891}$	$\frac{800}{1053}$
3	1	$\frac{214}{891}$	$\frac{1}{33}$	$\frac{650}{891}$		$-\frac{1}{78}$

Some General Comments



- Accuracy high for the numerical solution
 - Step-sizes are small and order of the method higher
- Stability to be good
 - ► Smaller step-sizes and lower order in case of multi-step methods
 - Choice of an implicit scheme over explicit scheme
- It is to be noted that Runge-Kutta methods are better
 - Higher order is more stable and hence step-sizes can be larger
 - Only negative is the number of computations compared to multi-step methods
- There are problems with extremely different time-scales
 - ► This can cause variable step-size methods to become computationally expensive
 - Step-size dictated by the lowest time scale or largest frequency component

Cash-Karp Method



- Developed for problems where sharp changes/discontinuites in the variables occur
- Cash and Karp in fact suggest that R-K methods are more suitable in such situations
 - ► The advantage is the flexibility in step-size of R-K method
- But when steep changes occur a number of steps are rejected as the error tolerance is not met
- So they have proposed a 5th order R-K but with orders 1-4 imbedded in that
 - ► Various coefficients required given here CashKarp

Augmented form: Lagrange Multipliers

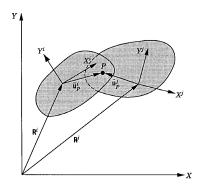


- ullet Consider body i rigidly connected to body j at P
 - $\mathbf{R}^i + \mathbf{A}^i \bar{\mathbf{u}}_P^i \mathbf{R}^j \mathbf{A}^j \bar{\mathbf{u}}_P^j = \mathbf{0}$
 - $\theta^i \theta^j = 0$
- The Jacobian matrix can be written as

$$lackbox{f C}_{f q} = egin{bmatrix} {f C}_{{f q}^i} & {f C}_{{f q}^j} \end{bmatrix}$$

$$\mathbf{C}_{\mathbf{q}^i} = \begin{bmatrix} \mathbf{I} & \mathbf{A}_{\theta}^i \bar{\mathbf{u}}_P^i \\ \mathbf{0} & 1 \end{bmatrix}$$

$$\mathbf{C}_{\mathbf{q}^j} = - \begin{bmatrix} \mathbf{I} & \mathbf{A}_{\theta}^j \bar{\mathbf{u}}_P^j \\ \mathbf{0} & 1 \end{bmatrix}$$



Courtesy: A. A. Shabana, 2010, *Computational Dynamics*, Third Edition, John Wiley & Sons.

Forces



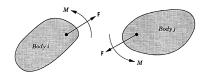
Let us define

$$\lambda = - \begin{Bmatrix} \mathbf{F} \\ M \end{Bmatrix}$$

- ullet Then the constraints force vector ${f F}^i = oldsymbol{\lambda}$ and ${f F}^j = oldsymbol{\lambda}$
- The generalized constraint forces at the reference points

$$\qquad \mathbf{Q}_c^i = \left\{ \begin{matrix} \mathbf{F} \\ M + (\mathbf{A}^i \bar{\mathbf{u}}_P^i \times \mathbf{F}) \cdot \mathbf{k} \end{matrix} \right\}$$

$$\qquad \qquad \mathbf{Q}_c^j = -\left\{ \begin{matrix} \mathbf{F} \\ M + (\mathbf{A}^j \bar{\mathbf{u}}_P^j \times \mathbf{F}) \cdot \mathbf{k} \end{matrix} \right\}$$



Courtesy: A. A. Shabana, 2010, *Computational Dynamics*, Third Edition, John Wiley & Sons.

Final Form



- Recall that
 - $\qquad \qquad (\mathbf{A}^i \bar{\mathbf{u}}_P^i \times \mathbf{F}) \cdot \mathbf{k} = \bar{\mathbf{u}}_P^{i^{\mathrm{T}}} \mathbf{A}_{\theta}^{i^{\mathrm{T}}} \mathbf{F}$
 - $\qquad \qquad (\mathbf{A}^j \bar{\mathbf{u}}_P^j \times \mathbf{F}) \cdot \mathbf{k} = \bar{\mathbf{u}}_P^{j^{\mathrm{T}}} \mathbf{A}_{\theta}^{j^{\mathrm{T}}} \mathbf{F}$
- Hence we can write

$$\mathbf{Q}_{c}^{i} = \begin{bmatrix} \mathbf{I} & \mathbf{0} \\ \bar{\mathbf{u}}_{P}^{i^{\mathrm{T}}} \mathbf{A}_{\theta}^{i^{\mathrm{T}}} & 1 \end{bmatrix} \begin{Bmatrix} \mathbf{F} \\ M \end{Bmatrix}; \ \mathbf{Q}_{c}^{j} = - \begin{bmatrix} \mathbf{I} & \mathbf{0} \\ \bar{\mathbf{u}}_{P}^{j^{\mathrm{T}}} \mathbf{A}_{\theta}^{j^{\mathrm{T}}} & 1 \end{bmatrix} \begin{Bmatrix} \mathbf{F} \\ M \end{Bmatrix}$$

- $ullet \ \ \mathbf{Q}_c^i = -\mathbf{C}_{\mathbf{q}^i}^{\mathrm{T}} oldsymbol{\lambda} \ \ ext{and} \ \ \mathbf{Q}_c^j = -\mathbf{C}_{\mathbf{q}^j}^{\mathrm{T}} oldsymbol{\lambda}$
- ullet λ is called as Lagrange Multipliers

Revolute Joint



- Suppose body i and j connected by a revolute joint
- Then the constraint equation for this joint

$$\mathbf{R}^i + \mathbf{A}^i \bar{\mathbf{u}}_P^i - \mathbf{R}^j - \mathbf{A}^j \bar{\mathbf{u}}_P^j = \mathbf{0}$$

The Jacobian is

$$\qquad \qquad \mathbf{C_q} = \begin{bmatrix} \mathbf{I} & \mathbf{A}_{\theta}^i \bar{\mathbf{u}}_P^i & -\mathbf{I} & -\mathbf{A}_{\theta}^j \bar{\mathbf{u}}_P^j \end{bmatrix}$$

• The generalized constraint force on body $\mathbf{Q}_c^i = -\mathbf{C}_{\mathbf{q}^i}^{\mathrm{T}} \boldsymbol{\lambda}$ and $\mathbf{Q}_c^j = -\mathbf{C}_{\mathbf{q}^j}^{\mathrm{T}} \boldsymbol{\lambda}$ are computed

Generalized Constraint Force: Body i



$$\mathbf{Q}_{c}^{i} = -\begin{bmatrix} 1 & 0 \\ 0 & 1 \\ -\bar{x}_{P}^{i} \sin \theta^{i} - \bar{y}_{P}^{i} \cos \theta^{i} & \bar{x}_{P}^{i} \cos \theta^{i} - \bar{y}_{P}^{i} \sin \theta^{i} \end{bmatrix} \begin{Bmatrix} \lambda_{1} \\ \lambda_{2} \end{Bmatrix}$$

$$= \begin{cases} -\lambda_{1} \\ -\lambda_{2} \\ \lambda_{1}(\bar{x}_{P}^{i} \sin \theta^{i} + \bar{y}_{P}^{i} \cos \theta^{i}) - \lambda_{2}(\bar{x}_{P}^{i} \cos \theta^{i} - \bar{y}_{P}^{i} \sin \theta^{i}) \end{Bmatrix}$$

Generalized Constraint Force: Body j



$$\begin{bmatrix}
\mathbf{Q}_{c}^{j} = -\begin{bmatrix} -1 & 0 \\ 0 & -1 \\ \bar{x}_{P}^{j} \sin \theta^{j} + \bar{y}_{P}^{j} \cos \theta^{j} & -\bar{x}_{P}^{j} \cos \theta^{j} + \bar{y}_{P}^{j} \sin \theta^{j} \end{bmatrix} \begin{Bmatrix} \lambda_{1} \\ \lambda_{2} \end{Bmatrix} \\
= \begin{Bmatrix} \lambda_{1} \\ \lambda_{2} \\ -\lambda_{1}(\bar{x}_{P}^{j} \sin \theta^{j} + \bar{y}_{P}^{j} \cos \theta^{j}) + \lambda_{2}(\bar{x}_{P}^{j} \cos \theta^{j} - \bar{y}_{P}^{j} \sin \theta^{j}) \end{Bmatrix}$$

Multiple Joints



- ullet Suppose body i is connected to other bodies by n_i joints
 - $\mathbf{C}_1(\mathbf{q},t) = \mathbf{0}; \ \mathbf{C}_2(\mathbf{q},t) = \mathbf{0}; \ \cdots; \ \mathbf{C}_{n_i}(\mathbf{q},t) = \mathbf{0}$

$$\mathbf{q} = \begin{bmatrix} \mathbf{q}^{1^{\mathrm{T}}} & \mathbf{q}^{2^{\mathrm{T}}} & \cdots & \mathbf{q}^{n_b^{\mathrm{T}}} \end{bmatrix}^{\mathrm{T}}$$

• The corresponding generalized constraint forces are

$$\qquad \qquad \mathbf{Q}_1^i = -(\mathbf{C}_1)_{\mathbf{q}^i}^{\mathrm{T}} \boldsymbol{\lambda}_1; \ \mathbf{Q}_2^i = -(\mathbf{C}_2)_{\mathbf{q}^i}^{\mathrm{T}} \boldsymbol{\lambda}_2; \cdots; \ \mathbf{Q}_{n_i}^i = -(\mathbf{C}_{n_i})_{\mathbf{q}^i}^{\mathrm{T}} \boldsymbol{\lambda}_{n_i}$$

ullet The overall generalized constraint force \mathbf{Q}_c^i

$$\qquad \qquad \mathbf{Q}_c^i = - \sum_{k=1}^{n_i} (\mathbf{C}_k)_{\mathbf{q}^i}^{\mathrm{T}} \boldsymbol{\lambda}_k$$

One can expand to include all joints to write

$$\mathbf{Q}_c^i = -\mathbf{C}_{\mathbf{q}^i}^{\mathrm{T}} oldsymbol{\lambda}$$

 \blacktriangleright Joint reactions not associated with body i will have rows with zeros in $\mathbf{C}_{\mathbf{q}^i}$

Multi-body System



• The total vector for n_b bodies becomes

$$\mathbf{Q}_c = \begin{bmatrix} \mathbf{Q}_c^1 & \mathbf{Q}_c^2 & \cdots & \mathbf{Q}_c^{n_b} \end{bmatrix}^{\mathrm{T}}$$

$$\mathbf{Q}_c = -egin{cases} \mathbf{C}_{\mathbf{q}^1}^{\mathrm{T}} oldsymbol{\lambda} \ \mathbf{C}_{\mathbf{q}^2}^{\mathrm{T}} oldsymbol{\lambda} \ dots \ \mathbf{C}_{\mathbf{q}}^{\mathrm{T}} oldsymbol{\lambda} \ dots \ \mathbf{C}_{\mathbf{q}} = egin{bmatrix} \mathbf{C}_{\mathbf{q}^n} oldsymbol{\lambda} \ \mathbf{C}_{\mathbf{q}} = egin{bmatrix} \mathbf{C}_{\mathbf{q}^1} & \mathbf{C}_{\mathbf{q}^2} & \cdots & \mathbf{C}_{\mathbf{q}^{n_b}} \end{bmatrix}$$

- Recall earlier we had $M\ddot{q} = Q_e + Q_c$
- ullet This now becomes $\mathbf{M}\ddot{\mathbf{q}} + \mathbf{C}_{\mathbf{q}}^{\mathrm{T}} oldsymbol{\lambda} = \mathbf{Q}_e$

Augmented Formulation



- Assuming that \mathbf{Q}_e is known then we have n equations with $n+n_c$ unknowns
 - n accelerations $\ddot{\mathbf{q}}$ and n_c Lagrange multipliers $\boldsymbol{\lambda}$
- So we introduce the following set of equations by differentiating the constraint equations

$$\mathbf{C}_{\mathbf{q}}\ddot{\mathbf{q}} = \mathbf{Q}_d$$

• This leads to the augmented form

$$\begin{bmatrix} \mathbf{M} & \mathbf{C}_{\mathbf{q}}^{\mathrm{T}} \\ \mathbf{C}_{\mathbf{q}} & \mathbf{0} \end{bmatrix} \begin{Bmatrix} \ddot{\mathbf{q}} \\ \boldsymbol{\lambda} \end{Bmatrix} = \begin{Bmatrix} \mathbf{Q}_e \\ \mathbf{Q}_d \end{Bmatrix}$$

Numerical Implementation



 Now that all the co-ordinates and velocities are available one can calculate the accelerations and Lagrange Multipliers

$$egin{bmatrix} \mathbf{M} & \mathbf{C}_{\mathbf{q}}^{\mathrm{T}} \ \mathbf{C}_{\mathbf{q}} & \mathbf{0} \end{bmatrix} egin{bmatrix} \ddot{\mathbf{q}} \ m{\lambda} \end{pmatrix} = egin{bmatrix} \mathbf{Q}_e \ m{\lambda} \end{pmatrix}$$

- The Lagrange multipliers provide the generalized reaction forces
- Now we integrate forward in time
 - Using the state-space form
- This will yield the velocities and co-ordinates for the next time step
- While doing this we hold the Lagrange multiplier values fixed

Constraint Stabilization Method



- One of the issues is the growth of the equation below with time
 - $\ddot{\mathbf{C}} = \mathbf{C}_{\mathbf{q}}\ddot{\mathbf{q}} \mathbf{Q}_d = \mathbf{0}$
 - While at t = 0 this is satisfied exactly with increasing t it is not
 - ▶ In fact the form of the solution is $\mathbf{q} = \mathbf{a}_1 t + \mathbf{a}_2$
- Baumgarte proposed a stabilization scheme
 - $\ddot{\mathbf{C}} + 2\alpha \dot{\mathbf{C}} + \beta^2 \mathbf{C} = \mathbf{0}$
 - Here $\alpha > 0$ and $\beta \neq 0$
 - ightharpoonup Solution form is $\mathbf{q} = \mathbf{a}_1 e^{s_1 t} + \mathbf{a}_2 e^{s_2 t}$
- This leads to the form
 - $\mathbf{C}_{\mathbf{q}}\ddot{\mathbf{q}} = \mathbf{Q}_d 2\alpha(\mathbf{C}_{\mathbf{q}}\dot{\mathbf{q}} + \mathbf{C}_t) \beta^2\mathbf{C}$
- This form is used in the augmented form to yield

Stabilization Scheme

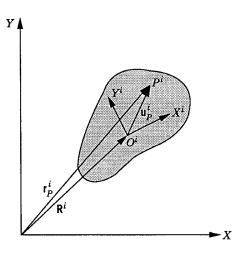


$$\begin{bmatrix} \mathbf{M} & \mathbf{C}_{\mathbf{q}}^{\mathrm{T}} \\ \mathbf{C}_{\mathbf{q}} & \mathbf{0} \end{bmatrix} \begin{Bmatrix} \ddot{\mathbf{q}} \\ \boldsymbol{\lambda} \end{Bmatrix} = \begin{Bmatrix} \mathbf{Q}_{e} \\ \mathbf{Q}_{d} - 2\alpha(\mathbf{C}_{\mathbf{q}}\dot{\mathbf{q}} + \mathbf{C}_{t}) - \beta^{2}\mathbf{C} \end{Bmatrix}$$

- All accelerations are integrated without partitioning into dependent and independent co-ordinates
- Gives good results but
- ullet No reliable method to choose lpha and eta

Reference Co-ordinates





Courtesy: A. A. Shabana, 2010, Computational Dynamics, Third Edition, John Wiley & Sons.





Cash-Karp R-K Formulas



0	0							
$\frac{1}{5}$ $\frac{3}{10}$	$\frac{1}{5}$	0						
$\frac{3}{10}$	$\frac{3}{40}$	$\frac{9}{40}$	0					
$\frac{3}{5}$	$\frac{3}{10}$	$-\frac{9}{10}$	$\frac{6}{5}$	0				
1	$-\frac{11}{54}$	$\frac{5}{2}$	$-\frac{70}{27}$	$\frac{35}{27}$	0			
$\frac{7}{8}$	$\frac{1631}{55296}$	$\frac{175}{512}$	$\frac{575}{13824}$	$\frac{44275}{110592}$	$\frac{253}{4096}$	0		(5
	$\frac{37}{378}$	0	$\frac{250}{621}$	$\frac{125}{594}$	0	$\frac{512}{1771}$	Order 5	
	$\frac{2825}{27648}$	0	$\frac{18575}{48384}$	$\frac{13525}{55296}$	$\frac{277}{14336}$	$\frac{1}{4}$	Order 4	
	$\frac{19}{54}$	0	$-\frac{10}{27}$	$\frac{55}{54}$	0	0	Order 3	
	$-\frac{3}{2}$	$\frac{5}{2}$	0	0	0	0	Order 2	
	1	0	0	0	0	0	Order 1	

Courtesy: J. R. Cash and A. H. Karp, 1990, A Variable Order Runge-Kutta Method for Initial Value Problems with Rapidly Varying Right-Hand Sides, ACM Transactions on Mathematical Software, pp. 201-222