# Challenges in estimating emissivity and surface temperature using flux tower measurements

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## **ABSTRACT**

Land surface temperature (LST) is a pre-eminent state variable that controls the energy and water exchange between the Earth's surface and the atmosphere. At landscape scale, LST is derived from thermal infrared radiance measured using space-borne radiometers. At the plot-scale, flux tower recorded longwave radiation components is inverted to retrieve LST. Since down-welling longwave component was not measured routinely for a long time a conventional practice is using Stephan Boltzman for the plot-scale LST retrieval (only up-welling longwave component). However, we hypothesize that ignoring downwelling longwave radiation for plot-scale LST estimations may lead to substantial discrepancies compared to the approach including downwelling longwave radiation. This has important implications for estimation of the correct surface emissivity using flux tower measurements, which are needed for plot-scale LST retrievals. The current work addresses these two important questions using observations at 10 eddy covariance towers operated on different land cover types. We found that the LST values obtained using up-welling and down-welling longwave component (long equation) are 0.5 to 1.5K lower compared to the equation only using up-welling longwave components(short equation). Plot-scale emissivity was estimated using observed sensible heat flux and estimated surface to air temperature difference. Plot-scale emissivity obtained using the complete equation was generally lower than if the short equation was used, for all land cover types. We also quantified the uncertainty in plot-scale LST and emissivity due uncertainty in measured fluxes. We found that despite additional input data for the long equation, the uncertainty in plot-scale LST was not greater than if short equation was used. Landscape-scale day-time LST obtained from satellite data (MODIS TERRA) were strongly correlated with our plot-scale estimates, but on average higher by several degrees, regardless of the estimation method. For most sites, the correspondence between MODIS LST and plot-scale LST estimates increased significantly if plot-scale emissivity was used instead of the landscape-scale emissivity obtained from satellite measurements. The results of this work have important implications on the combined use of aerodynamic and radiometric measurements (plot-scale emissivity estimation) for understanding of the interactions and feedbacks between LST and surface-atmosphere exchange processes.

## Introduction

The effects of global change are reflected in the surface temperature anomaly and its interannual variability<sup>1</sup>. Land surface temperature (LST) controls the magnitude and variability of the surface energy balance (SEB) components and simultaneously gets modulated by the SEB partitioning<sup>2,3</sup>. LST contains imprints of surface moisture and is immensely sensitive to evaporative cooling, which makes it a preeminent variable for studying evaporation and surface-atmosphere exchange<sup>4–6</sup>. The availability of a large network of eddy covariance measurements (FLUXNET) allows us to understand the interactions and feedbacks between the surface-atmosphere exchange processes such as evaporation, transpiration, and its control by the atmosphere and vegetation at the diurnal time scale. However, the unavailability of direct LST measurements at the same scale hinders a detailed understanding of the interactions and feedbacks between LST and surface-atmosphere exchange processes, which is of utmost importance to the climate modeling community (reference). Inversion of the longwave observations in FLUXNET data to obtain LST has been found to strongly depend on the emissivity of the underlying surface<sup>7</sup>, which is not available as routine measurement. Therefore, estimating in-situ LST is not straightforward due to the involvement of two unknowns (LST and emissivity) inside one measurement variable (upwelling longwave radiation). To circumvent this challenge, this study conducts simultaneous retrievals of LST and emissivity by exploiting the longwave radiation components in conjunction with associated SEB flux measurements.

The SEB components can be sub-divided into radiative components (often lumped in net radiation,  $R_{net}$ ) and thermodynamic

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components, including sensible, latent heat flux (H, LE respectively) and ground heat flux (G):

$$R_{net} = H + LE + G \tag{1}$$

As the surface to air temperature difference drives the exchange of sensible heat between surface and atmosphere, all components of Eq. (1) depend on the land surface temperature (LST). It directly affects the amount of emitted longwave radiation and it influences the saturation vapor pressure at the surface that drives latent heat flux. Thus, the ecohydrological functioning and carbon-water coupling are largely controlled by the surface temperature of the soil-vegetation system<sup>8</sup>.

Net radiation ( $R_{net}$ ) can be sub-divided into down-welling and up-welling components. Only a fraction of solar top-of-the-atmosphere radiation reaches the Earth's surface, as some is reflected back to space by clouds, some is absorbed by the atmosphere and emitted later as longwave radiation. The emitted longwave radiation as a function of surface temperature ( $T_s$ , K) and surface emissivity ( $\varepsilon$ ) is given by Stefan-Boltzmann (SB) equation:

$$R_{lem} = \varepsilon \sigma T_{\rm s}^4 \tag{2}$$

where  $\sigma$  ( $Wm^{-2}K^{-4}$ ) is the SB constant. Putting the radiative components together, we can sub-divide  $R_{net}$  into:

$$R_{net} = R_{sdwn} + R_{ldwn} - R_{sref} - R_{lref} - R_{lem} \tag{3}$$

Reflected shortwave in Eq. (3) is can be expressed as  $R_{sref} = \alpha R_{sdown}$ , where  $\alpha$  is the surface albedo. Similarly, reflected longwave is a function of down-welling longwave and surface emissivity ( $\epsilon$ ) and expressed as:

$$R_{lref} = (1 - \varepsilon)R_{ldwn} \tag{4}$$

LST or radiometric temperature can be estimated from the infrared radiance emanating from a given surface<sup>9</sup>. The emitted and down-welling longwave radiance are measured by an infrared radiometer at given angle within its instantaneous field of view (fov). The radiation received by a pyrgeometer or infrared sensor is a combination of the radiation emitted and reflected by the surfaces in the field of view. Thus, LST estimated from an infrared measurement is the "ensemble directional radiometric surface temperature", representing the ensemble of various surface elements present within the sensor's instantaneous fov<sup>10</sup>. A downward facing sensor relatively close to the surface (a few meters for an eddy covariance tower) essentially measures up-welling longwave radiation ( $R_{lup}$ ) from the surface, which is the sum of emitted and reflected longwave radiation:

$$R_{lup} = R_{lem} + R_{lref} \tag{5}$$

Substitution of Eqs. 4 and 2 into Eq. 5 yields  $R_{lup}$  as a function of emissivity, surface temperature and down-welling longwave radiation:

$$R_{lup} = \varepsilon \sigma T_s^4 + (1 - \varepsilon) R_{ldwn} \tag{6}$$

The above equation can be solved for surface temperature as a function of measured longwave and known surface emissivity:

$$T_s = \sqrt[4]{\frac{R_{ldwn}}{\sigma} - \frac{R_{ldwn}}{\varepsilon \sigma} + \frac{R_{lup}}{\varepsilon \sigma}}$$
 (7)

where,  $\varepsilon$  is the surface emissivity ranging between 0 and 1,  $\sigma$  ( $Wm^{-2}K^{-4}$ ) is the Stefan-Boltzman constant and  $T_s$  (K) is the LST. Emissivity is defined as efficiency of a surface to emit thermal energy relative to a perfect black body. It depends on soil type, vegetation cover, soil moisture, soil chemistry, roughness, spectral wavelength, temperature and view angle<sup>10</sup>. The radiative energy emitted at a specific wavelength is related to its temperature using Planck's function<sup>11</sup>. The emissivity of a land surface (composed of various grey body) is retrieved using radiative transfer models<sup>12–14</sup>. Thermal remote sensing (MODIS) is used to measure spectral emissivity through four channels (28, 29, 30, 31) at wavelengths ranging between 8-12  $\mu$ m<sup>15</sup> and the system of equations iteratively solved for a given range of wavelengths (9 - 12 $\mu$ m) to obtain  $\varepsilon$  and LST.

In last two decades, plot-scale radiometric data collected at eddy covariance sites (ECS) have gained popularity for in-situ LST retrieval due to its high spatial and temporal resolution<sup>?,16</sup>. In addition to this the land is also an important factor as the LST estimates relatively homogeneous footprint in comparison to MODIS pixels. These measurements are primarily use to assess the impacts and feedbacks of climate change on key ecosystem fluxes<sup>17</sup>. In order to invert LST as shown in Eq. (7),  $\varepsilon$  values are required. However, the surface emissivity of the tower footprint cannot be estimated using radiative transfer models as radiometers do not measure spectral bands separately to deduce emissivity directly.

By definition, LST is a thermodynamic temperature that can be felt or measured by an accurate thermometer at the land surface-atmosphere point-of-contact and is independent of wavelength<sup>18</sup>. The instantaneous value of LST is the result of

interplay between the net radiation at the surface, G,H and  $turLE^{19}$ . Thus, LST can also be used for the estimation of the sensible  $H^{20}$  and  $LE^{21}$  between the surface and the atmosphere. For plot-scale emissivity retreival we focus on H, which is defined as heat transfer driven by a surface to air temperature difference. It can be expressed mathematically in analogy to Ohm's law as:

$$H = \rho C_p (T_s - T_a) / r_a \tag{8}$$

$$H = m(T_s - T_a) \tag{9}$$

where  $T_a$  (K) is the temperature of the air measured at height above the surface,  $C_p(Jkg^{-1}K^{-1})$  is the specific heat capacity of air,  $\rho(kgm^{-3})$  is the air-density,  $r_a(sm^{-1})$  is the total resistance to heat transport from surface to the atmosphere.  $m(ms^{-1})$  is a proportionality constant obtained by lumping  $(\rho C_p/r_a)$  and can be broadly referred as heat transfer coefficient. m depends on surface characteristics and micro meteorology<sup>22</sup>. It is evident from Eq. (9), that for  $T_s - T_a = 0$  ( $\Delta T = 0$ ), H will be zero. This boundary condition and the linearity of H and  $\Delta T$  relationship is used to estimate  $\varepsilon$  at the plot scale from observed of H,  $T_a$  and estimated  $T_s$  using measured longwave radiation<sup>23,24</sup>. There is also other approach for plot-scale  $\varepsilon$  estimation which uses the linearity of  $H(\Delta T)$  at H close to zero will result into  $T_s = T_a^{25}$ .

However, due to surface heterogenity sparse canopies are prone to footprint mismatch between the aerodynamic (flux tower) footprint and radiometric (hemispherical) footprint<sup>26–28</sup> which can result into a different boundary condition i.e at  $\Delta T = 0, H \neq 0$  as shown in Eq. 10. Aerodynamic footprint represents the area contributing to measured sensible heat, air temperature and the hemispherical footprint is the area within radiometric FOV contributing to the measured longwave (used for  $T_s$  estimation).

$$H = m(T_s - T_a) \pm c(H) \tag{10}$$

H is representative of the sensible heat flux from the eddy covariance tower footprint,  $T_s$  is representative of all the radiating surfaces in the sensor's view. c is interpreted as the H from aerodynamic footprint which is not seen by the radiometer.

Plot-scale  $\varepsilon$  estimation using observed  $H, T_a, R_{lup}, R_{ldw}$  as described above, may be prone to substantial uncertainty. It is unclear how uncertainties in observed fluxes propagates into the uncertainty of estimated LST and  $\varepsilon$ . By design, IRT sensors only measure upwelling infrared radiance and therefore cannot explicitly account for the amount of reflected down-welling infrared radiation in the signal. For a long time, down-welling longwave  $R_{ldw}$  was not routinely observed at ECS<sup>29</sup> and was also considered to be the most poorly quantified component of the radiation budget<sup>30</sup>. Therefore, the second term in Eq. (6) is commonly omitted, arguing that  $\varepsilon \approx 1$ , and therefore Eq. (6) is simplified to Eq. (2)<sup>31</sup>:

$$R_{lup} \approx \varepsilon \sigma T_s^4$$
 (11)

Eq. 11 can be solved for  $T_s$  to yield what we will term the "short equation" (SEQ) for  $T_s$ :

$$T_s \approx \sqrt[4]{\frac{R_{lup}}{\varepsilon\sigma}}$$
 (12)

Note that the above derivation is actually flawed, as the second term of Eq. (6) was omitted arguing that  $\varepsilon \approx 1$ , and yet  $\varepsilon$  was retained in the first part of the equation. Nevertheless, even with the availability of down-welling longwave measurements<sup>32</sup>, the use of Eq. (11) is still a common practice<sup>7,33</sup>. This gives rise to the question if the short equation (Eq. 12) is adequate to estimate LST from ground-based measurements. In the remainder of this paper, we will refer to LST obtained using the complete or "long equation" (Eq. 7) as  $T_{leq}$  and to LST obtained using the simplified or short equation (Eq. 12) as  $T_{seq}$ .

In the present study, we evaluate and analyse the differences in plot-scale LST and emissivity estimated using the complete (Eq. 7) and short (Eq. 12) equations. The following research questions are addressed:

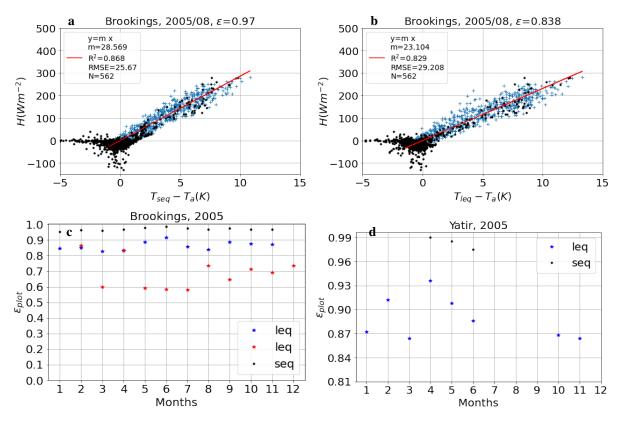
- 1. Is the short equation (Eq. 12), neglecting the reflected down-welling longwave radiation, adequate to estimate LST from ground-based measurements?
- 2. Does the estimation of plot-scale  $\varepsilon$  based on observed sensible heat flux H has an advantage over satellite derived  $\varepsilon$  for LST estimation?
- 3. How much uncertainity is introduced in plot-scale LST and  $\varepsilon$  due to errors in measured EC fluxes?

To answer these questions, we analysed data for 10 eddy covariance sites in different biomes and climates (see Table 2). Plot-scale broadband monthly emissivity was derived using observed H and estimated  $\Delta T$  as proposed by Holmes et al.<sup>23</sup>. Plot-scale LST was estimated using plot-scale or landscape-scale emissivity with (Eq. 7) and (Eq. 12). Plot-scale LST were compared with MODIS LST (TERRA satellite-sensed) for the times of satellite over-pass. Uncertainty in emissivity and LST due to error in observed fluxes was calculated using SOBOL based uncertainty analysis (SAlib)<sup>34</sup>. See the Methods section for more details.

## 1 Results

## 1.1 Plot-scale $\varepsilon$ using long and short equation

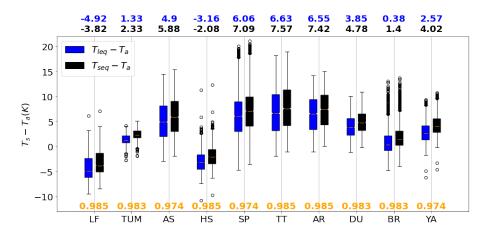
Following the method proposed by Holmes et al.  $^{23,35}$ , plot-scale monthly emissivity ( $\varepsilon$ ) was estimated at the study site by fitting  $\varepsilon$  to minimise the root mean square error (RMSE) of the regression between H and  $\Delta T = T_s - T_a$ . In Fig. 1a, c, and d, we reproduced the original data of Figs. 2a, 3C, and 3Q from Holmes et al.  $(2009)^{23}$  to validate our interpretation of their approach using the simplified equation (Eq. 12). We noted only marginal differences between the two results based on the short equation which are likely due to fitting algorithms. The replication of the  $H(\Delta T)$  plot using the long equation (Eq. 7) in their analysis is given in Fig. 1b and  $\varepsilon$  resulting is shown in Fig. 1c,d by blue stars. The use of complete equation (Eq. 7) resulted in substantially lower values of  $\varepsilon$  (10%) as compared the values estimated by Holmes et al  $^{23}$  for the common study sites (Brookings, Yatir Forest) (Fig. 1c and d). Note that the long equation also yielded  $H(\Delta T)$  relationship for many more months at Yatir Forest (blue star) than the short equation (black dots) as shown in Fig. 1d), as it resulted in achieving a strong correlation between H and  $\Delta T$  ( see methods section for details). The retrieved LST values were slightly higher using Eq. (7) (compare a and b in Fig. 1). The pattern of low  $\varepsilon$  and higher LST using the long equation compared to the short equation was confirmed for all 10 sites used in the present study (Table SI 2).



**Figure 1.** Reproduction of analysis presented in Figs. 2 (a) and 3 C and Q in Holmes et al.  $(2009)^{23}$ . (a) Sensible heat (H) vs.  $\Delta T = T_{seq} - T_a$  based on the short equation  $(T_{seq}, \text{Eq. 12})$ ; (b) H vs.  $\Delta T$  based on the complete equation  $(T_{leq}, \text{Eq. 7})$ . Both show data for August 2005 at Brookings. Blue crosses mark data points satisfying the filtering criteria while black dots mark points not considered in the analysis. N is the number of blue crosses used for regression (red line), m is the slope of regression, RMSE is the root mean square error and  $R^2$  is the coefficient of determination. The fitted  $\varepsilon$  value is reported in the title. (c) Optimised  $\varepsilon$  values at Brookings obtained for the months where  $R^2 > 0.5$  using the short equation (Eq. 11, black dots) and complete equation (Eq. 6, blue stars). (d) Same as (c), but for Yatir Forest. See Table 2 for site descriptions.

#### 1.2 Landscape scale vs plot-scale estimates of $\varepsilon$ and LST

At each site, LST was estimated using both the short equation ( $T_{seq}$ , Eq. 12) and the long equation ( $T_{leq}$ , Eq. 7). In the first step, tower-based longwave measurement and landscape-scale broadband  $\varepsilon$  from MODIS ( $\varepsilon_{MODIS}$ , Eq. 13) was used. The yearly daytime surface to air temperature difference for each study site is estimated and shown in Fig. 2. At all sites, Eq. 12 resulted in higher day-time plot-scale  $T_s$  estimates as compared to Eq. 7, when using  $\varepsilon_{MODIS}$ , with the medians of surface to air temperature differences ( $\Delta T$ ) differing by 0.8 to 1.5 K (Fig. 2). The difference in  $\Delta T$  using two equation is highest at the water limited sites e.g AS, YA. Note that for two sites (LF and HS), the median values of daytime  $\Delta T$  are negative.



**Figure 2.** Yearly distributions of half-hourly surface to air temperature differences ( $\Delta T = T_s - T_a$ ) for a representative year at each site. LST inverted using short and long equation (Eq. 12, Eq. 7) with landscape-scale emissivity ( $\varepsilon_{MODIS}$ ). The median values of  $\Delta T$  are shown at top of the plot and the emissivities used for the  $T_s$  retrieval are shown at the bottom in orange. See Table 2 for site abbreviations.

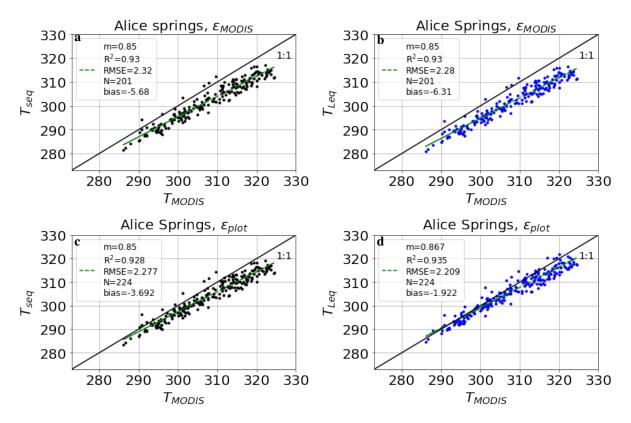
Comparison of estimated plot-scale LST using  $\varepsilon_{MODIS}$  at satellite overpass time with landscape-scale LST ( $T_{MODIS}$ ) revealed strong correlations between plot-scale and landscape-scale LST estimates as shown in Fig. 3a, b. Use of plot-scale  $\varepsilon_{plot}$  for plot-scale LST estimation ( $T_{seq}$  and  $T_{leq}$ ) resulted in substantial reduction of the bias as shown in Fig. 3c, d. This trend in bias reduction was similar to other sites (Table SI2 for details). The minimum bias is found at the Tumbarumba a closed canopy (eucalyptus forest) and highest bias was obtained at Litchfield, Howards Springs ecosystems with sparse canopies (woodland savanna). However, for some sites, weak correlation between satellite-derived and local LST estimates were also evident (at DU,  $R^2$  was reduced from 0.8 to 0.4, see Table SI2). Also, using plot-scale  $\varepsilon$  for LST estimation resulted in positive  $T_s - T_a$  at LF and HS as shown in SI3, Fig. (8).

#### 1.3 Plot-scale $\varepsilon$ estimation considering footprint mismatching of measurement devices

In order to account for the discrepancy in plot-scale  $\varepsilon$  estimation due to footprint mismatch of measuring devices the linear regression between  $H,\Delta T$  was allowed to accept an intercept (instead of forcing it to zero as in Fig. 1). As shown in Fig. 4 title the plot-scale  $\varepsilon$  using long equation with an intercept results into the conventional (realistic) range of emissivity which is in contrast to the estimates without intercept as shown in Table (1). The intercept value for Brookings, a grassland is 12% of maximum sensible heat flux ( $H_{max}$ ) and for a mulga shrubland like AS is 20% of  $H_{max}$ . The intercept value for HS is 70% of  $H_{max}$  and for Tum is 7% of  $H_{max}$ . The intercept value is lowest for a forest site (closed canopy) and highest for a HS, LF (sparse canopy: clumping of tree, grass and bare soil). The comparison of resulting plot-scale LST with landscape-scale LST shows an increase in bias with respect to the LST obtained using  $\varepsilon$  without an intercept as shown in Table (1).

## 1.4 Uncertainty in plot-scale $\varepsilon$ and LST due to measurement error

Each of the observed input variables used for the estimation of plot-scale  $\varepsilon$  and LST has a certain level of measurement uncertainty associated with it. To quantify the propagation of these uncertainties using the two equations(Eq. 12, Eq. 7) without and with an intercept in H versus  $\Delta$  T relationship, we plotted the ranges of results obtained by random perturbations of each input variable within its own measurement (uncertainty) bounds. Here we present results for Alice Springs, which showed the highest correlation between plot-scale and landscape-scale estimations of LST (Table SI2). The uncertainty in plot-scale  $\varepsilon$  estimated using Eq. 7 without allowing an intercept for  $H(\Delta T)$  ranged between 0.68 and 0.98 whereas  $\varepsilon$  estimated with allowing an intercept  $H(\Delta T + c)$  revealed very constrained values as shown by the green boxes in Fig. 5a. The short equation (Eq.12) also resulted in constrained values between 0.94 and 0.99 shown by the black boxes (Fig. 5a). The uncertainty in



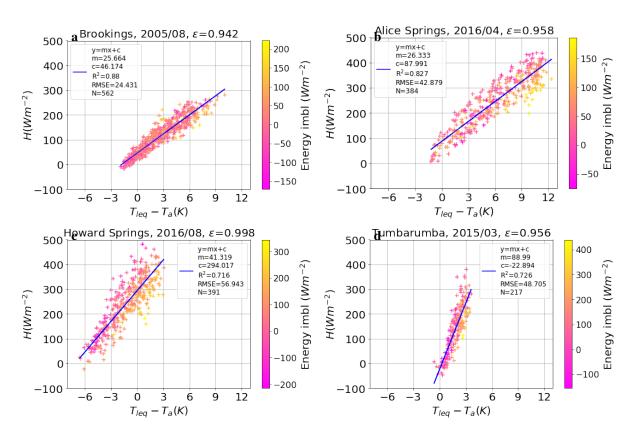
**Figure 3.** Landscape-scale LST ( $T_{MODIS}$  derived from MOD11A1) vs. plot-scale LST at Alice Springs for 2016-2018. (a)  $T_{seq}$  based on short equation (Eq. 12) and satellite-derived (MODIS) landscape-scale broadband emissivity; (b) Same as (a), but  $T_{leq}$  based on complete equation (Eq. 7); (c)  $T_{seq}$  based on short equation (Eq. 12) and monthly plot-scale emissivity; (d) Same as (c), but  $T_{leq}$  based on complete equation (Eq. 7). Bias is mean  $T_{seq} - T_{MODIS}$ , N is the number of daily overpasses of MODIS between 2016 and 2018, c is the intercept, m the slope, RMSE is the root mean square error and  $R^2$  is coefficient of determination. At each site, LST was estimated using both the short equation ( $T_{seq}$ , Eq. 12) and the long equation ( $T_{leq}$ , Eq. 7). In a first step, we used satellite-derived landscape-scale broadband emissivity from MODIS ( $\varepsilon_{MODIS}$ , Eq. 13) for estimating plot-scale LST from tower-based longwave measurements, and compared these with landscape-scale LST extracted from daily MODIS LST images ( $T_{MODIS}$ )

hourly  $T_s - T_a$  is estimated using uncertainty in plot-scale  $\varepsilon$  was quite similar for both complete and simplified equation (blue and black boxes in Fig. 5b, c. The uncertainty in  $T_s - T_a$  was higher when landscape-scale  $\varepsilon$  was used as shown by the orange boxes (Fig. 5b, c).

#### 2 Discussion

Our analysis revealed the deficiency in commonly used short equation for estimating plot-scale LST and  $\varepsilon$ . The use of complete equation for plot-scale  $\varepsilon$  estimation results into lower values for sparse canopy. In order to obtain realistic  $\varepsilon$  for sparse canopy conceding footprint mismatch between radiometric and aerodynamic footprint becomes important.

Depending on the equation (Eq. 7 or Eq. 12) used to estimate LST, small approximation (error) in  $\varepsilon$  can lead to large error in LST. The sensitivity of the long equation (Eq. 7) to  $\varepsilon$  is driven by the contrast between  $R_{lup}$  and  $R_{ldwn}$  whereas for short equation (Eq. 12) it is only driven by observed  $R_{lup}$ . For instance an error of 0.01 in  $\varepsilon$  at water limited sites like AS can cause an error of 0.17K using Eq. 7 and 0.79K using Eq. 12 respectively. This shows a clear advantage of Eq. 7 for plot-scale LST estimation since most of the time  $\varepsilon$  is used as an approximate value. Also plot-scale  $\varepsilon$  estimation using long equation (Eq. 7) results in  $H(\Delta T)$  relationship for more months (Fig. 1d). However,  $\varepsilon$  estimates using Eq. (7) are lower than Eq. (12) and in particular for sites like HS and LF the  $\varepsilon$  was unrealistic Table.**SI** 4 in comparison to the previously reported  $\varepsilon$  values for a soil-vegetation system<sup>36,37</sup>. The lower  $\varepsilon$  value using correct equation (Eq. 7, Fig. 1b) Table 1 suggests problem due to combining measurements coming from instruments characterized by different footprints<sup>27</sup>. The mismatch of source areas (footprint) of radiometric and aerodynamic flux measurements becomes important if the surface underlying the instruments is heterogeneous (e.g. AS,TT, YA, AS, HS, LF). The mismatch in footprint can further results into  $H \neq 0$  at  $\Delta T = 0$ . This

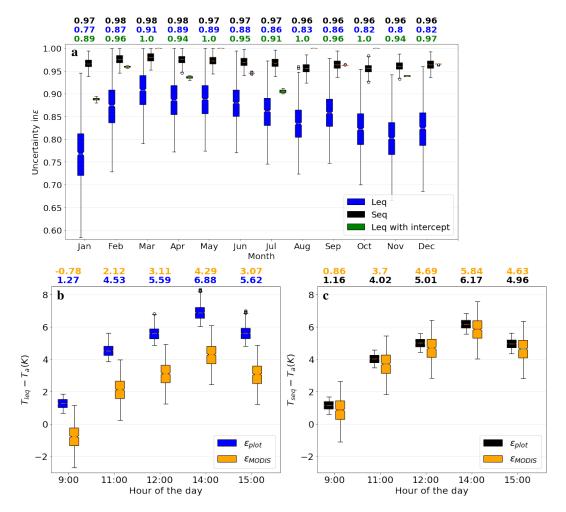


**Figure 4.** Using robust linear regression  $H = m(T_s - T_a) + c$ . The  $\varepsilon$  is estimated by minimising RMSE. The title of the plot shows site, year, month and the  $\varepsilon$ . The legend correspond to Fig. 1 and c is interpreted as the H from the aerodynamic footprint which is not seen by the radiometer.

Sites	Landscape-scale $arepsilon$			Plot-scale $\varepsilon$			Plot-scale $\varepsilon$				
Sites				$H=m*(\Delta T)$			$H = m * (\Delta T) + c$				
	$arepsilon_{land}$	$R^2$	bias (K)	$arepsilon_{plot}$	$R^2$	bias (K)	$arepsilon_{plot}$	$R^2$	bias (K)	$c (Wm^{-2})$	
SP	0.974	0.81	-4.61	0.85	0.82	-1.91	0.92	0.774	-2.563	18.12	
AS	0.974	0.93	-6.24	0.82	0.93	-1.92	0.993	0.915	-4.884	72.46	
TT	0.974	0.57	-8.30	0.80	0.52	-4.02	0.939	0.521	-7.466	58.70	
HS	0.985	0.16	-9.90	0.6	0.22	-2.47	0.949	0.18	-10.45	237.29	
LF	0.985	0.41	-11.0	0.6	0.41	-2.57	0.968	0.378	-11.47	258	
AR	0.985	0.27	-3.51	0.960	0.252	-2.98	0.996	0.27	-3.567	14.72	
DU	0.985	0.81	4.61	0.985	0.425	-3.926	0.994	0.405	-4.603	-8.11	
TUM	0.983	0.84	-2.10	0.97	0.89	-1.93	0.955	0.85	-1.696	-24.24	
BR	0.983	0.937	-0.195	0.82	0.895	2.72	0.919	0.906	1.662	17.72	
YA	0.974	0.855	-3.45	0.93	0.793	-0.582	0.873	0.826	0.073	-22.95	

**Table 1.** Comparison of daytime landscape-scale LST ( $T_{MODIS}$ ) with plot-scale LST ( $T_s$ ) estimated at TERRA time of pass using long equation. The emissivity used to retrieve plot scale LST is derived using relationship without intercept  $H(\Delta T)$  and with intercept ( $H((\Delta T) + c(H))$ )at study sites are reported. The reported plot-scale  $\varepsilon$  and intercept (c) are median values and landscape emissivity are MODIS based. Bias is defined as mean of  $T_s - T_{MODIS}$ ,  $R^2$  is coeffecient of determination between plot-scale LST in comparison to landscape-scale LST. The site acronyms can be found in Table 2.

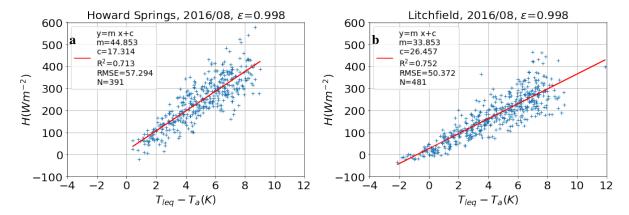
problem was not detected for Holmes et.al<sup>38</sup> as short equation (Eq. 12) was used and due to its high sensitivity to  $\varepsilon$  (SI Fig. 10 a) even with small reduction in  $\varepsilon$  the offset in  $H(\Delta T)$  was corrected (Fig. 1a). Whereas due to lower sensitivity of Eq. 7 to  $\varepsilon$  (SI Fig. 10 a) larger reduction is required to get rid of intercept resulting into lower  $\varepsilon$  as shown in Fig. 1b. We proposed an improvement to the method by allowing an intercept in  $H(\Delta T)$  regression (robust regression). Considering aerodynamic footprint to be greater than radiometric footprint<sup>27</sup> the offset was interpreted as the H from aerodynamic footprint which is not



**Figure 5.** Uncertainty in monthly emissivity for 2017 and hourly  $T_s - T_a$  for 15 August 2017. (a) Uncertainty in monthly optimised emissivity due to uncertainty in H,  $R_{lup}$ ,  $R_{ldw}$ ,  $T_a$  at using Eq. (6) and Eq. (11) shown in Blue and black for 2017 at AS. (b) Hourly daytime uncertainty in  $T_s - T_a$  due to perturbed fluxes and optimum  $\varepsilon$  using Eq. (6) in green and perturbed fluxes with MODIS based  $\varepsilon$  in orange(Fig. (2). (c) Hourly daytime uncertainty in  $T_s - T_a$  due to perturbed fluxes and optimum  $\varepsilon$  using Eq. (11) in black and perturbed fluxes with MODIS  $\varepsilon$  in orange. Unperturbed optimum  $\varepsilon$  values and  $T_s - T_a$  values correspond to the median of perturbed values

seen by the radiometer. And it resulted into realistic  $\varepsilon$  (Fig 4a, b, c, d) but intercept was very high for sites like HS and LF Table. 1. Looking closely at  $H(\Delta T)$  plots for HS and LF and placing high value of intercept and lower daytime  $T_s - T_a$  into perspective we could hypothesize an underestimation of  $R_{lup}$ . Testing the hypothesis for HS we found that adding roughly 40 Wm-2 (approx 8% of observed  $R_{lup}$ ) in observed  $R_{lup}$  led to significant reduction in the intercept from 294Wm-2 (Fig. 4 c) to 17Wm - 2 with positive  $T_s - T_a$  as shown in Fig. 6 a). The other linear regression parameter (m,  $R^2$ , RMSE) do not remained constant for Fig. 6a and Fig. 6c. Reminding ourselves the hemispherical view of the radiometers looking down at the canopy makes it a possible scenario a sparse canopy like HS and LF<sup>27,39</sup>. And radiometer will see more tree crowns and less soil which can lead to appoximately 5-10% underestimation of  $R_{lup}$  (SI 6, Fig. 12). As shown in Fig. 4 the offset was proportional to the maximum observed H for each and an illustration plot for 36 months for AS is shown in Fig. 4 (SI 7 Fig. 13). The ratio between  $H_{max}$  and intercept (c) varies between 0% to 30% for AS. The seasonality of ratio can be attributed to the change in solar angle and the dominant wind direction. Surface heterogeneity has also been recognized as one of the potential causes of the lack of energy balance closure observed at most ECS<sup>40,41</sup> at diurnal scales which further leads to underestimation of turbulent fluxes (H, LE). Therefore as a prerequisite the observed turbulent fluxes are corrected using an energy balance closure scheme 42 and used. However, in our analysis the use of energy balance closure scheme (based on Bowen ratio) led to much lower value of plot-scale  $\varepsilon$  using Holmes approach<sup>23,35</sup>. Similar studies on plot-scale  $\varepsilon$  estimation have used the observed fluxes without correction  $^{25,38,43,44}$ . The use of energy balance closure scheme for the for plot-scale  $\varepsilon$  estimation led to an increase in positive intercept (sparse canopy, Fig. 4). We also looked into the intercept of  $H(\Delta T)$  at minimum range of energy imbalance also results into high intercept as shown in Fig 6 b. Our analysis shows that although surface heterogenity is one of the common

reason for the intercept in  $H(\Delta T)$  and lack of energy balance closure but the relation between the two is non conclusive.



**Figure 6.**  $Hvs\Delta T$  using reboust linear regression, the crosses (see Fig. 4 c ) are color coded according to the energy imbalance  $(R_n - H - LE - G)$  (a) By adding 40  $(Wm^{-2})$  to measured  $R_{lup}$  led to significant reduction in the intercept from 294Wm - 2 (Fig. 4 c) to 17Wm - 2 with increase in  $T_s - T_a$ . (b) Choosing data points where energy imbalanceis between -50 to 50  $(Wm^{-2})$  intercept is increased from 294Wm - 2 (Fig. 4 c) to 345Wm - 2. The legends crossponds to Fig. 4 Comparison of plot-scale LST with MODIS LST resulted into very high bias for sparse canopy which is in agreement with

Comparison of plot-scale LST with MODIS LST resulted into very high bias for sparse canopy which is in agreement with previous studies where the bias for sparse canopy was high upto  $12K^{45}$ . Plot-scale LST estimate using plot-scale  $\varepsilon$  decreased the bias in comparsion to MODIS LST as shown in Table 4. The use of plot-scale  $\varepsilon$  also reduces the uncertainty in diurnal LST in comparsion to landscape-scale  $\varepsilon$  (Fig. 5 b,c). However LST estimated using plot-scale  $\varepsilon$  considering foot print mismatch  $(H = m * \Delta T + c)$  resulted into increase in bias for most of the study site as shown in Table. 1. The increase in bias could be attributed to implicit consideration of spatial heterogenity (by accepting intercept) at plot-scale whereas, MODIS LST (landscape-scale) are integrated pixel information considering radiating signal from a homogeneous surface.

The fluxes observed at a soil-vegetation system is representative of the composite signal from both soil and vegetation which typically have a different range of surface temperatures and emissivities  $^{15}$ . The  $\varepsilon$  of soil strongly depends on soil moisture content<sup>46</sup>. Whereas the emissivity of a canopy depends on its structure attributes and leaf area index, the latter of which varies strongly at the seasonal scale<sup>47</sup>. For example, laboratory-measured directional  $\varepsilon$  for various canopy elements (bark, leaf and its arrangement, stem wood) ranged between 0.9 to 1<sup>48</sup>. Laboratory measurements of thermal infrared reflectance spectra suggests that the  $\varepsilon$  variation from structural unknowns, such as leaf orientation, is more significant than the differences in leaf component emissivity among plant species<sup>37</sup>. Consequently it is clear that the  $\varepsilon$  of a surface is a function of many factors and detailed study of all these factors is out of scope of the present study. Thus, derivation of landscape-scale broadband  $\varepsilon_{land}$  from narrowband spectral emissivity is considered as a first-order approximation for capturing the integrated effects of land cover from MODIS spectral bands<sup>15</sup>. However, due to general lack of emissivity information for natural ecosystem at ECS, remotely sensed emissivity serves as plausible estimates and are valuable input for plot-scale LST estimation. Plot-scale  $\varepsilon$  by fitting observed H and estimated  $\Delta T$  without an intercept results into a lower value and is statistically questionable<sup>49</sup>. Since fluxes observed at ECS contain composite (soil and vegetation) signal and is represented as area weighted average plot-scale  $\varepsilon$  by allowing an intercept  $H(\Delta T)$  is a statistically correct and results into realistic  $\varepsilon$  for sparse canopy. Estimation of an intrinsic property like surface emissivity using time varying fluxes can be questioned however considering the spatial and temporal limitation of satellite derived data the plot-scale  $\varepsilon$  estimation is a good alternative.

Our result shows that under no condition for grey bodies the use of simplified equation and complete equation can lead to equal values of  $\varepsilon$  and LST and thus the use of the SEQ is not recommended. Considering the widely spread use of both ECS recorded fluxes and MODIS LST in LSMs applications, the use of plot-scale  $\varepsilon$  can be advantageous as it reduces the bias between plot-scale and MODIS LST. Therefore, the approach proposed in this work has the potential to provide a more robust benchmark information for model calibration and validation. Plot-scale  $\varepsilon$  reduces also the uncertainty in diurnal LST due to measurement errors. Overall, the implications of our study are of particular relevance for the research community interested in understanding diurnal and seasonal feedback in soil-vegetation systems based on fluxes, emissivity and LST estimated at a consistent scale.

## 3 Methods

#### 3.0.1 Research data

**Tower data** ECS collect micro-meteorological measurements above the surface (vegetation canopy) using towers (flux tower), following common measurement protocols<sup>17</sup>. The towers are commonly equipped with an instrument made up of

pyrgeometers or radiometers to measure up-welling and down-welling shortwave and longwave radiation, which is further used to calculate net radiation (Eq. 3). Besides radiative fluxes, measurement at EC towers also include sensible and latent heat fluxes, net carbon dioxide exchange and a range of meteorological variables, such as air temperature, humidity and wind speed.  $T_a$  is the air temperature measured at a reference height above the canopy. Each flux measurement is accompanied by a flagging system and follows the agreement with second CarboEurope-IP QA/QC workshop<sup>50</sup>. Flag 0 is designated as high data quality, and have been used in our current work. For the analysis 10 sites were selected to represent a variety of land cover types and climates (Table 2). 8 sites belong to the North Australian Tropical Transect (NATT) and 2 sites (Yatir Forest, Brookings) are chosen to replicate result from Holmes et.al<sup>38</sup>. Eddy covariance Level 3 data is obtained from http://data.ozflux.org.au/portal/pub/listPubCollections.jspx for Australian sites. The data for Brookings in obtained from ameriflux and for Yatir Forest the data was obtained using personal communication to Prof Dan Yakir's lab group as the data used in Holmes work was from  $2005^{23}$  and the current available data was an updated version of the fluxnet data.

**MODIS data** Landscape- scale emissivity and LST data (MODIS product MOD11A1) was downloaded from NASA earth data . It is a level 3 daily LST product gridded in the sinusoidal projection at a spatial resolution of 0.928 km by 0.928 km. A tile contains 1200 x 1200 grids in 1200 rows and 1200 columns<sup>51</sup>. The daily LST pixel values in each granules is retrieved by the generalized split-window algorithm under clear-sky conditions and LST values are averaged by overlapping pixels in each grid with overlapping areas as weight. The downloaded data are in Hierarchical Data Format (hdf) which were converted into tagged image file format (tiff) using a python package called PyModis<sup>52</sup>. From tiff the files are converted into csv format. The details of data extraction and conversion can be found at https://renkulab.io/projects/gitanjali.thakur/modislstfpar/. The dataset columns used to compare plot-scale LST and emissivity are: day time daily LST, local view time, night view time, channel 31, 32 spectral  $\varepsilon$  respectively.

Site-specific approach This approach was initially proposed by Holmes<sup>38</sup> to estimate plot-scale  $\varepsilon$  using short equation with  $H, R_{lup}, T_a$ . In present work we have used both long equation (Eq. 7) and short equation (Eq. 12) to estimate  $\varepsilon$  The prime variable used in the study are  $H, R_{lup}, R_{ldwn}, T_a$  and the ancillary variable are  $R_n$  and wind speed (Ws) are used to filter the data for analysis. The filtering condition were  $R_n > 25Wm^-2$ ) and wind speed  $(Ws > 2ms^-1)^{23}$ . Plot scale emissivity is derived by segregating each month data satisfying the filtering criteria. For each month observed H,  $R_{lup}$ ,  $R_{ldwn}$ ,  $T_a$  using filtered measurements (based on net radiation and wind speed). For estimating  $T_s$  from longwave measurement( $R_{lup}$ ,  $R_{ldwn}$ ) a predefined range of  $\varepsilon$  is chosen starting from 0.998 and is reduced at a step size of 0.002. For each month data  $H(T_s - T_a)$  linear regression is performed using scipy https: //docs.scipy.org/doc/scipy - 0.14.0/reference/generated/scipy.stats.linregress.html. The  $R^2$  is checked and months where  $R^2 > 0.5$  (i.e substantial part of variance in H is explained by  $T_s - T_a$ ) is chosen for  $\varepsilon$  estimation. To obtain epsilon regression line of  $H, \Delta T$  is forced through origin (at  $T_s - T_a = 0$  H = 0) by minimising the RMSE. An illustration plot for RMSE and  $\varepsilon$  is shown in SI5. The  $\varepsilon_{plot}$  is obtained for each month using both long Eq.(7) and short equation Eq.(12) and termed as  $\varepsilon_{seq}$  and  $\varepsilon_{leq}$  as shown in Fig. 7b. In the next step instead of forcing the  $Hvs\Delta T$  plot through origin robust linear regression is used, which allowed an offset in monthly  $H(\Delta T)$  plots. The plot-scale  $\varepsilon$  was obtained by minimising RMSE as explained before. Plot-scale LST estimated using plot-scale  $\varepsilon$  was compared to landscape-scale MODIS LST. Monthly tower based longwave measurement at study sites corresponding to TERRA daily time of pass was obtained using linear interpolation. TERRA (satellite) overpasses at local solar time between 10:30 am to 12 pm in ascending mode  $^{18}$ . Plot-scale LST ((  $T_{leg}$  and  $T_{sea}$ ) was obtained using interpolated day time longwave measurement and crossponding month plot-scale  $\varepsilon$  using Eq. (7), Eq. (12). Plot-scale daily LST is compared to MODIS LST in terms of mean, bias, RMSE and  $R^2$  using a robust linear regression model(scipy stat model) as shown in Fig. 7a. The goodness of fit between plot-scale and landscape-scale LST was determined by looking at  $(R^2)$  as shown in Fig. 7b. The bias is estimated as mean of deviation between daily MODIS LST and ground based  $T_s$  ( $T_{leq,seq} - T_{MODIS}$ ). See SI table for data sources and acronyms in SI1

**General Approach** In this approach we estimate Landscape-scale  $\varepsilon$  (Broadband) is estimated using MODIS spectral  $\varepsilon$  as shown<sup>53</sup>

$$\varepsilon_{MODIS} = 0.4587\varepsilon_{31} + 0.5414\varepsilon_{32} \tag{13}$$

Tower based longwave measurement ( $R_{lup}$ ,  $R_{ldwn}$ ) passing the filtering criteria (as mentioned in site- specific approach) along with MODIS based  $\varepsilon$  was used to invert LST using Eq. (7) and Eq. (12). Plot-scale LST estimated using MODIS based  $\varepsilon$  was compared to landscape-scale MODIS LST using a robust linear regression as done in site specific approach as shown in Fig. (7a).

Uncertainty estimation The turbulent fluxes of H measured at the ECS are averaged over smaller time (30 or 60 minutes) and the averaging over this time interval is typically dominated by random errors<sup>54,55</sup>. The other source of error is the radiation measurement which can vary between sites, due to the difference in accuracy and precision of the in-situ radiometers or pyrgeometer, differences in field of view (FOV), wavelength bands, and spectral response functions of the sensors. For reference, the accuracy of dedicated infrared temperature (IRT) sensors is given by their manufacturers as 0.5K, but it is not clear what sources of uncertainty are considered in this estimation<sup>56</sup>. The uncertainty in plot-scale  $\varepsilon$  due to error is observed fluxes is

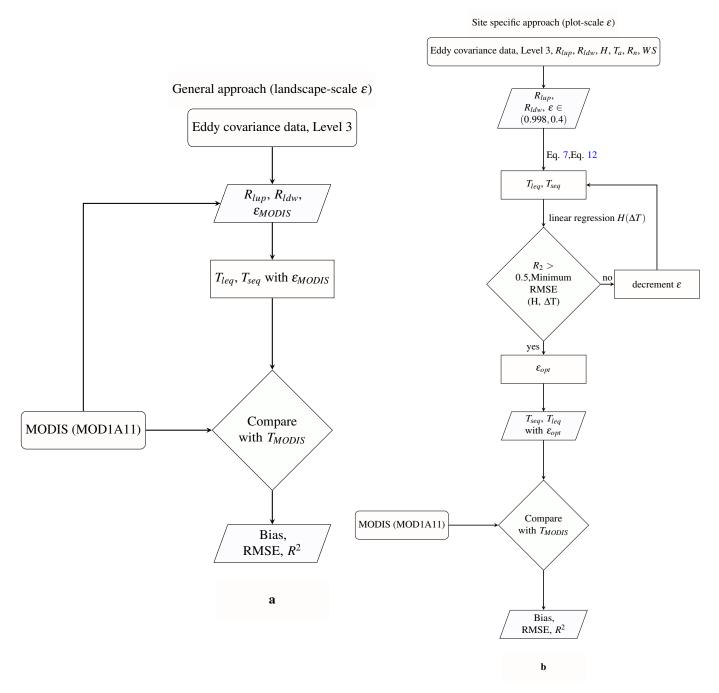
**Table 2.** Description of study sites

Study site	latitute, longitude	Landcover	Time- period	longwave sensors	Sensor instal- lation height (m)
Sturt Plains (SP)	-17.1507, 133.3502	Mitchell Grass	2016- 2019	CG-2	4.8
Alice Springs (AS)	-22.2828, 133.2493	Mulga wood- land, hummock grassland, river red gum forest	2016- 2018	CNR1	12.2
Ti Tree East (TT)	-22.2870, 133.6400	Grassy mulga woodland, Corym- bia/Triodia savanna	2016- 2018	CNR1	9.9
Howard Springs (HS)	-12.4943, 131.1523	Open woodland savanna	2016- 2018	CM-7B, CG-2	23
Litchfield (LF)	-13.1790, 130.7945	Tropical sa- vanna	2016- 2018	CNR4	31
Adelaide River (AR)	-13.0769, 131.1178	Savanna dominated by Eucalyptus tectifica and Planchonia careya	2006- 2009	CNR1	15
Daly Uncleared (DU)	-14.1592, 131.3881	Woodland savanna	2016- 2018	NRlite	21
Tumbarumba (TUM)	-35.6566, 148.1517	Wet sclerophyll	2015- 2018	CM3 and CG3	70
Brookings (BR)	44.352, 96.840	Cropland	2005	pyrgeometers <sup>18</sup>	NA
Yatir Forest (YF)	35.052, 31.345	Evergreen needleleaf forest	2005	pyrgeometers <sup>18</sup>	NA

estimated. In first step, the error bounds of each input variables  $(H, R_{lup}, R_{ldw}, T_a)$  is assumed. The error bounds for  $R_{lup}R_{ldwn}$  are -5 to 5 W m<sup>-230</sup>, for H is -20 to 20 W m<sup>-2</sup> and for  $T_a$  it is -1 to 1 K<sup>42</sup>. Uniformly distributed samples within error bounds are generated using saltelli sampling scheme<sup>57</sup> using python based package name SALIB. Each error samples are added to the monthly segregated measured fluxes. The plot-scale  $\varepsilon$  is estimated for the fluxes with added uncertainty (error sample) as explained in site specific approach. The obtained range of plot-scale  $\varepsilon$  is reported as uncertainty in emissivity . The plot-scale  $\varepsilon$  uncertainty with uncertainty in  $R_{lup}, R_{ldw}$  is used to calculate the uncertainty in hourly LST using short equation Eq. 12 and long equation Eq. 7 as shown in Fig. 5b,c. .

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**Figure 7.** Schematic representation of steps followed for plot scale LST retreival using eddy covariance measurement(a) Landscape emissivity and longwave measurement and compared to Landscape-scale LST( $T_{MODIS}$ ).(b) Plot-scale emissivity estimation using observed H,  $R_{ldwn}$ ,  $R_{lup}$  and plot-scale LST is estimated using plot-scale  $\varepsilon$  is compared to landscape-scale emissivity. The  $R^2$ , RMSE, Bias are mentioned in Fig. (3)

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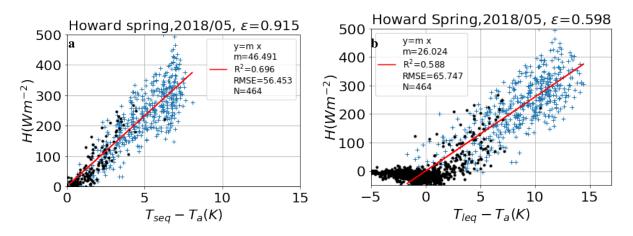
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# **5 Supplementary Information**

- SI1. Abbreviation list
- SI2. Comparison table of plot-scale LST with landscape LST using landscape and plot-scale arepsilon
- SI3. Emissivity estimation at Howards spring and positive  $T_s T_a$



**Figure 8.**  $Hand\Delta T$  plots at Howards Spring with optimised emissivity

#### SI4. $T_s$ sensitivity to $\varepsilon$ at Alice spring and Tumbarumba

In general the broadband emissivity range for a land cover can vary between 0.87 to 0.98<sup>58</sup>. The noontime measured longwave for two contrasting landcover types (semi-arid mulga, tropical savanna woodland) are used to quantify the sensitivity of LST to emissivity as shown in Fig. (10).

SI5.RMSE and Epsilon SI6.Energy imbalance and observed fluxes SI7.H and intercept

Table 3. Abbreviation list

Symbol	Description	Unit
$R_{net}$	Net radiation	$\mathrm{W}~\mathrm{m}^{-2}~\mathrm{m}$
H	Sensible heat flux	$ m W~m^{-2}$
LE	Latent heat flux	$\mathrm{W}~\mathrm{m}^{-2}$
G	Ground heat flux	$\mathrm{W}~\mathrm{m}^{-2}$
$R_{lem}$	Emittied longwave ra-	$ m W~m^{-2}$
	diation	
ε	Surface emissivity	(-)
σ	Stefan-Boltzmann	${ m W} \ { m m}^{-2} { m K}^{-4}$
	constant	
$T_s$	Surface temerature	K
$R_{sdwn}$	Down-welling short-	$\mathrm{W}~\mathrm{m}^{-2}$
	wave	
$R_{ldwn}$	Down-welling long-	${ m W~m^{-2}}$
	wave	
$R_{sref}$	Reflected shortwave	${ m W~m^{-2}}$
α	Surface albedo	(-)
m	Aerodynamic conduc-	(m/s)
	tance to heat transport	
ε 31	Spectral emissivity	(-)
	for wavelength of 11	
	$\mu$ m	
$\varepsilon$ 32	Spectral emissivity	(-)
	for wavelength of 12	
	$\mu$ m	
BADAM	Ameriflux dataset	(-)
TERRA	NASA scientific re-	(-)
	search satellite	
NATT	North Australian	(-)
	Tropical Transect	

Sites	Landscape-scale $arepsilon$					Plot-scale $\varepsilon$					
	ε	seq		eq seq			leq				
		$R^2$	bias	$R^2$	bias	opt $arepsilon$	$R^2$	bias	opt $arepsilon$	$R^2$	bias
SP	0.974	0.80	-3.67	0.81	-4.61	0.96	0.81	-3.0	0.85	0.82	-1.91
AS	0.974	0.93	-4.78	0.93	-6.31	0.96	0.93	-3.4	0.82	0.93	-1.92
TT	0.974	0.55	-6.76	0.57	-8.30	0.95	0.58	-5.06	0.80	0.52	-4.02
HS	0.985	0.16	-8.89	0.16	-9.90	0.92	0.21	-4.78	0.6	0.22	-2.47
LF	0.985	0.40	-10.0	0.41	-11.0	0.92	0.40	-4.41	0.6	0.41	-2.57
AR	0.985	0.18	-2.61	0.27	-3.51	0.997	0.23	-2.93	0.96	0.252	-2.98
DU	0.985	0.80	-3.67	0.81	-4.61	0.99	0.428	-3.682	0.985	0.425	-3.926
TUM	0.983	0.82	-2.27	0.84	-2.10	0.99	0.89	0.99	0.97	0.89	1.93
BR	0.983	0.937	0.525	0.937	-0.195	0.98	0.917	1.87	0.82	0.895	2.72
YA	0.974	0.855	-2.081	0.855	-3.45	0.97	0.522	-4.517	0.93	0.793	-0.582

**Table 4.** Comparison of plot-scale LST with landscape-scale daytime LST (MODIS, MODA001) at TERRA daily time of pass. Plot scale LST is obtained using landscape-scale emissivity (MODIS  $\varepsilon$ ) and plot-scale emissivity (Optimum  $\varepsilon$ ) at study sites. The reported plot-scale emissivity are median values and landscape emissivity are MODIS based. Bias is defined as mean of  $T_s - T_{MODIS}$ ,  $R^2$  is coefficient of determination between plot-scale LST in comparison to landscape-scale LST. The site acronyms can be found in Table 2

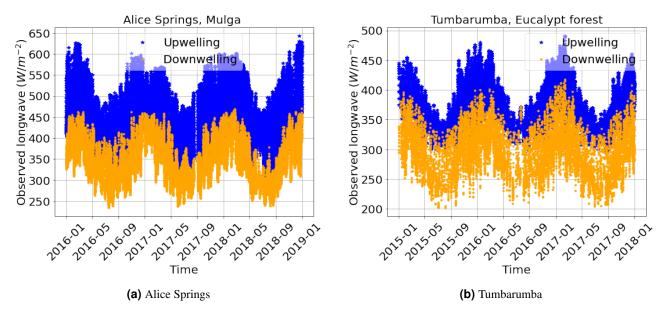
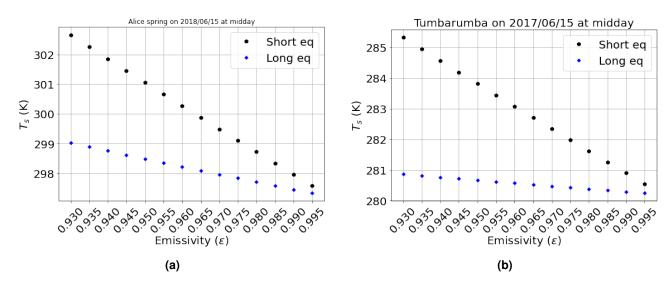


Figure 9. Measured up-welling and down-welling longwave timeseries



**Figure 10.** Sensitivity of LST estimated using two equations to the range of Broadband emissivity The black dots and blue Stars depicts LST using simplified (Eq. 11) and complete (Eq. (6). Midday longwave measurement for 15June, 2016 at Alice Springs and Tumbarumba is used



Figure 11. RMSE and emissivity curve



Figure 12. Measured fluxes and estimated energy imbalance at Howard springs for 2016/08



**Figure 13.** Monthly H  $\Delta$ T plots showing offset at  $\Delta$ T=0 (a) using simplified equation (Eq. 11). (b) using complete equation (Eq. 6)(c). Estimated intercept (c)for each month using  $H = m(T_s - T_a) + c$ . Red line is intercept and blue line shows monthly  $\varepsilon$  obtained by minimising RMSE at Alice Springs (2016 - 2018).