### 02562 Rendering - Introduction

Measuring light (radiometry and photometry)

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## Physics of light

- ▶ The three basic actions of quantum electrodynamics:
  - 1. A photon goes from place to place.
  - 2. An electron goes from place to place.
  - 3. An electron emits or absorbs a photon.
- $\triangleright$  The probability for a photon to be at a position x follows a time-varying wave function.
- ▶ The energy E of the photon is related to the angular frequency  $\omega$  of this wave:

$$E=\hbar\,\omega=hf=rac{hc}{\lambda}\,,\quad \hbar=rac{h}{2\pi}\,,\quad f=2\pi\omega\,,\quad h=6.626\cdot 10^{-34}\,\,\mathrm{Js}\,,$$

where  $\lambda$  is the wavelength and  $c = 2.9979 \cdot 10^8$  m/s is the speed of light.

- Light can be a vector field if we average over a large number of photons.
- ► The number of photons arriving per second determines the magnitude of the energy flux (J/s or W) in an electromagnetic field.
- We can measure such radiant energy.

Physics	Radiometry			Photometry	
Energy	Radiant Energy	Q	joule J $(kg m^2/s^2)$	Luminous Energy	talbot T (Im s)
Flux (Power)	Radiant Flux	$\Phi = \frac{dQ}{dt}$	watt W (J/s)	Luminous Flux	lumen Im (T/s)
Flux Density (incoming)	Irradiance	$E = \frac{d\Phi}{dA}$	W/m <sup>2</sup>	Illuminance	lux (lm/m <sup>2</sup> )
Flux Density (outgoing)	Radiant Exitance	$M = \frac{d\Phi}{dA}$	W/m <sup>2</sup>	Luminous Exitance	lux (lm/m <sup>2</sup> )
Flux Density (outgoing)	Radiosity (diffuse)	$B = \frac{d\Phi}{dA}$	W/m <sup>2</sup>	Luminosity	lux (lm/m <sup>2</sup> )
Angular Flux Density	Radiance	$L = \frac{\mathrm{d}\Phi}{\mathrm{d}A^{\perp}\mathrm{d}\omega}$	$W/(m^2 sr)$	Luminance	nit (cd/m²)
Intensity	Radiant Intensity	$I = \frac{d\Phi}{d\omega}$	W/sr	Luminous Intensity	candela cd (lm/sr)

#### Light emission

▶ An electric light emits light according to its power P and emissivity/efficiency  $\varepsilon$ :

$$\Phi = \varepsilon P = \varepsilon VC,$$

where V is voltage (volts) and C is current (amps).

If the source emits light isotropically in the solid angle  $\Omega$ , its intensity is

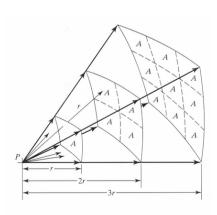
$$I = \frac{\mathsf{d}\Phi}{\mathsf{d}\omega} = \frac{\Phi}{\Omega} \,.$$

If the source emits light homogeneously across a surface, the radiant exitance is

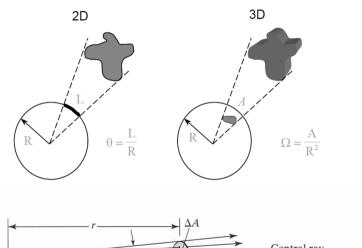
$$M = \frac{\mathrm{d}\Phi}{\mathrm{d}A} = \frac{\Phi}{A}$$
.

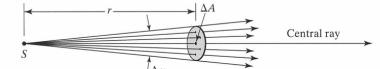
► The radiance from an isotropic point light reaching an observer at the distance *r* is

$$L_i = \frac{I}{r^2}$$

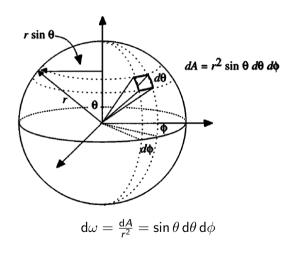


# Solid angles (radians vs. steradians)





# Calculating solid angles using spherical coordinates



Full sphere:

$$\Omega = \int_0^{2\pi} \int_0^{\pi} \sin \theta \, d\theta \, d\phi$$

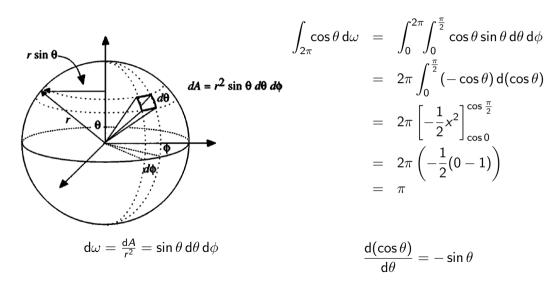
$$= \int_0^{2\pi} d\phi \int_{\cos \pi}^{\cos 0} d(\cos \theta)$$

$$= 2\pi (1 - (-1))$$

$$= 4\pi$$

$$\frac{\mathsf{d}(\cos\theta)}{\mathsf{d}\theta} = -\sin\theta$$

## Integrating over a cosine-weighted hemisphere



#### A diffuse emitter

Total emitted radiant flux:

$$\Phi = \int_{A} \int_{2\pi} L \cos \theta \, d\omega \, dA_{o}$$

$$= L \int_{A} dA_{o} \int_{0}^{2\pi} d\phi \int_{0}^{\frac{\pi}{2}} \cos \theta \sin \theta \, d\theta$$

$$= LA 2\pi \frac{1}{2} = LA\pi.$$

► Radiosity:

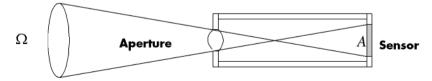
$$B = \int_{2\pi} L \cos \theta \, d\omega$$
$$= L 2\pi \int_{0}^{\frac{\pi}{2}} \cos \theta \sin \theta \, d\theta = L\pi.$$

### The measurement equation

- ► Radiance is flux per projected area per solid angle:  $L = \frac{d\Phi}{dA^{\perp}d\omega} = \frac{d\Phi}{\cos\theta \, dA \, d\omega}$
- ▶ Suppose we let *R* denote the relative responsivity of the sensor.
- ightharpoonup The energy received per second by a sensor of area A and solid angle  $\Omega$  is then

$$\Phi_i(A,\Omega) = \int_A \int_\Omega R(\mathbf{x}_i,\vec{\omega}_i) L_i(\mathbf{x}_i,\vec{\omega}_i) (\vec{\omega}_i \cdot \vec{n}) d\omega_i dA_i$$

where  $\vec{n}$  is the surface normal of the sensor area  $(\cos \theta = \vec{\omega}_i \cdot \vec{n})$ .



## Irradiance due to a source of a given intensity

 $\triangleright$  Element of solid angle subtended by an element of surface area at distance r:

$$d\omega = \frac{dA^{\perp}}{r^2} = \frac{dA\cos\theta}{r^2} \quad \Leftrightarrow \quad \frac{d\omega}{dA} = \frac{\cos\theta}{r^2} \,,$$

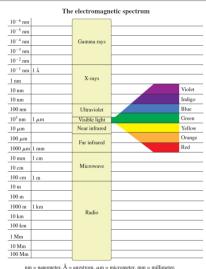
where  $\theta$  is the angle between the surface normal and the central ray of the solid angle.

▶ Relation between irradiance and source intensity:

$$E = \frac{d\Phi}{dA} = \frac{d\Phi}{d\omega} \frac{d\omega}{dA} = I \frac{d\omega}{dA} = I \frac{\cos\theta}{r^2}.$$

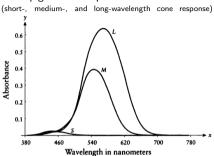
- ▶ Intensity of an isotropic point light:  $I = \frac{\mathrm{d}\Phi}{\mathrm{d}\omega} = \frac{\Phi}{\Omega} = \frac{\Phi}{4\pi}$ .
- ► Irradiance due to an isotropic point light:  $E = I \frac{\cos \theta}{r^2} = \frac{\Phi}{4\pi} \frac{\cos \theta}{r^2}$ .

## Colour perception

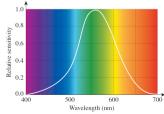


nm = nanometer,  $\mathring{A}$  = angstrom,  $\mu$ m = micrometer, mm = millimeter, cm = centimeter, m = meter, km = kilometer, Mm = Megameter

#### Photopigment absorption



#### Luminous efficiency



### **Photometry**

- For every spectral radiometric quantity  $f_r(\lambda)$ , there is a related photometric quantity  $f_p$  measuring the "usefulness" of this quantity to a human observer.
- Conversion formula:

$$f_p = 683 \frac{\text{Im}}{\text{W}} \int_{380 \text{ nm}}^{800 \text{ nm}} \overline{y}(\lambda) f_r(\lambda) \, d\lambda$$

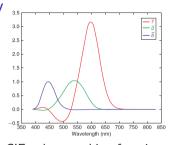
where  $\overline{y}$  is the luminous efficiency function (sometimes called V).

Example: radiance to luminance

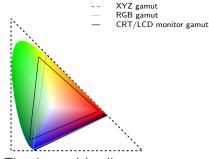
$$Y = 683 \frac{\text{Im}}{\text{W}} \int_{380 \text{ nm}}^{800 \text{ nm}} \overline{y}(\lambda) L(\lambda) \, d\lambda,$$

- Luminance Y is also one of three components in the XYZ color space, and  $\overline{y}$  is one of three color matching function:  $\overline{x}$ ,  $\overline{y}$ ,  $\overline{z}$ .
- ▶ Similarly, there are color matching functions for the RGB color space:  $\overline{r}$ ,  $\overline{g}$ ,  $\overline{b}$ .

## Colorimetry



CIE color matching functions



The chromaticity diagram

$$R = \int_{\mathscr{V}} C(\lambda) \bar{r}(\lambda) \, d\lambda$$

$$G = \int_{\mathscr{V}} C(\lambda) \bar{g}(\lambda) \, d\lambda$$

$$B = \int_{\mathscr{V}} C(\lambda) \bar{b}(\lambda) \, d\lambda ,$$

where  $\mathscr V$  is the interval of visible wavelengths and  $C(\lambda)$  is the spectrum that we want to transform to RGB.