



LISA Analysis Tools Workshop

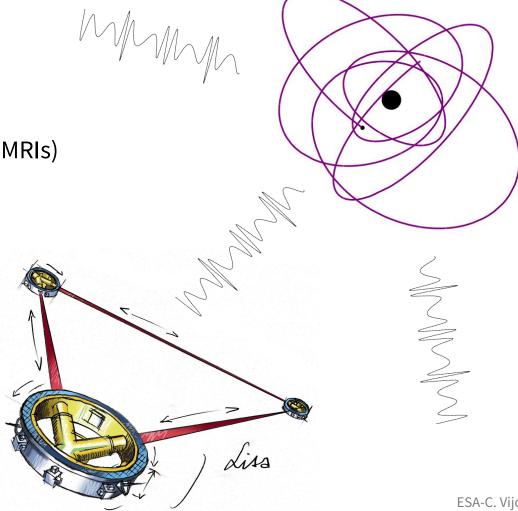
Day 1 - Afternoon Session

EMRI waveforms and the LISA response function

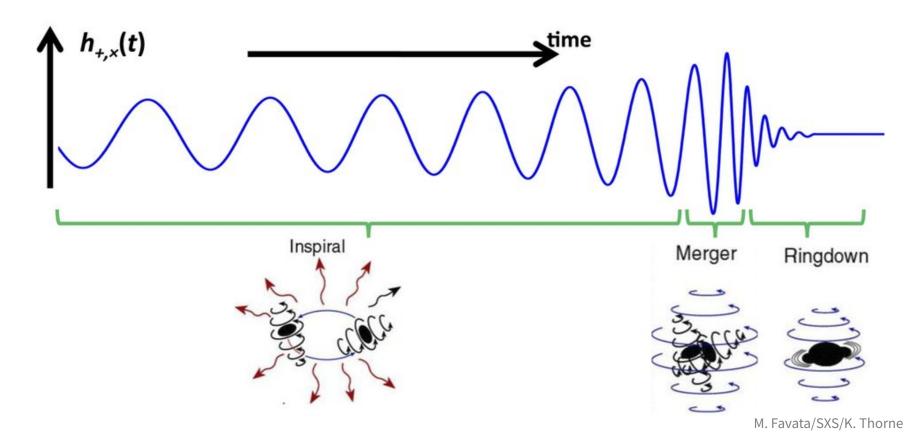
Christian Chapman-Bird Institute for Gravitational Research, University of Glasgow

Outline

- Extreme-mass-ratio inspirals (EMRIs)
- FastEMRIWaveforms (**few**)
- The LISA response function
- fastlisaresponse

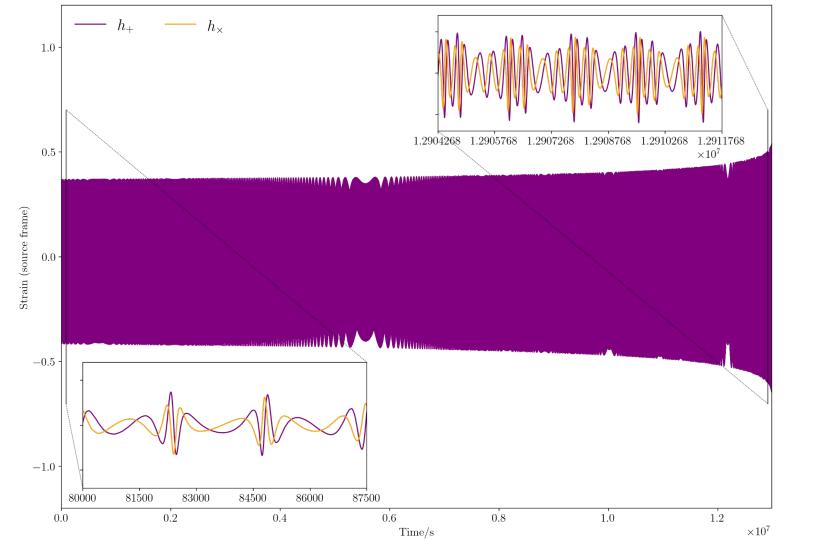


Compact binary coalescences

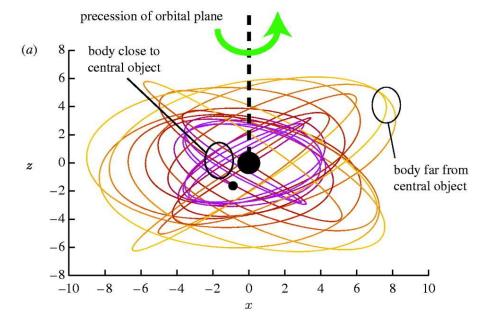


Extreme-mass-ratio inspirals (EMRIs)

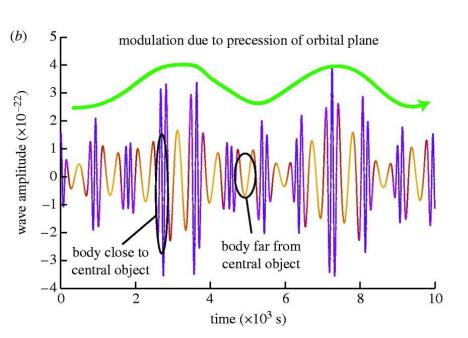
- Compact binary coalescence: massive black hole (MBH) + compact object (CO)
- Characterised by mass ratio $q \equiv \mu/M \lesssim 10^{-4}$, two main consequences:
 - \circ CO completes $\mathcal{O}(10^5)$ orbital cycles **in the LISA band** before merger
 - Highly sensitive to initial conditions
 - Orbits retain significant eccentricity and orbital inclination
 - Waveform contains many higher-order modes

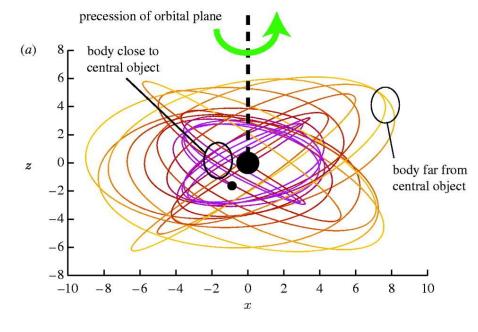


We can identify features in the inspiral trajectory....

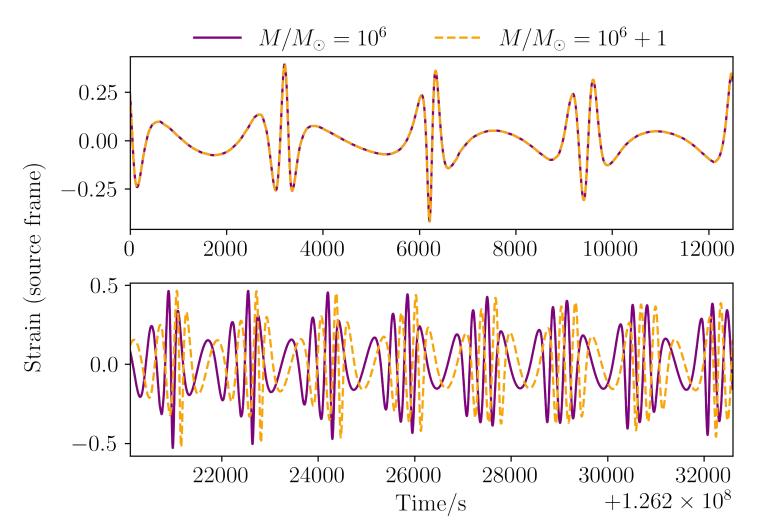


We can identify features in the inspiral trajectory....





... and their impact on the shape of the waveform.



(few)

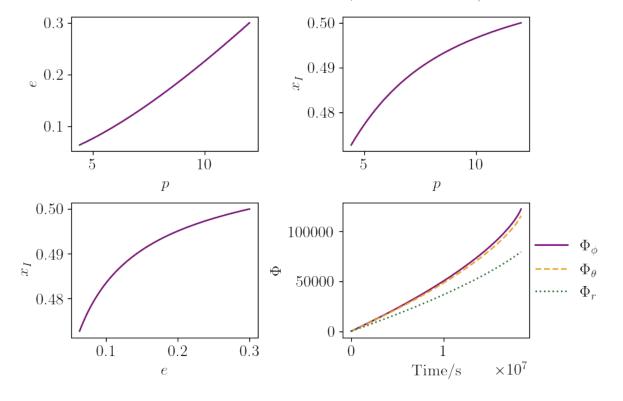
A part of the Black Hole Perturbation Toolkit bhptoolkit.org

$$h(t) = \sum_{lmkn} H_{lmkn}(p(t), e(t), x_I(t), \theta, \phi) e^{-i\Phi_{mkn}(p(t), e(t), x_I(t))}$$

$$(\Phi_{mkn} = m\Phi_{\phi} + k\Phi_{\theta} + n\Phi_r)$$

$$h(t) = \sum_{lmkn} H_{lmkn} \left(p(t), e(t), x_I(t), \theta, \phi \right) e^{-\left[i\Phi_{mkn}(p(t), e(t), x_I(t))\right]} \left(\Phi_{mkn} = m\Phi_{\phi} + k\Phi_{\theta} + n\Phi_r \right)$$

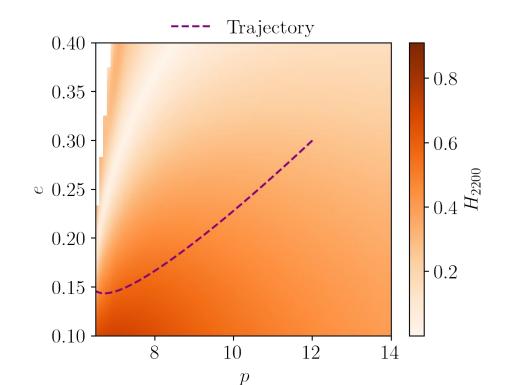
few.trajectoryIntegrate ODE system



$$h(t) = \sum_{lmkn} H_{lmkn}(p(t), e(t), x_I(t), \theta, \phi) e^{-i\Phi_{mkn}(p(t), e(t), x_I(t))} e^{-i\Phi_{mkn}(p(t), e(t), x_I(t))} e^{-i\Phi_{mkn}(p(t), e(t), x_I(t))}$$

few.trajectoryIntegrate ODE system

few.amplitudeFunctions of trajectory,
fold in viewing angles



$$h(t) = \sum_{lmkn} H_{lmkn}(p(t), e(t), x_I(t), \theta, \phi) e^{-i\Phi_{mkn}(p(t), e(t), x_I(t))}$$

 $(\Phi_{mkn} = m\Phi_{\phi} + k\Phi_{\theta} + n\Phi_{r})$

few.trajectory

Integrate ODE system

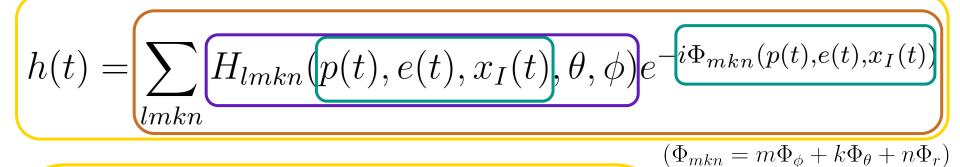


Functions of trajectory, fold in viewing angles



Sum over all modes to build waveform





few.trajectory
Integrate ODE system

few.amplitudeFunctions of trajectory,
fold in viewing angles

few.summationSum over all modes to

build waveform

few.waveform

Joins this all together!





$$h(t) = \sum_{lmkn} H_{lmkn}(p(t), e(t), x_I(t), \theta, \phi) e^{-i\Phi_{mkn}(p(t), e(t), x_I(t))}$$

 $(\Phi_{mkn} = m\Phi_{\phi} + k\Phi_{\theta} + n\Phi_{r})$

Challenges:

- Sum contains tens of thousands of terms (one per mode)
- For $10\,\mathrm{s}$ sampling cadence, $\mathcal{O}(10^7)$ data points per mode

This is a computationally expensive operation!



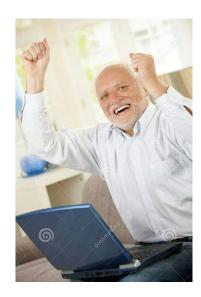
Performant

Interpolation, graphics processing unit (GPU) vectorisation



Performant

Interpolation, graphics processing unit (GPU) vectorisation



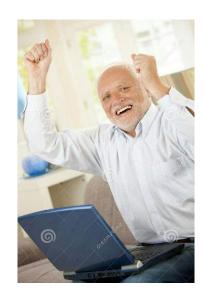
Easy to use

EMRI waveforms out of the box!



Performant

Interpolation, graphics processing unit (GPU) vectorisation



Easy to use

EMRI waveforms out of the box!



Modular

General framework for building custom EMRI waveforms

Currently supported waveform models:

- FastSchwarzschildEccentricFlux (Non-spinning black holes)
- Pn5AAKWaveform (Kludge amplitudes, generic PN inspirals)

Coming soon (names a WIP...):

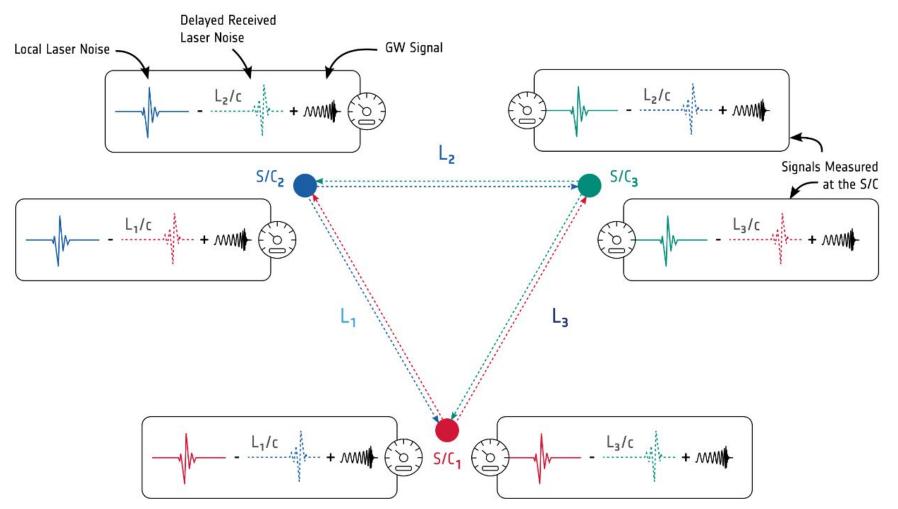
- KerrEccentricEquatorialFlux (Spinning MBH, equatorial inspirals)
- Pn5TrajPn5AdiabaticWaveform (Generic PN inspirals and amplitudes)

The LISA response function

LISA data will come in the form of time delay interferometry (TDI) observables

- Linear combinations of data streams delayed in time
- Necessary to suppress laser frequency noise
- Delays are multiples of light-travel time between spacecraft
 - Spacecraft orbits change with time, so the delays do too

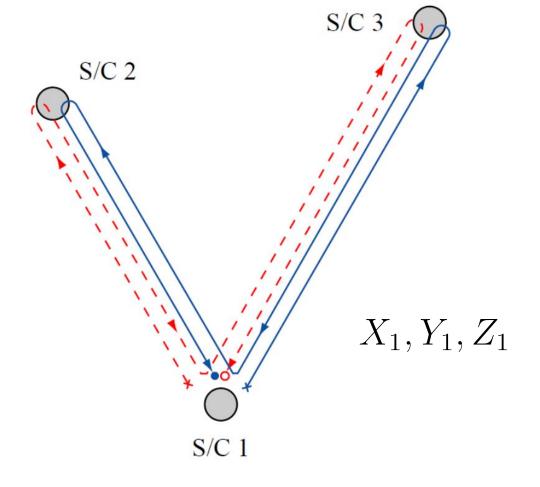
An efficient implementation of the LISA response **essential** for feasible analyses



LISA Definition Study Report

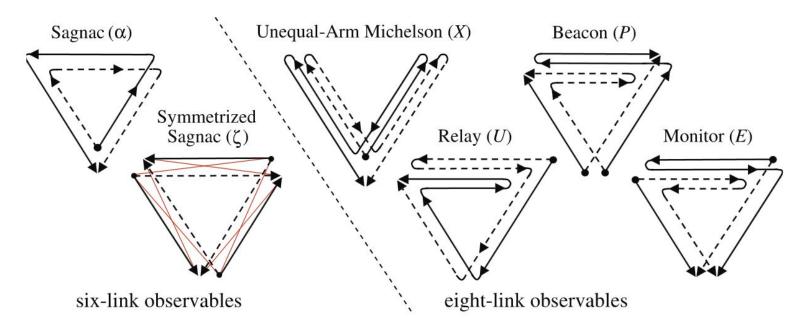
TDI Combinations

- Construct effective paths from delayed laser link measurements and subtract them
- First-generation TDI combinations account for unequal arms
- Second-generation TDI combinations (constructed from first-gen combinations) account for linear changes in arm-lengths with respect to time



Time-delay interferometry

Many combinations are useful...



fastlisaresponse () lisa-on-gpu)

One-stop shop for time-domain LISA response:

- GPU vectorisation for millisecond response evaluation
- Link-based TDI generation: construct any combination!
- Commonly-used TDI combinations (XYZ, AET) are included **out-of-the-box**

Interfaces with lisaorbits via lisatools. detector:

- Generic orbital interface
- Sample orbit files provided for synthetic data studies

fastlisaresponse () lisa-on-gpu)

Waveform model

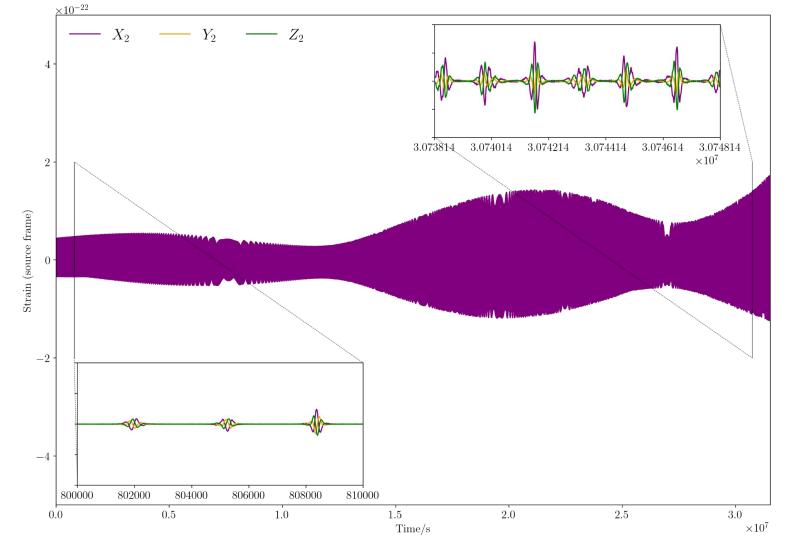
fastlisaresponse.pyResponseTDI

Takes waveform strain and produces TDI combinations

Spacecraft orbits (lisaorbits)

fastlisaresponse.ResponseWrapper

Single waveform generator object



Conclusions

- EMRI waveforms are complicated and computationally expensive
- few:
 - Out-of-the-box EMRI waveforms
 - Highly-optimised GPU methods: millisecond waveform generation
 - Modular: easily implement your own waveform models
- LISA data analysis requires flexible, efficient response implementation
- fastlisaresponse
 - Time-domain response, fast via GPU vectorisation
 - Support for generic orbits
 - Build any TDI combination: base support for XYZ/AET

Bonus slides

