O código utilizado no trabalho foi criado na linguagem Python3 utilizando as bibliotecas Numpy, Argparse e os. O Algoritmo é apresentado a seguir:

```
from argparse import ArgumentParser, FileType
import os
import numpy as np
def get args():
    parser = ArgumentParser(description='Program to calculate \
                                           the period of a star \
                                           using the Conditional \
                                           Entropy method')
    parser.add argument('-i', type=str, default=None,
                          help='Location or file of stellar data')
    parser.add argument('-o', type=str, default=None,
                          help='Location for output files')
    \verb|parser.add_argument| ( \text{'-minp'}, \text{ '--min-period'}, \text{ } \mathbf{type} = \mathbf{float} \;,
                          default= 0.1,
                          help='Minimun Period to search'
                                   '(default = 0.1)')
    parser.add_argument('-maxp', '--max-period', type=float,
                          default= 32.0,
                          help='Maximum Period to search'
                                   '(default = 32.0)')
    parser.add argument('-precision', type=float,
                          default = 0.0001,
                          help='Step between periods'
                                   '(default = 0.0001)')
    parser.add argument('-p bins', '--phase-bins', type=int,
                          default=10,
                          help='Quantity of phase bins'
                                   '(default = 10)')
    parser.add argument ('-m bins', '--mag-bins', type=int,
                          default=5,
                          help='Quantity of magnitude bins'
                                   '(default = 5)')
```

```
args = parser.parse args()
    return args
def get_files(data):
    Get the correct path of file or files inside a directory
    if os.path.isfile(data):
        return [data]
    elif os.path.isdir(os.path.abspath(data)):
        path = os.path.abspath(data)
        os.chdir(path)
        list of files = os.listdir(path)
        return sorted (list of files)
def out_dir(out):
    , , ,
    check if the output path is relative or not
    and return the absolute path
    if out == None:
        return os.getcwd()
    elif os.path.isdir(out):
        return os.path.abspath(out)
def rephase(data, period, col=0, copy=True):
    transform the time of observations to the phase space
    rephased = np.ma.array(data, copy=copy)
    rephased [:, col] = get_phase(rephased [:, col], period)
    return rephased
def get_phase(time, period):
    divide the time of observations by the period
    return (time / period) % 1
```

```
def normalization (data):
    Normalize the magnitude of the star
    norm = np.ma.copy(data)
    \operatorname{norm}[:,1] = (\operatorname{norm}[:,1] - \operatorname{np.min}(\operatorname{norm}[:,1])) \setminus
         / (\operatorname{np.max}(\operatorname{norm}[:,1]) - \operatorname{np.min}(\operatorname{norm}[:,1]))
    return norm
def periods_array(min_period, max_period, step):
    Creates a period array from min period to max period with a
    step equals to the third argument
    period = np.arange(min period, max period+step, step)
    return period
def cond_entropy(period, data, p bins=10, m bins=5):
    Compute the conditional entropy for the
    normalized observations
    if period \leq 0:
         return np.PINF
    r = rephase(data, period)
    bins, *_{\_} = np.histogram2d(r[:,0], r[:,1], [p_bins, m_bins],
                                    [[0,1], [0,1]]
    size = r.shape[0]
    if size > 0:
         divided bins = bins / size
         arg positive = divided bins > 0
         column_sums = np.sum(divided_bins, axis=1)
         column sums = np.repeat(np.reshape(column sums,
                                    (p bins, 1), m bins, axis=1
         select divided bins = divided bins [arg positive]
         select column sums = column sums arg positive
         A = np.empty((p_bins, m_bins), dtype=float)
        A[ arg positive] = select divided bins \
```

```
* np.log(select column sums \
                         / select_divided_bins)
       A[^{\sim} arg positive] = 0
        return np.sum(A)
    else:
        return np.PINF
def main():
    args = get args()
    files = get files(args.i)
    out = out dir(args.o)
    ce = []
    periods = periods array(args.min period, args.max period,
                            args.precision)
    for star in files:
        ent data = []
        ce period = []
        data = np.ma.array(np.loadtxt(star), mask=None,
                            dtype=float)
        norm = normalization(data)
        for p in periods:
            ent data.append(cond entropy(p, norm,
                                            args.phase bins,
                                            args.mag bins))
            ce_period.append(p)
        np.savetxt(os.path.join(out,
                'entropies_'+os.path.basename(star)+'.txt'),
                np.dstack((ce period, ent data))[0],
                fmt = \%s,
        right_period = star, ce_period[np.argmin(ent_data)]
        ce.append(right_period)
        np.savetxt(os.path.join(out, 'results.dat'),
                    ce, fmt='%s')
exit (main())
```

```
Algoritmo 1: Esquema básico do algoritmo em português estruturado
   Entrada: Tempo e Magnitude
   {\bf Saída} \colon {\sf Período}\ P que minimiza a entropia
 1 Início
       Leitura dos dados de entrada como vetores;
 \mathbf{2}
       Cria um vetor com n períodos sendo \mathbf{P}=(p_1\;,\,p_2,\,\cdots,\,p_n);
 3
       Normalização da magnitude;
 4
       para cada p_i em P faça
 \mathbf{5}
          Transformar o tempo para o espaço de fase;
 6
          Faz as repartições e contabiliza os pontos;
 7
          Calcula a entropia de Shannon condicional;
 8
          Armazena a entropia calculada para o período p_i
 9
       _{
m fim}
10
       Achar o valor mínimo de entropia: E_{min} = \min(\text{Entropia})
11
       Achar o período que minimiza a entropia: P_{E_{min}}=P[min(entropia)]
12
13 Fim
14 retorna P_{E_{min}}
```