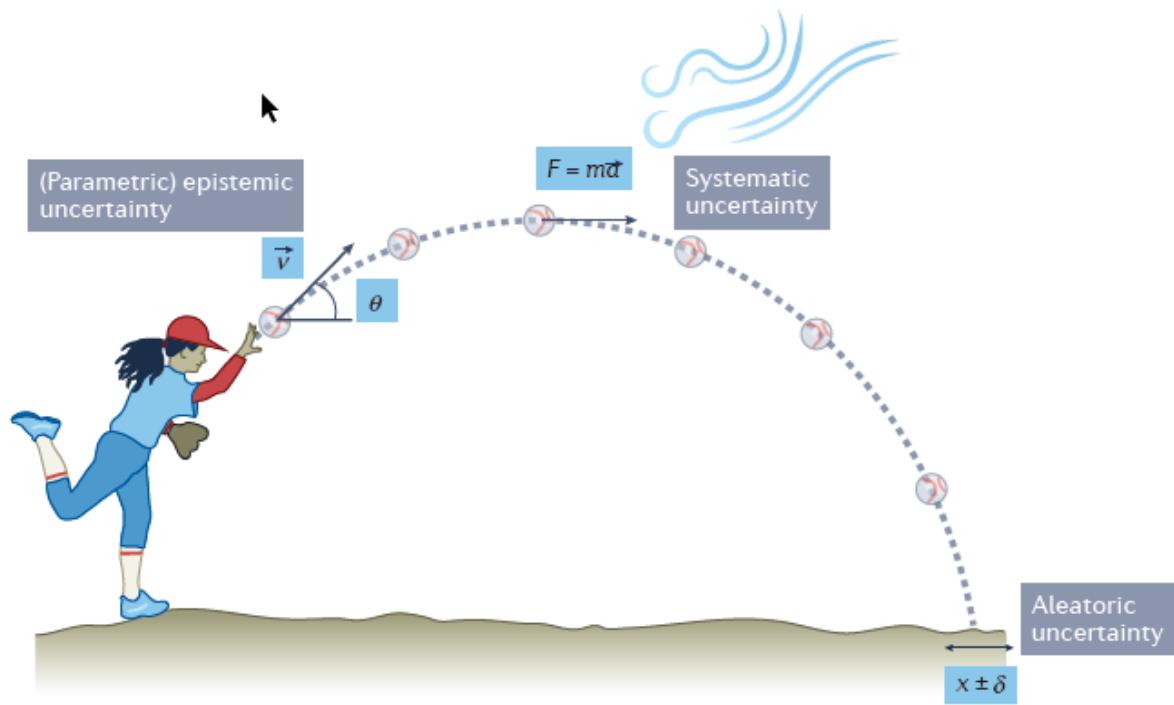


# An introduction to simulation-based inference

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[g.louppe@uliege.be](mailto:g.louppe@uliege.be)



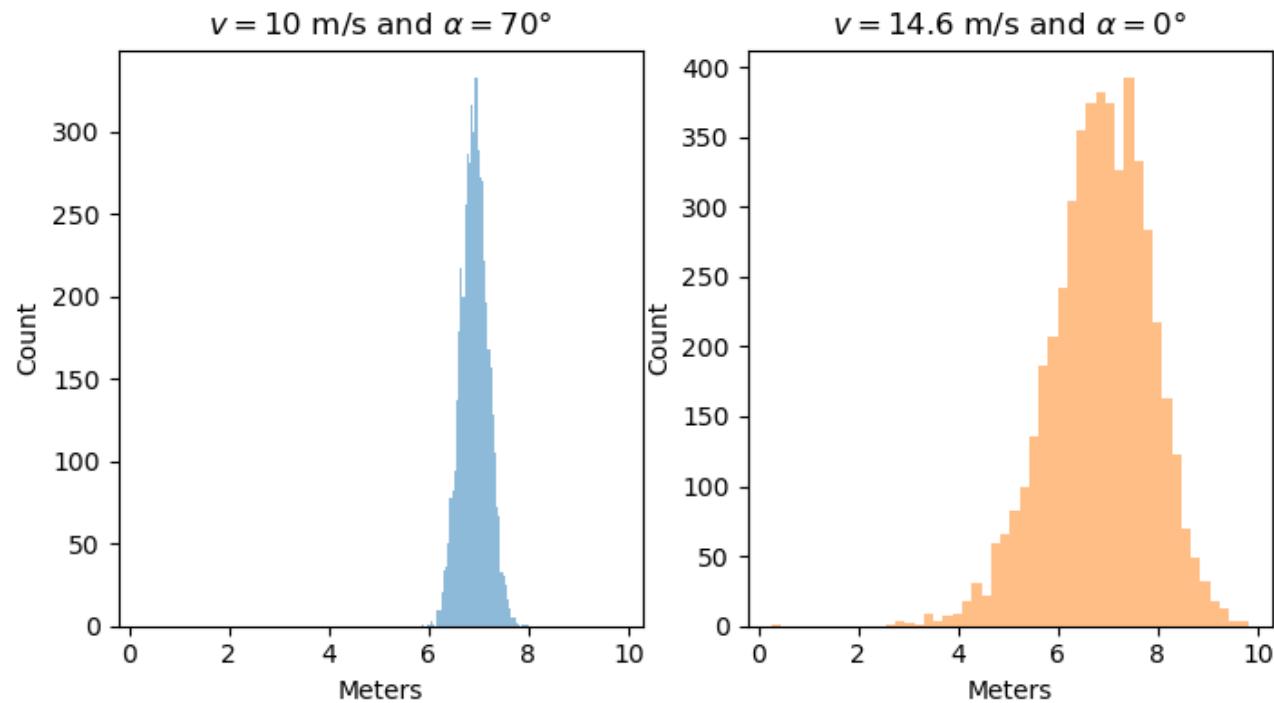
$$v_x = v \cos(\alpha), \quad v_y = v \sin(\alpha),$$

$$\frac{dx}{dt} = v_x, \quad \frac{dy}{dt} = v_y, \quad \frac{dv_y}{dt} = -G.$$

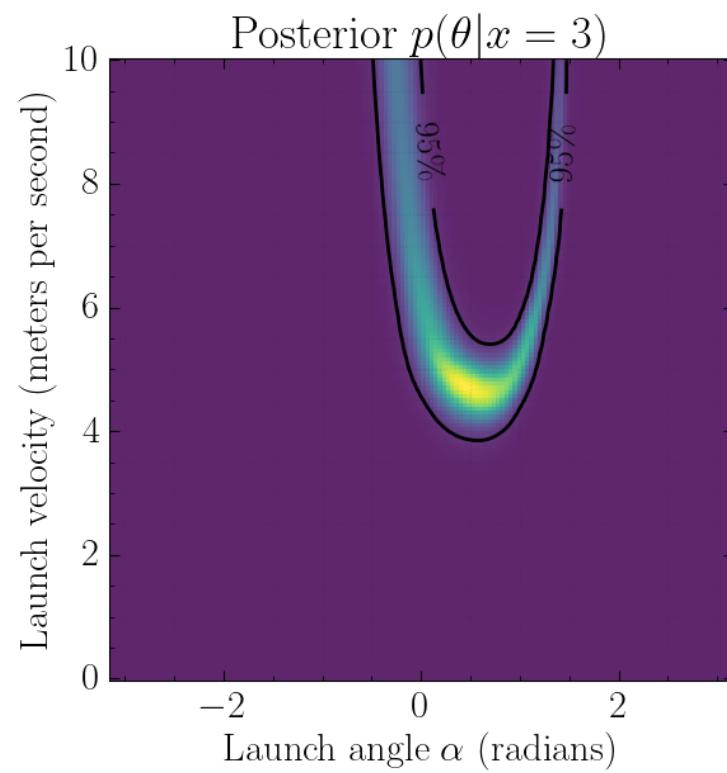
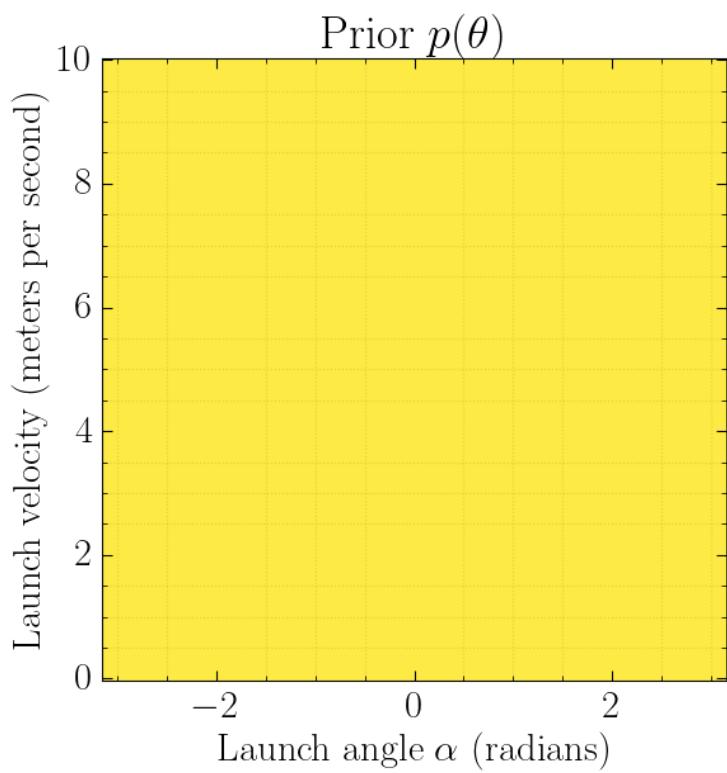
```
def simulate(v, alpha, dt=0.001):
    v_x = v * np.cos(alpha) # x velocity m/s
    v_y = v * np.sin(alpha) # y velocity m/s
    y = 1.1 + 0.3 * random.normal()
    x = 0.0

    while y > 0: # simulate until ball hits floor
        v_y += dt * -G # acceleration due to gravity
        x += dt * v_x
        y += dt * v_y

    return x + 0.25 * random.normal()
```



What parameter values  $\theta$  are the most plausible?

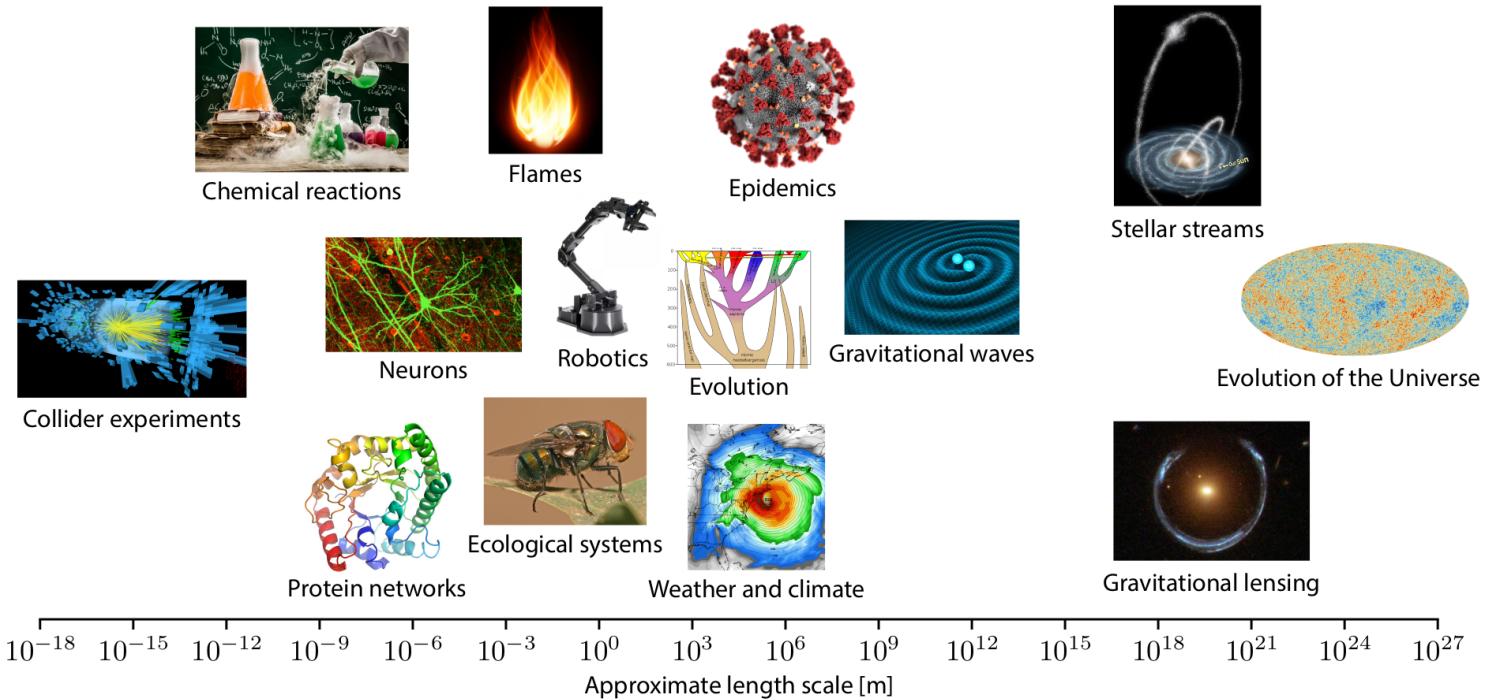


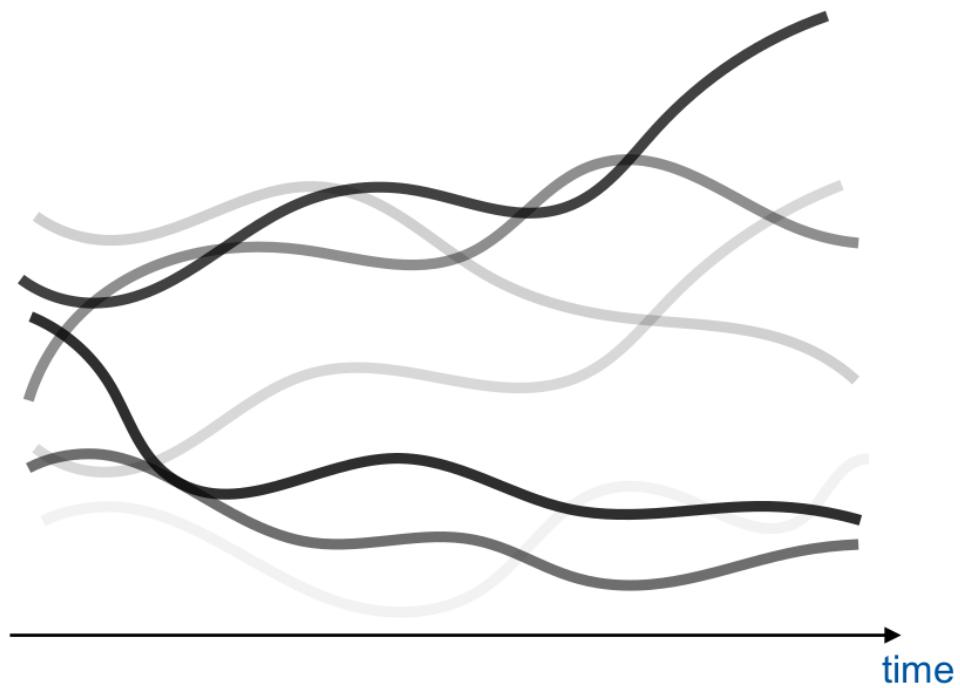
# Outline

1. Simulation-based inference
2. Algorithms
  - Neural ratio estimation
  - Neural posterior estimation
  - Neural score estimation
3. Diagnostics

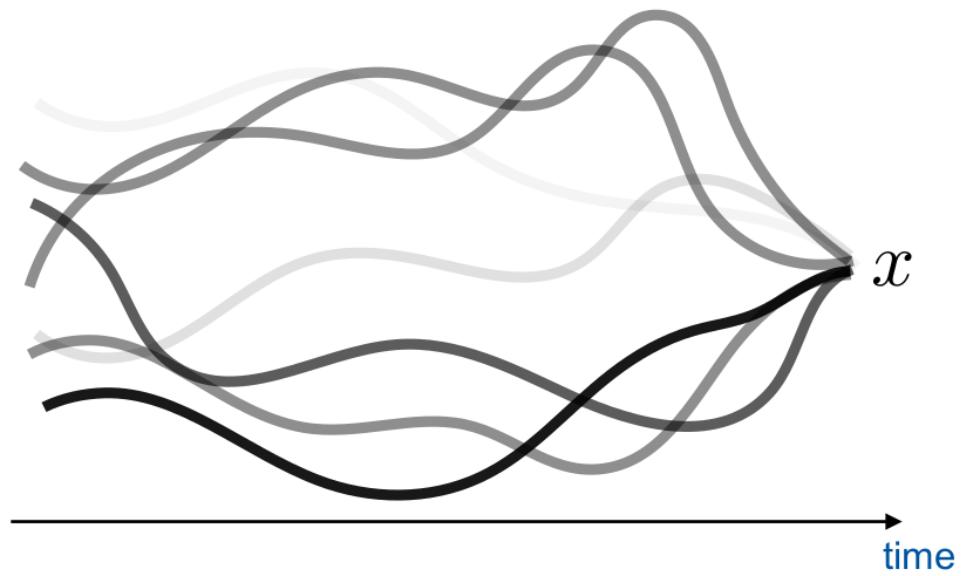
# **Simulation-based inference**

# Scientific simulators





$$\theta, z, x \sim p(\theta, z, x)$$



$$\theta, z \sim p(\theta, z|x)$$

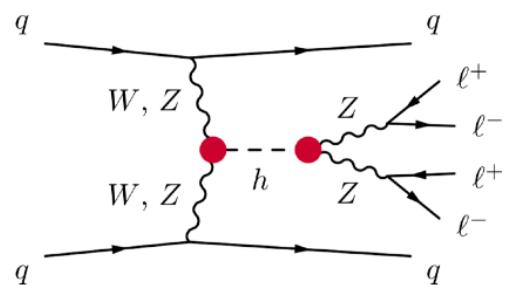
Latent variables

Parameters  
of interest

Parton-level  
momenta

Theory  
parameters

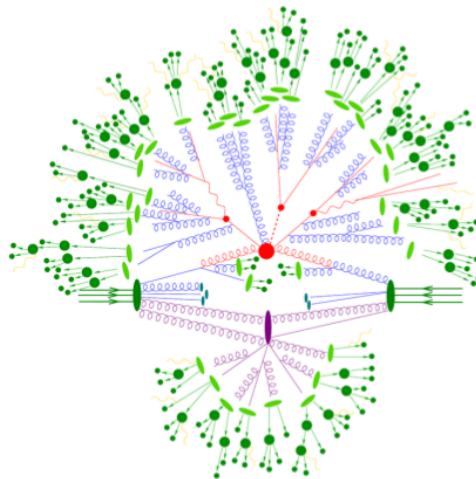
$$z_p \leftarrow \theta$$

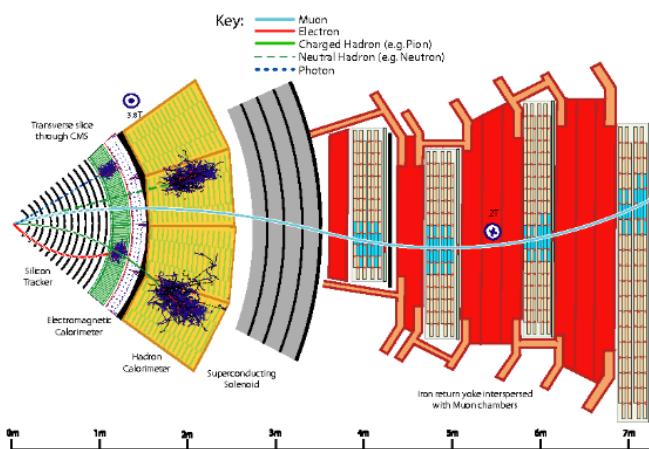
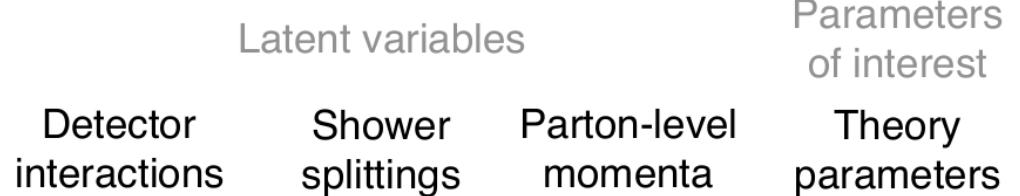


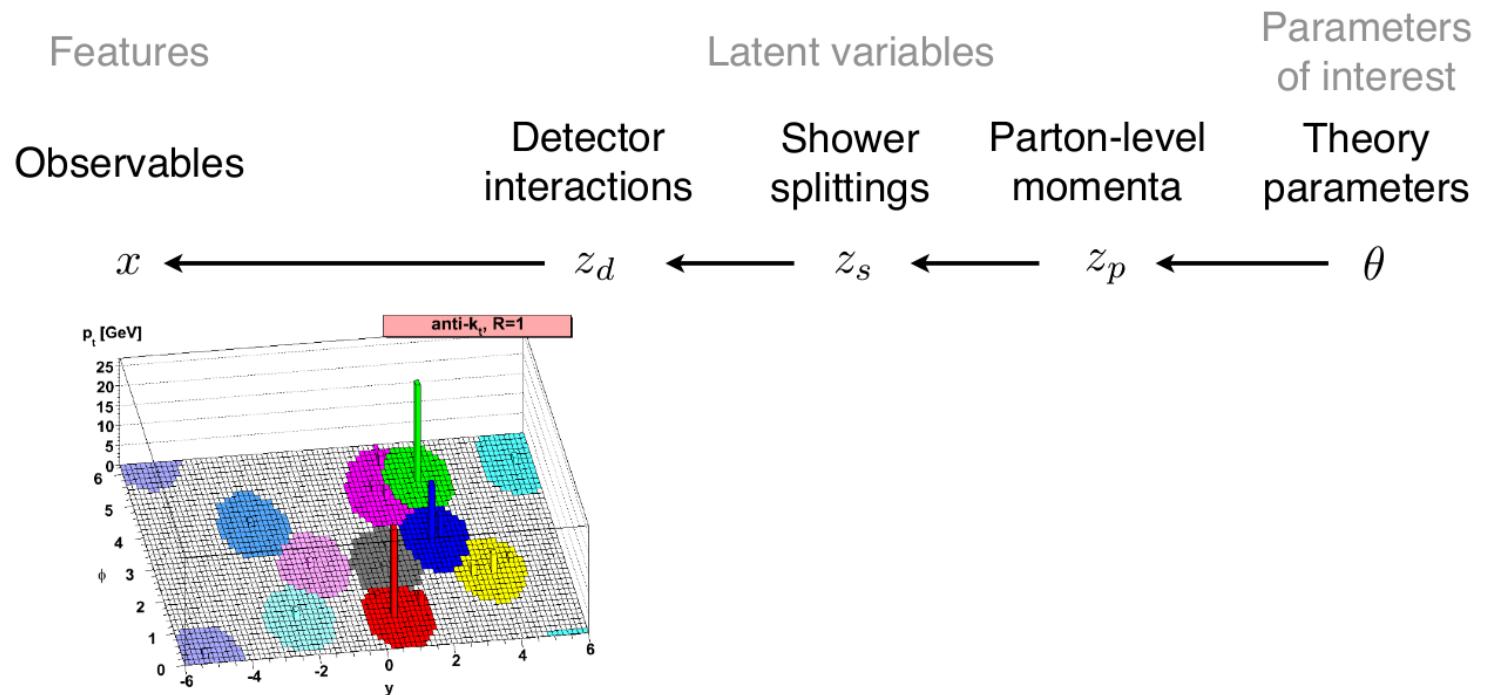
Latent variables      Parameters  
of interest

Shower      Parton-level      Theory  
splittings      momenta      parameters

$$z_s \leftarrow z_p \leftarrow \theta$$







[Image source: M. Cacciari,  
G. Salam, G. Soyez 0802.1189]

$$p(x|\theta) = \underbrace{\iiint}_{\text{yikes!}} p(z_p|\theta)p(z_s|z_p)p(z_d|z_s)p(x|z_d)dz_p dz_s dz_d$$

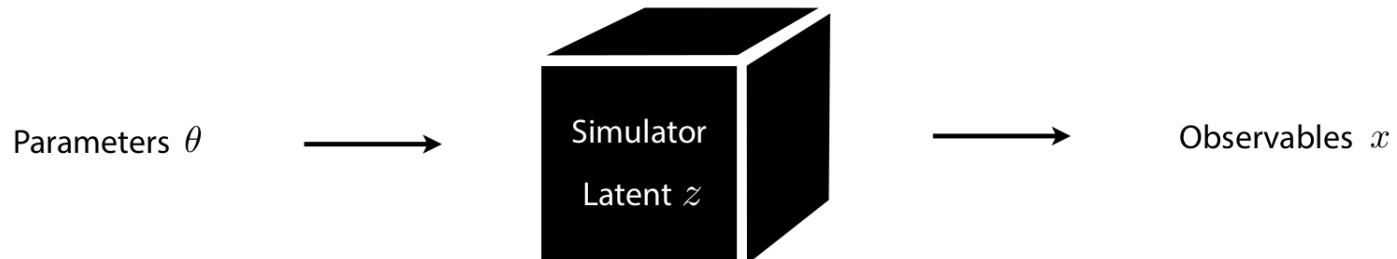
## Bayesian inference

Start with

- a simulator that can generate  $N$  samples  $x_i \sim p(x_i | \theta_i)$ ,
- a prior model  $p(\theta)$ ,
- observed data  $x_{\text{obs}} \sim p(x_{\text{obs}} | \theta_{\text{true}})$ .

Then, estimate the posterior

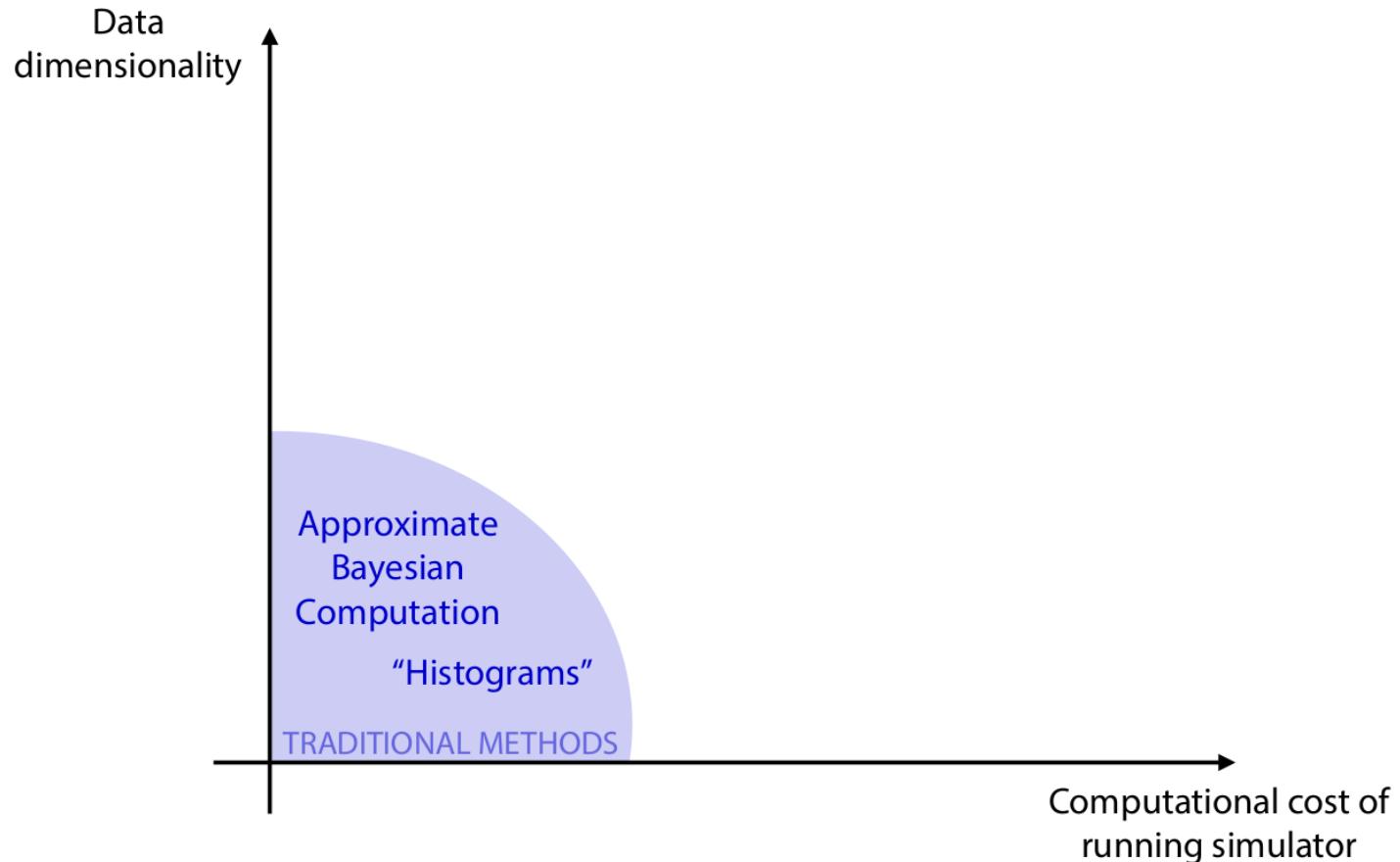
$$p(\theta | x_{\text{obs}}) = \frac{p(x_{\text{obs}} | \theta)p(\theta)}{p(x_{\text{obs}})}.$$



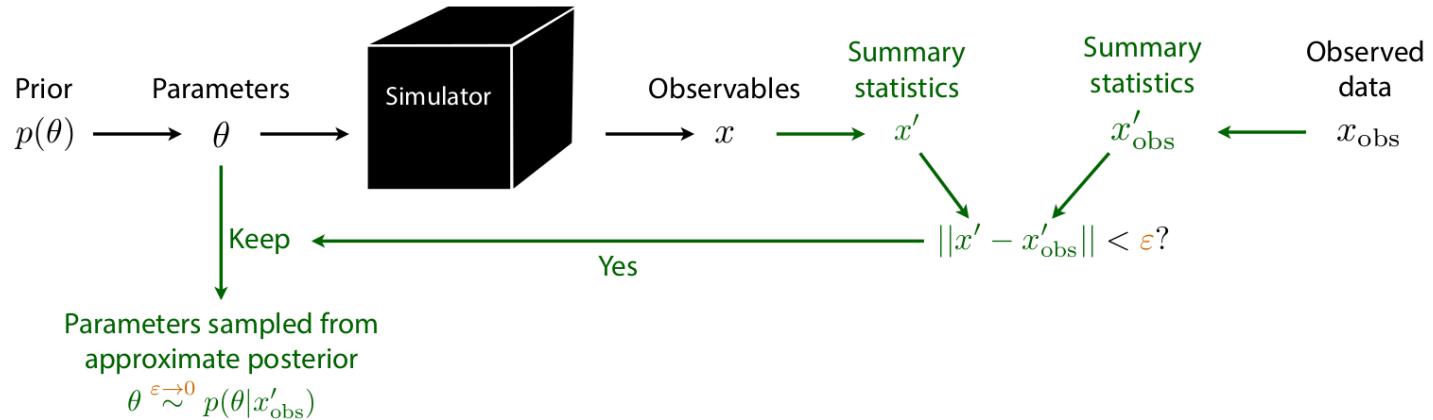
- Prediction:
- Well-motivated mechanistic, causal model
  - Simulator can generate samples  $x \sim p(x|\theta)$

- Inference:
- Interactions between low-level components lead to challenging inverse problems
  - Likelihood  $p(x|\theta) = \int dz p(x, z|\theta)$  is intractable

# Algorithms

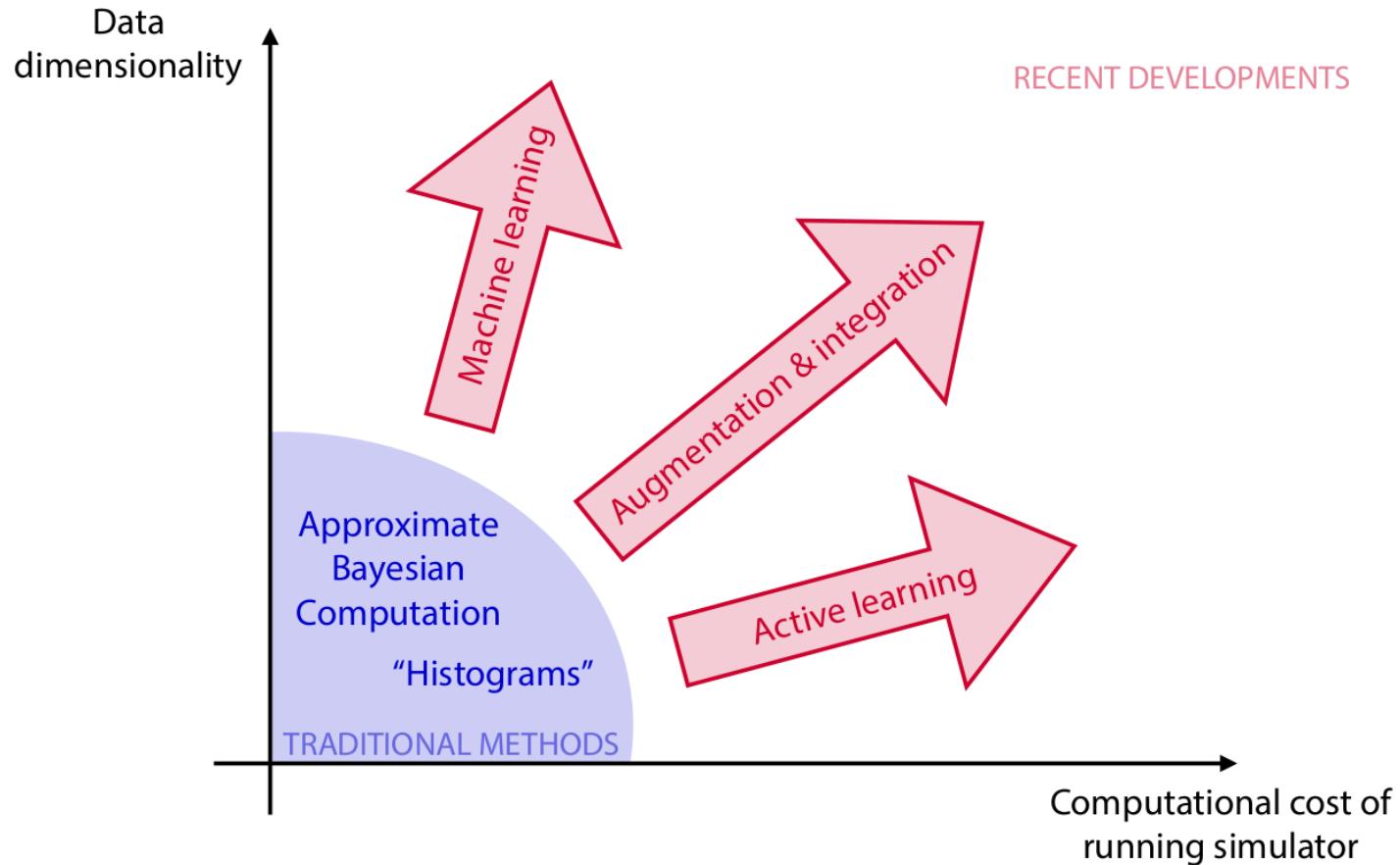


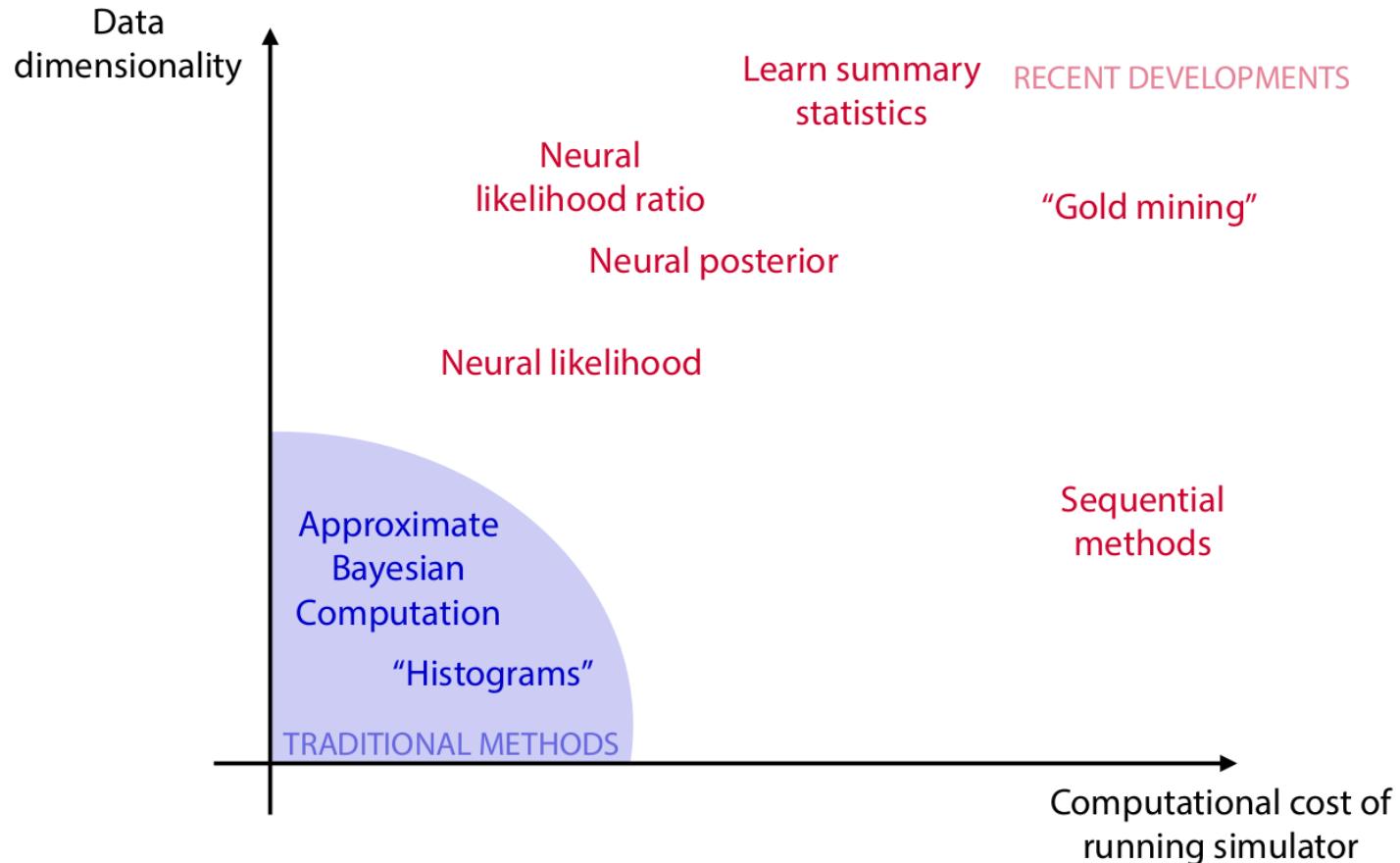
# Approximate Bayesian Computation (ABC)



Issues:

- How to choose  $x'?$   $\epsilon?$   $|| \cdot ||?$
- No tractable posterior.
- Need to run new simulations for new data or new prior.

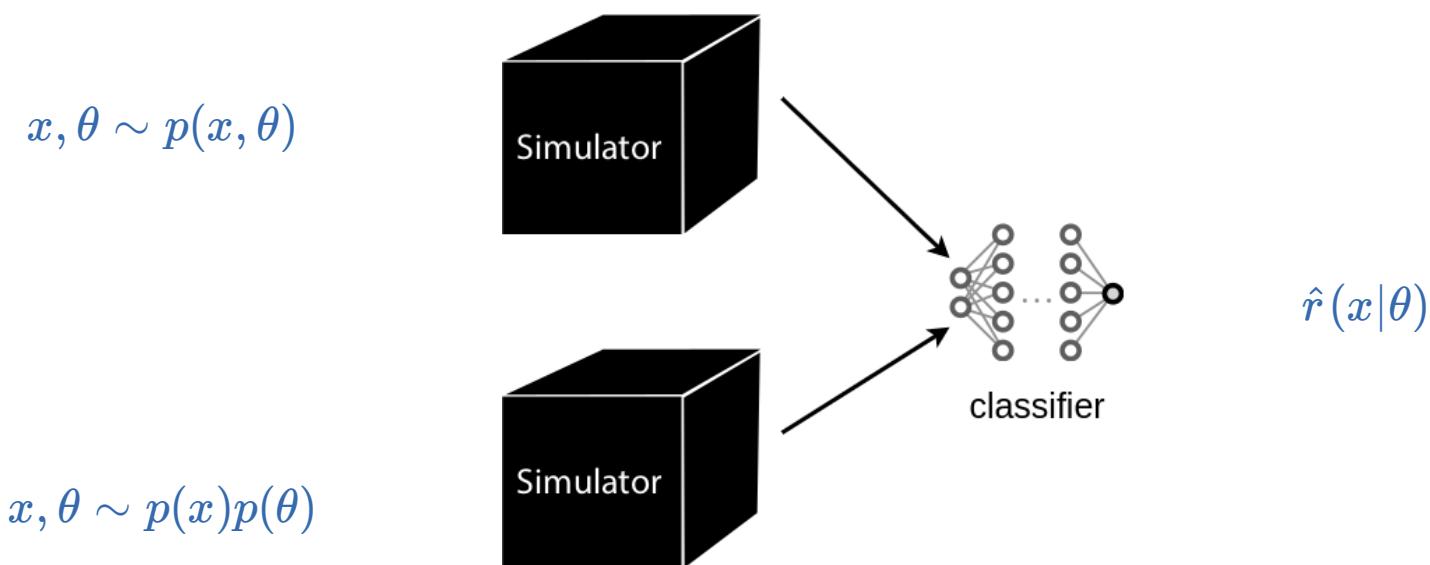




# Neural ratio estimation



The likelihood-to-evidence  $r(x|\theta) = \frac{p(x|\theta)}{p(x)} = \frac{p(x,\theta)}{p(x)p(\theta)}$  ratio can be learned, even if neither the likelihood nor the evidence can be evaluated:



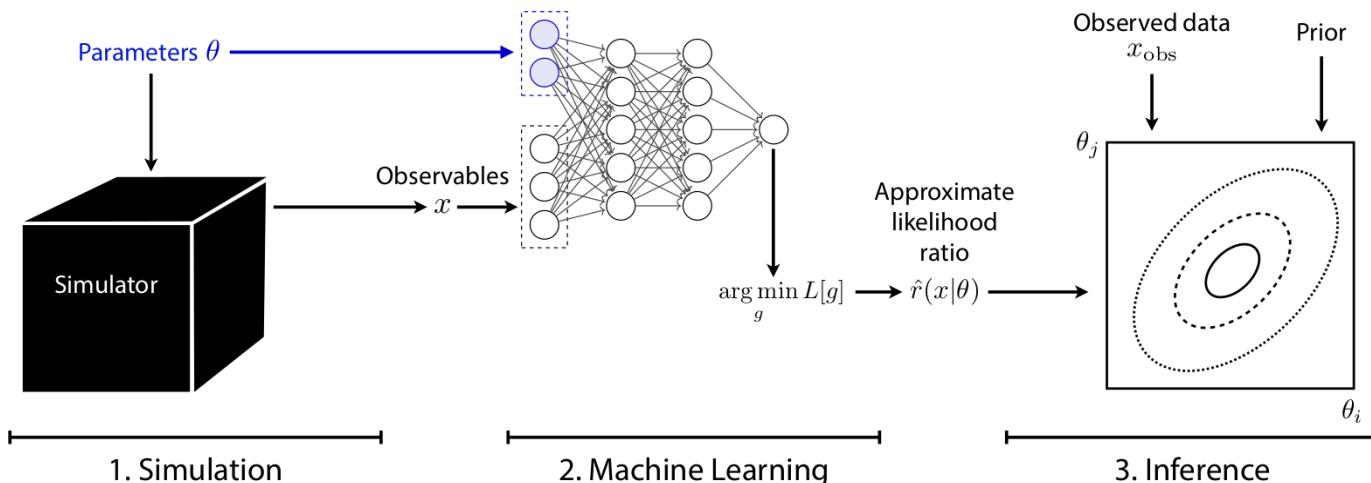


The solution  $\mathbf{d}$  found after training approximates the optimal classifier

$$d(x, \theta) \approx d^*(x, \theta) = \frac{p(x, \theta)}{p(x, \theta) + p(x)p(\theta)}.$$

Therefore,

$$r(x|\theta) = \frac{p(x|\theta)}{p(x)} = \frac{p(x, \theta)}{p(x)p(\theta)} \approx \frac{d(x, \theta)}{1 - d(x, \theta)} = \hat{r}(x|\theta).$$



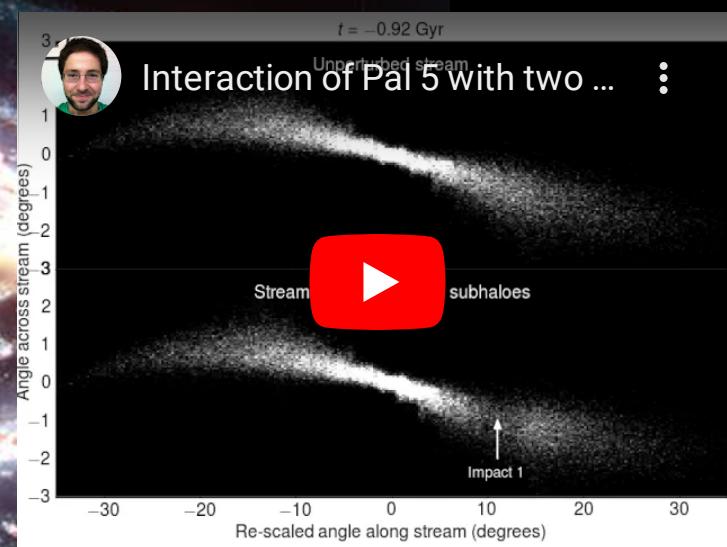
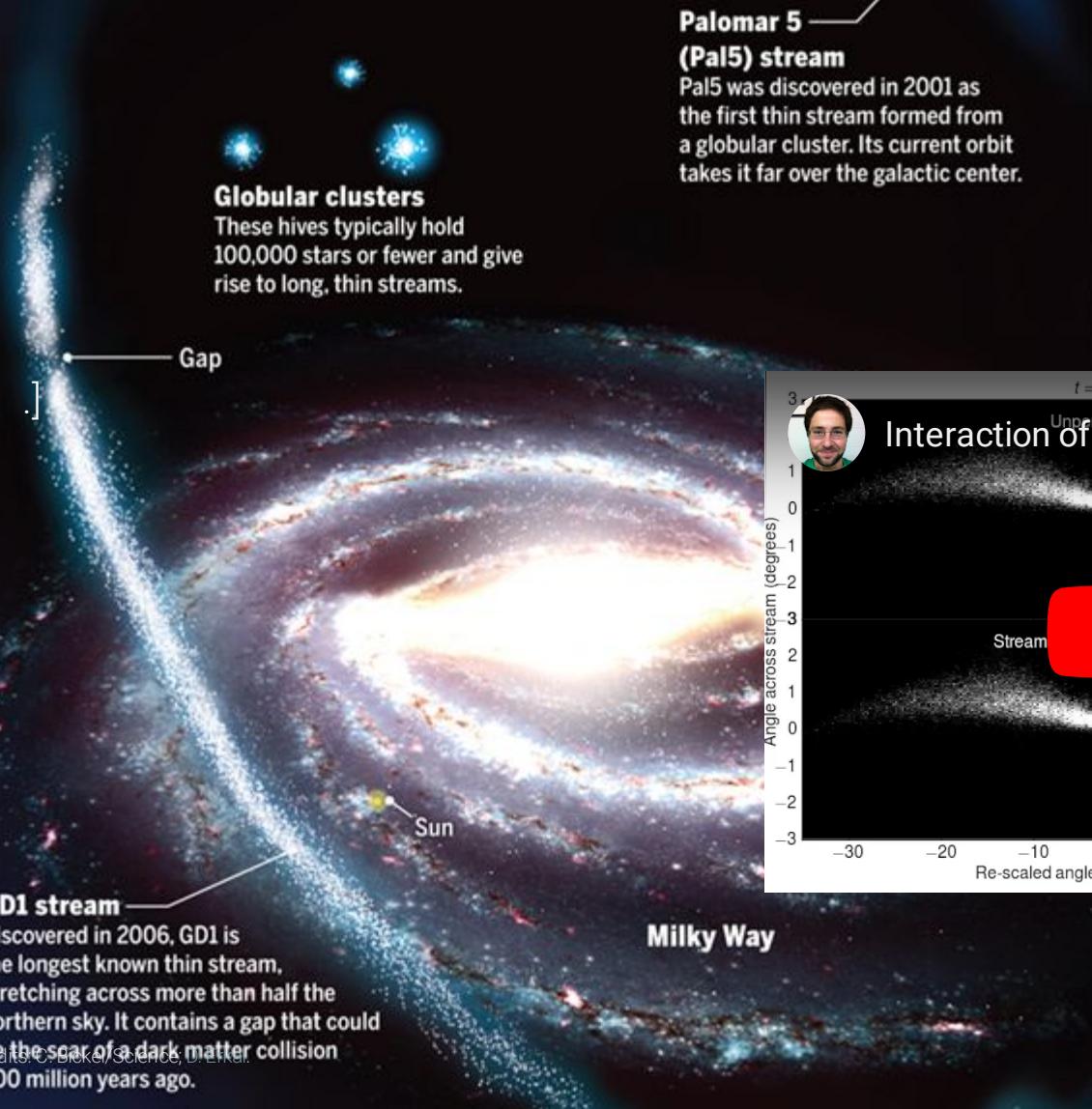
Run simulator and save data

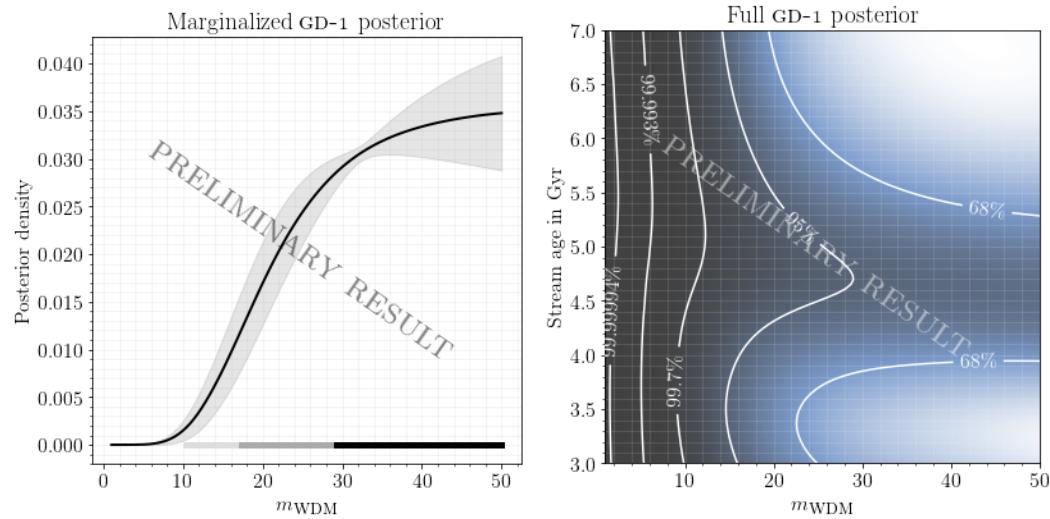
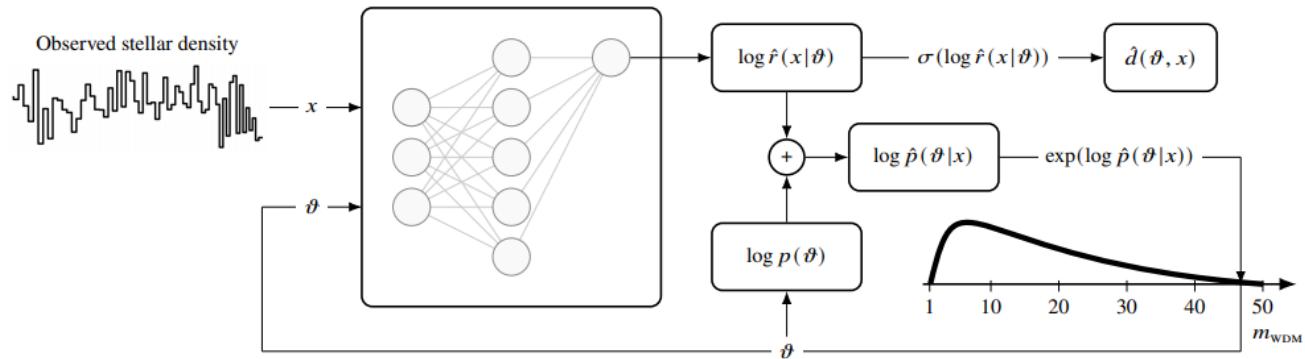
Train NN classifier, interpret as likelihood ratio estimator

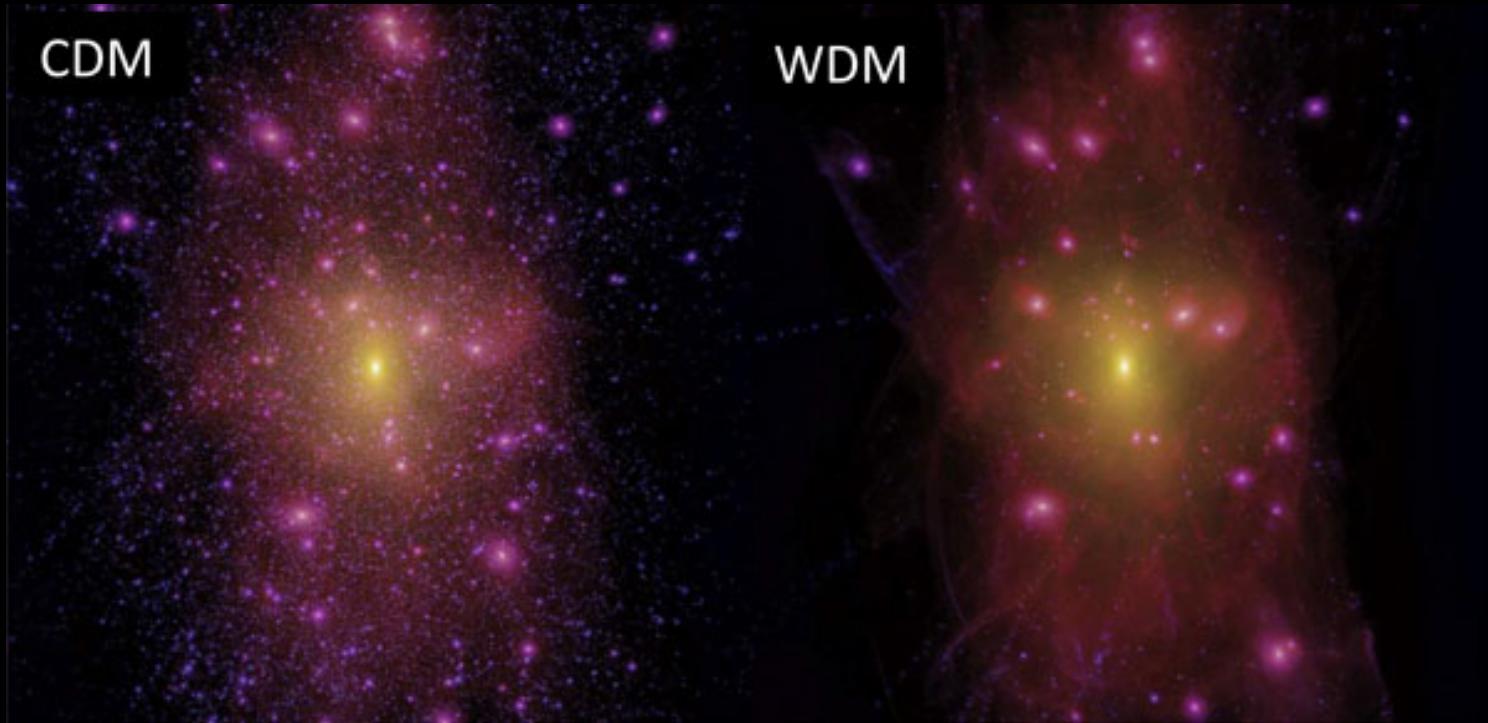
Amortized: cheap to repeat for new data

$$p(\theta|x) \approx \hat{r}(x|\theta)p(\theta)$$

# Constraining dark matter with stellar streams

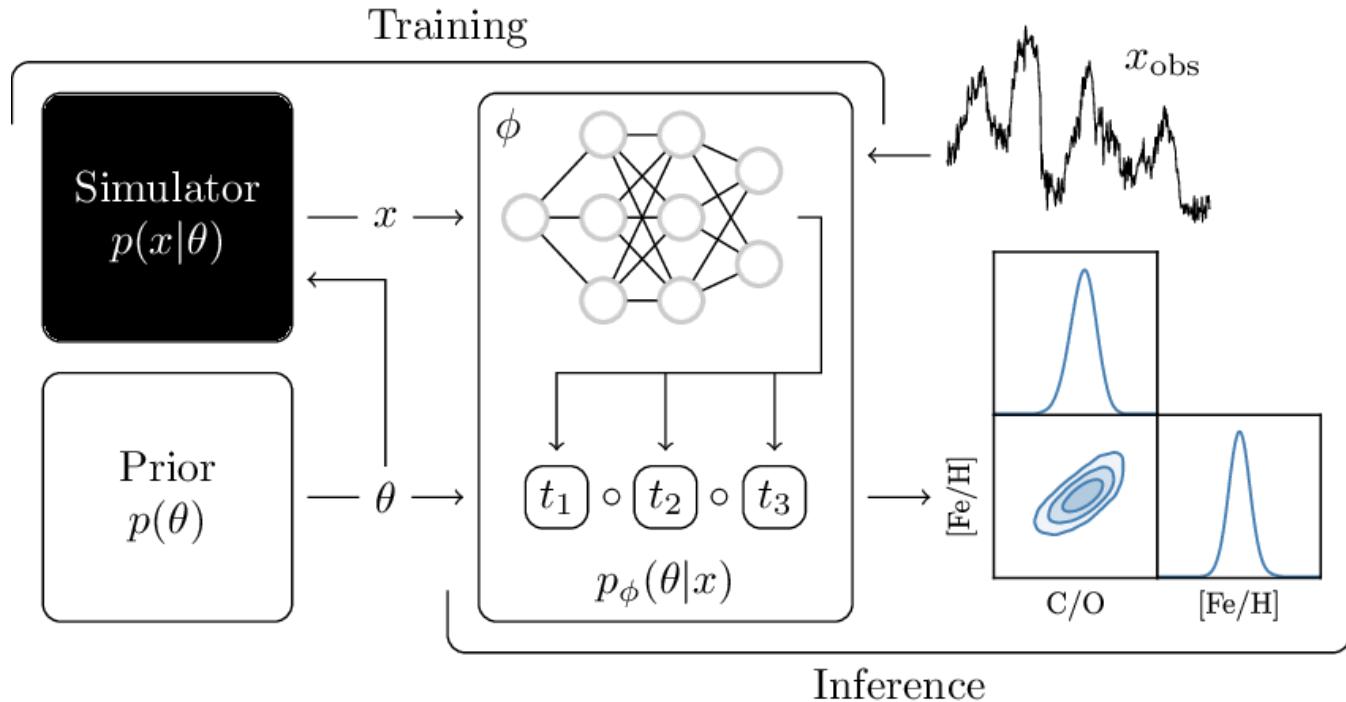






Preliminary results for GD-1 suggest a **preference for CDM over WDM**.

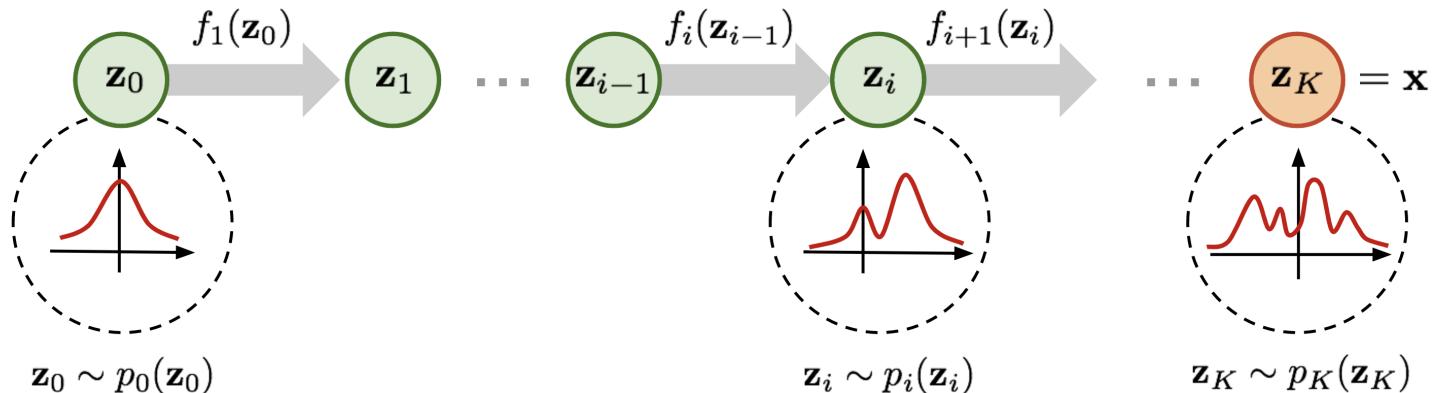
# Neural Posterior Estimation



$$\min_{q_\phi} \mathbb{E}_{p(x)} [\text{KL}(p(\theta|x) || q_\phi(\theta|x))]$$

## Normalizing flows

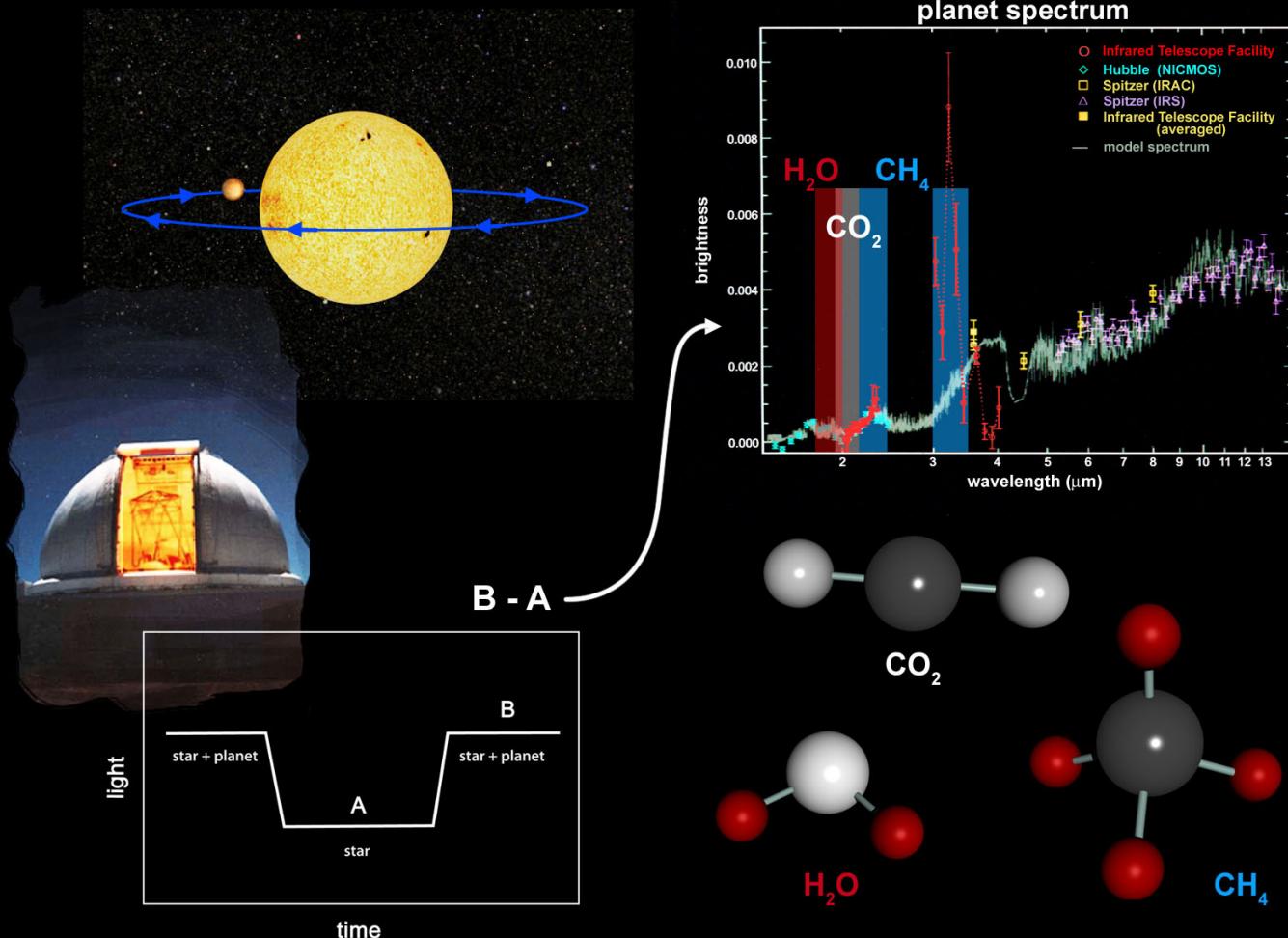
A normalizing flow is a sequence of invertible transformations  $f_k$  that map a simple distribution  $p_0$  to a more complex distribution  $p_K$ :

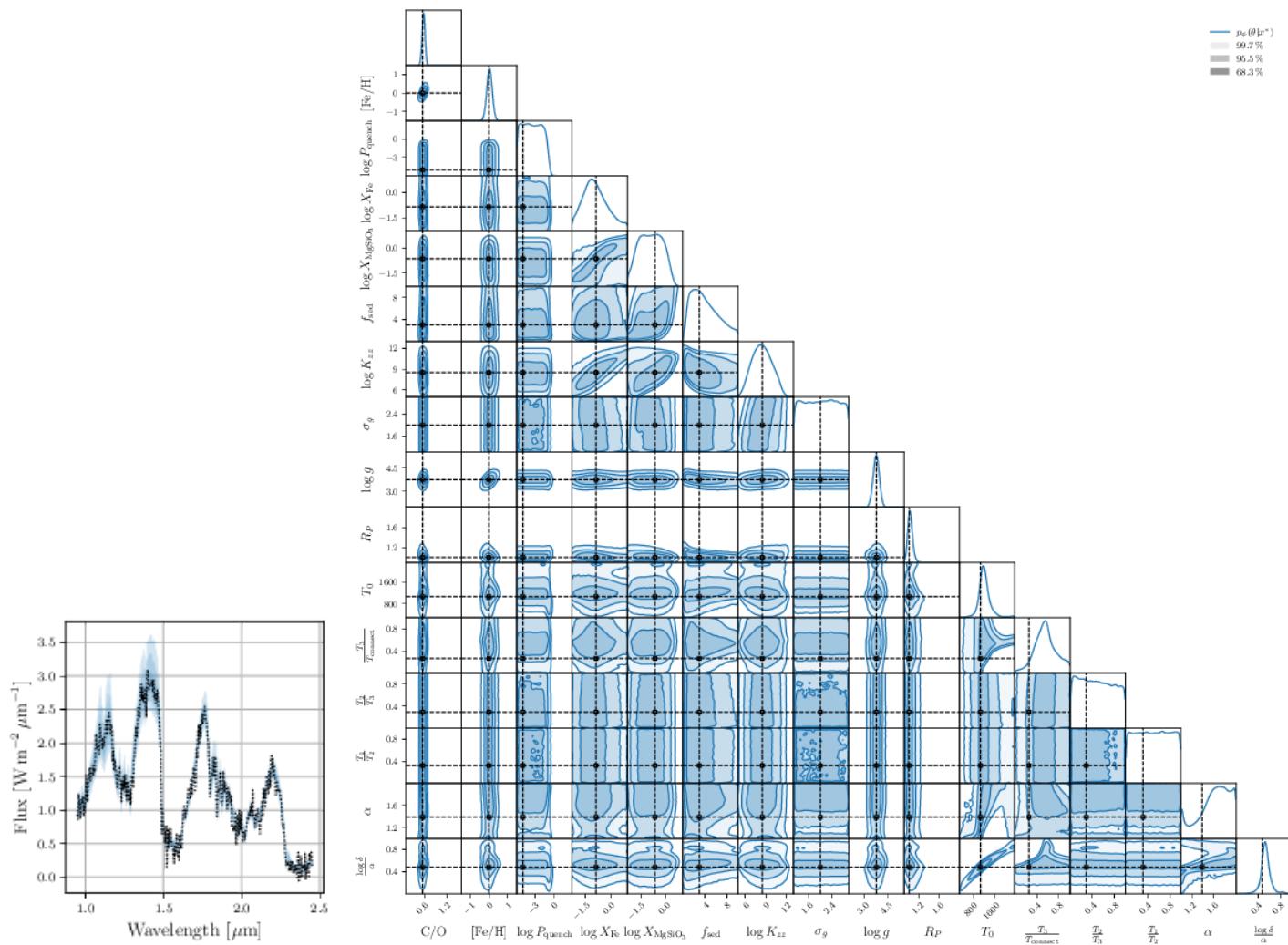


By the change of variables formula, the log-likelihood of a sample  $\mathbf{x}$  is given by

$$\log p(\mathbf{x}) = \log p(\mathbf{z}_0) - \sum_{k=1}^K \log |\det J_{f_k}(\mathbf{z}_{k-1})|.$$

# Exoplanet atmosphere characterization

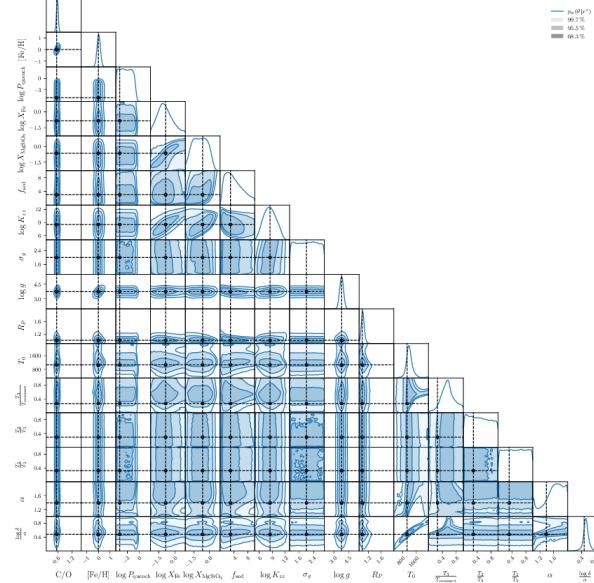




# Diagnostics

$$\hat{p}(\theta|x) = \text{sbi}(p(x|\theta), p(\theta), x)$$

We must make sure our approximate simulation-based inference algorithms can (at least) actually realize faithful inferences on the (expected) observations.



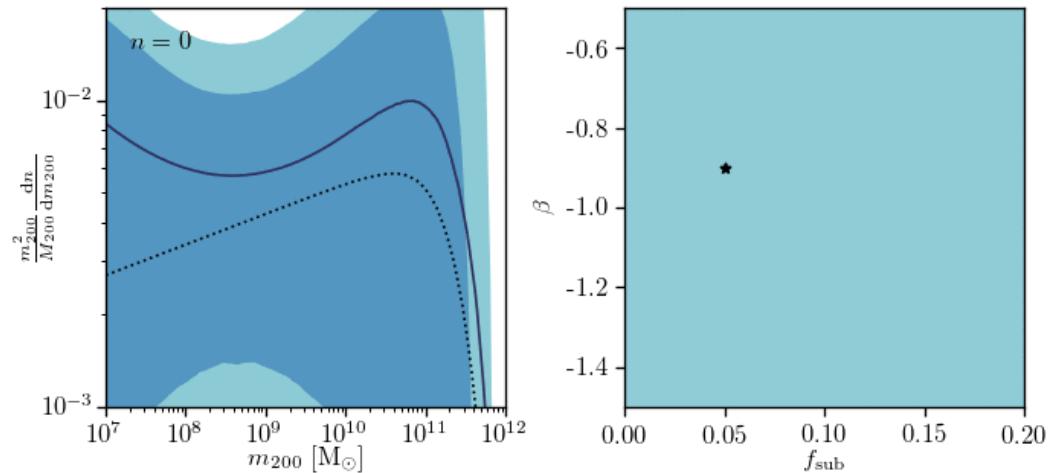
*How do we know this is good enough?*



## Mode convergence

The maximum a posteriori estimate converges towards the nominal value  $\theta^*$  for an increasing number of independent and identically distributed observables  $x_i \sim p(x|\theta^*)$ :

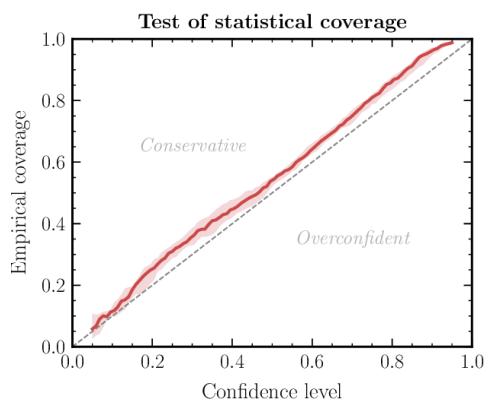
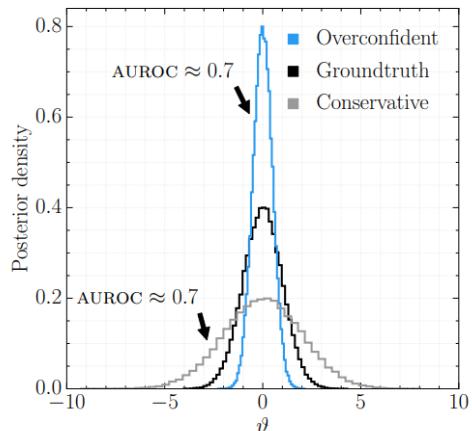
$$\begin{aligned} & \lim_{N \rightarrow \infty} \arg \max_{\theta} p(\theta | \{x_i\}_{i=1}^N) \\ &= \lim_{N \rightarrow \infty} \arg \max_{\theta} p(\theta) \prod_{x_i} r(x_i | \theta) = \theta^* \end{aligned}$$

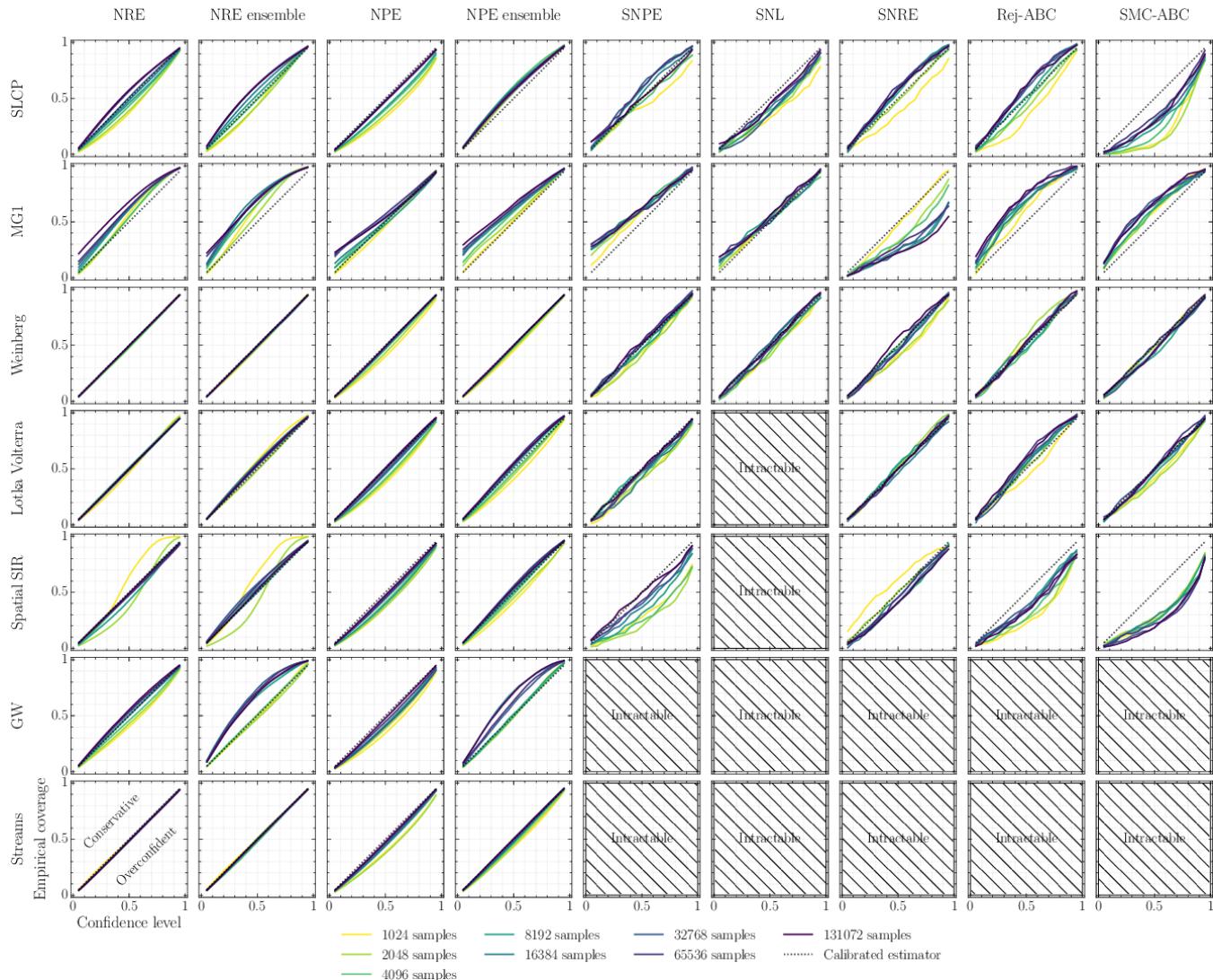




## Coverage diagnostic

- For  $x, \theta \sim p(x, \theta)$ , compute the  $1 - \alpha$  credible interval based on  $\hat{p}(\theta|x)$ .
- If the fraction of samples for which  $\theta$  is contained within the interval is larger than the nominal coverage probability  $1 - \alpha$ , then the approximate posterior  $\hat{p}(\theta|x)$  has coverage.





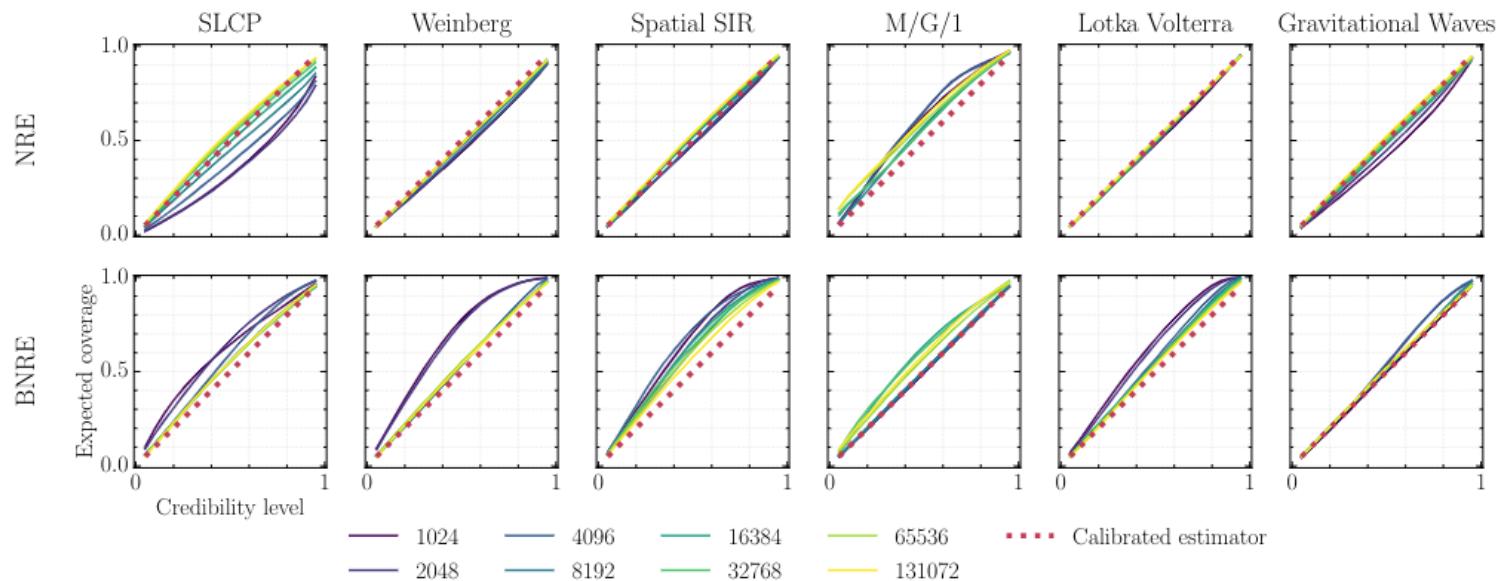
What if diagnostics fail?

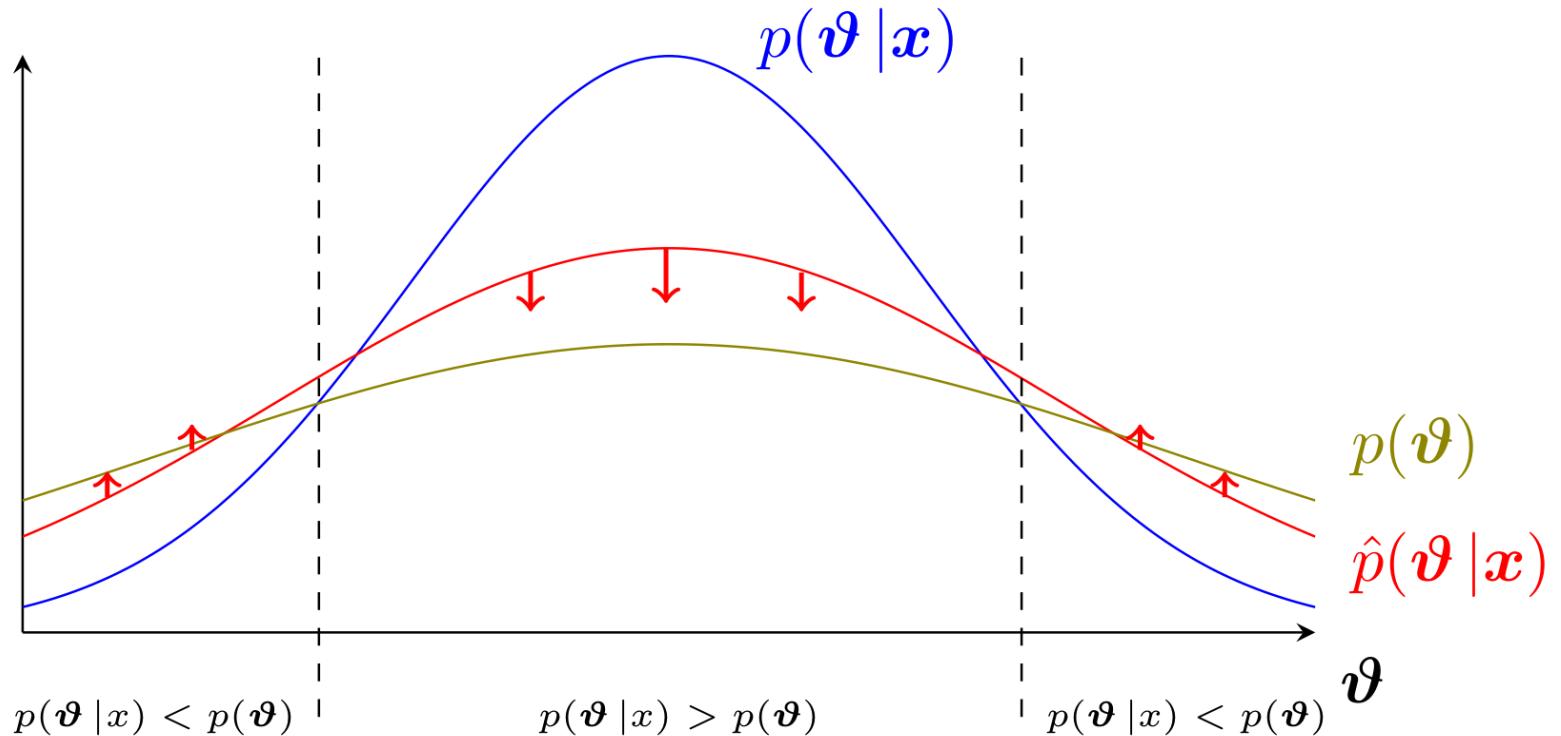
# Balanced NRE



Enforce neural ratio estimation to be **conservative** by using binary classifiers  $\hat{d}$  that are balanced, i.e. such that

$$\mathbb{E}_{p(\theta, x)} \left[ \hat{d}(\theta, x) \right] = \mathbb{E}_{p(\theta)p(x)} \left[ 1 - \hat{d}(\theta, x) \right].$$





# Summary

Advances in deep learning have enabled new approaches to statistical inference.

This is a major evolution in the statistical capabilities for science, as it enables the analysis of complex models and data without simplifying assumptions.

Inference remains approximate and requires careful validation.

Obstacles remain to be overcome, such as the curse of dimensionality and the need for large amounts of data.

