

Near-inertial waves and turbulence driven by the growth of swell

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1 Introduction

Winds flowing over the ocean excite a spectrum of surface gravity waves ranging from ripples to kilometer-long swells. Beneath persistent winds, the statistics of the shortest gravity waves saturate in a stationary balance between wind input, nonlinear interactions between wave components, and dissipation and breaking (Phillips, 1985). The saturated fraction of the wave spectrum is called the ‘equilibrium range’. Longer waves, or ‘swell’, are out of equilibrium by definition and therefore grow along the fetch of the wind.

Much of the momentum input into the ocean is transferred via form stress acting on the faces of equilibrium range waves (Grare *et al.*, 2013; Melville, 1996). The shortness of equilibrium range waves and the accompanying effects of wave breaking motivate the parameterization of air-sea momentum transfer as an effective stress imposed at the air-sea interface. Yet some fraction of the total momentum transferred between the atmosphere and ocean — perhaps as small as 5% in conditions typical to the laboratory and field experiments (Melville, 1996), or as large as 25% beneath hurricane-strength winds (Fan *et al.*, 2009) — is not transferred to the ocean at or just beneath the surface, but is instead *distributed in depth* by pressure gradients associated with the growth of swell. In this paper, we address the effects of the ‘swell-mediated’ fraction of the total momentum flux on turbulent ocean surface boundary layers.

Swell mediates momentum transfer between the atmosphere and ocean when swell is resonantly excited by coherent viscous stresses (Longuet-Higgins, 1969) or a coherent component of the atmospheric pressure spectrum. We illustrate this process by considering the oceanic response to the traveling atmospheric pressure disturbance

$$p_a(x, t) = \rho_0 P \cos \left(kx - \underbrace{\sqrt{gk} t}_{\stackrel{\text{def}}{=} \sigma} \right), \quad (1)$$

where ρ_0 is an ocean reference density, P is the ocean-side kinematic pressure, and g is gravitational acceleration. Over an ocean at rest, the pressure field in (1) excites an infinite, monochromatic surface wave field with wave number k and frequency $\sigma = \sqrt{gk}$ with surface displacement

$$s(x, t) = \underbrace{\frac{P}{2\sigma}}_{\stackrel{\text{def}}{=} a(t)} t \sin(kx - \sigma t). \quad (2)$$

In (2) we define the time-dependent swell amplitude, $a(t)$. If p_a vanishes after some time $t = T$, the outcome is a infinite, steadily-propagating wave field with amplitude $a(T)$. This outcome was first investigated in a seminal paper by Stokes (1847).

The depth-integrated, swell-averaged x -momentum budget for a non-rotating and inviscid ocean reveals the air-sea momentum transfer generated by (1),

$$\partial_t \left\langle \int_{-h}^s u dz \right\rangle = \frac{1}{\rho_0} \langle p_a \partial_x s \rangle, \quad (3)$$

where h is a depth at which momentum fluxes vanish, and the angle brackets $\langle \cdot \rangle$ denote a simultaneous average over horizontal directions and a time-average over the rapid oscillations of swell. The time-derivative in (3) is the rate of change of depth-integrated momentum over time-scales much longer than the oscillation of the surface wave field.

The term $\langle p_a \partial_x s \rangle$ on the right of (3) describes atmospheric pressure forces impacting the tilted ocean surface and is called ‘form stress’. With the pressure field in (1) and surface displacement in (2), we find that momentum is transferred to the ocean by form stress at the rate

$$\frac{1}{\rho_0} \langle p_a \partial_x s \rangle = \frac{(Pk)^2}{4\sigma} t. \quad (4)$$

A remarkable aspect of this scenario, first described by Stokes (1847), is that the form stress in (4) forces an irrotational current whose average velocity at a fixed position — the Eulerian-mean velocity $\langle u \rangle$ — is zero. In other words, (4) acting on a non-rotating ocean forces a current with $\langle u \rangle = 0$, and

$$\partial_t \left\langle \int_{-h}^s u dz \right\rangle = \partial_t \int_{-h}^0 u^S dz = \frac{(Pk)^2}{4\sigma} t, \quad \text{where} \quad u^S(z, t) = e^{2kz} a(t)^2 \sigma k \quad (5)$$

is called the ‘Stokes drift’ (Bühler, 2014). The rate-of-change of the depth-integrated Stokes drift (5) is equal to the air-sea momentum transfer due to form stress in (4).

In this paper, we numerically simulate the response of a rotating, stratified, turbulent boundary layer to the resonant atmospheric pressure forcing (1), as well as mixtures of effective surface stress and swell-mediated form stress with similar total momentum input. In these more general scenarios, the Eulerian-mean velocity is non-zero. However, the rate of change of the Stokes drift in (5) still measures the part of the air-sea momentum transfer rate associated with swell-coherent form stress. To form the total momentum budget, we define the complexified total boundary layer momentum and depth-integrated Stokes drift,

$$U^L(t) \stackrel{\text{def}}{=} \left\langle \int_{-h}^s u + iv dz \right\rangle, \quad \text{and} \quad U^S(t) \stackrel{\text{def}}{=} \int_{-h}^0 u^S + iv^S dz, \quad (6)$$

and parameterize the net air-sea momentum transfer rate with $\tau + \partial_t U^S$, where τ is a complexified effective surface stress that models the net effect of momentum transfer through swell-incoherent viscous stresses, form stress acting on equilibrium range waves, and the concomitant effects of wave breaking. The superscript ‘L’ in (6) stands for ‘Lagrangian-mean’, whose agency is discussed in section 2. The governing equation for the total momentum U^L , which may be derived either from the rotating Navier-Stokes equations beneath a free surface, or from the Craik-Leibovich wave-averaged Boussinesq equations (9)–(11) that form the basis for the simulations in this paper (Kukulka *et al.*, 2010), is

$$\partial_t U^L + if U^L = \tau + \partial_t U^S, \quad (7)$$

where f is the Coriolis parameter.

Ursell & Deacon (1950), Hasselmann (1970), Pollard (1970), and equation (7) show that, in the presence of rotation, the current that arises beneath a growing monochromatic wave field is not steady as in Stokes (1847)’s problem, but instead rotates inertially. A surprising result of our numerical simulations is that, due to three-dimensionality, stratification, and pre-existing turbulence, swell-transmitted flows both rotate inertially and transfer energy to turbulence via shear production, thus mixing and deepening the boundary layer. This is in striking contrast to the laminar, unidirectional flow that arises in the irrotational case analyzed by Stokes (1847) and exemplified by (5). The evolution of a turbulent boundary layer in our most basic case beneath a growing surface wave field with $\tau = 0$ is depicted in figure 1. Setting $\tau = 0$ distinguishes our work from Kukulka *et al.* (2010), Sullivan *et al.* (2012) and Large *et al.* (2019) that investigate time-dependent mixing processes due to a combination of time-dependent τ and $\partial_t U^S$, and which do not separately investigate the effects of the ‘additional’ momentum input by non-zero $\partial_t U^S$.

Section 2 introduces the wave-averaged Boussinesq equations and discusses some of their basic properties. In section 3, we continue beyond the scenario depicted in figure 1 to investigate the qualitative differences between the mixing and deepening of turbulent boundary layers forced by an effective surface

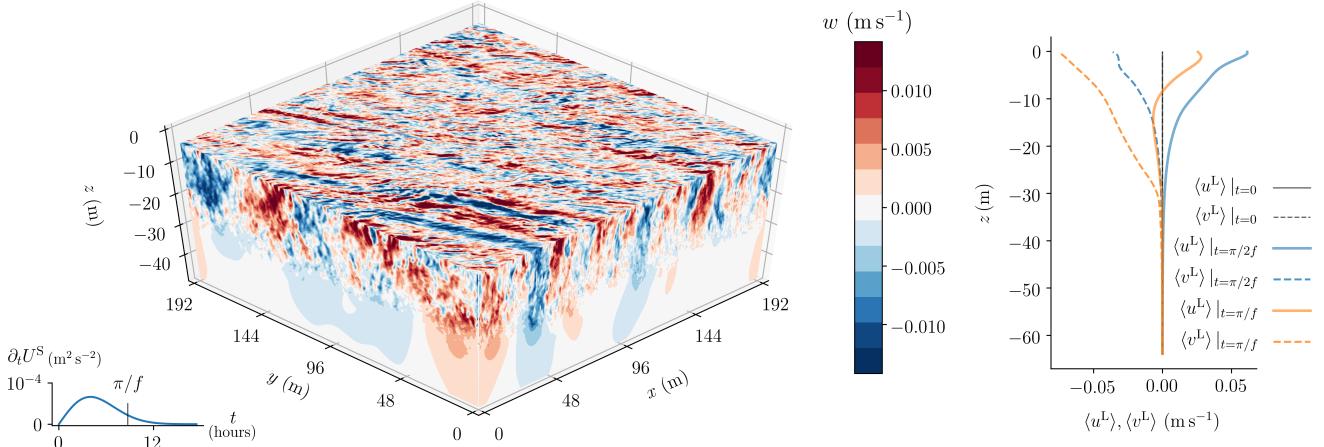


Figure 1: Currents and turbulence beneath forced surface waves with the time-dependent Stokes drift $\mathbf{u}^S = e^{2kz} a^2 k \sqrt{gk} \left(1 - e^{-t^2/2T_w}\right) \hat{\mathbf{x}}$ with wavenumber $k = 2\pi/100 \text{ m}^{-1}$, growth time scale $T_w = 4 \text{ hours}$, equilibrium amplitude $a = 2 \text{ meters}$, and gravitational acceleration $g = 9.81 \text{ m s}^{-2}$. Contours of w^L (left) and the horizontal-averaged horizontal velocity components $\langle u^L \rangle$ and $\langle v^L \rangle$ (right) are shown after half an inertial period at $t = \pi/f$. The tendency of the depth integral of the Stokes drift, $\partial_t U^S = \int_{-h}^0 u^S dz$, is a stress accompanying the growth of swell that forces a vertically-sheared inertial oscillation and drives turbulent mixing and boundary layer deepening. More details about the physics, numerics, and software can be found in sections 2–3, table 1 and at <https://github.com/glwagner/WaveTransmittedTurbulence.jl>.

stress τ and swell growth via $\partial_t U^S$. We conclude that depth-distributed forcing by $\partial_t U^S$ produces *less mixing* than forcing by τ , because $\partial_t U^S$ drives currents with weaker shear that relinquish less energy to turbulence. In shallow boundary layers, moreover, some of the swell-transmitted atmospheric forcing acts directly on laminar near-inertial motions below the base of the boundary layer. We emphasize that these observations pertain to the *interior* effects of boundary layer forcing mechanisms associated with the shear production of turbulence augmented by Langmuir turbulence phenomena, rather than the surface-concentrated effects of wave breaking, which are neglected in our large eddy simulations.

Langmuir turbulence is not the main focus of this work. Nevertheless, the wave-catalyzed organization of turbulent motions into the coherent structures of Langmuir turbulence (Sullivan & McWilliams, 2010; D’Asaro *et al.*, 2014) features prominently in our numerical solutions. In section 3.5, we observe that the coherent Langmuir turbulence structures tend to align with the Lagrangian-mean shear, consistent with Sullivan *et al.* (2012)’s results beneath realistic hurricane winds and waves. After the surface stress dies down, however, the orientation of the coherent structures decouples from the weakening Lagrangian-mean shear and locks onto $\hat{\mathbf{y}}$, the axis of the effective background vorticity associated with the $\hat{\mathbf{x}}$ -propagating swell. We suggest an explanation for this phenomenon in the context of the Lagrangian-mean formulation of the wave-averaged Boussinesq equations.

Section 4 investigates the importance of swell growth rate in determining initial conditions for large eddy simulations, and subsequent deepening of turbulent boundary layers under steady surface stress and steady surface waves. Section 5 concludes by discussing how our results may corroborate, *a posteriori*, some assumptions that underpin the parameterization of atmosphere-ocean momentum transfer in general circulation models. Appendix A describes the large eddy simulation software “Oceananigans.jl”, appendix B provides the vector calculus identities required to manipulate the Eulerian-mean Craik-Leibovich equations into their Lagrangian-mean form, and appendix C produces a table of the particular large eddy simulations referenced in this work.

2 The surface-wave-averaged Boussinesq equations

The Lagrangian-mean velocity beneath surface gravity waves is

$$\mathbf{u}^L \stackrel{\text{def}}{=} \mathbf{u}^E + \underbrace{\overline{\tilde{\boldsymbol{\xi}} \cdot \nabla \tilde{\mathbf{u}}}}_{\stackrel{\text{def}}{=} \mathbf{u}^S}, \quad (8)$$

where $\tilde{\mathbf{u}}$ is oscillatory velocity associated with waves, $\tilde{\boldsymbol{\xi}} = (\tilde{\xi}, \tilde{\eta}, \tilde{\zeta})$ is the wavy zero-mean particle displacement defined via $\partial_t \tilde{\boldsymbol{\xi}} = \tilde{\mathbf{u}}$, the overline ($\overline{}$) is a running Eulerian time-average — a ‘wave average’ — over surface wave oscillations, $\mathbf{u}^E \stackrel{\text{def}}{=} \bar{\mathbf{u}}$ is the Eulerian-mean velocity, and \mathbf{u}^S is the Stokes drift. The Lagrangian-mean velocity $\mathbf{u}^L(x, y, z, t)$ advects mass, momentum, and vorticity, and its variance $\frac{1}{2}|\mathbf{u}^L|^2$ is the wave-averaged kinetic energy.

The wave-averaged Boussinesq equations in an f -plane tangent to and rotating with the ocean surface and cast in terms of \mathbf{u}^L are (Craik & Leibovich, 1976; Huang, 1979; Leibovich, 1980; Holm, 1996; Suzuki & Fox-Kemper, 2016; Seshasayanan & Gallet, 2019)

$$\partial_t \mathbf{u}^L + (\mathbf{u}^L \cdot \nabla) \mathbf{u}^L + (f \hat{\mathbf{z}} - \nabla \times \mathbf{u}^S) \times \mathbf{u}^L + \nabla p = b \hat{\mathbf{z}} - \nabla \cdot \boldsymbol{\mathcal{T}} + \partial_t \mathbf{u}^S, \quad (9)$$

$$\partial_t b + \mathbf{u}^L \cdot \nabla b = -\nabla \cdot \mathbf{q}, \quad (10)$$

$$\nabla \cdot \mathbf{u}^L = 0 \quad (11)$$

where p is Eulerian-mean kinematic pressure, b is Eulerian-mean buoyancy, f is the Coriolis parameter, and $\boldsymbol{\mathcal{T}}$ and \mathbf{q} are the stress tensor and buoyancy flux due either to molecular diffusion or a subfilter turbulent diffusion model for large eddy simulation. We show how (9)–(11) are derived from the Eulerian-mean form of the Craik-Leibovich equations, and how they are connected to the generalized Lagrangian-mean equations derived by Andrews & McIntyre (1978) in appendix B.

Equations (9)–(11) are an asymptotic approximation of the Navier-Stokes equations beneath a small amplitude and weakly modulated surface wave field. In particular, (11) neglects a term relevant for time-dependent swell and associated with divergence of the vertical component of the Stokes drift. If accounted for, this vertical divergence would lift the mean position of each fluid parcel by $\frac{1}{2}a^2k e^{2kz}$ as the swell amplitude a increases (see McIntyre (1981) and equation 3.7 in Longuet-Higgins (1986)). In the cases considered in this paper, however, the vertical velocity associated with this mean vertical displacement is minuscule: for example, the growth of $a = 2$ m amplitude swell with wavenumber $k = 2\pi/100$ m $^{-1}$ lifts the sea surface by just $\frac{1}{2}a^2k = 0.12$ m over a period of 4 hours. Because the effect is so small, we neglect the vertical component of the Stokes drift and prescribe $w^L = 0$ at $z = 0$.

The momentum equation (9) is often written as a prognostic equation for the Eulerian-mean velocity, \mathbf{u}^E , both for analysis (Craik & Leibovich, 1976; Suzuki & Fox-Kemper, 2016) and large eddy simulation (Skillingstad & Denbo, 1995; McWilliams *et al.*, 1997; Noh *et al.*, 2004; Polton & Belcher, 2007; Harcourt & D'Asaro, 2008; Yang *et al.*, 2015). We use the Lagrangian-mean velocity \mathbf{u}^L instead as our prognostic variable for both numerical simulations and analysis. Using \mathbf{u}^L explicitly identifies the role of atmospheric momentum forcing transmitted via growing swell, and thus $\partial_t \mathbf{u}^S$, in driving the turbulent evolution of the simulated surface boundary layers in section 3.

Prescribing $\mathbf{u}^S(\mathbf{x}, t)$ in (9) determines the effects of swell on the evolution of the Lagrangian-mean momentum, \mathbf{u}^L . Vertical fluxes of horizontal momentum into \mathbf{u}^L through saturated surface waves in the equilibrium range are prescribed through stress boundary conditions on \mathcal{T}_{xz} and \mathcal{T}_{yz} at $z = 0$. Buoyancy fluxes are prescribed through $q_z|_{z=0}$. The complexified *downwards* surface stress in (7) is $\tau \stackrel{\text{def}}{=} -\mathcal{T}_{xz} - i\mathcal{T}_{yz}|_{z=0}$.

2.1 Wave-averaged effective background vorticity

The term

$$\underbrace{(\underbrace{f\hat{z} - \nabla \times \mathbf{u}^S}_{\stackrel{\text{def}}{=} \Omega^\dagger}) \times \mathbf{u}^L}_{(12)},$$

appears in the momentum equation (9), where $f\hat{z}$ is the planetary vorticity and $\nabla \times \mathbf{u}^S$ is the surface wave pseudovorticity. As discussed by Bühler (2014) in their section 11.3.2, (12) means that the *effective* background vorticity Ω^\dagger is advected by the Lagrangian-mean velocity \mathbf{u}^L . In other words, the total vorticity is

$$\Omega \stackrel{\text{def}}{=} \Omega^\dagger + \nabla \times \mathbf{u}^L = f\hat{z} + \nabla \times \underbrace{(\mathbf{u}^L - \mathbf{u}^S)}_{\stackrel{\text{def}}{=} \mathbf{u}^E}, \quad (13)$$

in the sense that $\nabla \times$ (9) with $b = 0$ and $\mathcal{T} = 0$ is

$$\partial_t \Omega + (\mathbf{u}^L \cdot \nabla) \Omega = (\Omega \cdot \nabla) \mathbf{u}^L, \quad (14)$$

the usual vorticity equation describing the three-dimensional advection of vortex lines (see again Bühler (2014), section 11.3.2). Thus while \mathbf{u}^L transports mass, momentum, and vorticity, the wave-averaged deviation from planetary vorticity is $\nabla \times \mathbf{u}^E$, rather than $\nabla \times \mathbf{u}^L$. This altered relationship between momentum and vorticity encapsulates the dynamical effect of surface waves on the evolution of \mathbf{u}^L . We use the interpretation of Ω^\dagger as an effective background vorticity to explain some of the phenomena observed in our large eddy simulations in section 5.

2.2 Wave-averaged kinetic energy

The conservation law for volume-integrated wave-averaged kinetic energy, $\frac{1}{2}|\mathbf{u}^L|^2$, in unstratified flow with $b = 0$ and inviscid flow with $\mathcal{T} = 0$ follows from $\int \mathbf{u}^L \cdot (9) dV$,

$$\frac{d}{dt} \int \frac{1}{2} |\mathbf{u}^L|^2 dV = \int \mathbf{u}^L \cdot \partial_t \mathbf{u}^S dV, \quad (15)$$

where we have assumed there are no momentum fluxes across the boundary of V . The total mean kinetic energy $\int \frac{1}{2} |\mathbf{u}^L|^2 dV$ is therefore conserved beneath steady surface waves in unstratified, inviscid flow: there is no energy exchange between \mathbf{u}^L and steady surface waves. Equation (15) shows that forced surface waves with non-zero $\partial_t \mathbf{u}^S$ are a source of oceanic momentum and kinetic energy.

Substituting $\mathbf{u}^E = \mathbf{u}^L - \mathbf{u}^S$ into equation (15) yields a formula for the volume-integrated Eulerian-mean kinetic energy $\frac{1}{2}|\mathbf{u}^E|^2 = \frac{1}{2}|\mathbf{u}^L|^2 + \frac{1}{2}|\mathbf{u}^S|^2 - \mathbf{u}^L \cdot \mathbf{u}^S$,

$$\frac{d}{dt} \int \frac{1}{2} |\mathbf{u}^E|^2 dV = - \int \mathbf{u}^S \cdot \partial_t \mathbf{u}^E dV. \quad (16)$$

Due to the term on the right of (16), Eulerian-mean kinetic energy is not conserved in unstratified, inviscid flow beneath steady surface waves.

3 Large eddy simulations beneath growing swell

We consider the forced growth of two deep water swells with amplitudes $a = 1$ meter and $a = 2$ meters. Both swells grow over a time scale of $T_w = 4$ hours, have 100 meter wavelength, 8 second period, and the horizontal Stokes drift profiles

$$u^S(z, t) = \underbrace{e^{2kz} a^2 k \sqrt{gk}}_{\stackrel{\text{def}}{=} u_{\text{eq}}^S(z)} \left(1 - e^{-t^2/2T_w^2} \right), \quad (17)$$

where $g = 9.81 \text{ m s}^{-2}$. $k = 2\pi/100 \text{ m}^{-1}$ and $a = 1, 2 \text{ m}$ in (17) are chosen (*i*) to illustrate the swell-mediated transmission of momentum below a shallow boundary layer; and (*ii*) to illustrate how stronger and thus deeper turbulent mixing increases the effectiveness with which the swell-transmitted momentum mixes the boundary layer.

The Stokes acceleration $\partial_t u^S$ and total stress $\partial_t U^S = \int_{-h}^0 \partial_t u^S dz$ associated with (17) are

$$\partial_t u^S(z, t) = \frac{t e^{-t^2/2T_w^2}}{T_w^2} u_{\text{eq}}^S(z), \quad \text{and} \quad \partial_t U^S = \frac{t e^{-t^2/2T_w^2}}{T_w} \underbrace{\frac{a^2 \sqrt{gk}}{2T_w}}_{\stackrel{\text{def}}{=} \tau_w}. \quad (18)$$

In (18), the equilibrium Stokes profile $u_{\text{eq}}(z)$ is defined in (17), and τ_w scales the stress exerted on the ocean by the atmosphere via forced swell. The maximum vertically-integrated momentum forcing of the boundary layer due to (17) is

$$\max(\partial_t U^S) = \partial_t U^S|_{t=T_w} = \frac{\tau_w}{\sqrt{e}} \approx \begin{cases} 1.65 \times 10^{-5} \text{ m}^2 \text{s}^{-2} & \text{for } a = 1 \text{ m}, \\ 6.61 \times 10^{-5} \text{ m}^2 \text{s}^{-2} & \text{for } a = 2 \text{ m}, \end{cases} \quad (19)$$

similar to the stress exerted by 12-hour-long wind pulses with maximum speed 4 m s^{-1} and 8 m s^{-1} .¹

3.1 Numerical methods and software

Our large eddy simulations solve (9)–(11), where \mathbf{u}^L is the *resolved* Lagrangian-mean velocity filtered to remove scales smaller than the numerical grid scale. Spatial filtering motivates down-gradient approximations for the subfilter stress tensor and diffusive flux,

$$\mathcal{T}_{ij} = -2\nu_e \Sigma_{ij}^L, \quad \text{and} \quad q_i = -\kappa_e \partial_i b, \quad (20)$$

in terms of the filtered strain tensor $\Sigma_{ij}^L = \frac{1}{2} (\partial_i u_j^L + \partial_j u_i^L)$ and filtered buoyancy gradient $\partial_i b$, where the indices $i = (1, 2, 3)$ correspond to the Cartesian directions (x, y, z) . The eddy viscosity ν_e and eddy diffusivity κ_e in (20) are modeled with the anisotropic minimum dissipation (AMD) formalism introduced by Rozema *et al.* (2015) and Abkar *et al.* (2016), refined by Verstappen (2018), and validated and described in detail for ocean-relevant scenarios by Vreugdenhil & Taylor (2018). Additional details about (20) are given in appendix A.

We solve (9)–(11) numerically with Oceananigans, a software package developed by the authors in the Julia programming language that runs on Graphics Processing Units (GPUs) (Bezanson *et al.*, 2012; Besard *et al.*, 2018). The simulations in this paper use second-order finite volume spatial discretization, second-order Adams-Bashforth time discretization, a pressure projection method to ensure $\nabla \cdot \mathbf{u}^L = 0$, and a fast method based on the fast Fourier transform to solve the pressure Poisson equation discretized with second-order differences on a regular grid (Schumann & Sweet, 1988). Oceananigans code and documentation are hosted at <https://github.com/CliMA/Oceananigans.jl>.

Our simulations are performed in two rectangular domains: one ‘regular’ size domain with dimensions $128 \times 128 \times 64 \text{ m}$ in x, y, z , grid spacings $0.5 \times 0.5 \times 0.25 \text{ m}$, and resolution 256^3 , and a second ‘large’ domain of dimension $192 \times 192 \times 96 \text{ m}$ with $0.75 \times 0.75 \times 0.25 \text{ m}$ grid spacing and resolution $256 \times 256 \times 384$. The domains are horizontally periodic in x, y and have rigid top and bottom boundaries, where we impose $w^L = 0$. To absorb internal waves radiated downwards from the turbulent surface boundary layer, we implement bottom sponge layers of the form

$$F_\phi = e^{-(z+H)/\delta} \mu (\phi^\dagger - \phi), \quad (21)$$

¹With drag coefficient $C_d = 10^{-3}$, air density $\rho_a = 1.225 \text{ kg m}^{-3}$, seawater density $\rho_w = 1035 \text{ kg m}^{-3}$, and ocean-side kinematic stress parameterized with $|\tau| \approx \frac{\rho_a}{\rho_w} C_d u_a^2$, we find $\tau = 1.65 \times 10^{-5} \text{ m}^2 \text{s}^{-2}$ for $u_a = 3.7 \text{ m s}^{-1}$.

for each variable $\phi \in (u, v, w, b)$, where H is the depth of the domain, $\mu = \frac{1}{60} \text{ s}^{-1}$, $\delta = 4 \text{ m}$, $b^\dagger = N^2 z$, and $u^\dagger = v^\dagger = w^\dagger = 0$. Each F_ϕ is added to its corresponding equation in (9)–(10).

The simulations reported in this section and the rest of this paper are listed in table 1.

3.2 Generation of a weakly turbulent initial condition

We use preliminary simulations to generate a shallow, weakly turbulent boundary layer initial condition. The spin-up simulations are forced by weak cooling associated with an upwards flux of buoyancy at the surface,

$$q_z|_{z=0} = 5 \times 10^{-10} \text{ m}^2 \text{ s}^{-3}. \quad (22)$$

The spin-up simulations are run for half an inertial period until $t_{\text{spin}} = \pi/f$, where $f = 10^{-4} \text{ s}^{-1}$ is the Coriolis parameter. The boundary layer depth at the end of the spin up is approximately 8 meters. The horizontally-averaged buoyancy and velocity at $t_{\text{spin}} = \pi/f$ are shown in figure 2 and serve as initial conditions for subsequent simulations reported in this section, which use the same domain, resolution, and sponge layer configuration.

3.3 Near-inertial waves, turbulence, and boundary layer deepening

Growing swells with the increasing Stokes drift profile in (17) and equilibrium amplitudes $a = 1 \text{ m}$ and $a = 2 \text{ m}$ are imposed on the weakly-turbulent end state of the spin-up simulation. The simulation time is reset to $t = 0$ in (17), so that $\mathbf{u}^S|_{t=0} = 0$. We impose free slip conditions on the horizontal velocity components and no normal flow on the vertical velocity component at top and bottom boundaries,

$$w^L = \partial_z u^L = \partial_z v^L = 0 \quad \text{at } z = 0 \quad \text{and } z = -H. \quad (23)$$

The boundary conditions (23) ensure there is no source of momentum from the boundaries. We impose a no-flux condition on buoyancy at the top boundary,

$$\partial_z b|_{z=0} = 0. \quad (24)$$

The bottom boundary condition $\partial_z b|_{z=-H} = N^2$ and the bottom sponge layer in (21) restore the near-bottom buoyancy profile to $N^2 z$.

Figure 1 plots contours of vertical velocity after half an inertial period $w^L|_{t=\pi/f}$ and depth profiles of the $\langle u^L \rangle$ and $\langle v^L \rangle$ for the strong swell case with equilibrium amplitude $a = 2 \text{ m}$. Near-inertial shear excited by the growing surface wave field drives turbulence that mixes and deepens the boundary layer. At the same time, mixing is enhanced by the organization of turbulent motions into the coherent structures of Langmuir turbulence, which manifest in the left panel of figure 1 as elongated rolls of alternating positive and negative vertical velocity oriented at roughly 10° angles from the y -axis. The penetration of vertical motions through the stratified base of the boundary produces smooth downwards-emitting bulbs of vertical velocity at around $z = -25 \text{ m}$. Turbulent motions at the base of the boundary layer excite downward-propagating internal waves. The right panel shows the inertial rotation and downwards turbulent penetration of the swell-transmitted current in time.

3.4 Surface-concentrated versus depth-distributed stress

To isolate the relative effect of distributed momentum forcing associated with $\partial_t \mathbf{u}^S$ in (18), we run three additional simulations for $a = 1 \text{ m}$ and $a = 2 \text{ m}$ each with the boundary condition

$$\mathcal{T}_{xz}|_{z=0} = -\partial_t U^S = -\tau_w \frac{t e^{-t^2/2T_w^2}}{T_w}, \quad (25)$$

where τ_w is defined in (18). The boundary condition (25) prescribes a 4 hour pulse of surface stress. We combine (25) with the Stokes drift profiles

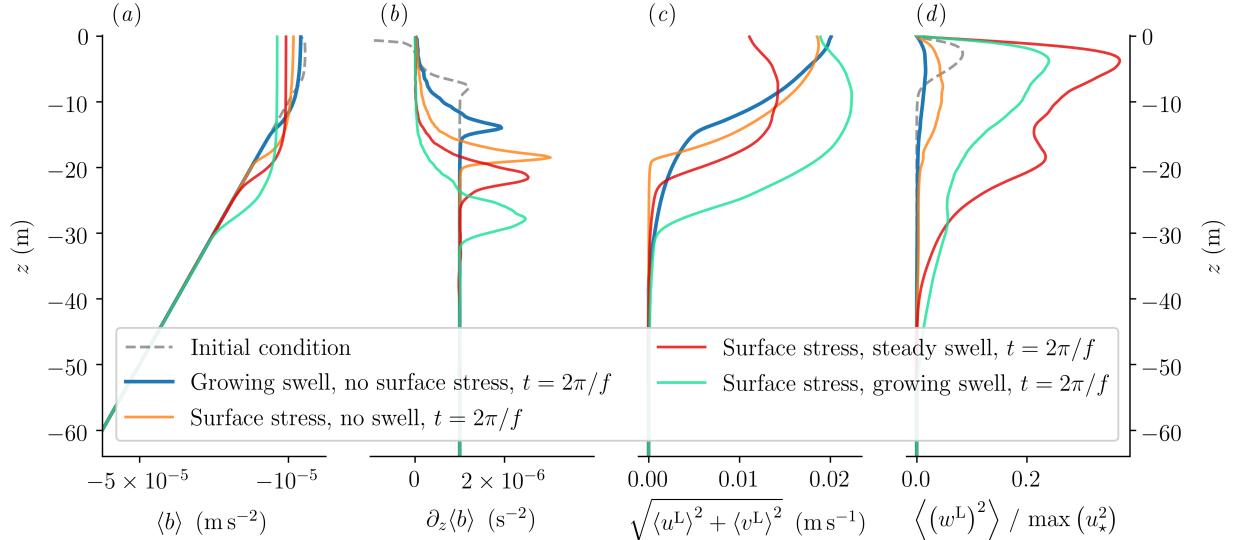


Figure 2: Horizontally-averaged fields in three LES forced by growing swell with equilibrium amplitude $a = 2$ m (solid lines), a pulse of surface stress with no waves (dashed lines) a pulse of surface stress beneath steady waves with amplitude $a = 1$ m. The panels a–d show horizontally-averaged buoyancy, buoyancy gradient, speed, and vertical velocity variance normalized by $\max(u_*^2) = 1.65 \times 10^{-5} \text{ m}^2 \text{s}^{-2}$.

1. $u^S = 0$ (surface stress, no swell), and
2. $u^S = u_{\text{eq}}^S(z)$ for $u_{\text{eq}}^S(z)$ in (17) (surface stress, steady swell).

For $a = 1$ m we also run a simulation forced by both the surface stress in (25) and growing swell with $\mathbf{u}^S(z, t)$ from (17), which we call ‘surface stress, growing swell’. The cases described in the previous section 3.3 are called the ‘growing swell, no surface stress’ cases.

Figure 2 compares horizontally-averaged profiles of buoyancy, buoyancy gradient, horizontal speed $\sqrt{\langle u^L \rangle^2 + \langle v^L \rangle^2}$, and vertical velocity variance $\langle (w^L)^2 \rangle$ after one inertial period at $t = 2\pi/f$ in the four simulations that correspond to an equilibrium wave amplitude of $a = 1$ m. The maximum value of the friction velocity,

$$\max(u_*) \stackrel{\text{def}}{=} \sqrt{\max(\partial_t U^S)} \approx 4 \text{ mm s}^{-1} \quad (\text{for } a = 1 \text{ m}) \quad (26)$$

is used to normalize speed and vertical velocity variance. Turbulent mixing produced by growing swell deepens the boundary layer from $z \approx -8$ meters at $t = 0$ to $z \approx -12$ meters at $t = 2\pi/f$. A substantial fraction of the total momentum is transmitted to laminar, non-turbulent near-inertial motions below $z \approx -12$, whose shear does not contribute to turbulent mixing. The fingerprint of momentum transmitted by swell below the boundary layer is the blue exponential tail in figure 2c below $z = -12$ meters. The boundary layers driven by surface stress are deeper, therefore, because more of the mean kinetic energy transmitted to the boundary layer is converted to turbulence.

Figure 2d shows the average vertical velocity variance $\langle (w^L)^2 \rangle$. The case forced by surface stress and beneath steady swell has the strongest vertical velocities. We hypothesize this is due to the formation of vigorous Langmuir structures beneath strong, steady swell in the presence of surface stress. Vertical velocities are somewhat weaker in the case with both surface stress *and* growing swell, perhaps because the Stokes drift is not as strong during active momentum forcing by surface stress. Nevertheless, the surface stress, growing swell case — which experiences the strongest horizontal momentum forcing by both surface stress and swell-transmitted stress — boasts the deepest boundary layer and the strongest horizontal velocities.

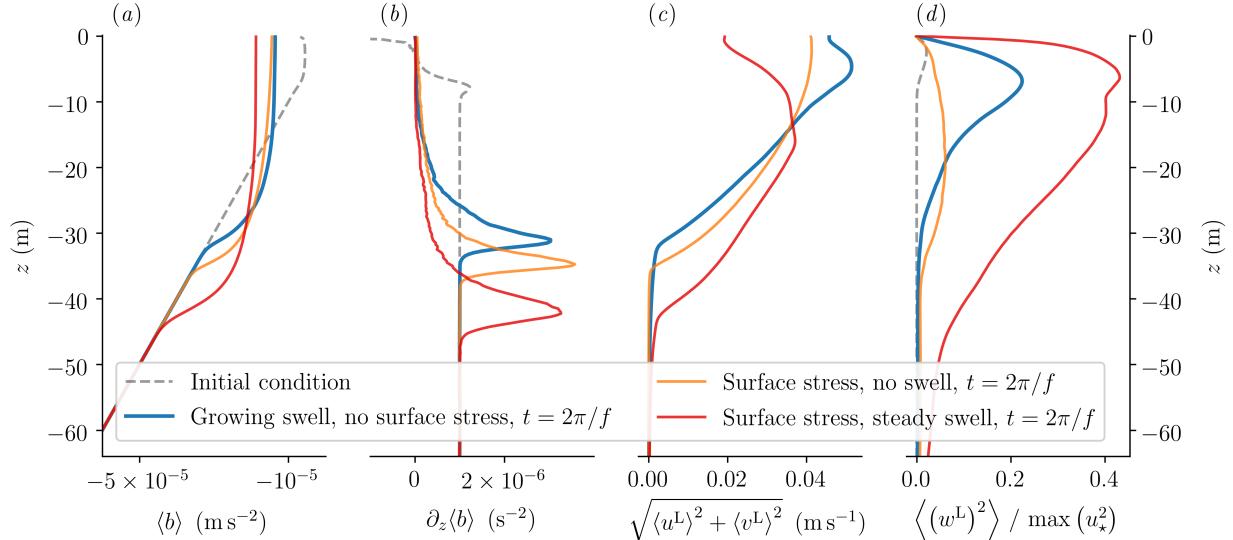


Figure 3: Horizontally-averaged fields in three LES forced by growing swell with equilibrium amplitude $a = 2$ m (solid lines), a pulse of surface stress with no waves (dashed lines) a pulse of surface stress beneath steady waves with amplitude $a = 2$ meters. The panels a-d show horizontally-averaged buoyancy, buoyancy gradient, speed, and vertical velocity variance normalized by $\max(u_*^2) = 6.61 \times 10^{-5} \text{ m}^2 \text{s}^{-2}$.

Figure 3 is similar to figure 2 but for equilibrium wave amplitude $a = 2$ meters. The maximum value of the friction velocity for $a = 2$ m is

$$\max(u_\star) = \sqrt{\max(\partial_t U^S)} \approx 8 \text{ mm s}^{-1}. \quad (27)$$

The crucial difference between the results depicted in figure 2 and the results in 3 is that turbulence in the growing swell simulation mixes the boundary layer to $z \approx -32$ meters. This greater deepening enables greater turbulence production by shear transmitted by the growing swell, and thus mixing rates more comparable to the boundary layers driven by surface stress. The dramatic differences in vertical velocity variance between the blue lines in figure 2d and 3d are evidence that Langmuir turbulence is more active and effective in the growing swell simulation with $a = 2$ meters than in the growing swell simulation with $a = 1$ meter. Despite strong vertical velocities beneath growing swell (blue line, figure 3d), which suggest the presence of Langmuir turbulence (McWilliams *et al.*, 1997; Sullivan & McWilliams, 2010), the boundary layer is deeper in the surface stress case with no swell (blue and orange lines, figure 3b).

Figure 4 illustrates the temporal evolution of the horizontally-averaged x -velocity $\langle u^L \rangle(z, t)$ and the horizontally-averaged turbulent kinetic energy

$$E(z, t) \stackrel{\text{def}}{=} \left\langle \frac{1}{2} |\mathbf{u}^L - \langle \mathbf{u}^L \rangle|^2 \right\rangle. \quad (28)$$

The total stress exerted on the boundary layer for the ‘growing waves’ and ‘surface stress, steady waves’ simulations, which are plotted in the top right and left panels of figure 4, are identical. Yet because surface stress is concentrated at the top of the domain, it drives faster, more strongly sheared currents that transfer more of their energy to turbulent kinetic energy. The turbulent kinetic energy with surface stress and steady wave overhead is roughly $3\times$ larger than with growing waves and no surface stress. The temporal structure is also different: a burst of turbulence appears beneath growing waves around $t = 8$ hours; while a turbulent layer steadily penetrates the stratified fluid below the surface stress over the 8 hour period of significant forcing.

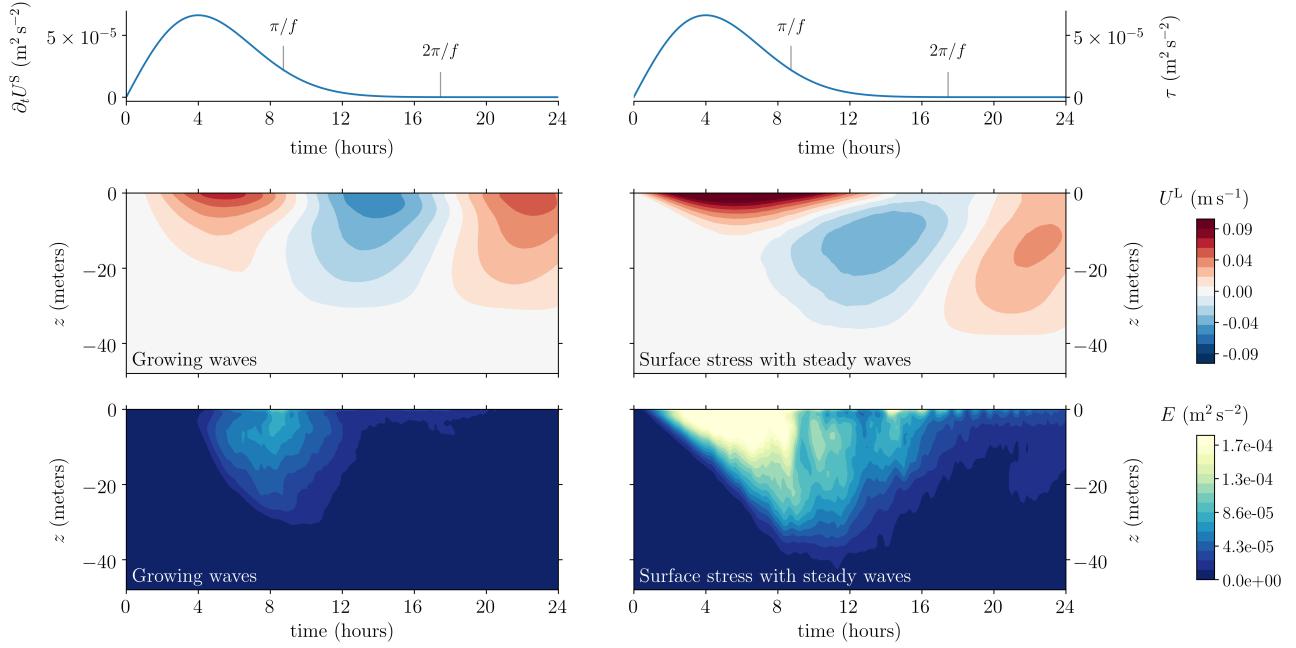


Figure 4: A depth-time of the horizontally-averaged x -velocity, $\langle u^L \rangle(z, t)$, and turbulent kinetic energy $E(z, t)$ defined in (28).

3.5 Alignment of freely-decaying coherent Langmuir turbulence structures perpendicular to Stokes drift

Figure 5 plots a time-series of normalized vertical velocity contours at $z = -2$ meters and $z = -8$ meters for the surface stress, steady waves case with $a = 2$ m. A black arrow in the center of each plot indicates the direction of the mean horizontal flow at each time and depth. The time series shows that the coherent Langmuir turbulence structures rotate and grow in size over time, especially after the forcing dies out after $t \approx \frac{3}{4} \times 2\pi/f$.

The rotation of the coherent Langmuir turbulence structures appears fixed to the mean flow direction at early times up to $t = \frac{1}{2} \times 2\pi/f$, when momentum forcing is significant. After $t = \frac{1}{2} \times 2\pi/f$, and as the mean shear weakens, the cell appear to rotate into alignment with $\hat{\mathbf{y}}$ at around $t = 2\pi/f$. $\hat{\mathbf{y}}$ is the axis of the background vorticity associated with the surface waves,

$$-\nabla \times \mathbf{u}^S = \partial_z u^S \hat{\mathbf{y}}. \quad (29)$$

Thus the coherent structures align with the direction of mean shear at early times and during active forcing, as in Sullivan *et al.* (2012), but rotate onto the axis of the surface wave pseudovorticity at later times, during free decay.

4 Turbulent mixing following rapid and gradual surface wave growth

In this section we reveal some differences between the turbulent evolution of boundary layers forced by the (*i*) rapid and (*ii*) gradual growth of a surface wave field. In short, when $f \neq 0$, the rapid growth of a surface wave field leads to the momentum distribution $\mathbf{u}^L = \mathbf{u}^S$ and thus $\mathbf{u}^E = 0$, while the gradual

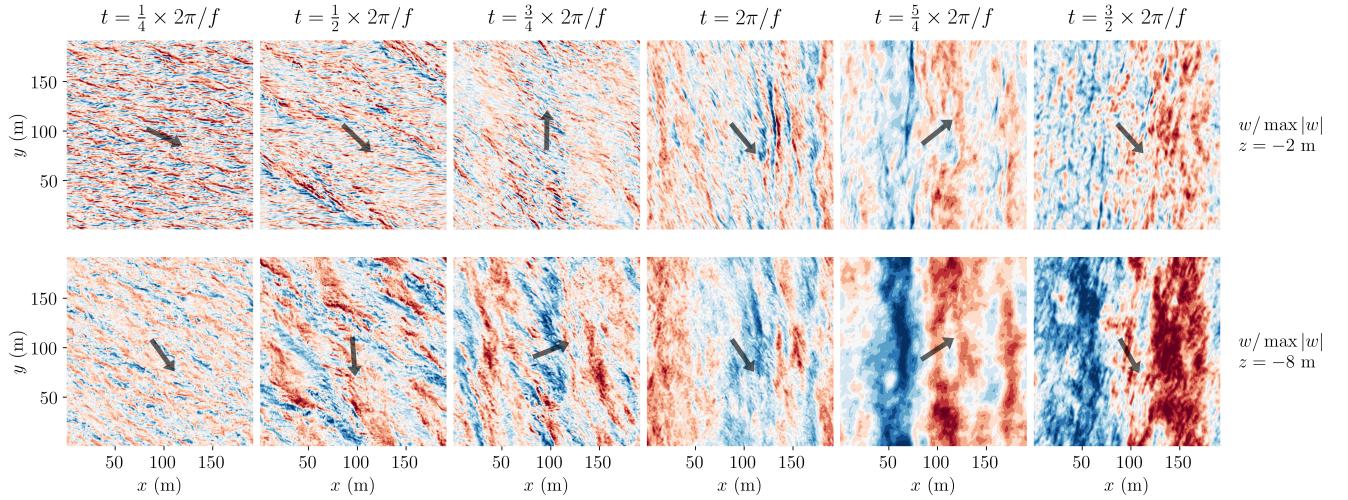


Figure 5: Contours of vertical velocity normalized by its maximum absolute value at $z = -2$ m and $z = -8$ m from the strong ‘surface stress with steady waves’ simulation (simulation 8 in table 1), showing the rotation of coherent Langmuir turbulence structures into \hat{y} . A thick arrow in the middle of the domain indicates the direction of the horizontally-averaged horizontal flow at the same depth as the vertical velocity. The simulation is actively forced for $t \approx 12$ hours; the flow begins to decay freely after $t \approx \frac{3}{4} \times 2\pi/f$. At early times, the coherent structures are roughly aligned with the direction of mean shear, which is aligned with \hat{x} , the direction of forcing. As the current rotates inertially, so do the coherent structures, though after $t \approx \pi/f$ the rotation rate of the structures begins to slow until $t \approx 2\pi/f$, at which point the structure axes are fixed to \hat{y} . Between $t = \frac{5}{4} \times 2\pi/f$ and $t = \frac{3}{2} \times 2\pi/f$, the mean shear rotates through \hat{x} , perpendicular to the orientation of the coherent structures, which destroys much of the structures’ coherence (not shown). The aftermath of this shearing event is apparent from comparing the vertical velocity at $z = -2$ m: much of the coherency evident at $t = \frac{5}{4} \times 2\pi/f$ is destroyed at $t = \frac{3}{2} \times 2\pi/f$. The strength and coherency of the Langmuir turbulence structures thus pulsates as the near-inertial current rotates, strengthening when the current is aligned or anti-aligned with \hat{y} , and shearing and disintegrating when the mean current is perpendicular to \hat{y} .

growth of a surface wave field leads to $\mathbf{u}^L = 0$. We explore the difference between these two initial conditions on turbulent mixing driven by surface stress using large eddy simulations.

4.1 Rapid and gradual surface wave growth over laminar boundary layers

If τ is a prescribed function of time, the solution to (7) is

$$U^L(t) = U^L(0) + e^{-ift} \int_0^t e^{ift'} (\tau + \partial_t U^S) dt', \quad (30)$$

In the limit that surface waves grow rapidly with $\tau = 0$, such that $\partial_t U^S = U_{\text{eq}}^S \delta(t)$ for some equilibrium Stokes transport U_{eq}^S , (30) is an inertial oscillation,

$$U^L(t) = U^L(0) + e^{-ift} U_{\text{eq}}^S. \quad (31)$$

If surface waves grow over much longer time-scales than $1/f$, on the other hand, then $U^L \approx 0$.

A generalization of the vertically-integrated result in (31) occurs beneath surface waves with the time-dependent Stokes drift field

$$\mathbf{u}^S(z, t) = \mathbf{u}_{\text{eq}}^S(z) H(t), \quad (32)$$

where $H(t)$ is the Heaviside function and $\mathbf{u}_{\text{eq}}^S(z)$ is the equilibrium Stokes drift. In this case, the Stokes drift tendency is $\partial_t \mathbf{u}^S = \mathbf{u}_{\text{eq}}^S \delta(t)$ and Lagrangian-mean velocity field just after $t = 0$, when the surface wave Stokes drift profile is steady, is

$$\lim_{t \rightarrow 0^+} \mathbf{u}^L = \mathbf{u}_{\text{eq}}^S(z). \quad (33)$$

This flow has zero Eulerian-mean current, and corresponds to an ocean boundary layer forced by the rapid growth of a surface wave field. If stable, the initial condition (33) develops into the perfect inertial oscillation

$$u^L(z, t) + i v^L(z, t) = e^{-ift} u_{\text{eq}}^S(z), \quad (34)$$

with constant mean kinetic energy

$$\frac{1}{2} |\mathbf{u}^L|^2 = \frac{1}{2} (u_{\text{eq}}^S)^2, \quad (35)$$

and oscillatory Eulerian-mean kinetic energy

$$\frac{1}{2} |\mathbf{u}^E|^2 = (1 - \cos ft) (u_{\text{eq}}^S)^2. \quad (36)$$

In the absence of sources of vorticity such as rotation, stratification, or viscous stresses, the flow $\mathbf{u}^E = 0$ is stable and cannot transition to turbulence.² In stratified surface-wave-averaged flows, however, stability is guaranteed only when the Lagrangian-mean Richardson number is greater than $\frac{1}{4}$ (Holm, 1996),

$$\text{Ri}^L \stackrel{\text{def}}{=} \frac{\partial_z b}{(\partial_z u^L)^2 + (\partial_z v^L)^2} > \frac{1}{4}. \quad (37)$$

4.2 Large eddy simulations of turbulent mixing driven by surface stress following rapid and gradual surface wave growth

We use large eddy simulations to analyze the turbulent mixing of the ocean surface boundary layer following either rapid surface wave growth leading to (33), or gradual surface wave growth leading to

$$\mathbf{u}^L|_{t=0} = 0. \quad (38)$$

²While the wave-averaged momentum is \mathbf{u}^L , the wave-averaged vorticity is $\nabla \times \mathbf{u}^E$ (see Bühler (2014) or chapter 2.4 in Wagner (2016)). Thus without a source of vorticity, a flow with $\mathbf{u}^E = 0$ is irrotational and cannot transition to turbulence.

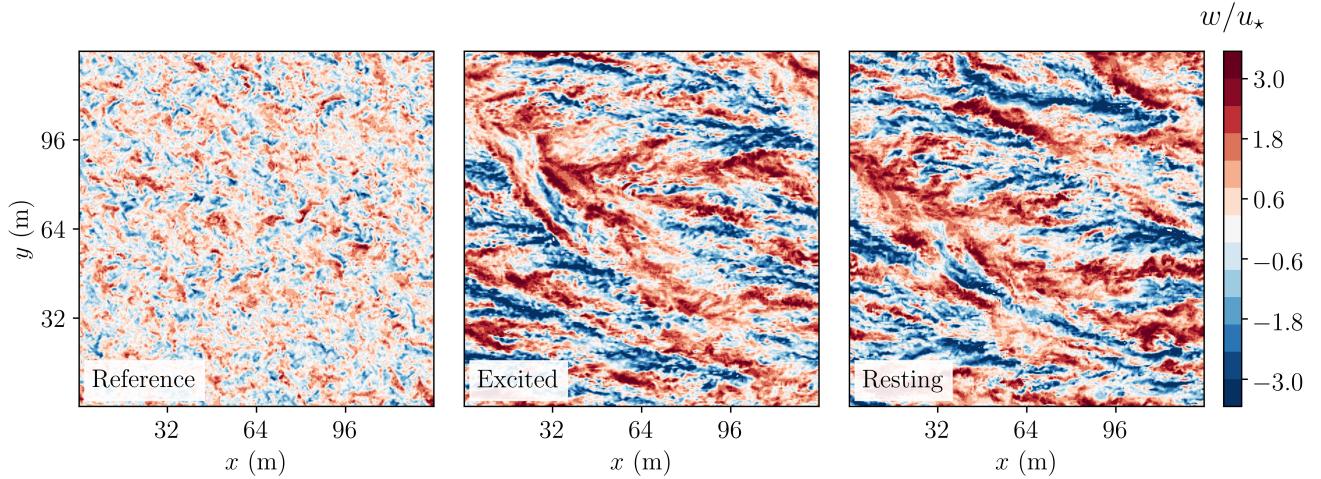


Figure 6: Contours of vertical velocity in the xy -plane and at a depth $z = -4$ meters in the reference, ‘1× excited’, and ‘1× resting large eddy simulations. Simulation parameters are detailed in table 1.

We refer to (33) as the ‘excited’ state, and (38) as the ‘resting’ state.

We choose parameters such that our results resemble those of McWilliams *et al.* (1997). McWilliams *et al.* (1997) uses the Stokes drift

$$\mathbf{u}^S(z) = e^{2kz} a^2 k \sqrt{gk} \hat{\mathbf{x}}. \quad (39)$$

corresponding to a monochromatic deep water wave with amplitude $a = 0.8$ m and wavenumber $k = 0.105 \text{ m}^{-1}$. The surface stress and surface buoyancy flux are

$$\mathcal{T}_{xz}|_{z=0} = 3.72 \times 10^{-5} \text{ m}^2 \text{s}^{-2}, \quad (40)$$

$$q_z|_{z=0} = -2.31 \times 10^{-8} \text{ m}^2 \text{s}^{-3}. \quad (41)$$

Turbulent mixing is mostly driven by the surface stress in (40). The buoyancy flux in (41), which has a negligible impact on overall turbulent mixing, serves mainly to counteract the spurious laminarization of the near-surface velocity field in (McWilliams *et al.*, 1997). Our initial buoyancy profile is

$$b|_{t=0} = N^2 z \quad \text{with} \quad N^2 = 1.94 \times 10^{-5} \text{ s}^{-2}, \quad (42)$$

superposed with small random noise. With $\partial_z b = N^2$ in (42) and $\partial_z u^S = 2e^{2kz}(ak)^2\sqrt{gk}$, $a = 0.8$ m, $k = 0.105 \text{ m}^{-1}$, and $g = 9.81 \text{ m s}^{-2}$, $\text{Ri}^L \approx 0.09$ at $t = 0$ and $z = 0$ such that the sufficient condition for instability in (37) is met. The simulations are performed in a rectangular domain with dimensions $128 \times 128 \times 64$ m, grid spacing $0.5 \times 0.5 \times 0.25$ m, resolution 256^3 , and with sponge layers of the form (21).

We refer to cases with $a = 0.8$ m in (39) ‘1×’ cases, and run two additional cases with excited and resting initial conditions and the same boundary conditions, except with 4× stronger Stokes drift fields

$$\mathbf{u}_{4\times}^S = 4\mathbf{u}^S, \quad (43)$$

corresponding to surface waves with twice the amplitude. We finally run a ‘reference’ simulation with $\mathbf{u}^S = 0$. These five simulations are labeled A–E in table 1. A visualization of vertical velocity in the xy -plane after two inertial periods such that $t = 2 \times 2\pi/f$, at a depth of $z = -4$ meters, and for the reference case, 1× excited case, and 1× resting case, are shown in figure 6. Both the 1× excited and 1× resting cases exhibit the coherent structures of well-developed Langmuir turbulence (Sullivan & McWilliams, 2010).

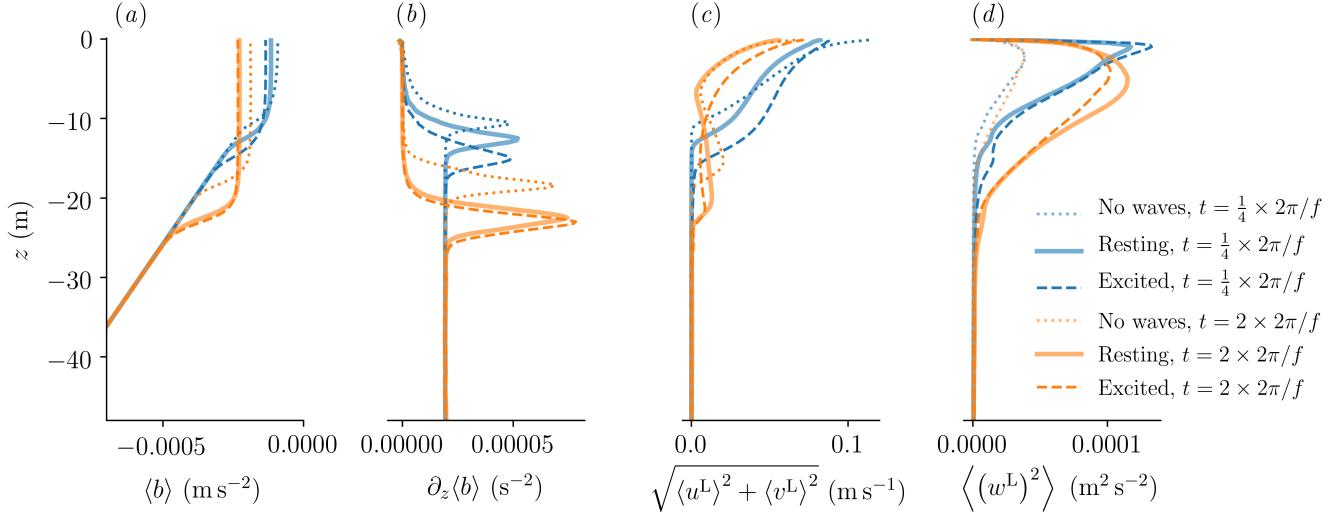


Figure 7: Horizontally-averaged fields in ‘1×’ large eddy simulations. (a) buoyancy, (b) buoyancy gradient, (c) speed, and (d) vertical velocity variance. ‘Excited’ initial conditions (dashed lines) cause an initial mixing event to deepen the boundary layer compared to ‘resting’ initial conditions (solid lines). This is most evident in the buoyancy gradient profiles at $t = \frac{1}{4} \times 2\pi/f$ (blue lines in b). By $t = 2 \times 2\pi/f$, memory of the initial condition is lost and the buoyancy, buoyancy gradient, and vertical velocity variance profiles are similar between the excited and resting case (orange dashed and solid lines in a, b, and d). The horizontal velocities in the excited case still exhibit faster speeds (orange dashed and solid lines in c) due to the excitation of a strong inertial oscillation at $t = 0$. The differences between the ‘no waves’ and ‘excited’ cases are discussed in McWilliams *et al.* (1997). Figure 10 reproduces some of the plots from McWilliams *et al.* (1997). Simulation parameters are detailed in table 1.

The main result of this section is that the ‘excited’ initial conditions in (33) provide a reservoir of near-inertial shear that quickly transitions to turbulence and rapidly deepens the boundary layer at early times. This rapid early deepening, compared to the more gradual evolution of boundary layers initialized with the resting state (38), is illustrated in figure 7 and 8, which depict the horizontally-averaged buoyancy, vertical buoyancy gradient, and Lagrangian-mean velocity fields $\langle u^L \rangle$ and $\langle v^L \rangle$ after a quarter of an inertial period (blue lines, $t = \frac{1}{4} \times 2\pi/f$) and after two inertial periods (orange lines, $t = 2 \times 2\pi/f$) in the $1\times$ and $4\times$ cases, respectively. The effect of $1\times$ excited initial conditions is modest — the boundary layer is approximately 20% deeper than the $1\times$ resting case at $t = \frac{1}{4} \times 2\pi/f$ — and the two boundary layer depths become similar by $t = 2 \times 2\pi/f$ as boundary layer deepening slows.

Because the surface wave field is stronger in the $4\times$ cases, the effect of initial enhanced shear in the ‘excited’ case is more dramatic: after $\frac{1}{4}$ of an inertial period, the boundary layer is almost twice as deep in the excited simulation as in the ‘resting’ and ‘reference’ simulations. The imprint of the initial excited state is still evident even at $t = 2 \times 2\pi/f$, at which point the excited boundary layer is still deeper than the boundary layer spun up from a resting state. We conclude with the visualization of the evolution of vertical velocity in figure 9 from the $4\times$ resting and excited experiments at the early time $t = \frac{1}{4} \times 2\pi/f$. In the $4\times$ excited simulation, turbulence penetrates deeper and regions of organized vertical velocity are stronger, reflecting both the greater energy available for turbulent mixing in the excited cases, and the organization of the more energetic turbulence into stronger coherent structures.

5 Discussion

The simulation illustrated by figure 1 shows that the growth of surface waves can excite sheared near-inertial waves and drive turbulence and mixing that deepens the ocean surface boundary layer. The simulations in section 3 show that, more generally, the partitioning of ocean momentum forcing into a surface-concentrated component and a depth-distributed component mediated by growing swell impacts

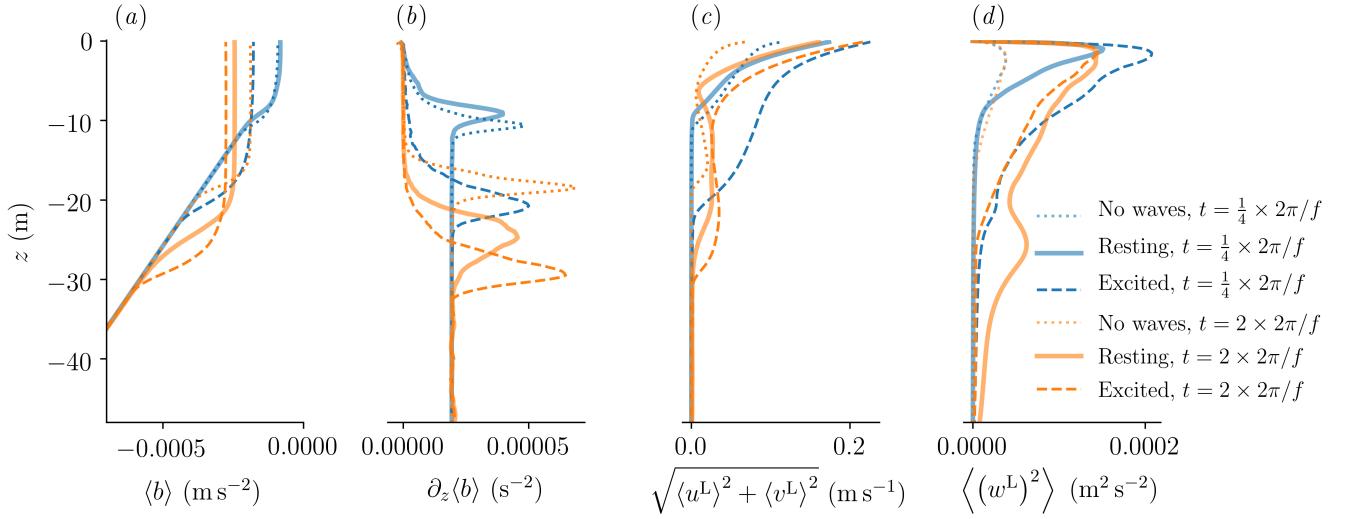


Figure 8: Like figure 7, except for ‘4×’ large eddy simulations with 4× stronger wave fields and excited initial conditions. The difference between the boundary layer depth at $t = \frac{1}{4} \times 2\pi/f$ for the resting and excited cases is dramatic, as evidenced by the blue dashed and solid lines in panel *b*. Unlike figure 7, memory of the initial mixing event persists at $t = 2 \times 2\pi/f$. The ‘resting’ case exhibits stronger vertical velocities than the excited case at $t = 2 \times 2\pi/f$, especially at the base of the boundary layer. The reason for this is unknown.

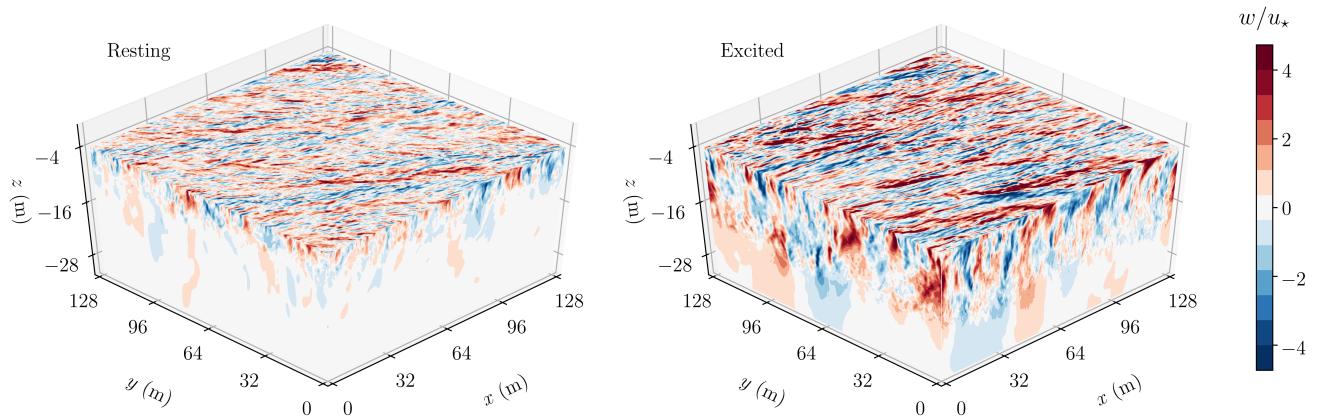


Figure 9: Comparison of the vertical velocity at $t = \frac{1}{4} \times 2\pi/f$ in two large eddy simulations with ‘resting’ and ‘excited’ initial conditions.

the ensuing mean currents and turbulent evolution and deepening of the ocean surface boundary layer. The simulations in section 4 show that the rate at which swell grows also affects the evolution of the boundary layer, a fact that is well-appreciated in the context of surface-concentrated momentum forcing, but less so for swell-mediated momentum forcing.

The distinction between momentum transfer via equilibrium range waves and viscous stress near or at the surface, and depth-distributed momentum transfer via the growth of swell may be important when $\partial_t U^S$ comprises a significant fraction of the total water-side stress $\tau + \partial_t U^S$. For example, Fan *et al.* (2009) find that $\partial_t U^S$ rises to 25% of the total stress in strongly-forced hurricane conditions. However, depending on the quantity of interest, the differences between boundary layers driven by τ and boundary layers driven by $\partial_t U^S$ are probably small in typical scenarios when $\partial_t U^S$ is small compared to τ . This is especially true in boundary layers deeper than the longest swell components, where preexisting turbulence capably converts swell-deposited kinetic energy to turbulent kinetic energy.

The *unimportance* of the depth-dependence of atmospheric momentum forcing under typical conditions has implications for parameterizations of atmospheric momentum forcing in general circulation models. Current general circulation models do not partition atmospheric momentum forcing into the two components in (7). Instead, general circulation models impose atmospheric momentum forcing through a surface stress. This approximation, which is sensible due to the fact that 90–95% of the stress exerted on the ocean by the atmosphere enters via equilibrium range waves with very short decay scales (Melville, 1996), is *further justified* by our results, which suggest that the depth-dependence of the remaining 5–10% of the stress exerted on the ocean is relatively unimportant except in very shallow boundary layers. More realistic observations and modeling are warranted to further evolution of boundary layers forced by both surface stress and swell growth.

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A Subfilter fluxes in wave-averaged large eddy simulations

Our large eddy simulations of (9)–(11) use a downgradient hypothesis to model the subfilter stress tensor \mathcal{T}_{ij} appearing in (9),

$$\mathcal{T}_{ij} = -2(\nu_e + \nu) \Sigma_{ij}^L + \nu \delta_{3j} \partial_z u^S, \quad (44)$$

where ν is the molecular viscosity, ν_e is an eddy viscosity that is a nonlinear function of resolved velocity field, δ_{ij} is the Kronecker delta, and

$$\Sigma_{ij}^L \stackrel{\text{def}}{=} \frac{1}{2} (\partial_i u_j^L + \partial_j u_i^L), \quad (45)$$

is the Lagrangian-mean rate of strain tensor. The divergence $\nabla \cdot \mathcal{T}$ in (9) is written $\partial_j \mathcal{T}_{ij}$ in indicial notation. The subfilter flux of buoyancy in (10) is, similarly,

$$\mathbf{q} = -\kappa_e \nabla b, \quad (46)$$

where κ_e is the eddy diffusivity of buoyancy.

A.1 The Lagrangian-mean strain tensor and kinetic energy dissipation

Typical models for large eddy simulations of (9)–(11) (Skyllingstad & Denbo, 1995; Polton & Belcher, 2007; Yang *et al.*, 2015; McWilliams *et al.*, 1997; Noh *et al.*, 2004; Harcourt & D’Asaro, 2008) use a subfilter stress tensor \mathcal{T}_{ij} proportional to the Eulerian-mean rate of strain. Our subfilter model, however, uses the Lagrangian-mean rate of strain Σ_{ij}^L in (44). We note in advance that the following discussion above may not have significant practical import: the agreement shown in figure 10 between our simulations and the simulations of McWilliams *et al.* (1997), which use Σ_{ij}^E , suggests that the differences between models using Σ_{ij}^L and Σ_{ij}^E are small when resolved turbulent fluxes dominate over turbulent fluxes modeled by \mathcal{T}_{ij} .

The first important assumption we make is that the terms proportional to molecular viscosity ν in (44) are negligible. This is justified after the fact by the results of our large eddy simulations, where ν_e is roughly 10^2 – 10^3 times larger than ν within the bulk of the boundary layer, and where $|\partial_z \mathbf{u}^S|$ is largely similar to $|\partial_z \mathbf{u}^L|$. Note that neglecting terms proportional to molecular viscosity also means we neglect the term $\nu \partial_z u^S$ in (44) associated with the molecular dissipation of the surface wave field. This term is crucial for describing streaming flows in viscous boundary layers and other viscous surface wave phenomena at low Reynolds number (see, for example, Longuet-Higgins (1953)). In this paper, we assume that molecular dissipation of the surface wave field has a negligible effect on boundary layer evolution. This assumption is almost always justified at the high Reynolds numbers of typical ocean surface boundary layers.

The second important assumption is that the Lagrangian-mean kinetic energy, $\frac{1}{2} |\mathbf{u}^L|^2$, undergoes a Kolmogorovian forward cascade through a spectral inertial range en route to the small, unresolved scales of molecular dissipation. The assumption of an inertial range for spectral fluxes of Lagrangian-mean kinetic energy follows from the conservation of Lagrangian-mean kinetic energy in (15) in the absence of stratification, dissipation, or forcing by nonzero $\partial_t \mathbf{u}^S$. We note that an alternative hypothesis that the Eulerian-mean kinetic energy, $\frac{1}{2} |\mathbf{u}^E|^2$, undergoes a forward cascade through an inertial range is more difficult to justify because $\frac{1}{2} |\mathbf{u}^E|^2$ is not conserved. (For example, equation (36) shows that $\frac{1}{2} |\mathbf{u}^E|^2$ oscillates between 0 and $2 |\mathbf{u}^L|^2$ in an adiabatic inertial oscillation depending on whether \mathbf{u}^L is aligned or anti-aligned with \mathbf{u}^S .) The ‘turbulence-induced anti-Stokes’ flow observed in Pearson (2018)’s large eddy simulations is evidence that turbulent momentum fluxes tend down Lagrangian-mean gradients, and thus tend to dissipate Lagrangian-mean kinetic energy.

Inserting $\mathcal{T}_{ij} = -2\nu_e \Sigma_{ij}^L$ into (9), and maintaining $b = \partial_t \mathbf{u}^S = 0$, we find that $\int \mathbf{u}^L \cdot (9) dV$ yields

$$\frac{d}{dt} \int \frac{1}{2} |\mathbf{u}^L|^2 dV = - \int 2\nu_e \Sigma_{ij}^L \Sigma_{ij}^L dV. \quad (47)$$

The modeled dissipation rate of Lagrangian-mean kinetic energy on the right side of (47) is positive-definite because $\nu_e > 0$ by its definition below in (48). As a result, \mathcal{T}_{ij} in (44) models a forward cascade of turbulent energy across the filter scale and toward molecular dissipation.

A.2 Eddy viscosity and eddy diffusivity

We use the Anisotropic Minimum Dissipation model (Rozema *et al.*, 2015; Abkar *et al.*, 2016; Verstappen, 2018) described by Vreugdenhil & Taylor (2018) for the eddy viscosity ν_e in (44) and eddy diffusivity κ_e in (46).³ ν_e and κ_e are defined to be strictly non-negative,

$$\nu_e \stackrel{\text{def}}{=} \max(\nu_e^\dagger, 0), \quad \text{and} \quad \kappa_e \stackrel{\text{def}}{=} \max(\kappa_e^\dagger, 0), \quad (48)$$

³See https://github.com/CliMA/Oceananigans.jl/blob/v0.27.2/src/TurbulenceClosures/turbulence_closure_implementations/verstappen_anisotropic_minimum_dissipation.jl for the implementation in software.

where the nonlinear functions ν_e^\dagger and κ_e^\dagger are

$$\nu_e^\dagger \stackrel{\text{def}}{=} -C\Delta^2 \frac{(\hat{\partial}_k \hat{u}_i)(\hat{\partial}_k \hat{u}_j)\hat{\Sigma}_{ij}^L}{(\hat{\partial}_\ell \hat{u}_m)^2}, \quad \text{and} \quad \kappa_e^\dagger \stackrel{\text{def}}{=} -C\Delta^2 \frac{(\hat{\partial}_k \hat{u}_i)(\hat{\partial}_k b)\hat{\partial}_i b}{(\hat{\partial}_\ell b)^2}. \quad (49)$$

In (49), Δ is the filter width defined via

$$\frac{1}{\Delta^2} \stackrel{\text{def}}{=} \frac{1}{3} \left(\frac{1}{\Delta_x^2} + \frac{1}{\Delta_y^2} + \frac{1}{\Delta_z^2} \right), \quad (50)$$

in terms of the potentially anisotropic x, y, z grid spacings Δ_i . The hats in (49) denote the scaled quantities

$$\hat{\partial}_i \stackrel{\text{def}}{=} \Delta_i \partial_i, \quad \hat{u}_i \stackrel{\text{def}}{=} \frac{u_i^L}{\Delta_i}, \quad \text{and} \quad \hat{\Sigma}^L \stackrel{\text{def}}{=} \frac{1}{2} (\hat{\partial}_j \hat{u}_i + \hat{\partial}_i \hat{u}_j). \quad (51)$$

C in (49) is a model constant. For simulations A–E in table 1, which are reported in section 4, we set $C = \frac{1}{12}$ following Verstappen (2018) and Vreugdenhil & Taylor (2018). For simulations 1–8 in table 1 and reported in 3, we implement⁴ a model that increases C from $\frac{1}{12}$ to $\frac{2}{3}$ at $z = 0$ over a scale of $4\Delta_z = 1$ m,

$$C(z) = C_I + e^{z/d}(C_0 - C_I), \quad (52)$$

where $C_I = \frac{1}{12}$, $C_0 = \frac{2}{3}$, and $d = 1$ meter. The model constant enhancement in (52) is necessary for obtaining smooth buoyancy profiles near the surface during the spin-up simulations 1 and 6 in table 1. We find that without an enhancement of the kind in (52), the eddy diffusivity κ_e is too small near $z = 0$ during free convection, which prevents a smooth transition between the boundary-adjacent cells where the unresolved diffusive flux \mathbf{q} dominates, and the turbulent interior where advective fluxes \mathbf{ub} in (10) control the evolution of the buoyancy distribution.

Our use of u_i^L in (51) is justified by deriving ν_e and κ_e from the Lagrangian-mean velocity gradient energy equation (Rozema *et al.*, 2015). In this derivation, the term $-(\nabla \times \mathbf{u}^S) \times \mathbf{u}^L$ in (9) arises as a ‘transport term’ in the velocity gradient energy equation, and thus, similar to Coriolis accelerations, does not affect the form of ν_e (Abkar *et al.*, 2016).

A.3 Model validation

Figure 10 plots the horizontally-averaged velocity and vertical variance from the $1 \times$ excited case, reproducing parts of figures 2 and 6 from McWilliams *et al.* (1997). Our results are similar despite the differences between our subfilter flux model and McWilliams *et al.* (1997)’s.

B Lagrangian-mean form of the Craik-Leibovich equations

The Eulerian-mean form of the Craik-Leibovich momentum equation (Leibovich, 1977; Suzuki & Fox-Kemper, 2016) is

$$\partial_t \mathbf{u}^E + (\mathbf{u}^E \cdot \nabla) \mathbf{u}^E - \mathbf{u}^S \times (\nabla \times \mathbf{u}^E) + f \hat{z} \times \mathbf{u}^L + \nabla \phi = b \hat{z}, \quad (53)$$

where

$$\phi \stackrel{\text{def}}{=} p + \frac{1}{2} |\mathbf{u}^S|^2 + \mathbf{u}^S \cdot \mathbf{u}^E, \quad (54)$$

p is the Eulerian-mean pressure, and $\mathbf{u}^S \times (\nabla \times \mathbf{u}^E)$ is the “vortex force”. Some algebraic gymnastics lead from (53) to (9). Starting with the vector identity

$$\mathbf{u}^L \times (\nabla \times \mathbf{u}^S) + \mathbf{u}^S \times (\nabla \times \mathbf{u}^L) = \nabla \cdot (\mathbf{u}^S \cdot \mathbf{u}^L) - (\mathbf{u}^S \cdot \nabla) \mathbf{u}^L - (\mathbf{u}^L \cdot \nabla) \mathbf{u}^S, \quad (55)$$

⁴See <https://github.com/glwagner/WaveTransmittedTurbulence/blob/master/src/les.jl>.

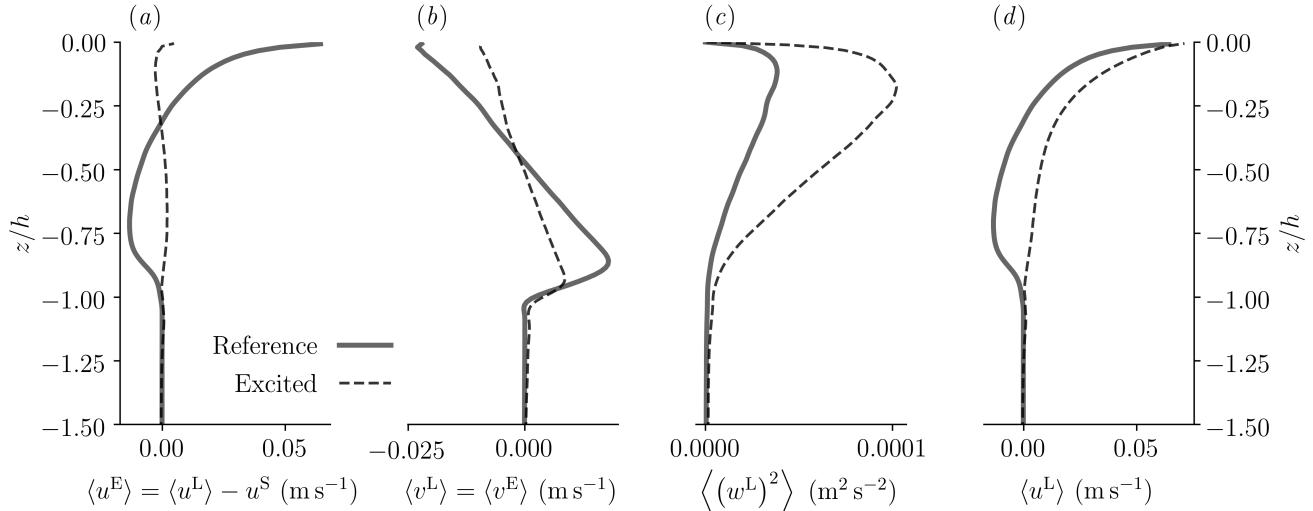


Figure 10: Horizontally-averaged Eulerian-mean x -velocity $\langle u^E \rangle$ (a), horizontally-averaged y -velocity $\langle v^E \rangle = \langle v^L \rangle$ (b), horizontally-averaged vertical variance $\langle (w^L)^2 \rangle$ (c), and horizontally-averaged Lagrangian-mean $\langle u^L \rangle$ (d). $\langle v^E \rangle = \langle v^L \rangle$ because the Stokes drift is in x , such that $\mathbf{u}^S = u^S \hat{x}$. The vertical coordinate z/h is height normalized by h , the depth of the maximum horizontally-averaged buoyancy gradient such that $\partial_z \langle b \rangle(h) = \max(\partial_z \langle b \rangle)$. Panels a and b reproduce figure 2 in McWilliams *et al.* (1997). Panel c reproduces figure 6 in McWilliams *et al.* (1997).

and subtracting both $\mathbf{u}^S \times (\nabla \times \mathbf{u}^L) = \mathbf{u}^S \times (\nabla \times \mathbf{u}^E) + \mathbf{u}^S \times (\nabla \times \mathbf{u}^S)$ and

$$\mathbf{u}^S \times (\nabla \times \mathbf{u}^S) = \nabla \left(\frac{1}{2} |\mathbf{u}^S|^2 \right) - (\mathbf{u}^S \cdot \nabla) \mathbf{u}^S, \quad (56)$$

yields the identity

$$\mathbf{u}^L \times (\nabla \times \mathbf{u}^S) + \mathbf{u}^S \times (\nabla \times \mathbf{u}^E) + (\mathbf{u}^S \cdot \nabla) \mathbf{u}^E + (\mathbf{u}^L \cdot \nabla) \mathbf{u}^S - \nabla \left(\frac{1}{2} |\mathbf{u}^S|^2 + \mathbf{u}^S \cdot \mathbf{u}^E \right) = 0. \quad (57)$$

Equation (57) then implies that

$$(\mathbf{u}^E \cdot \nabla) \mathbf{u}^E - \mathbf{u}^S \times (\nabla \times \mathbf{u}^E) + \nabla \phi = (\mathbf{u}^L \cdot \nabla) \mathbf{u}^L - (\nabla \times \mathbf{u}^S) \times \mathbf{u}^L + \nabla p \quad (58)$$

where we have used (54).

With the identity (58), $\partial_t \mathbf{u}^E = \partial_t \mathbf{u}^L - \partial_t \mathbf{u}^S$, and the assumption $\nabla \cdot \mathbf{u}^S = 0$ (valid for weakly modulated waves, as discussed in section 2), we can convert the Eulerian-mean form of the Craik-Leibovich equation (53) and the continuity equation $\nabla \cdot \mathbf{u}^E = 0$ into their Lagrangian-mean counterparts (9) and (11).

Finally, we note that (53) is derived by Leibovich (1980) from the Generalized Lagrangian-Mean momentum equation presented in Theorem I of Andrews & McIntyre (1978). Thus Leibovich (1980) provides a link between Andrews & McIntyre (1978) and our (9)–(11). As discussed by Leibovich (1980), in the scenarios we consider with relatively slow background rotation and slowly-modulated waves, we obtain a near-equivalence of the pseudomomentum that appears in Theorem I of Andrews & McIntyre (1978) and the surface wave Stokes drift \mathbf{u}^S .

C Table of large eddy simulations

Table 1 provides a list and details about the large eddy simulations reported in this paper. Additional information, including instructions for reproducing the simulations and figures in this paper, are hosted at <https://glwagner.github.io/WaveTransmittedTurbulence>.

ID	Simulation name	Stokes drift	Surface conditions	Domain size	Grid spacing
1	weak spin-up	$u^S = 0$	$q_z = 5 \times 10^{-10} \text{ m}^2 \text{s}^{-3}$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
2	weak growing waves	$u^S = u_{\text{eq}}^S(z) \left(1 - e^{-t^2/2T_w^2}\right)$	free slip, insulating	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
3	weak surface stress, no waves	$u^S = 0$	$\mathcal{T}_{xz} = -\tau_w e^{-t^2/2T_w^2} t/T_w$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
4	weak surface stress, steady waves	$u^S = u_{\text{eq}}^S(z)$	$\mathcal{T}_{xz} = -\tau_w e^{-t^2/2T_w^2} t/T_w$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
5	weak surface stress, growing waves	$u^S = u_{\text{eq}}^S(z) \left(1 - e^{-t^2/2T_w^2}\right)$	$\mathcal{T}_{xz} = -\tau_w e^{-t^2/2T_w^2} t/T_w$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
6	strong spin-up	$u^S = 0$	$q_z = 5 \times 10^{-10} \text{ m}^2 \text{s}^{-3}$	large	$0.75 \times 0.75 \times 0.25 \text{ m}$
7	strong growing waves	$u^S = 4u_{\text{eq}}^S \left(1 - e^{-t^2/2T_w^2}\right)$	free slip, insulating	large	$0.75 \times 0.75 \times 0.25 \text{ m}$
8	strong surface stress, no waves	$u^S = 0$	$\mathcal{T}_{xz} = -4\tau_w e^{-t^2/2T_w^2} t/T_w$	large	$0.75 \times 0.75 \times 0.25 \text{ m}$
9	strong surface stress, steady waves	$u^S = 4u_{\text{eq}}^S(z)$	$\mathcal{T}_{xz} = -4\tau_w e^{-t^2/2T_w^2} t/T_w$	large	$0.75 \times 0.75 \times 0.25 \text{ m}$
A	reference	$u^S = 0$	$\mathcal{T}_{xz} = \tau_{\text{MSM}}$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
B	1× excited	$u^S = e^{2kz} u_0^S$	$\mathcal{T}_{xz} = \tau_{\text{MSM}}$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
C	1× resting	$u^S = e^{2kz} u_0^S$	$\mathcal{T}_{xz} = \tau_{\text{MSM}}$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
D	4× excited	$u^S = e^{2kz} 4u_0^S$	$\mathcal{T}_{xz} = \tau_{\text{MSM}}$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$
E	4× resting	$u^S = e^{2kz} 4u_0^S$	$\mathcal{T}_{xz} = \tau_{\text{MSM}}$	regular	$0.5 \times 0.5 \times 0.25 \text{ m}$

Table 1: Simulations 1–8 are reported in section (3) and simulations A–E are reported in section (4). Simulations 1–5 and A–E use $128 \times 128 \times 64 \text{ m}$ domains with $256 \times 256 \times 256$ grid points and uniform $0.5 \times 0.5 \times 0.25 \text{ meter}$ grid spacing in x, y, z . Simulations 6–8 use $192 \times 192 \times 96 \text{ m}$ domains with $256 \times 256 \times 384$ grid points and uniform $0.75 \times 0.75 \times 0.25 \text{ meter}$ grid spacing in x, y, z . The effective wave-forced stress during the growth of 100 meter wave length, 1 meter amplitude deep water waves on a time-scale of $T_w = 4$ hours is $\tau_w = 2.72 \times 10^{-5} \text{ m}^2 \text{s}^{-2}$. The wind stress prescribed by McWilliams *et al.* (1997) is $\tau_{\text{MSM}} = -3.72 \times 10^{-5} \text{ m}^2 \text{s}^{-2}$. Simulation 1 is initialized with $\mathbf{u}^L = 10^{-6} \times e^{z/2} \xi_u(x, y, z) (\Delta_z q_z|_{t=0})^{1/3}$ and $b = N^2 z + 10^{-6} \times e^{z/2} \xi_b(x, y, z) N^2 \Delta_z$, where each ξ_ψ is a Gaussian-distributed, unit standard deviation random field for each $\psi \in (u^L, v^L, w^L, b)$, $\Delta_z = 0.25 \text{ m}$ is the vertical grid spacing, and $N^2 = 10^{-6} \text{ s}^{-2}$ is the initial buoyancy gradient. Simulations 2–7 are initialized from simulation 1 at $t = \pi/f$ with Coriolis parameter $f = 10^{-4} \text{ s}^{-1}$. Simulations A–E use $u_0^S = 0.068 \text{ m s}^{-1}$ and $k = 0.105 \text{ m}^{-1}$. The simulation data was generated with Oceananigans (<https://github.com/CliMA/Oceananigans.jl>). Instructions for reproducing the results are at <https://github.com/glwagner/WaveTransmittedTurbulence.jl>.

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