# Numerical solutions of the Swift-Hohenberg equation

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#### 1 Introduction

In this manuscript I will be solving the Swift-Hohenberg equation with the form:

$$\frac{\partial \Psi}{\partial t} = \epsilon - (1 + \Delta)^2 \Psi + g \Psi^2 - \Psi^3 \tag{1}$$

This equation was first considered for instabilities in the Raleigh convection giving rise to pattern formation in fluids, explored in [Swift and Hohenberg, 1977]. As discussed by [Sánchez Pérez-Moreno et al., 2013], this equation gives rise to complex behaviour that could be found in the more complex Navier-Stokes equation. The physical interpretation of the equation for  $\Psi$ , is the vertical fluid flow (out of the page) at the center of two heated plates. As described by [Sánchez Pérez-Moreno et al., 2013], the parameter  $\epsilon$ , represents the reduced Rayleigh number and g the strength of the symmetry break in the equation.

#### 2 Numerical Method

In this equation, the domain will be expanded for x and y, with periodic boundary conditions on x and no-slipping Dirichlet boundary condition on y. With:

$$0 < x < L \quad 0 < y < L \quad \Psi(x, y = 0, L) = 0 \tag{2}$$

with:

$$\begin{cases} \partial_y \Psi - \Psi_y + \tau_1 U_{M-1}(y) = 0 \\ \partial_y \Psi_y - \Psi_{yy} + \tau_2 U_{M-1}(y) = 0 \\ \partial_y \Psi_{yy} - \Psi_{yyy} + \tau_3 U_{M-1}(y) = 0 \\ \frac{\partial \Psi}{\partial t} + (1 - \epsilon) \Psi + 2 \Psi_{yy} + 2 \partial_x^2 \Psi + 2 \partial_x^2 \Psi_{yy} + \partial_x^4 \Psi + \partial_y \Psi_{yyy} + \tau_4 U_{M-1}(y) = g \Psi^2 - \Psi^3 \end{cases}$$

Which corresponds to generalized  $\tau$  method of the first order formulation to satisfy the boundary conditions in y, as described by [Burns et al., 2020]. This

system will be solved using a pseudo-spectral method using Fourier series and Chebyshev polynomials. There is no need to split the partial derivatives on x as this will be expanded using a Fourier series, therefore derivatives are simply calculated.

Then the Galerkin method is used to discretize each one of the variables, in this case each one of the  $\Psi$  and thee three first partials of y and the four  $\tau$ 's.

This have the form of:

$$\Psi = \sum_{n,m}^{N} a_{n,m}(t) cos(\frac{2\pi}{L}nx) T_m(my) - \sum_{n,m}^{N} b_{n,m}(t) sin(\frac{2\pi}{L}x) T_m(my)$$
 (3)

$$\tau_1 = \sum_{n=1}^{N} \tau_{a_1,n}(t) T_m(my) - \sum_{n=1}^{N} \tau_{b_1,n}(t) (T_m(my))$$
(4)

Therefore we will need to keep track of the a's and b's for the  $\Psi$  vectors and its derivatives. Also one  $\tau_a$  and  $\tau_b$  for each of the 4  $\tau$ 's.

Using this projections, it gives rise to the system:

$$W = -2Cn^2 + Cn^4 + (1 - \epsilon)C$$

Where  $T_m$  are the translated Chebyshev polynomials of the first kind with  $\frac{L}{2}(T_m(my)) + \frac{L}{2}$ . Here D is a derivative matrix for Chebyshev polynomials of the second kind, C is a conversion matrix from Chebyshev polynomials of the first kind to second kind an n is a vector with the Fourier wave numbers  $\frac{2\pi n_j}{L}$ . Z is a matrix of zeros. Each one of the matrices has size  $N \times N$  and the vector  $N \times 1$ . In addition, at the right of the matrix there are 8 columns of zeros, with a 1 in the  $(i^{th}(N) - 1, j^{th},)$  entry for each one of the  $\tau$ 's coefficients. On the bottom, to force Dirichlet (No slipping) boundary conditions, there is going to be a row of ones, in the form  $(i^{th}, 0: N) = (-1)^{j^{th}}$  and  $(i^{th} + 1, 0: N) = 1$  for each one of the  $4\tau$ 's coefficients corresponding columns.

Then we have:

$$M = \begin{pmatrix} \mathbf{0} & | \mathbf{0} \\ C & Z & | \mathbf{0} \\ Z & C & | \mathbf{0} \end{pmatrix}$$

(6)

$$X = (a, b, a', b', a'', b'', a''', b''', \tau_{(a_1, n)}, \tau_{(b_1, n)}, ..., \tau_{(a_4, n)}, \tau_{(b_4, n)})$$
(7)

With size  $(N+8) \times (N+8)$  for  $M, N \times 1$  for the a's and b's and one single entry for the  $\tau$ 's

Then:  $M\partial_t X + LX = F$ . With F, vector of size  $(8N + 8) \times 1$ , being all zeros except for the the rows corresponding to a''' and b'''. Which will be equal to the projection of the right hand side of the  $4^{th}$  equation of 2. This will be done in grid space to ensure a faster convergence of the solution. A scale of order 2 will be used to prevent aliasing on the non-linearity.

## 3 Resolution study

In this section, there will be a resolution study to ensure the exponential convergence of the algorithm. Instead of using an analytic solution, I will use a high resolution solution (N=512) as the true solution and compare it to lower order values of N ranging from (16,256) in multiples of 2. For this case, I will be using uniform initial conditions as in later simulations to show the behaviour of equation I will be uniformly randomly distributed initial conditions. In this case L=50 with a time step of  $10^{-3}$ . In Fig. 1, we can see the resolution study and how the system converges exponentially until N=128 where the error stagnates. In addition for the  $\Delta t$  value,  $\Delta X=L/N$  with minimum value of 0.1 when N=512. Then  $\Delta t \leq \frac{\Delta x}{32} \approx 0.003$  as seen in [wen Xi et al., 1991] for stability.

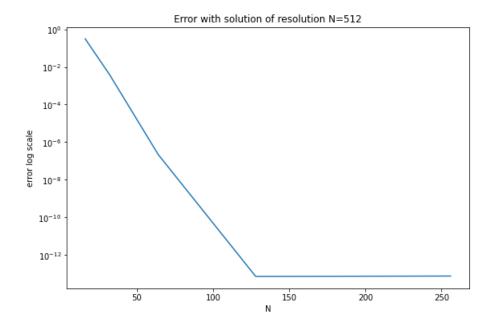


Figure 1: This figure represents the resolution study for the system with uniform initial conditions,  $\Delta t = 10^{-3}$ , L = 5. The tested N's ranged from (16,256) in multiples of 2, The y-axis is in log scale, we can see how is linear from (16,128) showing that the error is exponentially decaying. There is no much benefit in increasing N to reduce the error after 128. Error corresponds for 1000 iterations.

### 4 Simulations

In this section, I will be showing example simulations for the equation with L=50, N=256 and  $\Delta t=5\times 10^{-3}$ . I will be varying the values of  $\epsilon$  and g. For this simulation the code in **SwiftPDE.ipynb** is used. This code saves the time-step graphs in the folder where the code is run as images. The code also allows to save the data as a GIF video.

#### 4.1 Case with constant $\epsilon$

The initial conditions will be uniform randomly distributed with a constant seed in the range (0,0.3) for this subsection. This would mean an initial general random upward flow.

In Fig. 2, we can see the case when  $\epsilon=1$  and g=0.1. In this case linear formations take longer to progress and there is a larger ensemble of spots in intermediate steps.

In contrast, in Fig. 3, we can see the case when  $\epsilon = 1$  and g = 1. Here linear formations in the grid are favoured after 10 seconds of simulation. This

exemplifies how the parameter g determines the braking of symmetry.

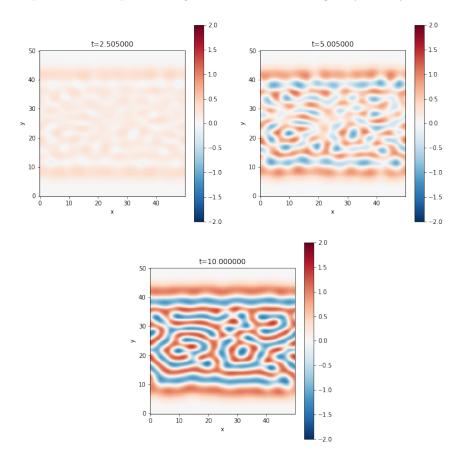


Figure 2: In this test, g=0.1 and  $\epsilon=1$  with random initial conditions. It is noticeable that with g=0.1, the formation of linear patterns and spots are mixed. With a slower progression of the line formation. The simulation is run for 2000 iterations.

#### 4.2 Increasing the value of $\epsilon$

The initial conditions will be uniform randomly distributed with a constant seed in the range (-2.5, 2.5) for this subsection. This would mean an initial general non-dominant flow in any direction.

In Fig. 4 we can see the plots of this scenario with g=2 and  $\epsilon=2$ . It can be seen how the upward convection starts to dominate until eventually downward flow will disappear.

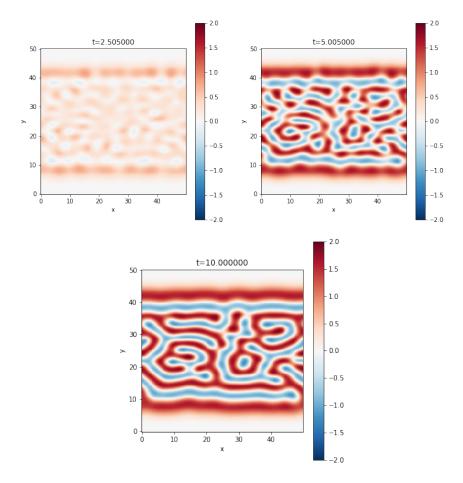


Figure 3: In this test, g=1 and  $\epsilon=1$  with random initial conditions. It is noticeable that with g=1, the formation of linear patterns is reinforced. The simulation is run for 2000 iterations.

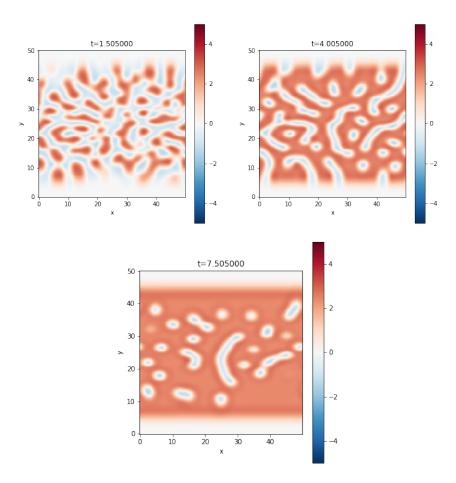


Figure 4: In this test, g=2 and  $\epsilon=2$  with random initial conditions. It is noticeable that by increasing  $\epsilon$ , the convection parameter starts to dominate over the pattern formation and stripes start to disappear. The simulation is run for 1500 iterations.

# References

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