```
clear
close all
clc
format long g
```

## Problem 1- i) - Acceleration of asteroid relative to Sun

```
% Semi-major axis of Jupiters orbit
                 = 778279959;
                                              % [km]
a
% Gravitational Parameters: mu = G*m [km^3/s^2]
                = 132712440017.99;
mu sun
mu_asteroid
                = 10;
mu_jupiter
                = 126712767.8578;
% Define 2-D position vectors [x-y] plane
% Assume equilateral triangle of 60 deg
r sun jupiter = [a;0];
                                              % [km]
r_{sun_asteroid} = [a*cosd(60); a*sind(60)]; % [km]
% Dominant acceleration of asteroid wrt sun [km/s^2]
a_sa_dominant
                    = -(mu_sun + mu_asteroid)*r_sun_asteroid...
                   /(norm(r sun asteroid)^3)
a sa dominant = 2 \times 1
     -1.0954938498391e-07
    -1.89745100730055e-07
% Magnitude of dominant acceleration of asteroid wrt sun
a_sa_dominant_mag = norm(a_sa_dominant)
a sa dominant mag =
      2.1909876996782e-07
% Position vector from asteroid to Jupiter [km]
r asteroid_jupiter = r_sun_jupiter - r_sun_asteroid;
% Direct perturbing acceleration due to Jupiter on asteroid wrt sun [km/s^2]
a_sa_direct
                       = mu_jupiter*r_asteroid_jupiter/(norm(r_asteroid_jupiter)^3)
a_sa_direct = 2×1
       1.045968696341e-10
    -1.81167092518919e-10
% Magnitude of direct perturbing acceleration on asteroid wrt sun
                        = norm(a_sa_direct)
a_sa_direct_mag
a sa direct mag =
       2.091937392682e-10
% Indirect perturbing acceleration due to Jupiter on asteroid wrt sun [km/s^2]
                         = mu jupiter*r sun jupiter/(norm(r sun jupiter)^3)
a_sa_indirect
a sa indirect = 2 \times 1
       2.091937392682e-10
```

```
% Net perturbing acceleration on asteroid wrt sun
                            = a sa direct - a sa indirect
 a sa pertubing
 a sa pertubing = 2 \times 1
       -1.045968696341e-10
      -1.81167092518919e-10
 % Magnitude of perturbing acceleration on asteroid wrt sun
 a sa pertubing mag
                             = norm(a sa pertubing)
 a_sa_pertubing_mag =
        2.091937392682e-10
 % Net acceleration of asteroid wrt sun
                             = a_sa_dominant + a_sa_pertubing
 a_sa_net
 a sa net = 2 \times 1
      -1.09653981853544e-07
      -1.89926267822574e-07
 % Magnitude of net acceleration of asteroid wrt sun
                              = norm(a_sa_net)
 a_sa_net_mag
 a_sa_net_mag =
      2.19307963707088e-07
Problem 1- ii) - Acceleration of asteroid relative to Jupiter
 % Relevant position vectors
 r asteroid sun
                           = -r_sun_asteroid;
 r jupiter sun
                           = -r sun jupiter;
 r_jupiter_asteroid
                           = -r_asteroid_jupiter;
 % Dominant acceleration of asteroid wrt Jupiter [km/s^2]
                           = -(mu_jupiter + mu_asteroid)*r_jupiter_asteroid...
 a_ja_dominant
                              /(norm(r jupiter asteroid)^3)
 a ja dominant = 2 \times 1
      1.04596877888743e-10
      -1.81167106816381e-10
 % Magnitude of dominant acceleration of asteroid wrt Jupiter
 a_ja_dominant_mag
                           = norm(a ja dominant)
 a ja dominant mag =
       2.09193755777486e-10
 % Direct perturbing acceleration due to sun on asteroid wrt Jupiter [km/s^2]
                           = mu_sun*r_asteroid_sun/(norm(r_asteroid_sun)^3)
 a_ja_direct
 a_ja_direct = 2×1
      -1.09549384975655e-07
      -1.89745100715758e-07
 % Magnitude of direct perturbing acceleration on asteroid wrt Jupiter
                          = norm(a ja direct)
 a ja direct mag
```

```
% Indirect perturbing acceleration due to sun on asteroid wrt Jupiter [km/s^2]
a_ja_indirect
                          = mu sun*r jupiter sun/(norm(r jupiter sun)^3)
a ja indirect = 2 \times 1
    -2.19098769951311e-07
% Net perturbing acceleration on asteroid wrt Jupiter
                           = a_ja_direct - a_ja_indirect
a ja pertubing
a_{ja_pertubing} = 2 \times 1
     1.09549384975655e-07
    -1.89745100715758e-07
% Magnitude of perturbing acceleration on asteroid wrt Jupiter
a ja pertubing mag
                            = norm(a_ja_pertubing)
a ja_pertubing_mag =
     2.19098769951311e-07
% Net acceleration of asteroid wrt Jupiter
                             = a_ja_dominant + a_ja_pertubing
a_ja_net
a ja net = 2 \times 1
     1.09653981853544e-07
    -1.89926267822574e-07
% Magnitude of net acceleration of asteroid wrt Jupiter
a_ja_net_mag
                             = norm(a_ja_net)
a_ja_net_mag =
     2.19307963707088e-07
```

## Problem 3b)

```
% Radius of Earth [km]
r earth
            = 6378.1363;
altitude
            = 8560;
                                 % [km]
% Position vector from Earth CG to spacecraft (polar coordinates)
r_earth_sc = [r_earth + altitude;0;0];
% Inertial velocity of spacecraft relative to earth (polar coordinates)
v_earth_sc = [-2.11;4.89;0];
                                 % [km/s]
                                 % Spacecraft mass [kg]
m_sc
            = 300;
% Gravitational constant km<sup>3</sup>/kg-s<sup>2</sup>
            = 6.6743e-11*(1/1000)^3;
            = 398600.4415/G;
                                 % Earth mass
m earth
% Total system angular momentum
C3
            = (m_earth*m_sc)*cross(r_earth_sc,v_earth_sc)...
              /(m_earth + m_sc) % [kg-km^2/s]
```

```
0
             21914245.9521
 % Specific angular momentum
 h_vec
               = cross(r_earth_sc,v_earth_sc) % [km^2/s]
 h_{vec} = 3 \times 1
                        0
                        a
              73047.486507
 % Total system kinetic energy
                = 1/2*(m_earth*m_sc)*dot(v_earth_sc,v_earth_sc)...
                 /(m_earth + m_sc) \% [kg-km^2/s^2]
 T =
                   4254.63
 % Gravitional potential
 U
                 = G*m_earth*m_sc/norm(r_earth_sc);
 % Total energy
 C4
                 = T - U
                                     % [kg-km<sup>2</sup>/s<sup>2</sup>]
 C4 =
          -3750.39352157545
 % Magnitude of angular momentum
                 = norm(h_vec);
 h
 % Gravitional parmeter
                 = G*(m_earth + m_sc);
 % Specific energy
 Е
                 = norm(v_earth_sc)^2/2 - mu/norm(r_earth_sc)
 E =
          -12.5013117385848
 % Areal velocity
 Adot
                   = h/2 \% [km^2/s]
 Adot =
             36523.7432535
Problem 3d)
 % Semi-latus rectum
                   = h^2/mu % [km]
 р
 p =
          13386.6768057515
 % Eccentricity vector
```

 $C3 = 3 \times 1$ 

e\_vec

= cross(v\_earth\_sc,h\_vec)/mu -...

```
r earth sc/norm(r earth sc);
% Eccentricity
                  = norm(e vec)
e
e =
        0.400383443095008
% Semi-major axis
                   = p/(1 - e^2) \% [km]
sma
sma =
        15942.3446849075
% Orbital Period
period
                    = 2*pi*sqrt(sma^3/mu)/3600 % [hours]
period =
        5.56463613853122
% Flight path angle
                    = acosd(v_earth_sc(2)/norm(v_earth_sc))
gamma
gamma =
        23.3398520112336
% True anomaly
                    = acosd((p/norm(r_earth_sc) - 1)/e) % [deg]
theta_star
theta star =
         105.03439032896
% Rotation matrix from Inertial p & e frame to polar frame
                    = [cosd(theta_star), sind(theta_star), 0;...
DCM_C_I
                        -sind(theta_star), cosd(theta_star), 0;...
                        0, 0, 1];
% Position vector in semi-latus rectum and eccentricity unit vectors
r_earth_sc_ep
                     = DCM_C_I'*r_earth_sc
r_earth_sc_ep = 3 \times 1
        -3874.93419372084
        14426.8084173774
% Velocity vector in semi-latus rectum and eccentricity unit vectors
                      = DCM_C_I'*v_earth_sc
v_earth_sc_ep
v_earth_sc_ep = 3 \times 1
        -4.17528537426884
        -3.30623532789436
```

## Problem 3e)

```
% Circular velocity
Vc = sqrt(mu/norm(r_earth_sc_ep))
```

Vc =

5.16559887511456

```
% Magnitude of relative velocity
v_mag = norm(v_earth_sc_ep)
```

v\_mag =

5.32580510345619