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```
clear
close all
clc
```

Problem 1a)

```
% Periapsis and Apoapsis
rp = 1.8; % Mars Radii
ra = 8; % Mars Radii

% Initial Mean Anomaly
M0 = -120; % [deg]
```

<u>Find</u>: $a, e, p, period, \epsilon, r_o, v_0, \theta_0^*, E_0, \gamma_0, t_0 - t_p$

```
% Calculate Semi major axis
a = (ra + rp)/2;
fprintf('The semi major axis %.1f Mars Radii', a)
```

The semi major axis 4.9 Mars Radii

$$a = \frac{r_a + r_p}{2} = 4.9 \ [R_{Mars}]$$

```
% Calculate Eccentricity
e = 1 - (rp/a);
fprintf('Eccentricity is %.3f', e)
```

Eccentricity is 0.633

$$r_p = a(1 - e)$$

$$e = 1 - \frac{r_p}{a} = 0.63$$

```
% Calculate semi latus rectum
p = rp*(1+e);
fprintf('Semi latus rectum is %.3f Mars Radii', p)
```

Semi latus rectum is 2.939 Mars Radii

$$r_p = \frac{p}{1+e}$$

 $p = r_p(1+e) = 2.94 \ [R_{Mars}]$

```
% Mars gravitional constant
mu_mars = 42828.314258067; % [km^3/s^2]

% Mars Radius at equator
R_mars = 3397; % [km]
```

```
% Calculate period [s]
tau = 2*pi*sqrt((a*R_mars)^3/mu_mars);
fprintf('The orbital period around mars is %.2f hours',tau/3600)
```

The orbital period around mars is 18.11 hours

$$period = 2\pi \sqrt{\frac{a^3}{\mu}} = 18.11 \ [hrs]$$

```
% Calculate specific energy
energy = -mu_mars/(2*a*R_mars);
fprintf('The specific energy is %.3f km^2/s^2', energy)
```

The specific energy is -1.286 km²/s²

$$\epsilon = \frac{-\mu}{2a} = -1.29 \ [km^2/s^2]$$

```
% Calculate eccentric anomaly via newton raphson
E0 = CalcEccentricAnomaly(e,M0*pi/180)*180/pi;
fprintf('The initial eccentric anomaly is %.2f deg',E0)
```

The initial eccentric anomaly is -142.21 deg

$$M = E - e \sin E$$
$$E_0 = -142.2 \left[deg \right]$$

```
% Calclate true anomaly from eccentric anomaly
ta0 = 2*atand(sqrt((1+e)/(1-e))*tand(E0/2));
fprintf('The initial true anomaly is %.2f deg',ta0)
```

The initial true anomaly is -161.56 deg

$$\tan \frac{\theta^*}{2} = \sqrt{\frac{1+e}{1-e}} \tan \frac{E}{2}$$

$$\theta_0^* = 2 \tan^{-1} \left(\sqrt{\frac{1+e}{1-e}} \tan \frac{E_0}{2} \right) = -161.6 \text{ [deg]}$$

```
% Calculate Initial radius
r0 = a*(1 - e*cosd(E0));
fprintf('The initial radius is %.4f Mars radii',r0)
```

The initial radius is 7.3499 Mars radii

$$r_0 = a(1 - e\cos E_0) = 7.35 [R_{Mars}]$$

```
% Calculate initial velocity
v0 = sqrt(2*(energy + mu_mars/(r0*R_mars)));
fprintf('The initial velocity is %.3f km/s', v0)
```

The initial velocity is 0.926 km/s

$$\epsilon = \frac{v^2}{2} - \frac{\mu}{r}$$

$$v_0 = \sqrt{2(\epsilon + \frac{\mu}{r_0})} = 0.93 \text{ [km/s]}$$

```
% Calculate angular momentum magnitude
h = sqrt(mu_mars*p*R_mars);

% Calculate angular velocity from angular momentum magntiude
ta0dot = h/((r0*R_mars)^2);

% Calculate radial component of velocity, use negative root due to sign of anomalys
r0dot = -sqrt(v0^2 - (r0*R_mars)^2*ta0dot^2);

% Calculate flight path angle
gamma0 = atan2d(r0dot,(r0*R_mars)*ta0dot);
fprintf('The initial flight path angle is %.2f deg', gamma0)
```

The initial flight path angle is -26.59 deg

$$h = \sqrt{\mu p} = r^2 \dot{\theta}$$

$$v^2 = \dot{r}^2 + r^2 \dot{\theta}^2$$

$$\tan \gamma = \frac{\dot{r}}{r \dot{\theta}} = \frac{v_r}{v_{\theta}}$$

$$\gamma_0 = \tan^{-1} \frac{\dot{r}}{r \dot{\theta}} = -26.6 \text{ [deg]}$$

```
% Calculate time since periapsis - convert M to degrees and convert M = -120 to 240
dtp = (M0*pi/180 + 2*pi)/(sqrt(mu_mars/(a*R_mars)^3));
fprintf('The time since the last periapsis is %.1f hours',dtp/3600)
```

The time since the last periapsis is 12.1 hours

$$M = \sqrt{\frac{\mu}{a^3}} (t - t_p)$$

$$(t_0 - t_p) = \frac{M_0}{\sqrt{\frac{\mu}{a^3}}} = 12.1 \text{ [hrs]}$$

Problem 1b)

<u>Find</u>: \overline{r}_0 , \overline{v}_0 , in \hat{r} , $\hat{\theta}$, and \hat{e} , \hat{p}

```
% Define r0 and v0 in orbit frame
v0_0 = [r0dot; (r0*R_mars)*ta0dot;0];
r0_0 = [r0;0;0];
```

```
\overline{v}_0 = r\hat{r} + r\dot{\theta}\hat{\theta} = -0.415 \ \hat{r} + 0.828 \ \hat{\theta} \ [km/s]
\overline{r}_0 = 7.35 \ \hat{r} \ [R_{mars}]
```

```
[r_0]^P = [PO][r_0]^O
[v_0]^P = [PO][v_0]^O
\overline{r}_0 = -6.97 \ \hat{e} - 2.32 \ \hat{p} \ [R_{mars}]
\overline{v}_0 = 0.655 \ \hat{e} - 0.654 \ \hat{p} \ [km/s]
```

Problem 1c)

<u>Find</u>: θ_1^* , r_1 , v_1 , E_1 , γ_1 , $(t_1 - t_p)$

```
% Mean motion
n = sqrt(mu_mars/(a*R_mars)^3);

% Change in time
dt = 0.75*tau;

% Calculate mean anomaly for time equal to 75% of period from t0
M1 = M0*(pi/180) + n*dt;

% Calculate eccentric anomaly from mean anomaly
E1 = CalcEccentricAnomaly(e,M1)*180/pi;
fprintf('The eccentric anomaly at new time is %.2f deg',E1)
```

The eccentric anomaly at new time is 161.50 deg

```
n = \sqrt{\frac{\mu}{a^3}}
M_1 = M_0 + n(t_1 - t_0)
M = E - e \sin E
E_1 = 161.5 \ [deg]
```

```
% Calclate true anomaly from eccentric anomaly
ta1 = 2*atand(sqrt((1+e)/(1-e))*tand(E1/2));
fprintf('The true anomaly at new time is %.2f deg',ta1)
```

The true anomaly at new time is 171.17 deg

$$\tan \frac{\theta^*}{2} = \sqrt{\frac{1+e}{1-e}} \tan \frac{E}{2}$$

$$\theta_1^* = 2 \tan^{-1} \left(\sqrt{\frac{1+e}{1-e}} \tan \frac{E_1}{2} \right) = 171.2 \ [deg]$$

```
% Calculate Initial radius
r1 = a*(1 - e*cosd(E1));
fprintf('The radius at new time is %.4f Mars radii',r1)
```

The radius at new time is 7.8398 Mars radii

$$r_1 = a(1 - e\cos E_1) = 7.84 [R_{Mars}]$$

```
% Calculate initial velocity
v1 = sqrt(2*(energy + mu_mars/(r1*R_mars)));
fprintf('The velocity at new time is %.3f km/s', v1)
```

The velocity at new time is 0.802 km/s

$$\epsilon = \frac{v^2}{2} - \frac{\mu}{r}$$

$$v_1 = \sqrt{2(\epsilon + \frac{\mu}{r_1})} = 0.80 \ [km/s]$$

```
% Calculate angular velocity from angular momentum magntiude
ta1dot = h/((r1*R_mars)^2);

% Calculate radial component of velocity, use positive root due to sign of anomalys
r1dot = sqrt(v1^2 - (r1*R_mars)^2*ta1dot^2);

% Calculate flight path angle
gamma1 = atan2d(r1dot,(r1*R_mars)*ta1dot);
fprintf('The flight path angle at the new time is %.2f deg', gamma1)
```

The flight path angle at the new time is 14.53 deg

$$h = r^{2}\dot{\theta}$$

$$v^{2} = \dot{r}^{2} + r^{2}\dot{\theta}^{2}$$

$$\tan \gamma = \frac{\dot{r}}{r\dot{\theta}} = \frac{v_{r}}{v_{\theta}}$$

$$\gamma_{1} = tan^{-1}\frac{\dot{r_{1}}}{r_{1}\dot{\theta_{1}}} = 14.5 \text{ [deg]}$$

```
% Calculate time since periapsis - convert M to degrees
dtp1 = M1/(sqrt(mu_mars/(a*R_mars)^3));
fprintf('The time since the last periapsis is %.1f hours',dtp1/3600)
```

The time since the last periapsis is 7.5 hours

$$M = \sqrt{\frac{\mu}{a^{3}}} (t - t_{p})$$

$$(t_{1} - t_{p}) = \frac{M_{1}}{\sqrt{\frac{\mu}{a^{3}}}} = 7.5 \text{ [hrs]}$$

```
% Calculate change in true anomaly and eccentric anomaly
dta = ta1 - ta0;
fprintf('The change in true anomaly is %.1f degrees', dta)
```

The change in true anomaly is 332.7 degrees

$$\Delta \theta^* = \theta_1^* - \theta_0^* = 332.7 \ [deg]$$

```
dE = E1 - E0;
fprintf('The change in eccentric anomaly is %.1f degrees', dE)
```

The change in eccentric anomaly is 303.7 degrees

$$\Delta E = E_1 - E_0 = 303.7$$
 [deg]

Problem 1d)

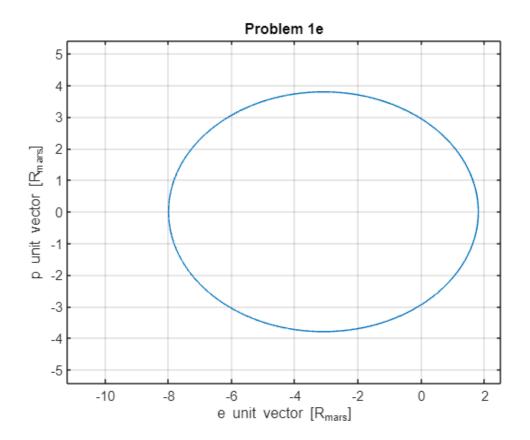
Find: \overline{r}_1 , \overline{v}_1

Note: f and \dot{g} are unitless, while g has units of [sec] and \dot{f} has units of [sec]⁻¹

```
\begin{split} f &= 1 - \frac{a}{r_0} \big[ 1 - \cos{(E_1 - E_0)} \big] \\ g &= (t_1 - t_0) - \sqrt{\frac{a^3}{\mu}} \big[ E_1 - E_0 - \sin{(E_1 - E_0)} \big] \\ \dot{f} &= -\frac{\sqrt{\mu a}}{r_1 r_0} \sin{(E_1 - E_0)} \\ \dot{g} &= 1 - \frac{a}{r_1} \big[ 1 - \cos{(E_1 - E_0)} \big] \\ \overline{r}_1 &= f \quad \overline{r}_0 + g \quad \overline{v}_0 = 0.703 \quad \overline{r}_0 \quad [R_{mars}] - 1.474 \times 10^4 \quad \overline{v}_0 \quad [km/s] = -7.75 \quad \hat{e} + 1.20 \quad \hat{p} \quad [R_{mars}] \\ \overline{v}_1 &= \dot{f} \quad \overline{r}_0 + \dot{g} \quad \overline{v}_0 = 3.34 \times 10^{-5} \quad \overline{r}_0 \quad [R_{mars}] + 0.722 \quad \overline{v}_0 \quad [km/s] = -0.32 \quad \hat{e} - 0.74 \quad \hat{p} \quad [km/s] \end{split}
```

Problem 1e) Plot Orbit

```
% Define time - increment by 10 second
time
        = 0:10:tau;
% Initialize Perifocal position and velocity vector
        = zeros(3,length(time));
% Calculate vectors in perifocal coordinates for each time step using f & g
% relations
for i = 1:length(time)
    % Calculate Mean Anonamly
                = (M0*pi/180) + n*(time(i) - time(1));
    % Calculate Eccentric anomaly [rad]
                = CalcEccentricAnomaly(e, M);
    % Calculate f & g relations
            = (1 - (a/r0)*(1 - cosd(E*(180/pi) - E0)));
            = (time(i) - time(1)) - sqrt((a*R_mars)^3/mu_mars)*(E - E0*(pi/180) - sind(E*(180/pi/180)))
    % Calculate perifocal position vector, convert r0 to km to be consistent with v0 in km/s
    r_P(:,i) = f*(r_0_P*R_mars) + g*v_0_P;
end
figure
plot(r_P(1,:)/R_mars,r_P(2,:)/R_mars)
xlim([-ra*1.4 rp*1.4])
axis equal
title('Problem 1e')
grid on
xlabel('e unit vector [R_{mars}]')
ylabel('p unit vector [R {mars}]')
```



```
function E = CalcEccentricAnomaly(e,M)
   % Input M in rad and output E in rad
   % Define error tolerance
      etol = 1e-8;
   % Initialize change in Eccentric Anomaly
          = 1;
   % Initialize Eccentric Anomaly and counter
      count= 0;
     Ε
          = M;
     while dE > etol
         % Increase counter
          count = count + 1;
         % Newton Raphson
         Enp1 = E - (E - e*sin(E) - M)/(1 - e*cos(E));
         % Assign new values and calculate delta
          dE
              = abs(Enp1 - E);
          Ε
              = Enp1;
          if count > 1000
              disp('Max iterations reached');
              break
```

end end

