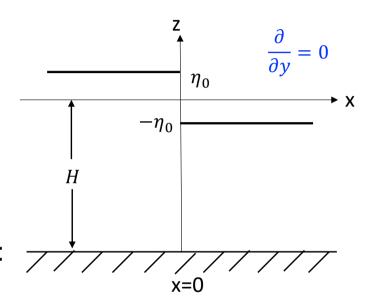
Geostrophic (Rossby) adjustment - barotropic

Initial state (unbalanced):

$$\eta = \begin{cases} \eta_0, & x < 0 \\ -\eta_0, & x > 0 \end{cases} \qquad \eta_0 << H$$

$$u = v = 0$$



For a steady final state (based on shallow water model):

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} - fv = -g \frac{\partial \eta}{\partial x}$$

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + fu = -g \frac{\partial \eta}{\partial y}$$

geostrophic flow

$$\frac{\partial \eta}{\partial t} + \frac{\partial hu}{\partial x} + \frac{\partial hv}{\partial y} = 0$$
 sufficient condition: $u = 0$ everywhere

For potential vorticity conservation:

x < 0:

$$\frac{f}{H + \eta_0} = \frac{f + \frac{\partial v}{\partial x}}{H + \eta} = \frac{f + \frac{g}{f} \frac{\partial^2 \eta}{\partial x^2}}{H + \eta}$$

$$\eta_0 << H$$

$$fH + f\eta = fH + f\eta_0 + \frac{gH}{f} \frac{\partial^2 \eta}{\partial x^2} + \frac{g\eta_0}{f} \frac{\partial^2 \eta}{\partial x^2}$$

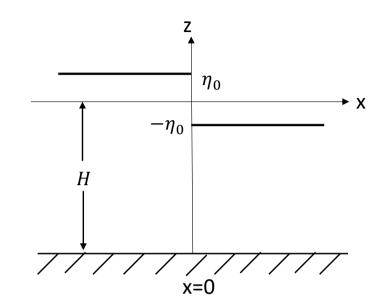
$$R^2 \frac{\partial^2 \eta}{\partial x^2} - \eta = -\eta_0 \qquad \qquad R = \frac{\sqrt{gH}}{f}$$

$$\eta = f(x) + \eta_0$$

$$f(x) = Ce^{\lambda x}$$
 \longrightarrow $R^2 \lambda^2 - 1 = 0$ \longrightarrow $\lambda = \pm \frac{1}{R}$ \longrightarrow $\lambda = \frac{1}{R}$

$$\eta = Ce^{\frac{x}{R}} + \eta_0 \qquad x = 0, \eta = 0, C = -\eta_0$$

$$\eta = -\eta_0 e^{\frac{x}{R}} + \eta_0$$



For PV conservation:

x > 0:

$$\frac{f}{H - \eta_0} = \frac{f + \frac{\partial v}{\partial x}}{H + \eta} = \frac{f + \frac{g}{f} \frac{\partial^2 \eta}{\partial x^2}}{H + \eta}$$

$$fH + f\eta = fH - f\eta_0 + \frac{gH}{f} \frac{\partial^2 \eta}{\partial x^2} - \frac{g\eta_0}{f} \frac{\partial^2 \eta}{\partial x^2}$$

$$R^2 \frac{\partial^2 \eta}{\partial x^2} - \eta = \eta_0$$

$$\eta = f(x) - \eta_0$$

$$f(x) = Ce^{\lambda x}$$

$$R^2\lambda^2 - 1 = 0 \longrightarrow$$

$$1 = 0 \longrightarrow \lambda = \pm$$

$$-\eta_0$$

$$e^{\frac{1}{R}x} \to \infty \ as \ x \to \infty$$

$$f(x) = Ce^{\lambda x}$$
 \longrightarrow $R^2\lambda^2 - 1 = 0$ \longrightarrow $\lambda = \pm \frac{1}{R}$ \longrightarrow $\lambda = -\frac{1}{R}$

$$\eta = Ce^{-\frac{x}{R}} - \eta_0$$
 $x = 0, \eta = 0, C = \eta_0$
 $\eta = \eta_0 e^{-\frac{x}{R}} - \eta_0$

$$\eta = \begin{cases} x \to -R, & \eta \to \eta_0 \\ -\eta_0 e^{\frac{x}{R}} + \eta_0, & x < 0 \end{cases}$$

$$\eta = \begin{cases} -\eta_0 e^{\frac{x}{R}} + \eta_0, & x < 0 \\ \eta_0 e^{-\frac{x}{R}} - \eta_0, & x > 0 \end{cases}$$

$$x \to R, & \eta \to -\eta_0 \end{cases}$$

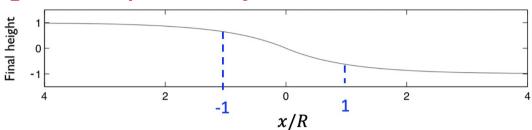
The adjustment spatial scale is the Rossby deformation radius $R = \frac{\sqrt{gH}}{f}$

$$v = \frac{g}{f} \frac{\partial \eta}{\partial x} = \begin{cases} -\sqrt{\frac{g}{H}} \eta_0 e^{\frac{x}{R}}, & x < 0 \\ -\sqrt{\frac{g}{H}} \eta_0 e^{-\frac{x}{R}}, & x > 0 \end{cases}$$

$$x \to R, \quad v \to 0$$

Energetics of geostrophic adjustment

For unit length in the y-direction:



Potential energy:

$$\int_0^{\eta} \int_0^1 \int_{x_1}^{x_2} \rho_0 gz dx dy dz = \int_0^{\eta} \int_{x_1}^{x_2} \rho_0 gz dx dz = \int_{x_1}^{x_2} \int_0^{\eta} \rho_0 gz dz dx = \frac{1}{2} \rho_0 g \int_{x_1}^{x_2} \eta^2 dx$$

PE change is limited to $(-R:R, -\eta_0: \eta_0)$

$$\begin{split} PE_I &= \frac{1}{2} \rho_0 g \int_{-R}^0 \eta_0^2 dx + \frac{1}{2} \rho_0 g \int_0^R (-\eta_0)^2 dx = \frac{1}{2} \rho_0 g R \eta_0^2 + \frac{1}{2} \rho_0 g R \eta_0^2 = \rho_0 g R \eta_0^2 \\ PE_F &= \frac{1}{2} \rho_0 g \eta_0^2 \int_{-R}^0 (1 - e^{\frac{x}{R}})^2 dx + \frac{1}{2} \rho_0 g \eta_0^2 \int_0^R (1 - e^{\frac{x}{R}})^2 dx \\ &= \frac{1}{2} \rho_0 g \eta_0^2 \left(x |_{-R}^0 - 2R e^{\frac{x}{R}}|_{-R}^0 + \frac{R}{2} e^{\frac{2x}{R}}|_{-R}^0 + x |_0^R + 2R e^{\frac{x}{R}}|_0^R - \frac{R}{2} e^{\frac{2x}{R}}|_0^R \right) \\ &= -\frac{1}{2} \rho_0 g R \eta_0^2 \end{split}$$

$$\Delta PE = -\frac{3}{2} \rho_0 g R \eta_0^2$$

$$\Delta PE = -\frac{3}{2} \rho_0 g R \eta_0^2$$

Kinetic energy:

$$KE_I = 0$$

$$\eta_0 \ll H$$

$$\int_0^0 \int_0^R 1$$

$$KE_F = \int_{-H}^{0} \int_{-R}^{R} \frac{1}{2} \rho_0 v^2 dx dz$$

$$= \int_{-H}^{0} \int_{-R}^{0} \frac{1}{2} \rho_0 \eta_0^2 \frac{g}{H} e^{\frac{2x}{R}} dx dz + \int_{-H}^{0} \int_{0}^{R} \frac{1}{2} \rho_0 \eta_0^2 \frac{g}{H} e^{\frac{2x}{R}} dx dz$$

$$= \frac{1}{2}\rho_0\eta_0^2 g(\int_{-R}^0 e^{\frac{2x}{R}} dx + \int_0^R e^{-\frac{2x}{R}} dx)$$

$$= \frac{1}{2}\rho_0 \eta_0^2 g(\frac{R}{2}e^{\frac{2x}{R}}|_{-R}^0 - \frac{R}{2}e^{-\frac{2x}{R}}|_0^R)$$

$$=\frac{1}{2}\rho_0 gR\eta_0^2$$

$$v = \begin{cases} -\sqrt{\frac{g}{H}} \eta_0 e^{\frac{x}{R}}, & x < 0 \\ -\sqrt{\frac{g}{H}} \eta_0 e^{\frac{x}{R}}, & x > 0 \end{cases}$$

$$\Delta PE = -\frac{3}{2}\rho_0 gR{\eta_0}^2$$

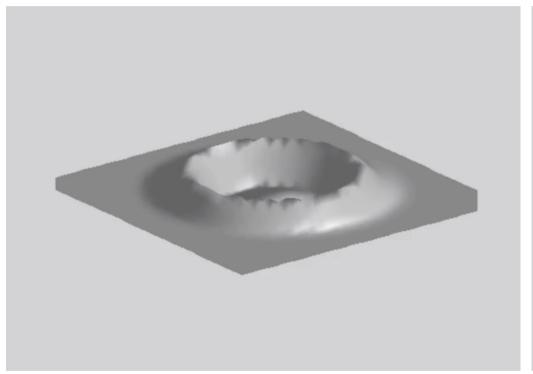
$$\frac{\Delta KE}{|\Delta PE|} = 1/3$$

Only 1/3 of released PE is converted into KE, and the rest of lost PE is radiated away by waves

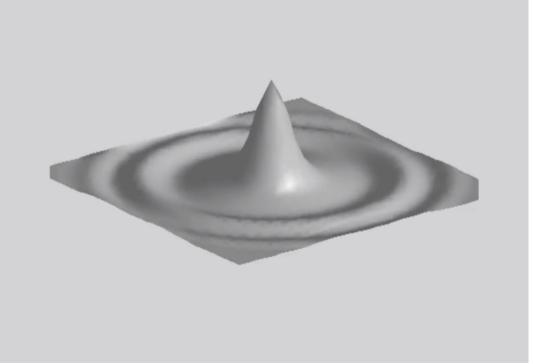
inertial-gravity waves

Adjustment of perturbations

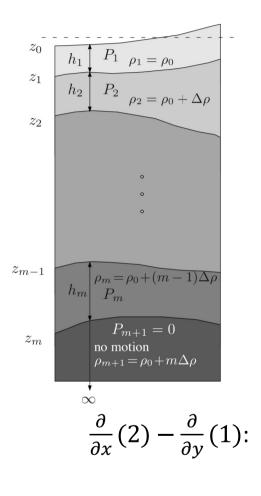
without rotation



with rotation



shallow-water reduced gravity model – one layer



$$z_1 = -h_1$$

$$z_1 = -h_1 \qquad \qquad P_1 = \rho_0 g' h_1$$

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} + v \frac{\partial u}{\partial y} - fv = -g' \frac{\partial h}{\partial x}$$
 (1)

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + v \frac{\partial v}{\partial y} + f u = -g' \frac{\partial h}{\partial y} \quad (2)$$

$$\frac{\partial h}{\partial t} + \frac{\partial hu}{\partial x} + \frac{\partial hv}{\partial y} = 0 \qquad \qquad \frac{dh}{dt} + h(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y}) = 0$$

$$\frac{dh}{dt} + h(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y}) = 0$$

$$\frac{\partial}{\partial x}(2) - \frac{\partial}{\partial y}(1): \qquad \frac{d(f+\zeta)}{dt} + \left(\frac{\partial u}{\partial x} + \frac{\partial v}{\partial y}\right)(f+\zeta) = 0 (1)$$

$$\frac{d}{dt} \left(\frac{f + \zeta}{h} \right) = 0$$
 PV conservation

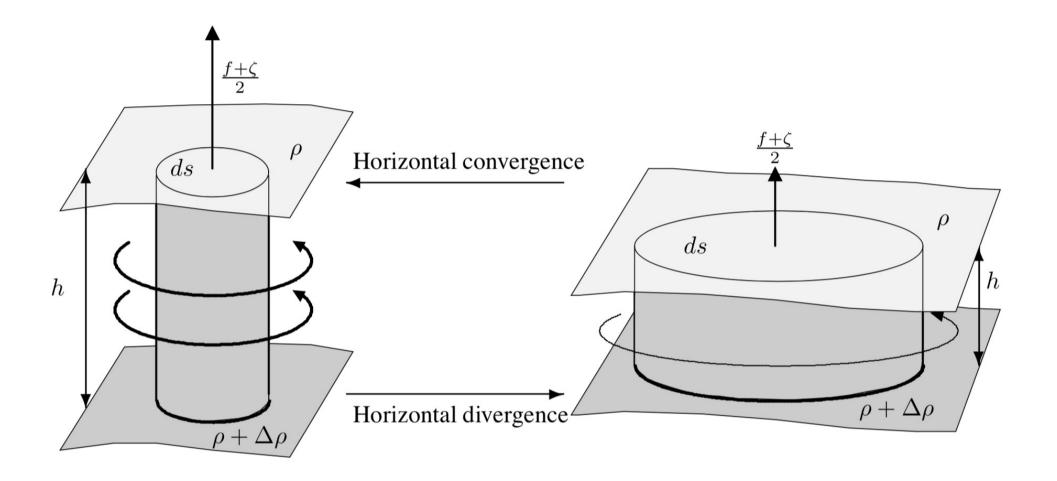
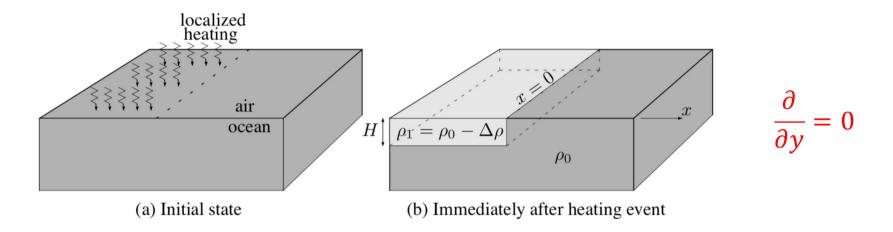


Figure 12-4 Conservation of volume and circulation in a fluid undergoing divergence (squeezing) or convergence (stretching). The products of h ds and $(f + \zeta)$ ds are conserved during the transformation, implying conservation of $(f + \zeta)/h$, too.

Geostrophic adjustment - baroclinic



For the lighter layer:

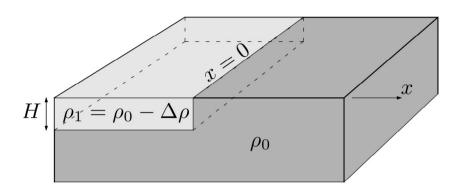
$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} - fv = -g' \frac{\partial h}{\partial x}$$

$$\frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + fu = 0$$

$$\frac{\partial h}{\partial t} + \frac{\partial}{\partial x} (hu) = 0$$

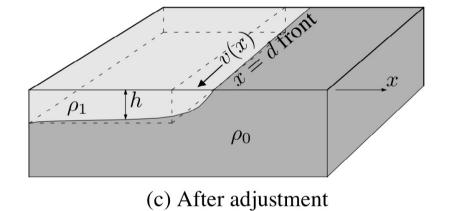
Initial state (unbalanced):

$$h = \begin{cases} H, & x < 0 \\ 0, & x > 0 \end{cases}$$
$$u = v = 0$$



Boundary conditions:

$$x \to -\infty$$
, $h \to H$, $u, v \to 0$
 $x \to d$, $h \to 0$



Final state (steady):
$$\frac{\partial}{\partial t} = 0$$

$$\frac{\partial hu}{\partial x} = 0$$

at
$$x = d$$
, $h = 0$, $hu = 0$: $hu = 0$ everywhere $\longrightarrow u = 0$

$$\frac{\partial u}{\partial t} + u \frac{\partial u}{\partial x} - fv = -g' \frac{\partial h}{\partial x} \qquad \frac{\partial v}{\partial t} + u \frac{\partial v}{\partial x} + fu = 0$$
$$-fv = -g' \frac{dh}{dx}$$

PV conservation:

$$\frac{f}{H} = \frac{f + \frac{\partial v}{\partial x}}{h} \longrightarrow \frac{f}{H} = \frac{f + \frac{g'}{f} \frac{d^2 h}{\partial x^2}}{h}$$

baroclinic deformation radius

$$R = \frac{\sqrt{g'H}}{f}$$

$$fh = fH + \frac{g'H}{f} \frac{d^2h}{\partial x^2}$$

$$R^2 \frac{d^2 h}{\partial x^2} - h + H = 0$$

$$h = f(x) + H$$

$$h = f(x) + H$$
 $f(x)$: $R^2 \frac{d^2 h}{\partial x^2} - h = 0$

$$x \to -\infty, h \to H$$

$$f(x) = Ae^{\lambda x} \longrightarrow R^2 \lambda^2 - 1 = 0 \longrightarrow \lambda = \pm \frac{1}{R} \longrightarrow \lambda = \frac{1}{R} \qquad f(x) = Ae^{x/R}$$

$$f(x) = Ae^{x/R}$$

$$x \to d, \qquad h \to 0: \qquad f(x) \to -H$$

$$f(x) = Be^{(x-d)/R} \qquad B = -H$$

$$h = H(1 - e^{\frac{x-d}{R}})$$

$$-fv = -g'\frac{dh}{dx}$$

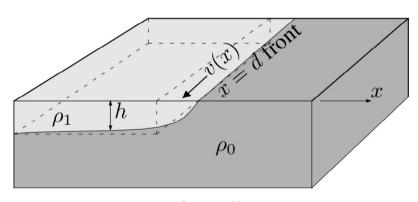
 $v = -\sqrt{g'H} e^{\frac{x-d}{R}}$

The depleted volume of light water in x < 0 should be equal to the volume of light water in x > 0:

$$\int_{-\infty}^{0} (H - h) dx = \int_{0}^{d} h dx$$

$$H \int_{-\infty}^{0} e^{\frac{x - d}{R}} dx = Hd - H \int_{0}^{d} e^{\frac{x - d}{R}} dx$$

$$Re^{\frac{x - d}{R}}|_{-\infty}^{0} = d - Re^{\frac{x - d}{R}}|_{0}^{d} \qquad d = R \quad \text{adjustment spatial scale is } R$$



(c) After adjustment

Energetics

Initial state:
$$KE_i = 0$$

$$KE_i = 0$$

$$PE_{i} = \frac{1}{2}\rho_{0} \int_{-\infty}^{0} g'H^{2}dx = \frac{1}{2}\rho_{0}g'H^{2}x|_{-\infty}^{0}$$

$$KE_f = \frac{1}{2}\rho_0 \int_{-\infty}^{R} hv^2 dx = \frac{1}{2}\rho_0 g' H^2 \int_{-\infty}^{R} (1 - e^{\frac{x - R}{R}}) e^{2\frac{x - R}{R}} dx$$

$$v = -\sqrt{g'H} e^{\frac{x-R}{R}}$$

$$h = H(1 - e^{\frac{x - R}{R}})$$

$$v = -\sqrt{g'H} e^{\frac{x-R}{R}} = \frac{1}{2}\rho_0 g' H^2 \left(\frac{R}{2} e^{2\frac{x-R}{R}}|_{-\infty}^R - \frac{R}{3} e^{3\frac{x-R}{R}}|_{-\infty}^R\right) = \frac{1}{12}\rho_0 g' H^2 R$$

$$PE_f = \frac{1}{2}\rho_0 \int_{-\infty}^{R} g'h^2 dx = \frac{1}{2}\rho_0 g'H^2 \int_{-\infty}^{R} (1 - e^{\frac{x - R}{R}})^2 dx$$

$$\begin{split} &= \frac{1}{2} \rho_0 g' H^2(x|_{-\infty}^R - 2Re^{\frac{x-R}{R}}|_{-\infty}^R + \frac{R}{2} e^{\frac{x-R}{R}}|_{-\infty}^R) \\ &= \frac{1}{2} \rho_0 g' H^2(x|_{-\infty}^R - \frac{3}{2}R) \end{split}$$

$$\Delta PE = PE_i - PE_f = \frac{1}{4}\rho_0 g' H^2 R \qquad \qquad \frac{\Delta KE}{\Delta PE} = 1/3$$