## The Moffat Point-Spread Function

- The PSF takes into account the spreading of the photons due to the atmospheric turbulence (seeing)
- It distributes the signal over an area larger than the real source ones (e.g., a star becomes a circle...)
- It can be analitically modelled.

Here we follow Trujillo et al., MNRAS, 2001, 328, 977.

Circular Moffat Point Spread Function:

$$PSF(r) = \frac{\beta - 1}{\pi \alpha^2} \left[ 1 + \left( \frac{r}{\alpha} \right)^2 \right]^{-\beta},$$

where r is the distance from the center of the source, beta is a parameter of the function and the Full-Width Half-Maximum is:

$$FWHM = 2\alpha\sqrt{2^{1/\beta} - 1},$$

Exercize: reproduce the upper panel of Fig. 1 in Trujillo 2001 (including the Gaussian)

Use alpha=2.1; for the gaussian, remember that FWHM=sigma, so mean is zero and sigma=1 in units of r/FWHM

Hint: on the x axis we have r/FWHM. So, generate r between a small number and 4, then multiply it by FWHM.

Now, substitute in Eq. 1 and produce a vector of r/f values as in Fig. 1. Note that the values of this vector do depend on beta.

Note that on the y-axis, you have PSF(r)/PSF(0)

Try to use functions: they will be useful in the following exercizes.

```
def PSF(r, alpha, beta):
    res= ....
    return res

N=200
beta=1.50
alpha=1.75
psf=np.zeros(N)
fwhm=(2.*alpha*np.sqrt(2.**(1/b)-1))
r=np.arange(0.01, 4, (4-0.01)/N)
for i,x in enumerate(r):
```

psf/= psf[0]

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You should get something like:

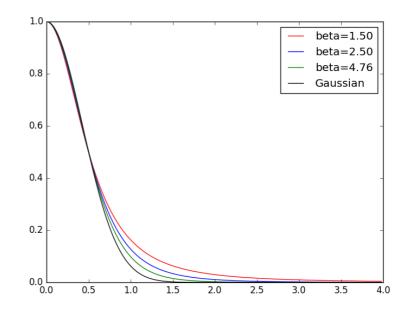
psf[i]=PSF(x\*fwhm,alpha,b)

Question: could you use a lambda function?

Question: what is the effect of varying alpha?

Exercize: produce the bottom

panel of Fig. 1



Final exercize: evaluate and plot the fraction of energy inside R, varying beta.

For doing this, we must do an area integral of PSF(r):

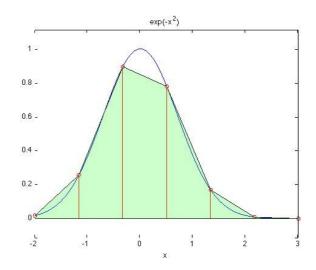
$$\int_{0}^{2\pi} d\varphi \int_{0}^{R} r PSF(r) dr = 2\pi \int_{0}^{R} r PSF(r) dr$$

This must be normalized to the same integral, done for  $R \rightarrow \infty$  (Numerically, R very large)

## Hints:

- write a function to evaluate PSF(r), with beta as a parameter
- write a function to evaluate the integral, at a given R, that must call the previous one
- write the main body of the code using such functions.
- do the integral using the trapezoidal rule

$$\begin{split} y &= \int_a^b f(x) dx \cong \frac{1}{2} \Bigg[ \sum_{i=1}^n f(x_{i-1}) \Delta x + \sum_{i=1}^n f(x_i) \Delta x \Bigg] \\ &= \frac{\Delta x}{2} \{ \big[ f(x_0) + f(x_1) \big] + \big[ f(x_1) + f(x_2) \big] + \\ &\dots + \big[ f(x_{n-1}) + f(x_n) \big] \} \\ &= \frac{\Delta x}{2} \big[ f(x_0) + 2 f(x_1) + \dots + 2 f(x_{n-1}) + f(x_n) \big] \end{split}$$



- redo the integral using scipy.integrate. For example, the integral of:

$$\int_{0}^{1} a x^{2} + bx dx$$

with a=2, b=1 can be computed by this way:

```
>>> from scipy.integrate import quad
>>> def integrand(x, a, b):
...    return a*x**2 + b
...
>>> a = 2
>>> b = 1
```

```
>>> I = quad(integrand, 0, 1, args=(a,b))
>>> I
(1.66666666666667, 1.8503717077085944e-14)
```

Note that there are a number of different integration methods in scipy.integrate!