

The Moffat Point-Spread Function

- The PSF takes into account the spreading of the photons due to the atmospheric turbulence (seeing)
- It distributes the signal over an area larger than the real source ones (e.g., a star becomes a circle...)
- It can be analytically modelled.

Here we follow Trujillo et al., MNRAS, 2001, 328, 977.

Circular Moffat Point Spread Function:

$$\text{PSF}(r) = \frac{\beta - 1}{\pi \alpha^2} \left[1 + \left(\frac{r}{\alpha} \right)^2 \right]^{-\beta},$$

where r is the distance from the center of the source, β is a parameter of the function and the Full-Width Half-Maximum is:

$$\text{FWHM} = 2\alpha \sqrt{2^{1/\beta} - 1},$$

Exercise: reproduce the upper panel of Fig. 1 in Trujillo 2001 (including the Gaussian)

Use $\alpha=2.1$; for the gaussian, remember that $\text{FWHM}=\sigma$, so mean is zero and $\sigma=1$ in units of r/FWHM

Hint: on the x axis we have r/FWHM . So, generate r between a small number and 4, then multiply it by FWHM .

Now, substitute in Eq. 1 and produce a vector of r/f values as in Fig. 1. Note that the values of this vector do depend on β .

Note that on the y-axis, you have $\text{PSF}(r)/\text{PSF}(0)$

Try to use functions: they will be useful in the following exercises.

```
def PSF(r, alpha, beta):
```

```
    res= ....
```

```
    return res
```

```
N=200
```

```
beta=1.50
```

```
alpha=1.75
```

```
psf=np.zeros(N)
```

```
fwhm=(2.*alpha*np.sqrt(2.**(1/b)-1))
```

```
r=np.arange(0.01, 4, (4-0.01)/N)
```

```
for i,x in enumerate(r):
```

```
    psf[i]=PSF(x*fwhm,alpha,b)
```

```
psf/= psf[0]
```

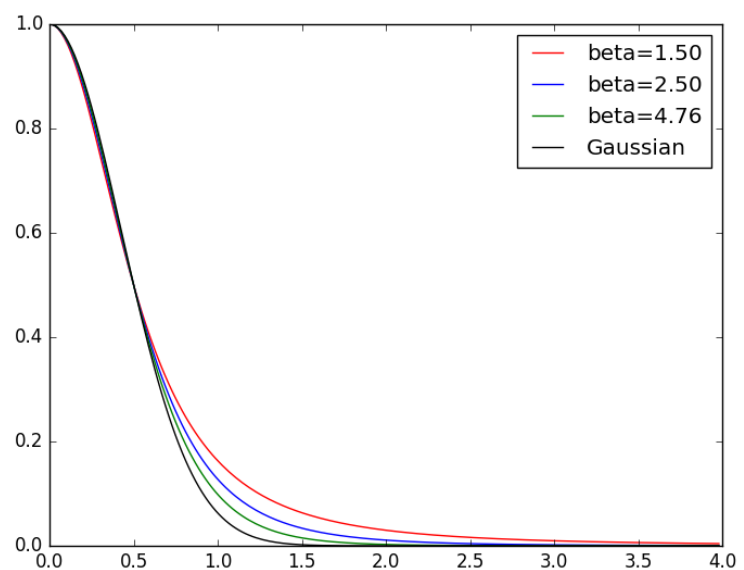
```
....
```

You should get something like:

Question: could you use a lambda function?

Question: what is the effect of varying alpha?

Exercise: produce the bottom panel of Fig. 1



Final exercise: evaluate and plot the fraction of energy inside R, varying beta.

For doing this, we must do an area integral of PSF(r):

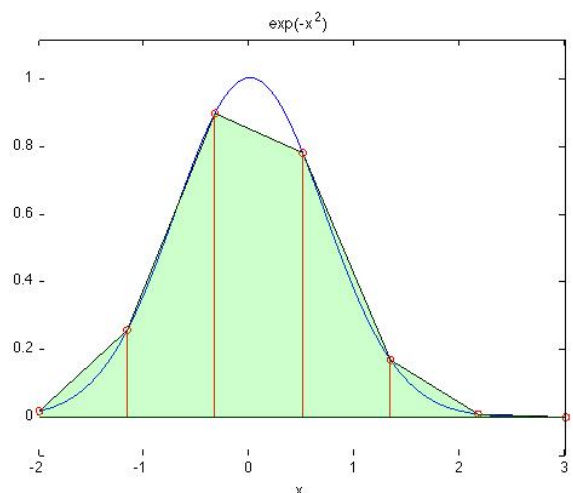
$$\int_0^{2\pi} d\varphi \int_0^R r \text{PSF}(r) dr = 2\pi \int_0^R r \text{PSF}(r) dr$$

This must be normalized to the same integral, done for $R \rightarrow \infty$ (Numerically, R very large)

Hints:

- write a function to evaluate PSF(r), with beta as a parameter
- write a function to evaluate the integral, at a given R, that must call the previous one
- write the main body of the code using such functions.
- do the integral using the trapezoidal rule

$$\begin{aligned} y &= \int_a^b f(x) dx \cong \frac{1}{2} \left[\sum_{i=1}^n f(x_{i-1}) \Delta x + \sum_{i=1}^n f(x_i) \Delta x \right] \\ &= \frac{\Delta x}{2} \{ [f(x_0) + f(x_1)] + [f(x_1) + f(x_2)] + \\ &\quad \dots + [f(x_{n-1}) + f(x_n)] \} \\ &= \frac{\Delta x}{2} [f(x_0) + 2f(x_1) + \dots + 2f(x_{n-1}) + f(x_n)] \end{aligned}$$



- redo the integral using scipy.integrate. For example, the integral of:

$$\int_0^1 a x^2 + b x dx$$

with a=2, b=1 can be computed by this way:

```
>>> from scipy.integrate import quad
>>> def integrand(x, a, b):
...     return a*x**2 + b
...
>>> a = 2
>>> b = 1
```

```
>>> I = quad(integrand, 0, 1, args=(a,b))
>>> I
(1.6666666666666667, 1.8503717077085944e-14)
```

Note that there are a number of different integration methods in `scipy.integrate`!

Adding a surface brightness profile

The surface brightness profile of an Elliptical galaxy can be described using a Sersic profile:

$$I(r) = I_0 10^{-b_n [(r/r_e)^{1/n}]}$$

where I_0 is the central intensity, r_e the effective radius, and n , b_n are parameters.

Taking

$$b_n = 0.868n - 0.142 \quad ,$$

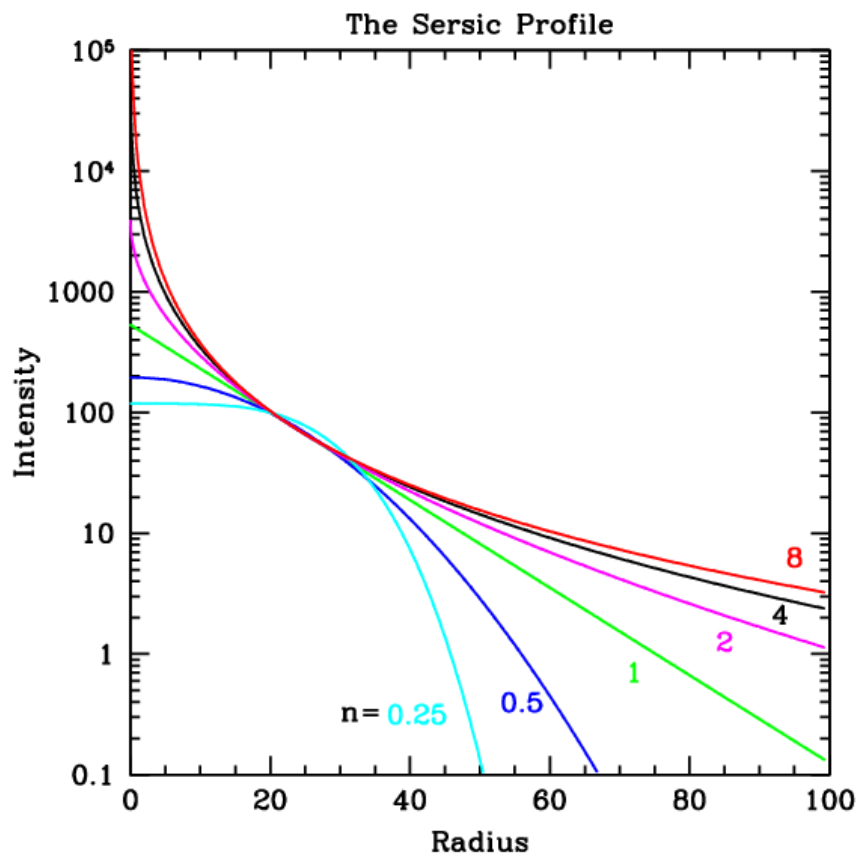
$$r_e = 2 \text{ FWHM} \quad , \text{ and}$$

$$I_0 = 1 \quad , \text{ plot the profiles}$$

for $n=1,2,4,8$

Exercise:

How would you compute the total intensity as a function of $n=1,2,4,8$, using the Moffat PSF?



Write a code that plots the total intensity as a function of r, n and β . Compare the total intensities with and without using the Sersic profile.