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EuclidAB	Calculation of the Euclidean Distance Among Two Sets of Points			

Description

Function EuclidAB calculates the Enclidean distance between two sets of points A and B.

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Usage

```
EuclidAB(a, b, squared = FALSE)
```

Arguments

a First matrix of cartesian coordinates.

b Second matrix of cartesian coordinates.

squared Should the squared Euclidean distances be returned (default: FALSE).

Details

The standard R function used to calculate the Euclidean distance, namely dist, does only allow one to calculate pairwise distances between the rows of a single matrix of cartesian coordinates and return a dist-class object, which is a one-dimensional array that is meant to be interpreted as a triangular matrix. Function Euclid allows the user to provide two data matrices a and b and output a rectangular Euclidean distance matrix.

Value

An matrix with as many rows as a and as many columns as there are rows in b.

Author(s)

Guillaume Guénard <guillaume.guenard@umontreal.ca>

References

```
[To be included...]
```

See Also

The dist-class and associated methods.

Examples

```
##
### First example: locations spread throughout the world
##
## [Examples here...]
```

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geodesics Calculation Of Geodesic Distances	geodesics	Calculation Of Geodesic Distances
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Description

Function geodesics carries out the calculation of pariwise geodesic distances within a set of coordinates or between two sets thereof using one of two calculation approaches.

Usage

```
geodesics(x, y, method = c("haversine", "Vincenty"), radius = 6371000,
  sma = 6378137, flat = 1/298.257223563, maxiter = 1024L,
  tol = .Machine$double.eps^0.75)
```

Arguments

X	$A \ set \ of \ geographic \ coordinates \ in \ the \ form \ of \ a \ two-column \ {\tt matrix} \ or \ {\tt data.frame}.$
У	An optional second set of coordinates in the same forms as x.
method	The calculation method used to obtain the distances (default: haversine method; see details).
radius	Radius of the planetary body (when assuming a sphere).
sma	Length of the semi-major axis of the planetary body (when assuming a revolution elipsoid).
flat	Flattening of the ellipsoid.
maxiter	Maximum number of iterations, whenever iterative calculation is involved.
tol	Tolerance used when iterative calculation is involved.

Details

When only one set of coordinates is given to the function (ie. when y is omitted), the function returns the pairwise distances in the form of a dist-class object representing a lower-triangle matrix. When the second coordinate set is given, the function calculates the distances between each coordinate of x and each coordinate of y and returns a matrix with one column for each coordinate in x and one row for each coordinate in y.

Two calculation methods are implemented. The first is the haversine formula, which assume the planetary body to be a sphere. The radius of that sphere is given to the function as its argument radius, with the default value being the mean radius of planet earth. Of the two methods implemented, the haversine formula is fastest but its precision depend on how well the planetary body match the sphericity assumption. The second method implemented is Vincenty's inverse formula, which assumes the the planetary body is a revolution elipsoid, which is expected for rotating semifluid such as planet earth. Argument sma, the length of the semi-major axis, corresponds to the radius of the circle obtained when the revolution elipsoid at the equator, whereas argument flat correspond to the compression of the sphere, along the diameter joining the poles, to form the ellipsoid of revolution. Their default values corresponds to parameters for planet Earth according to WGS84.

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These values, along with maxiter and tol, are ignored when using the haversine formula, as the value of argument radius is ignored when using Vincenty's inverse formula.

Vincenty's inverse formula is more precise (on the order of 0.5mm) on planet Earth than the haversine formula, but they involve more computation time and may sometimes fail to converge. This is more likely for pairs of locations that are nearly antipodal or both (numerically) very close to the equator. The results returned by the function when using Vincenty's inverse formula are given a niter attribute that gives the number of iterations that were necessary to acheive convergence. Numbers greater than argument maxiter are indicative of failed convergence; a warning is issued in such a circumstance.

Geodesic distance matrices are nonmetric.

Value

```
A dist-class object of a matrix.
```

Author(s)

Guillaume Guénard <guillaume.guenard@umontreal.ca>

References

Vincenty, T. 1975. Direct and Inverse Solutions of Geodesics on the Ellipsoid with application of nested equations. Survey Review XXIII (176): 88-93 doi:10.1179/sre.1975.23.176.88

Inman, J. 1835. Navigation and Nautical Astronomy: For the Use of British Seamen (3 ed.). London, UK: W. Woodward, C. & J. Rivington

See Also

The dist-class and associated methods.

Examples

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```
res_vif <- geodesics(coords,method = "Vincenty")
res_vif-res_hav  ## Absolute difference
200*(res_vif-res_hav)/(res_vif+res_hav) ## Relative difference are smaller
##</pre>
```

sp.cov

Calculation of Spatial Covariance Models

Description

Function sp. cov carries out the calculation of the distance-based spatial covariance models.

Usage

```
sp.cov(d, type = c("spherical", "exponential", "power", "hyperbolic",
    "superelliptic"), alpha = 1, beta = 1)
```

Arguments

d Pairwise distances among objects (see details).

type The type of covariance model (see details).

alpha Shape parameter of the covariance model.

beta Range parameter of the covariance model.

Details

Pairwise distances among objects can be an array with arbitrary dimensions or a dist-class object, in which case a one-dimensional array will be returned.

Argument type can be one of "spherical" (the default), "exponential", "power", "hyperbolic", "superelliptic", or any unambiguous abbreviation; and correspond to the particular function used to obtain the weights as a function of the distances. The function does not directly return a covariance matrix but weights that are used calculate the spatial covariances.

Distance-weight relationships are monotonically decreasing functions associating the level of ressemblance of two locations with the distance separating them. For all covariance models (argument type above), the weight is 1 at d=0 and 0 beyond the range (i.e., when d >= beta). The shape of the relationship is controlled by the shape parameter given as argument alpha. That parameter controls the gentleness of the slope of the distance-weight relationship. It is bounded between values of 0 (the steepest slope: a Kronecker's delta). and 1 (a straight line).

Value

An array with the same dimensions as argument d.

Author(s)

Guillaume Guénard <guillaume.guenard@umontreal.ca>

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References

```
[To be included...]
```

See Also

The dist-class and associated methods.

Examples

```
##
### First example: locations spread throughout the world
##
## [Examples here...]
```

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