

Measuring the Properties of Dwarf Streams

Principal Investigator: Prof. Mary E. Putman

Institution: Columbia University in the City of New York

Electronic Mail: mputman@astro.columbia.edu

Scientific Category: QUASAR ABSORPTION LINES AND IGM

Scientific Keywords: Chemical Abundances, Galaxy Formation And Evolution, Interacting And Merging Galaxies, Interstellar And Intergalactic Medium, Metal Absorption Systems

Instruments: COS

Proprietary Period: 12

Proposal Size: Small

UV Initiative: Yes

Orbit Request

Prime

Parallel

Cycle 21

8

0

Abstract

We propose to measure the metallicity and ionization conditions of a gaseous stream trailing behind a pair of dwarf galaxies in the local universe. The NGC 4532/DDO 137 system is a clear analog to the Milky Way's massive satellite galaxies, the Large and Small Magellanic Clouds; however, this system is not in close proximity to a massive spiral galaxy. Furthermore, it is the only other dwarf stream for which the metallicity and ionization conditions can be directly measured due to the fortuitous alignment of a bright background QSO with the HI stream. We will use the data to examine the properties and formation mechanisms of such streams and how they differ with environment. Such information can be used to develop a method to discriminate between cold accreting filaments and streams stripped from low mass galaxies at all redshifts.