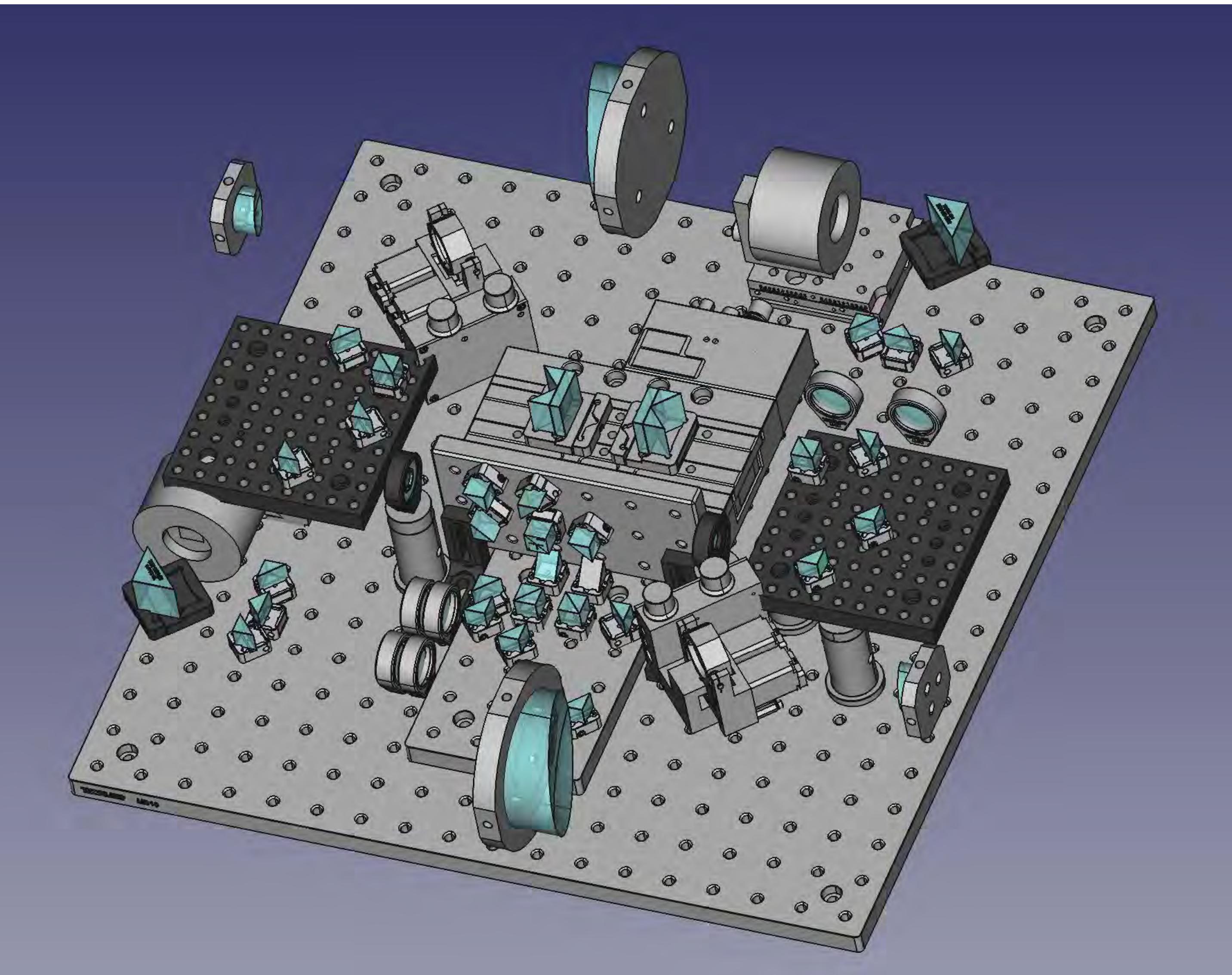


Optimast-SCI Technology: Precision In-Space Manufacturing for Structurally-Connected Interferometry

Simon Patane¹, Max Fagin¹, Daniel Riley¹, Thorin Tobiassen¹, John Kloske¹, Mike Snyder¹, Gerard T. van Belle²

¹Made in Space, Jacksonville FL; ²Lowell Observatory, Flagstaff AZ



Basic Optimast-SCI Parameters

Two-element interferometer

- 2 booms, 6-25m from 2 ESAMM units
 - Baselines selectable from 1 to 50m by running booms in & out
- Resolution: 2 to 6 ms (Rayleigh limit, better with superresolution modes)
- Spacecraft coherence time: 10sec (minimum)
- Sensitivity: $m_V \leq 12 - 16$
 - Out-performs all ground facilities by 2 – 6 magnitudes
- Two collecting apertures, each 2" – 6" (depending on mission)
- Bandpass: 0.4 – 1.0 μm

Precision tracking of outboard mirrors

- Enables optical interferometry without massive structures

Orbit: non-LEO, nominally Earth-Sun L2

- Thermally quiet

Spacecraft

- Mass: 300-400 kg



Extended
Structure
Additive
Manufacturing
Machine
(ESAMM)

At the Heart of Optimast-SCI: A Simple Michelson Combiner

- Simple, pupil-plane combiner with ~0.1 micron tolerances (not nanometer or picometer)
- 75mm range delay line
- Static ABCD fringe tracking similar to PRIMA FSU (Sahlmann 2007)
- Modest spectral resolution ($R \sim 100$)

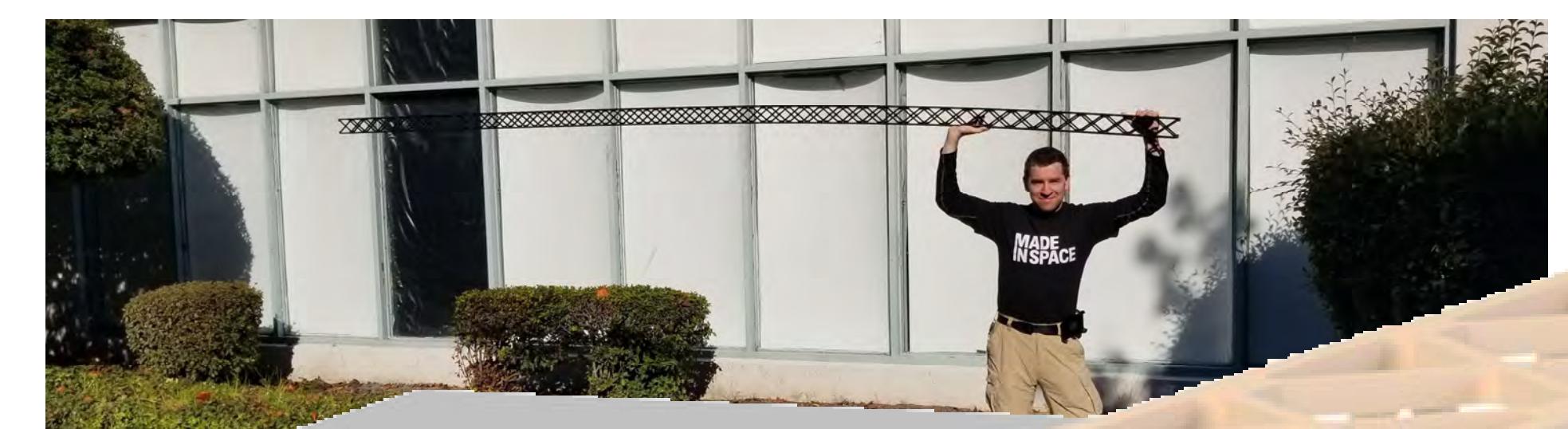
Optimast-SCI is currently a NASA SBIR-funded Phase 2 study of a space-based optical interferometer enabled by in-space manufacturing

SBIR Phase 2 Engineering demonstration units will be built this year for boom manufacturing, optics subsystems

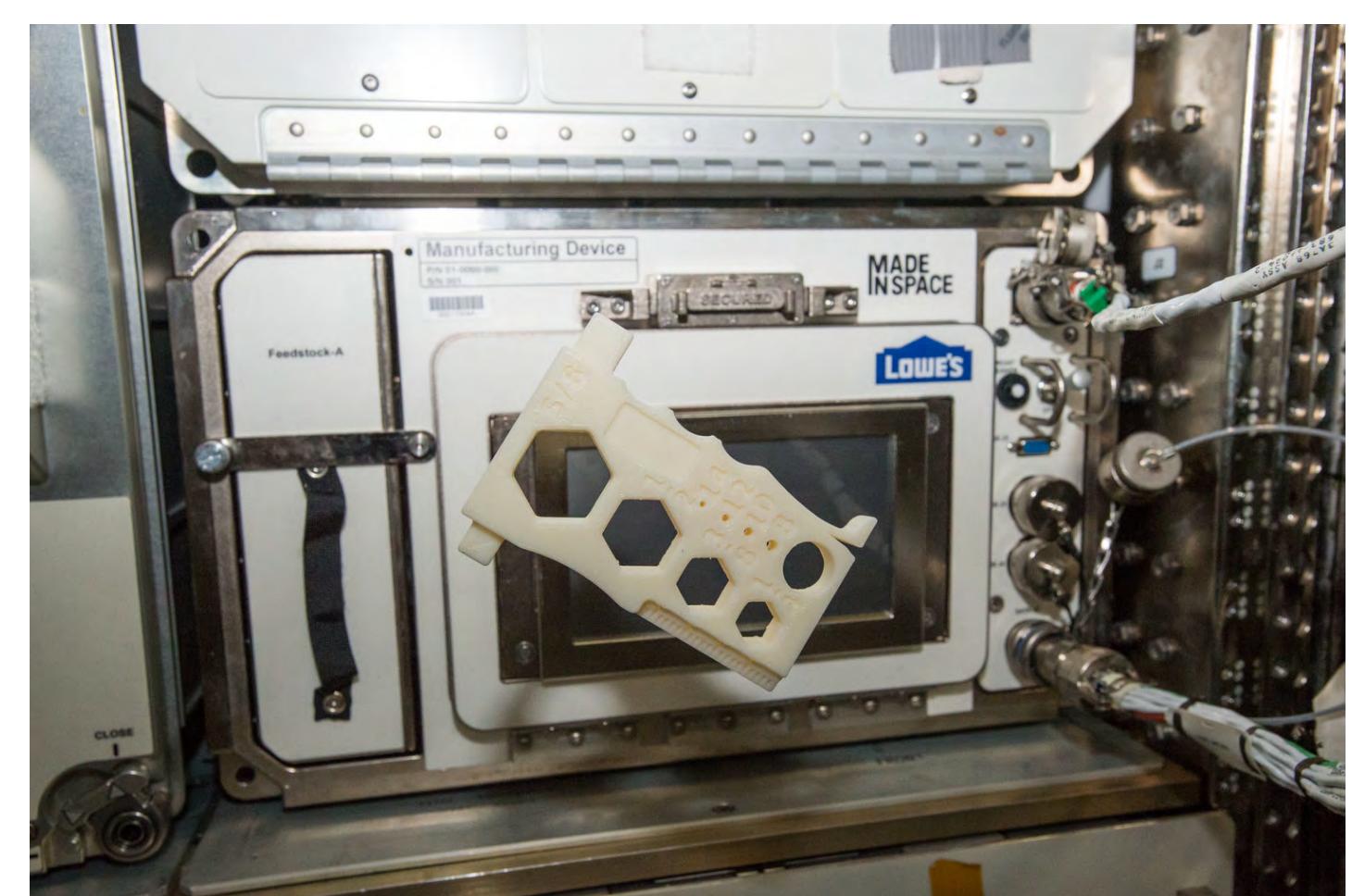
Manufactured Booms versus Free-Fliers

Booms are superior to free fliers

- One spacecraft versus three
- No consumables for pointing
- Outboard units are significantly simpler than free-fliers
- Short booms: single structure from a mechanical perspective
- Long booms (>100m): akin to tethers; outboard unit control is 2 DOF, not 6 DOF (as for free-flier)
 - Long boom case could be treated as simplified free-flier demo
- Failure modes are more failsafe / recoverable



Manufactured
Boom



Made In Space Flight Units for Zero-G 3D Printing

Flown Units aboard ISS Technology Demonstrator 3D Printer (2014)

- Demonstrated fused deposition modeling process in a microgravity environment

Additive Manufacturing Facility (2016)

- Permanent commercial manufacturing facility

MIS Fiber Optics (2017)

- Successfully pulled ZBLAN in microgravity

MIS Braskem Recycler (2019)

- Reuse of 3D objects into feedstock



Flight-Qualified: ESAMM

- Thermal-vac tested for flight: TRL 6
- Guinness World Record for longest single 3D printed piece: 37 meter boom (print terminated when shop space limit reached)

Selected: Archinaut One (2022)

- \$74M flight mission, printing solar array structures

Acknowledgements: This work has been supported by the NASA SBIR program, the Lowell Observatory, and Made In Space.

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Contact:
Simon Patane,
simon@madeinspace.us

Optimast-SCI Science: New Discovery Frontiers with Sensitive, Millarcsecond Resolution

Gerard T. van Belle¹, N. Moskovitz, Simon Patané², Max Fagin², D. Riley², John Kloske², Thorin Tobiassen², Mike Snyder²

¹Lowell Observatory, Flagstaff AZ; ²Made in Space, Jacksonville, FL



Optical interferometry from a space-based platform is freed from the limitations of the Earth's atmosphere.

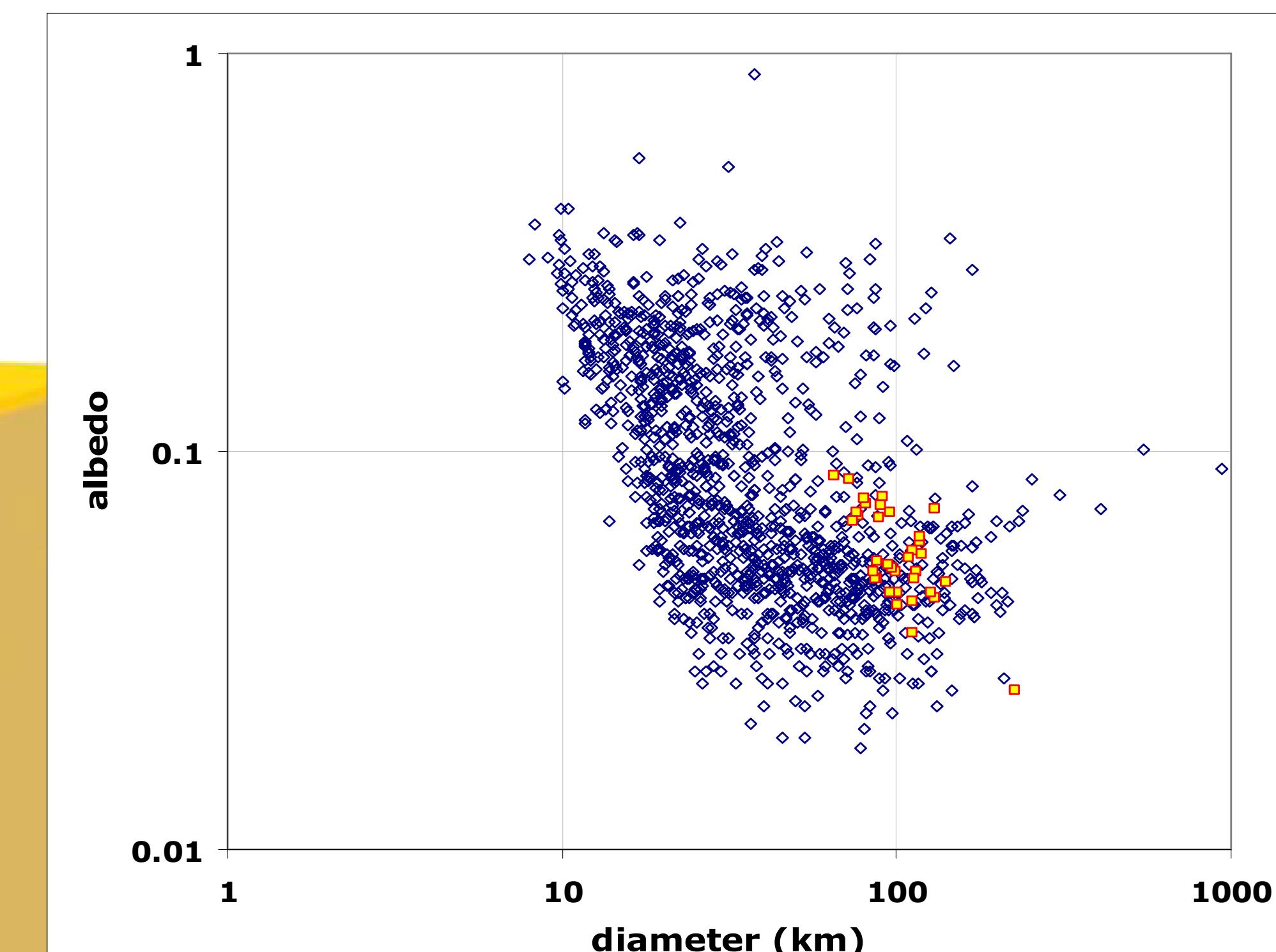
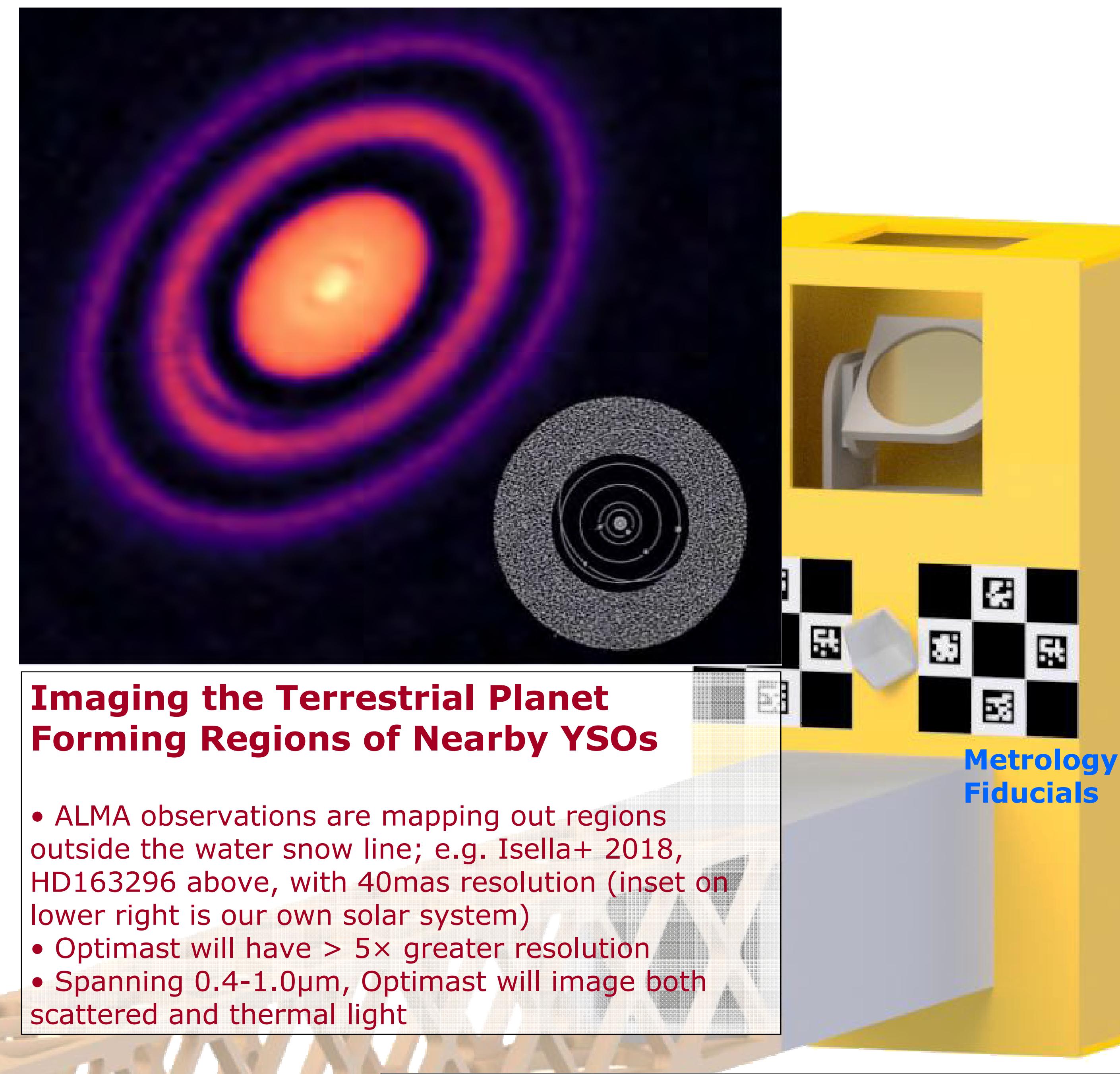
One fundamental benefit is **sensitivity**.

Gains of 100-1000× or more are possible, compared to ground-based optical interferometry.

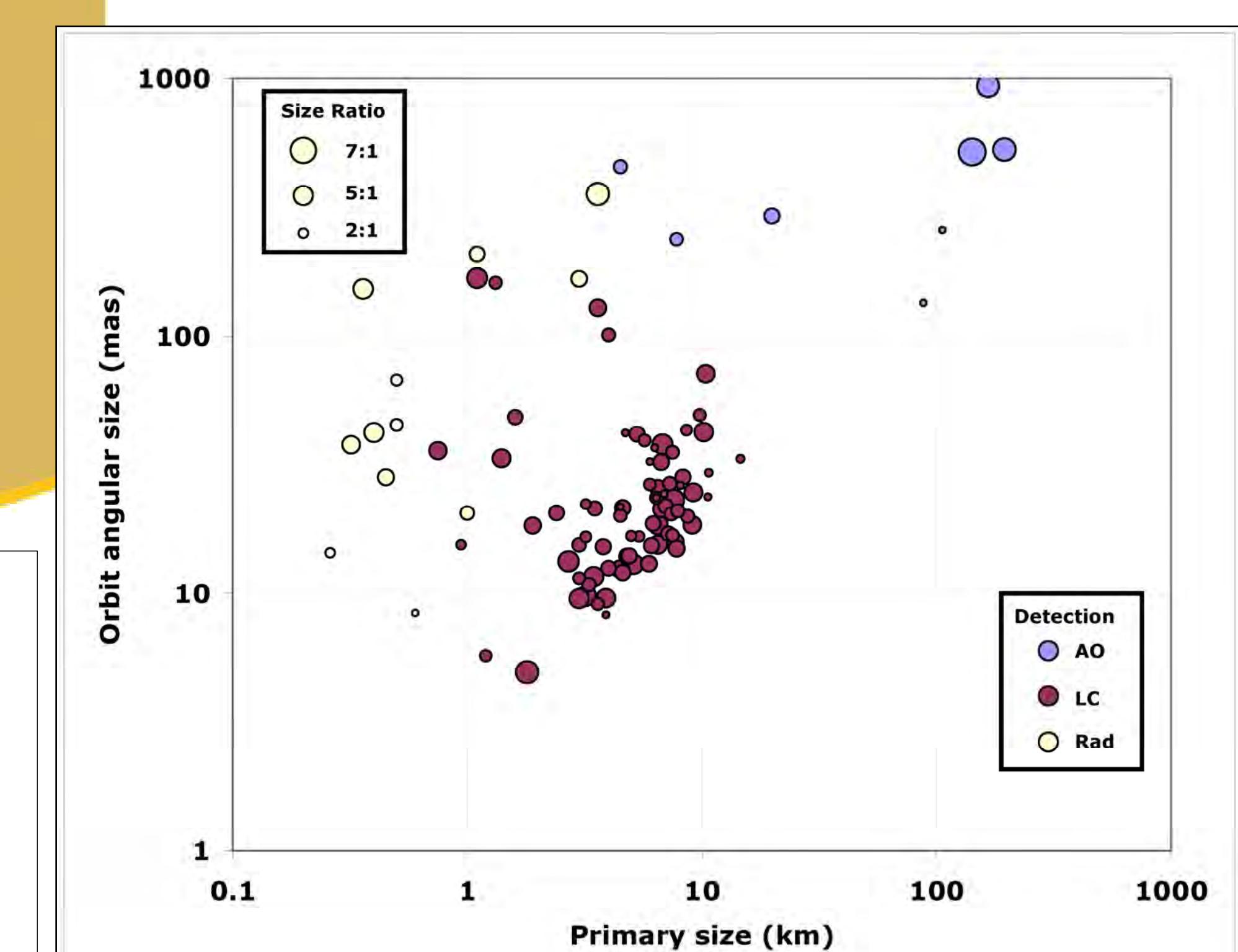
Fundamental Parameters of Low Mass Stellar Binaries

Follows ground-based work of bright stars.
 • The orbit of 12 Boo (Boden et al. 2005) measured by the Palomar Testbed Interferometer, which determined component masses to 0.3%.
 • Despite having a shorter base-line, with a shorter operational wavelength, Optimast will have similar angular resolution but for far fainter targets (e.g. low-mass stars).

Solar Panel



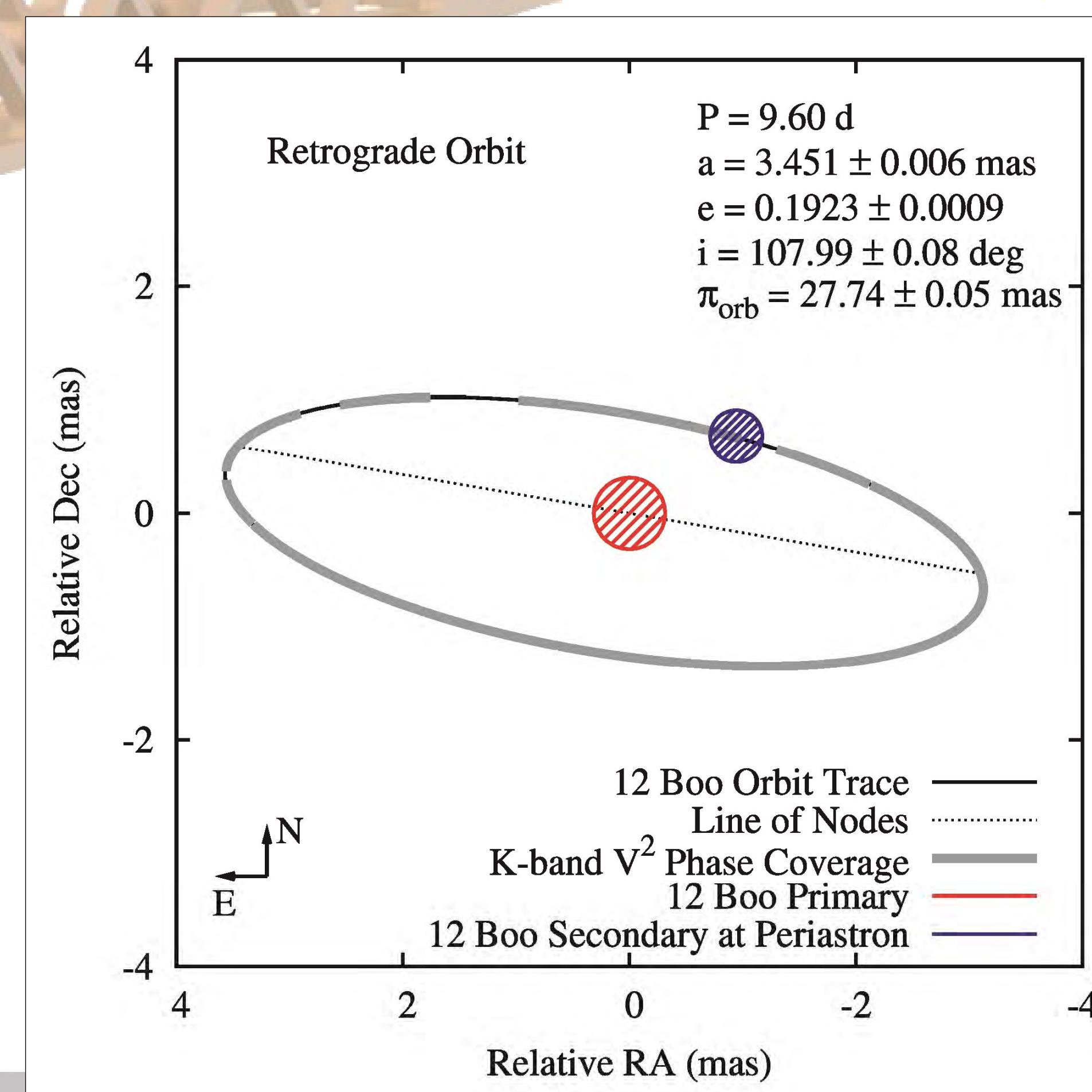
Single Object Targets Resolvable by Optimast (N>1,000)



Known Binary Targets for Orbit Mapping (N>100)

Asteroid exploration: Flyby quality data without the flyby

Extreme optical resolution (~2mas) from a two-element optical space interferometer:
 • Shapes, sizes of ~10³ asteroids
 • Orbit and mass determination for all known V<16 binaries
 • Spatially resolved spectra-photometry in >10 wavelength bands
 • Space-based operation enables **high sensitivity** for reflected light objects
 • Application of ground-based interferometry techniques (well developed but limited to emitted light objects, i.e. stars)



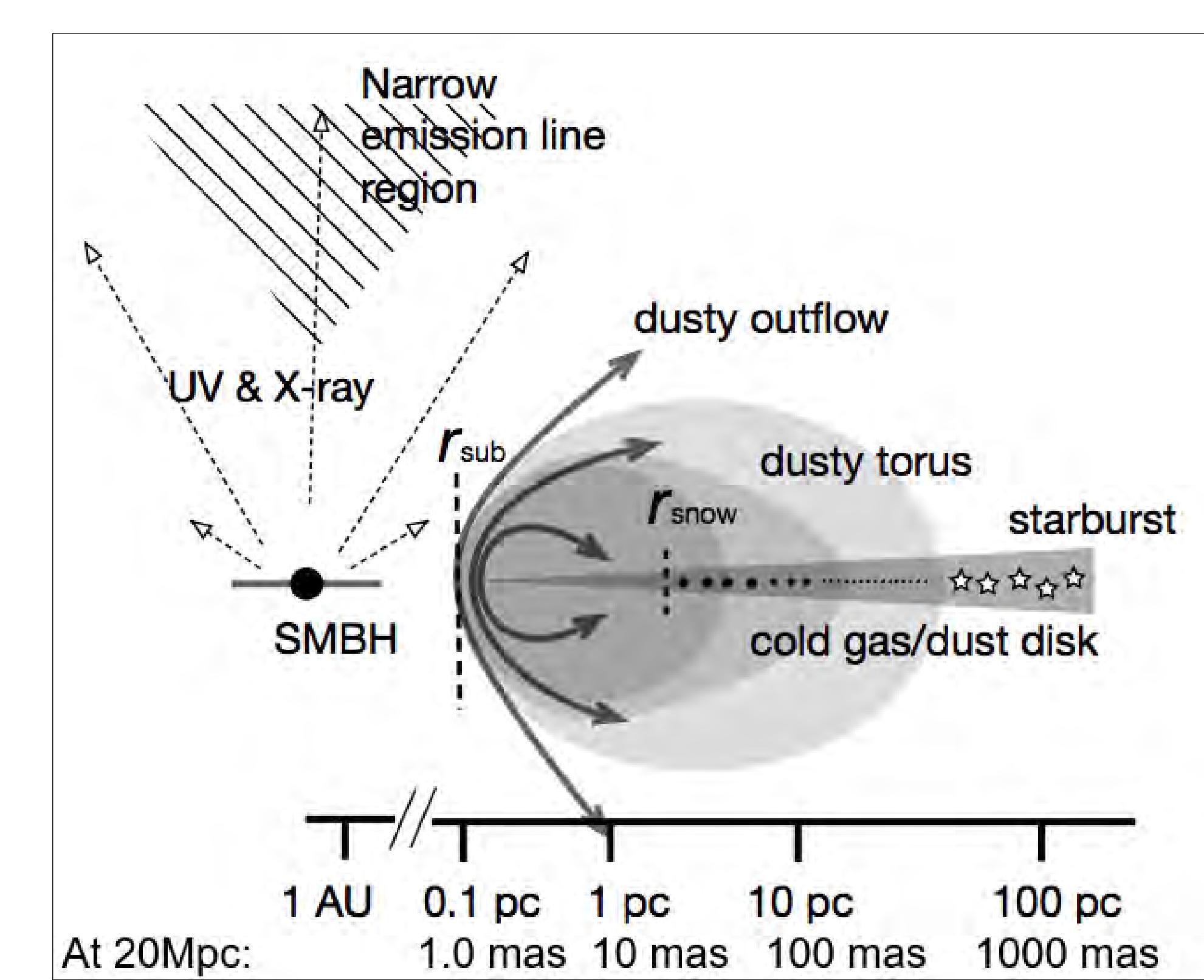
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Contact:
Gerard T. van Belle,
gerard@lowell.edu



Probing the Inner Core Architecture of Active Galactic Nuclei

- At 20 Mpc, 1mas resolution probes the inner edge of the AGN disk
- The feedback process between the dusty wind and host galaxy
- What fraction of supermassive black holes are binary?

Main Belt Asteroids
 • Sizes, shapes for any object > 10km (H<12.3)
 • Resolved surface mapping for > 30 km
 • Rotation > 6 hours (<5° 'smear' in 300sec)
 • Detection of binaries, Keplerian solutions for binary orbits
 • Hundreds of possible targets

Near-Earth Asteroids
 • Direct size determination for >10m objects (H<26)
 • Mapping of binary orbits

Jupiter Trojans
 • H<9.2 (~36 known targets)
 • Orbits / shapes for targets, binary detection

Additional targets: gas giant moons, ice dwarfs