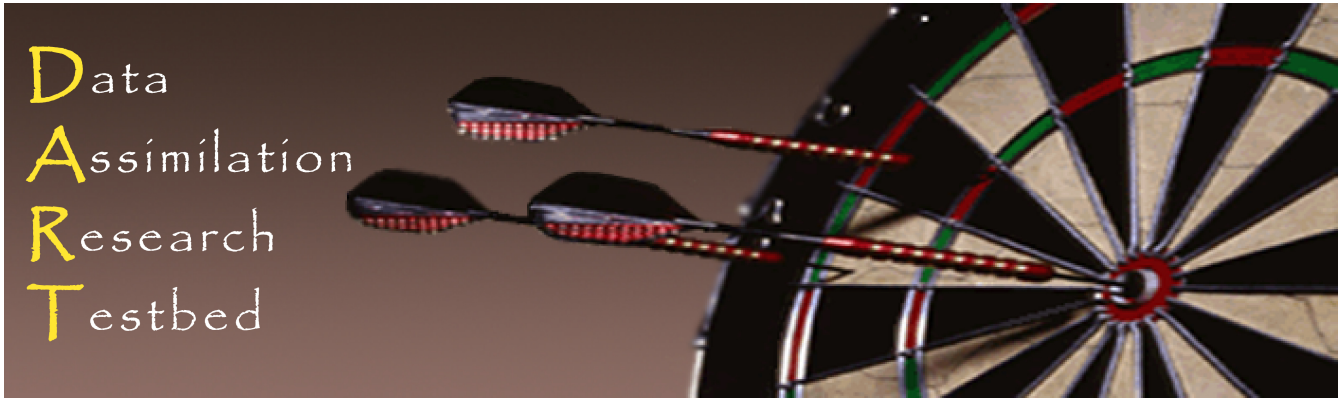
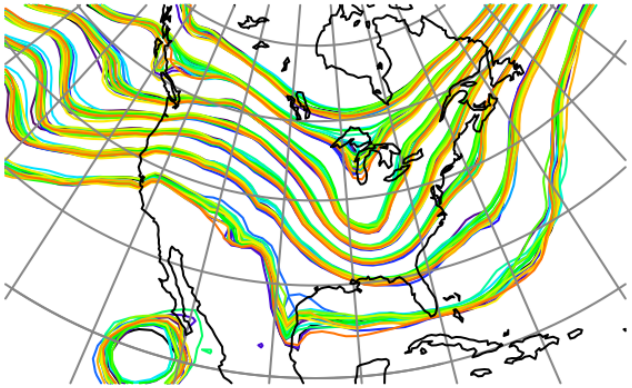


DART_LAB Tutorial Section 1: Ensemble Data Assimilation Concepts in 1D



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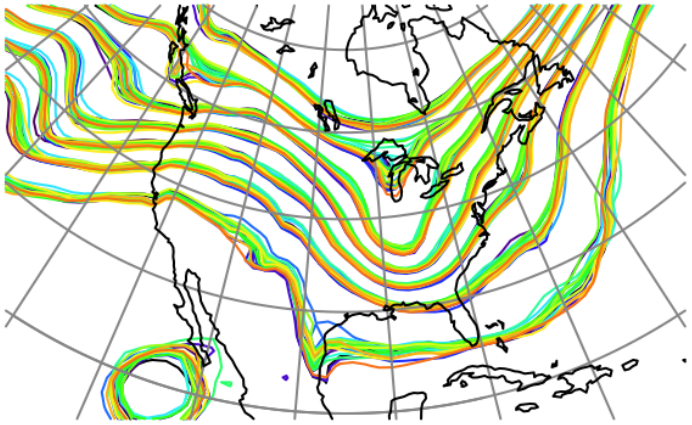
What is Data Assimilation?

Observations combined with a Model forecast …

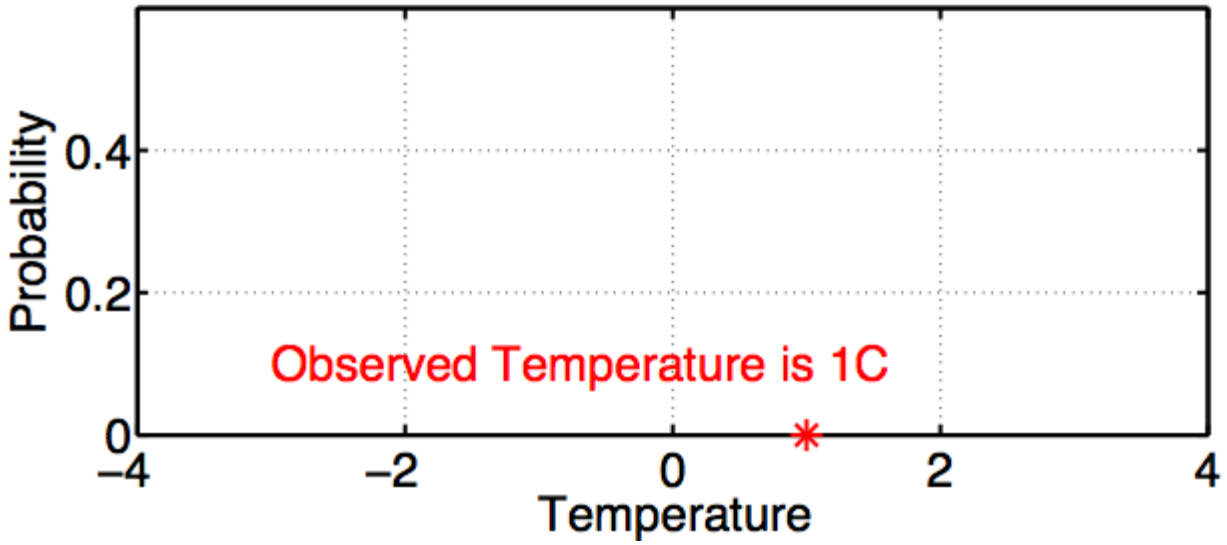
+



…to produce an analysis
(best possible estimate).

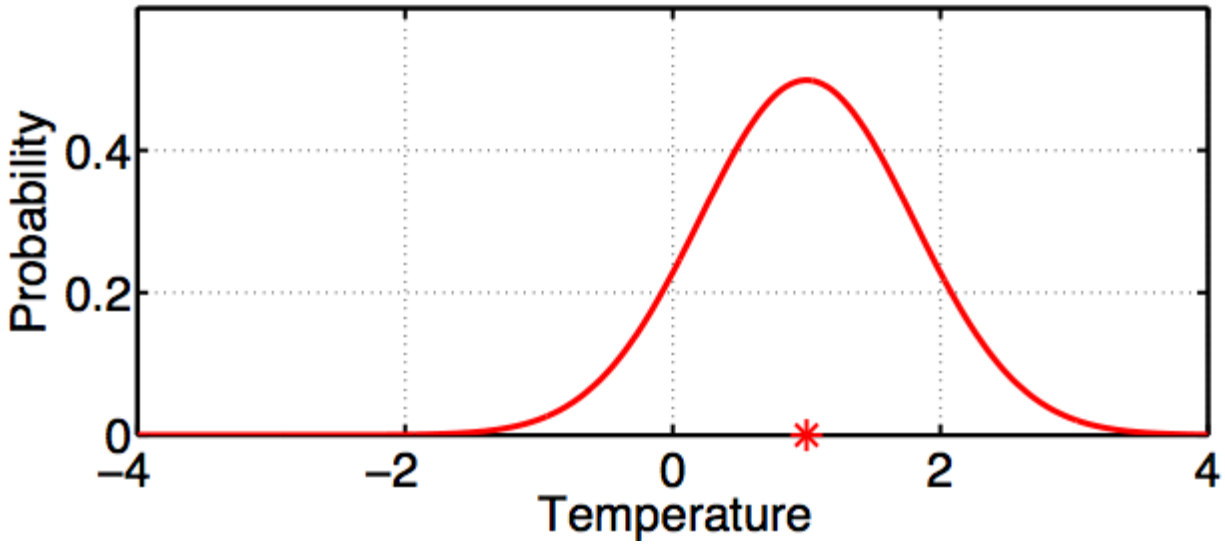
Example: Estimating the Temperature Outside

An observation has a value (*),



Example: Estimating the Temperature Outside

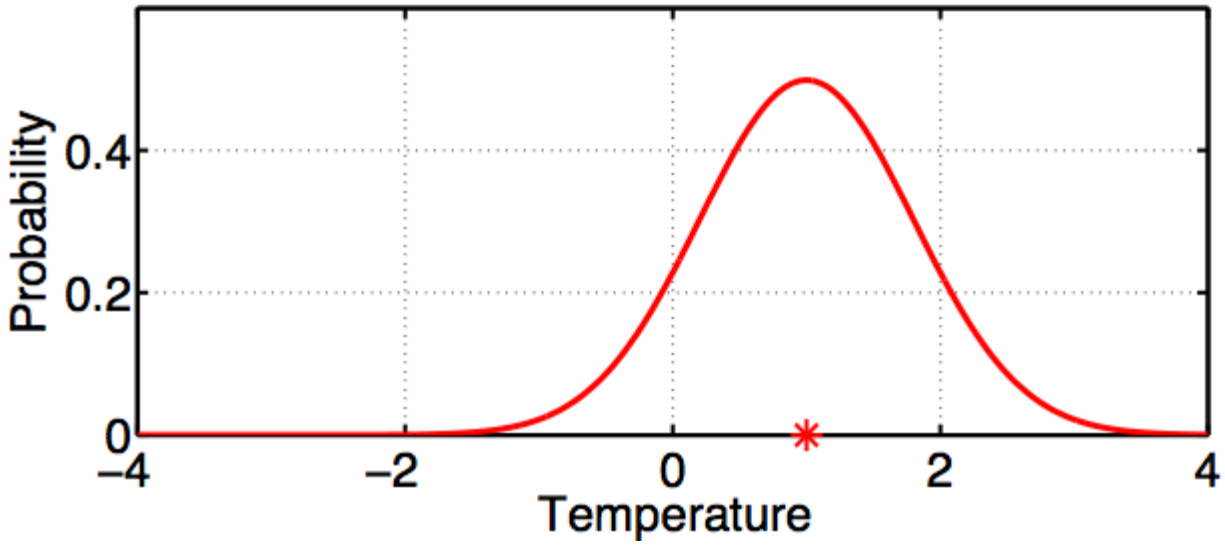
An observation has a value (*),



and an error distribution (red curve) that is associated with the instrument.

Example: Estimating the Temperature Outside

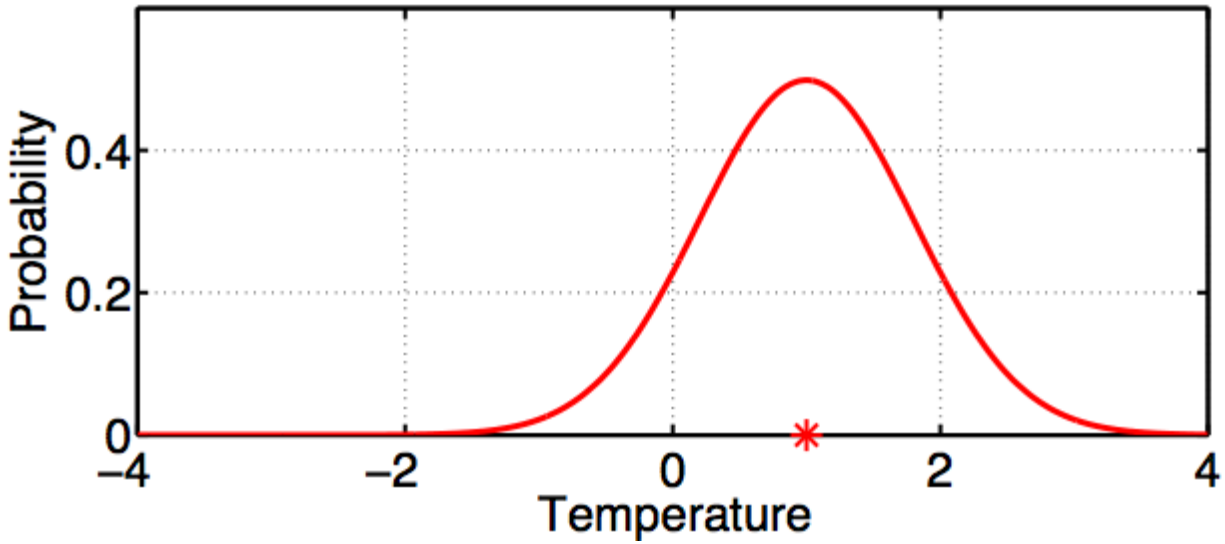
Thermometer outside measures 1° C.



Instrument builder says thermometer is unbiased with +/- 0.8° C gaussian error.

Example: Estimating the Temperature Outside

Thermometer outside measures 1° C.

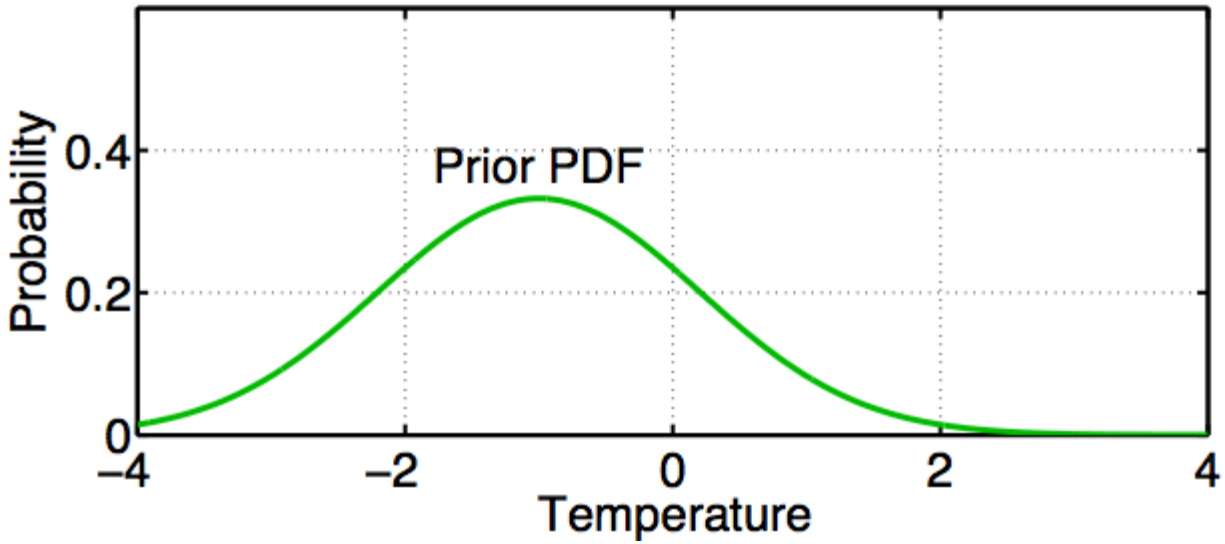


The red plot is $P(T \mid T_0)$;

probability of temperature *given* that T_o was observed.

Example: Estimating the Temperature Outside

We also have a prior estimate of temperature.



The green curve is $P(T \mid C)$;

probability of temperature given all available prior information C.

Example: Estimating the Temperature Outside

Prior information C can include:

1. Observations of things besides T;
2. Model forecast made using observations at earlier times;
3. *a priori* physical constraints ($T > -273.15^{o}$ C);
4. Climatological constraints (-30^{o} C $< T < 40^{o}$ C).

Combining the Prior Estimate and Observation

Bayes Theorem:

Likelihood: Probability that T_o is observed if T is true value and given prior information C.

$$P(T \mid T_o, C) = \frac{P(T_o \mid T, C)P(T \mid C)}{P(T_o \mid C)}$$

Prior

Posterior: Probability of T given observations and Prior. Also called **update** or **analysis**.

Combining the Prior Estimate and Observation

Rewrite Bayes as:

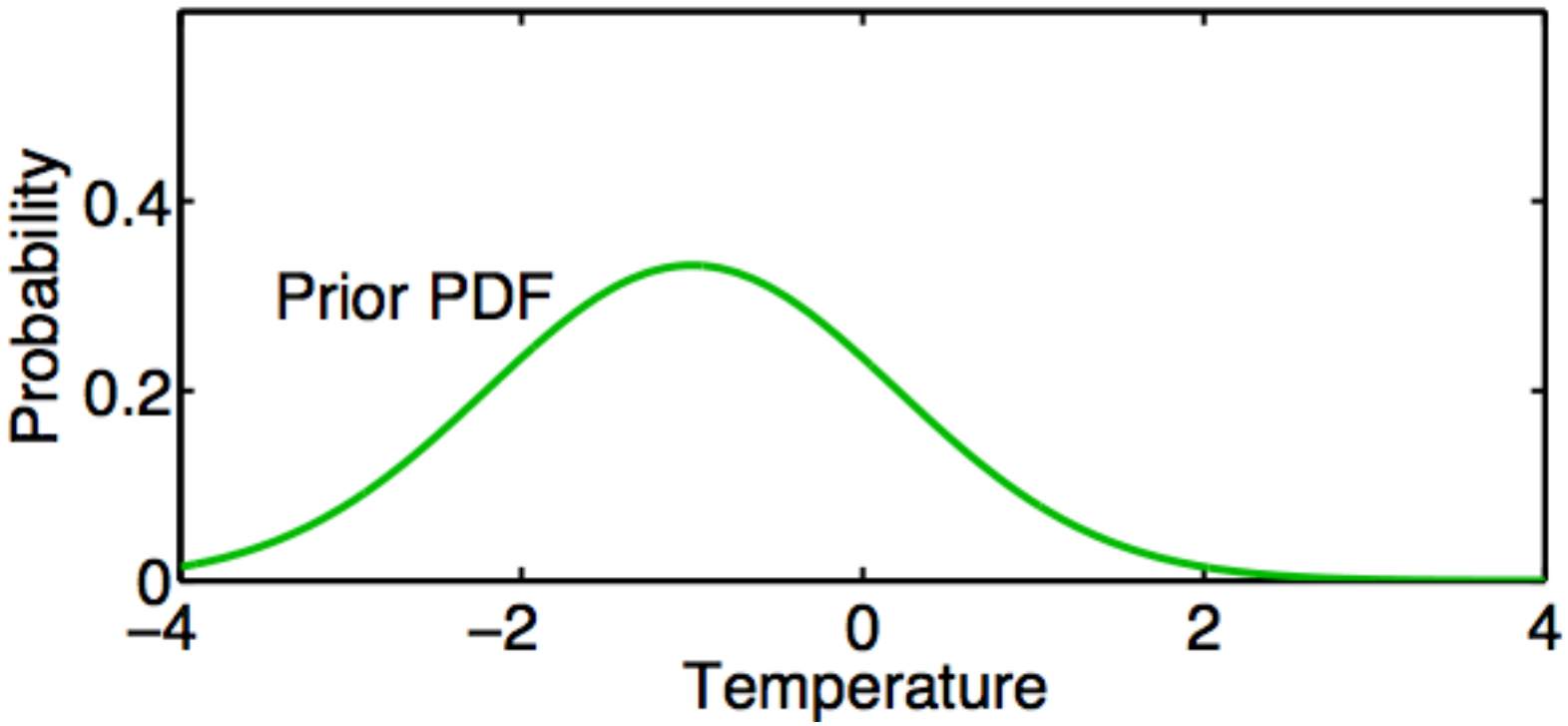
$$\frac{P(T_o \mid T,C)P(T \mid C)}{P(T_o \mid C)} = \frac{P(T_o \mid T,C)P(T \mid C)}{\int P(T_o \mid x)P(x \mid C)dx}$$

$$= \frac{P(T_o \mid T,C)P(T \mid C)}{normalization}$$

Denominator normalizes so Posterior is PDF.

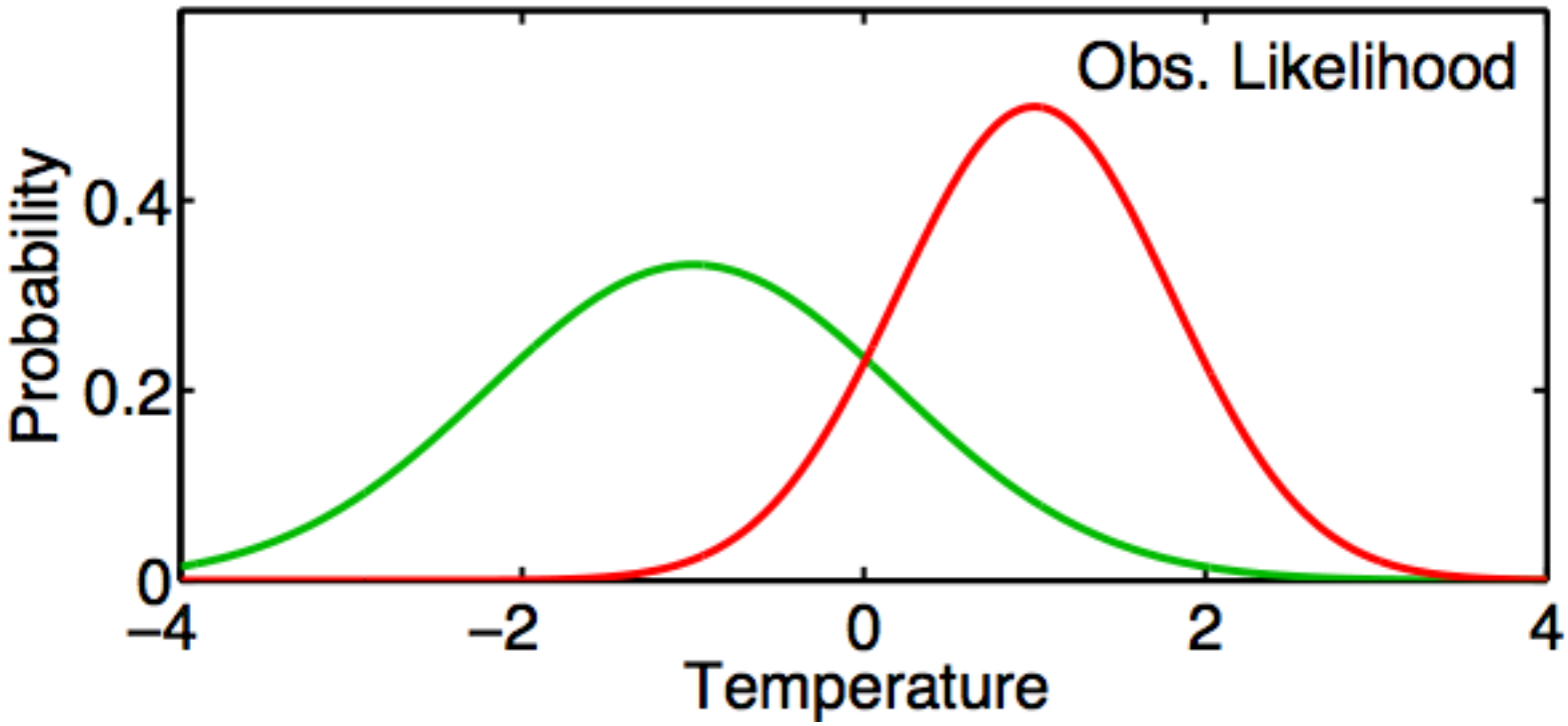
Combining the Prior Estimate and Observation

$$P(T \mid T_0, C) = \frac{P(T_0 \mid T, C)P(T \mid C)}{normalization}$$



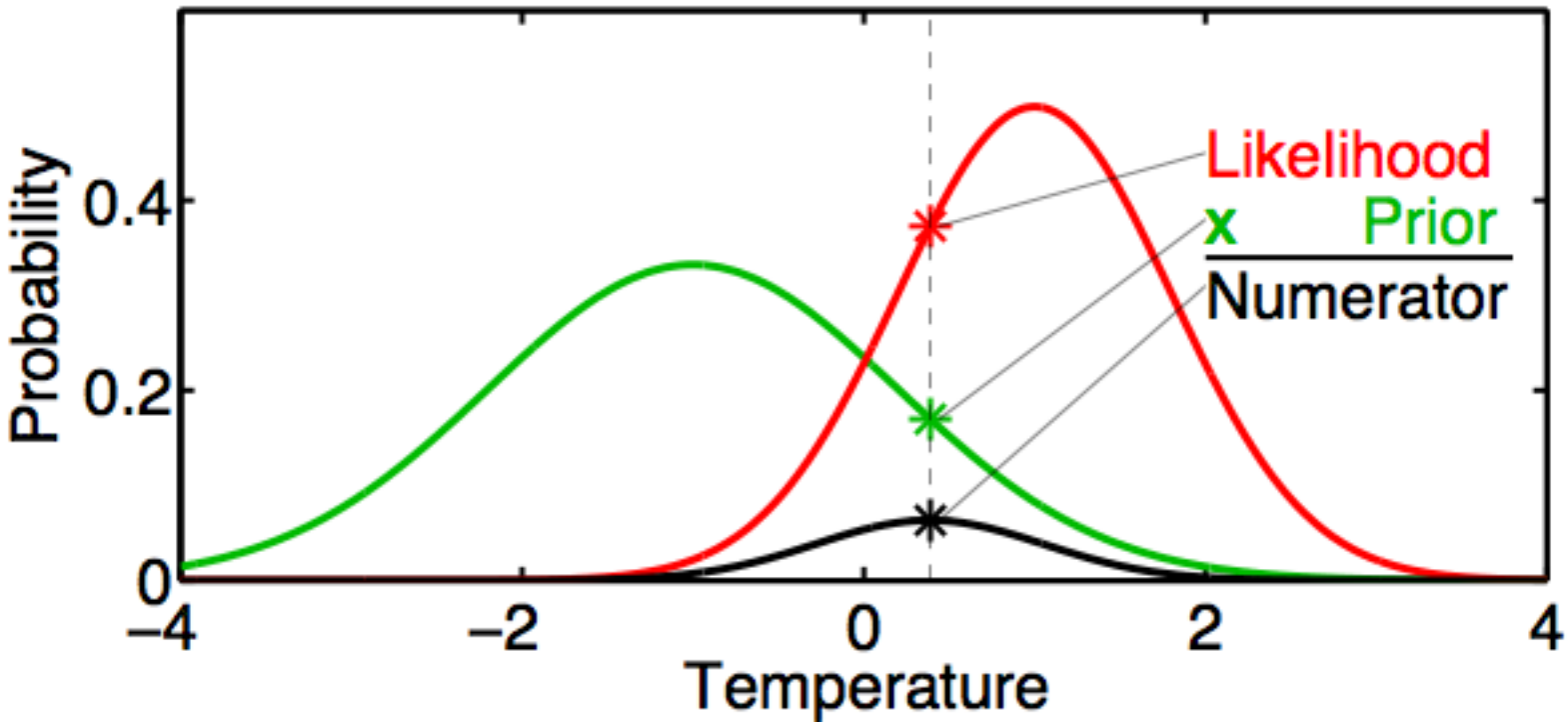
Combining the Prior Estimate and Observation

$$P(T \mid T_0, C) = \frac{P(T_0 \mid T, C) P(T \mid C)}{normalization}$$



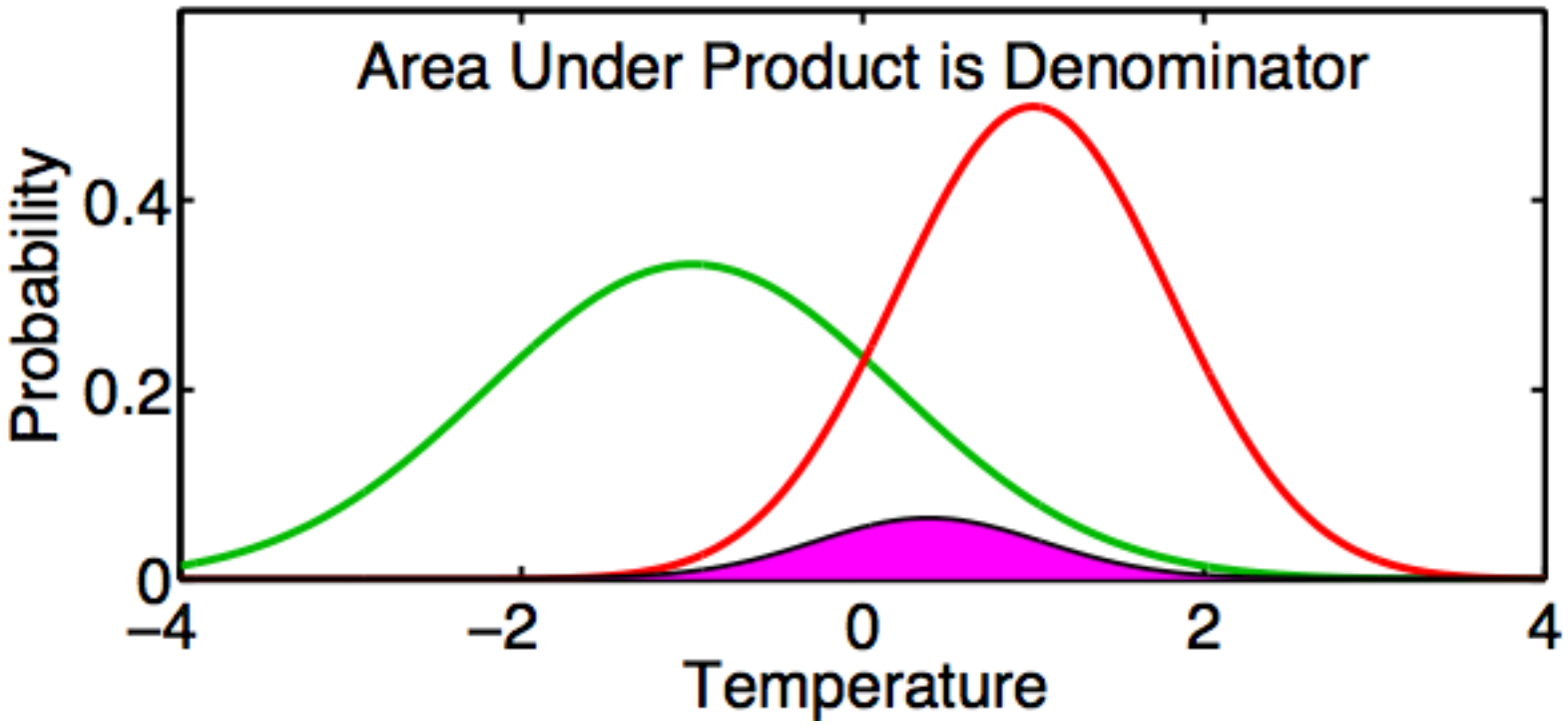
Combining the Prior Estimate and Observation

$$P(T \mid T_0, C) = \frac{P(T_0 \mid T, C)P(T \mid C)}{normalization}$$



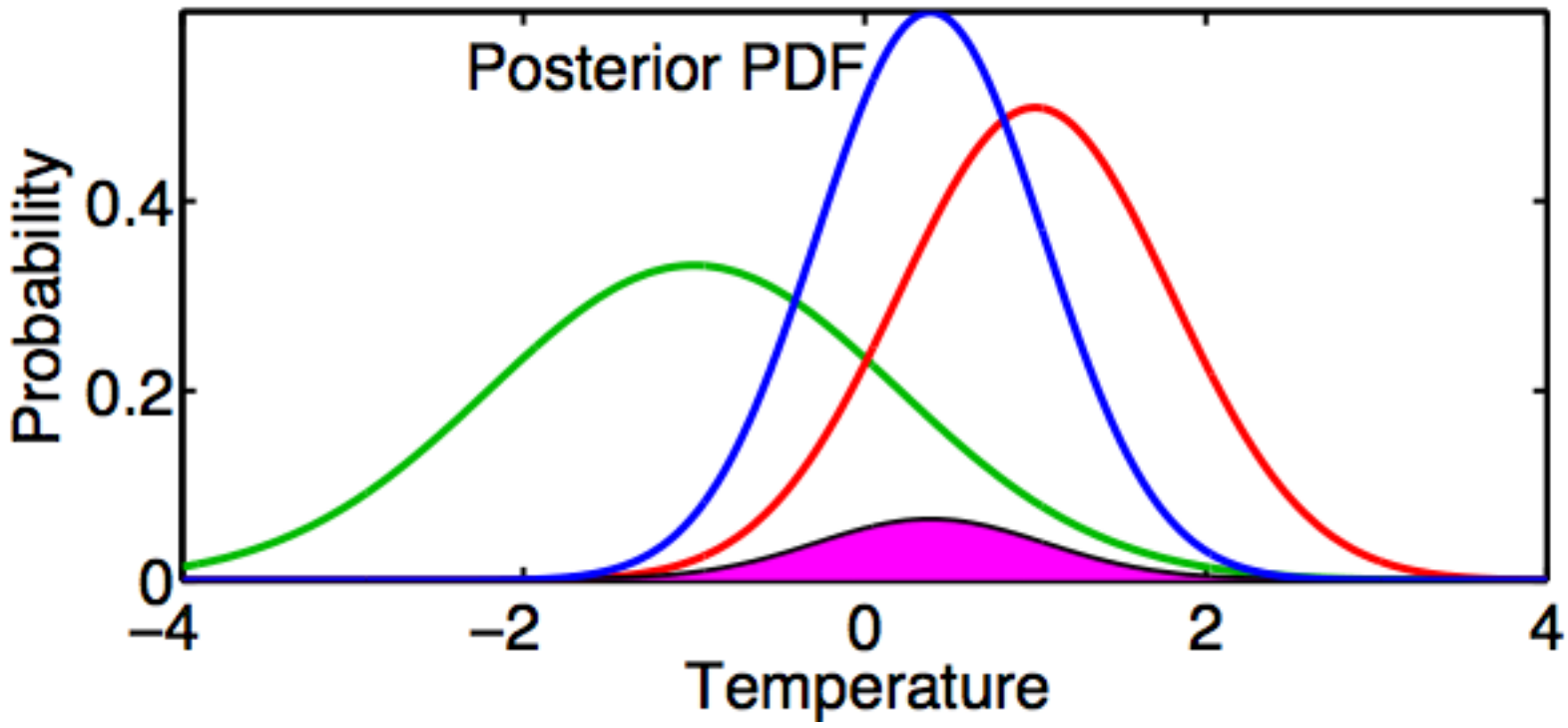
Combining the Prior Estimate and Observation

$$P(T \mid T_0, C) = \frac{P(T_0 \mid T, C)P(T \mid C)}{normalization}$$



Combining the Prior Estimate and Observation

$$P(T \mid T_0, C) = \frac{P(T_0 \mid T, C)P(T \mid C)}{normalization}$$



Consistent Color Scheme Throughout Tutorial

Green = Prior

Red = Observation

Blue = Posterior

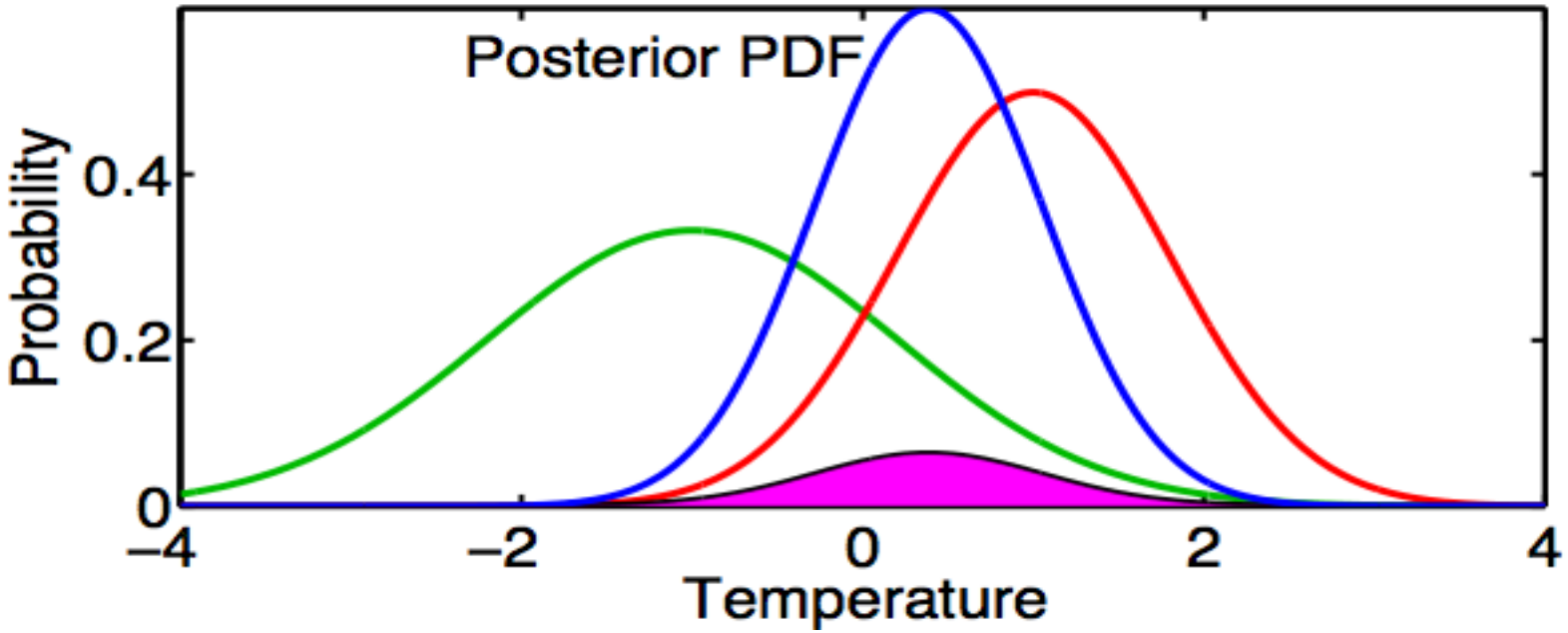
Black = Truth

(truth available only for ‘perfect model’ examples)

Combining the Prior Estimate and Observation

$$P(T \mid T_0, C) = \frac{P(T_0 \mid T, C) P(T \mid C)}{normalization}$$

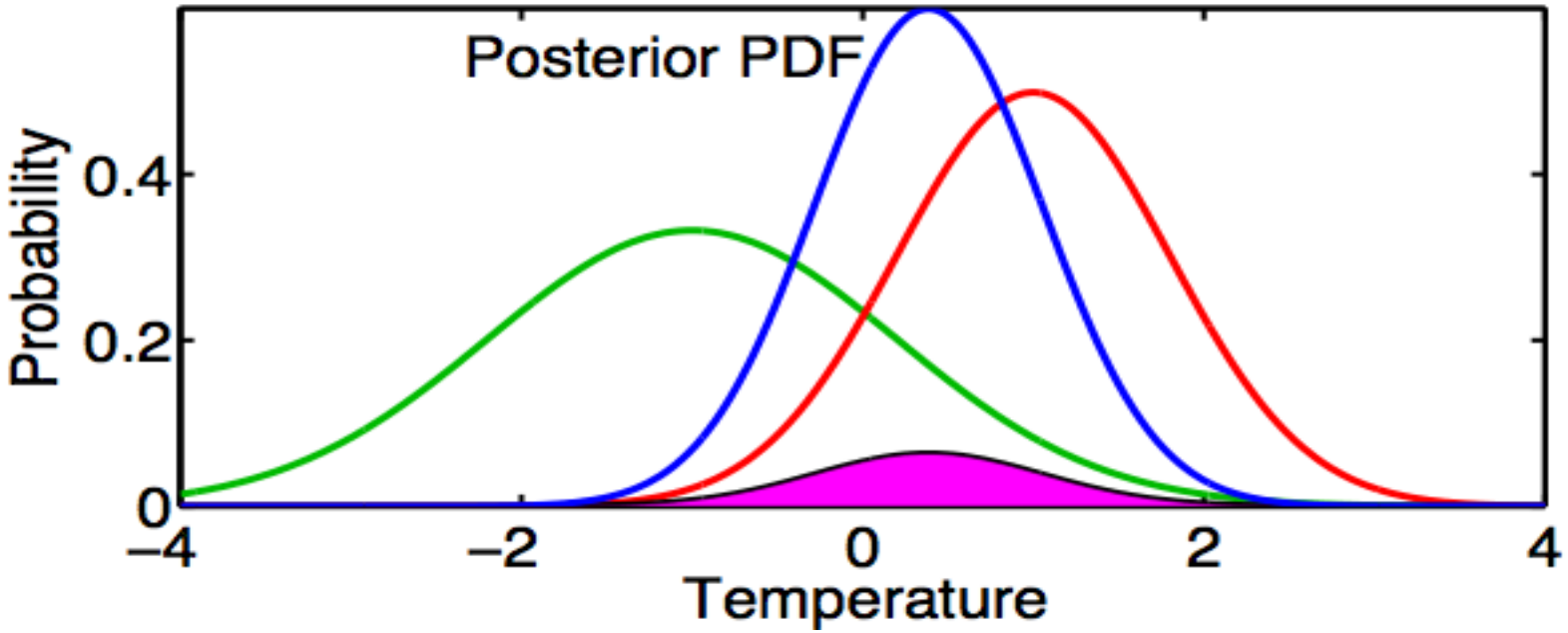
Generally no analytic solution for Posterior.



Combining the Prior Estimate and Observation

$$P(T \mid T_0, C) = \frac{P(T_0 \mid T, C) P(T \mid C)}{normalization}$$

Gaussian Prior and Likelihood -> Gaussian Posterior



Combining the Prior Estimate and Observation

For Gaussian prior and likelihood…

Prior $$P(T \mid C) = Normal(T_p, \sigma_p)$$

Likelihood $$P(T_o \mid T, C) = Normal(T_o, \sigma_o)$$

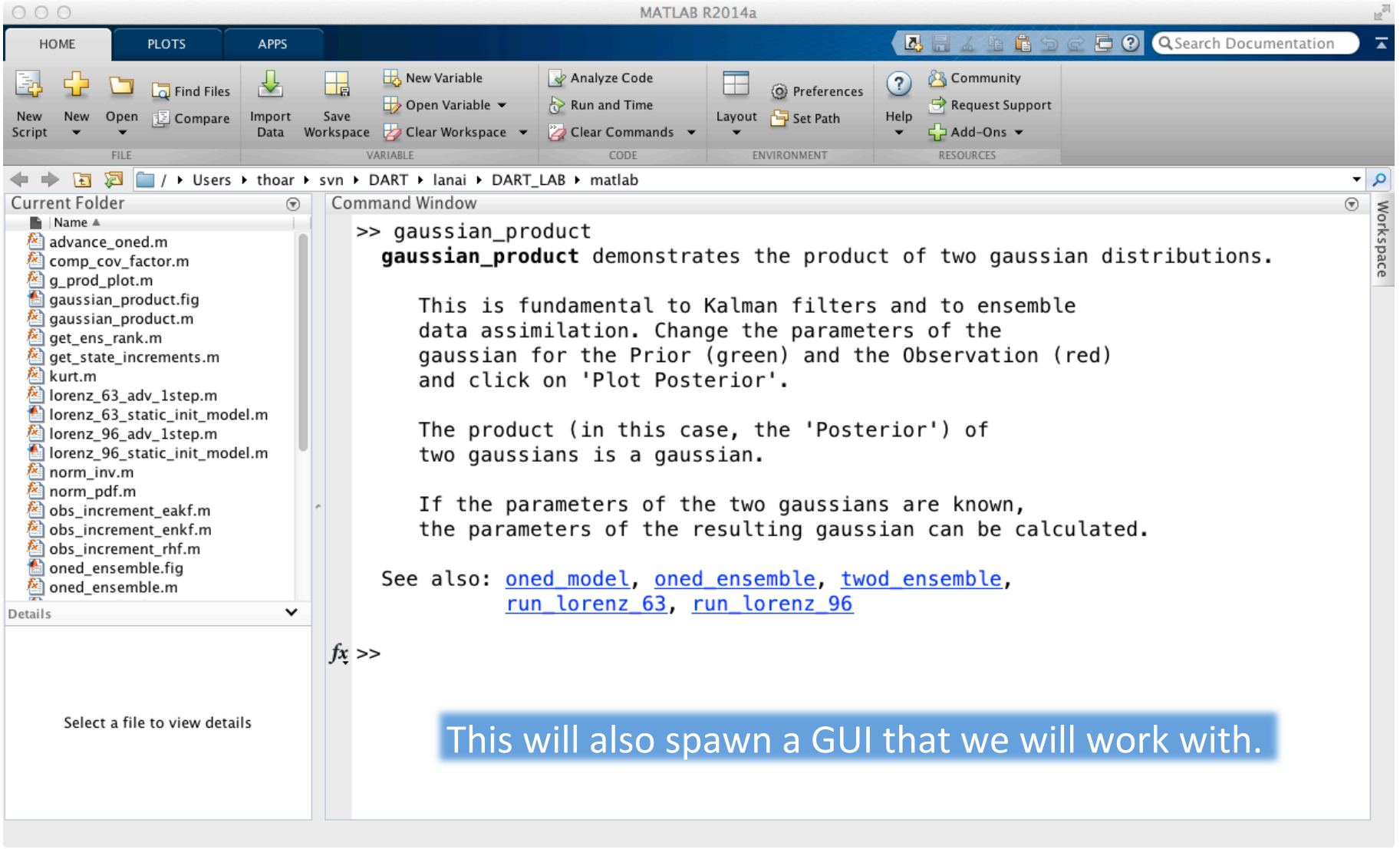
Then, Posterior $$P(T \mid T_o, C) = Normal(T_u, \sigma_u)$$

With

$$\sigma_u = \sqrt{\left(\sigma_p^{-2} + \sigma_o^{-2}\right)^{-1}}$$

$$T_u = \sigma_u^2 \left[\sigma_p^{-2} T_p + \sigma_o^{-2} T_o\right]$$

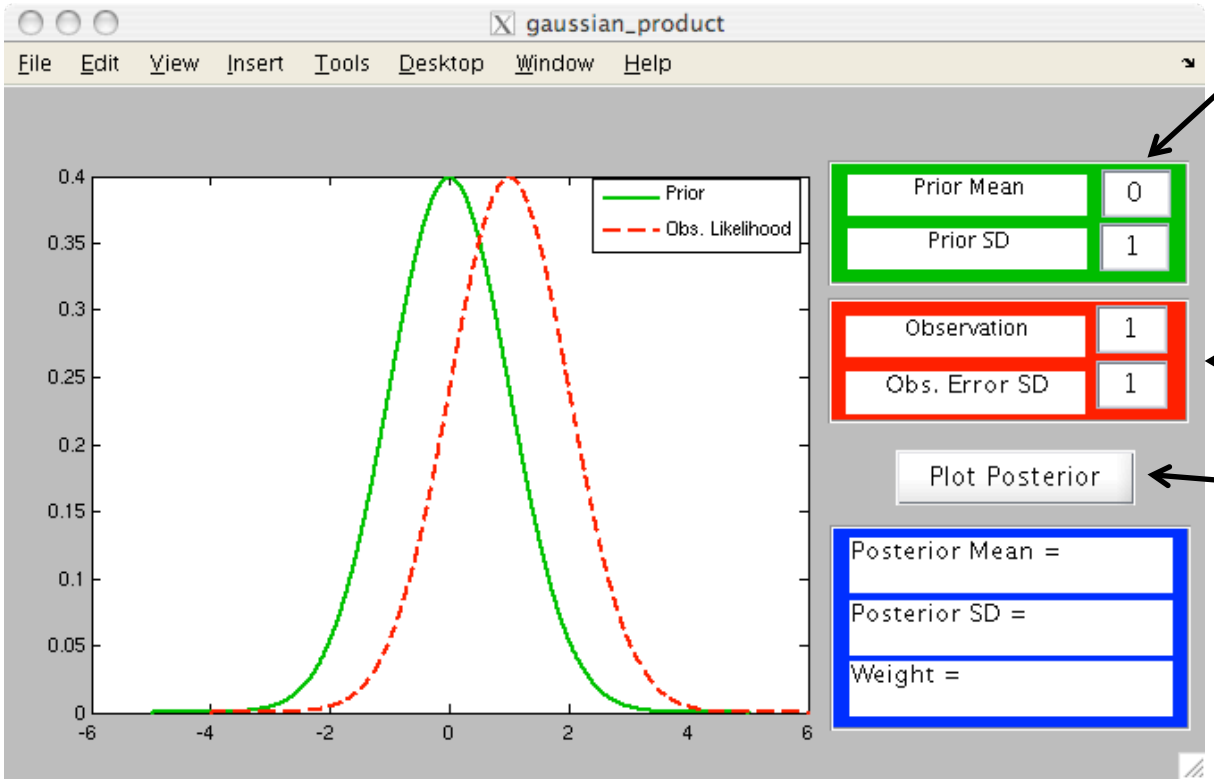
Matlab Hands-on: **gaussian_product**



This will also spawn a GUI that we will work with.

Matlab Hands-on: gaussian_product

Purpose: Explore the gaussian posterior that results from taking the product of a gaussian prior and a gaussian likelihood.



1) Set Prior Mean and Standard Deviation.

2) Set Observation Mean and Observation Error Standard Deviation.

3) Select Plot Posterior to Update the items in blue.

Matlab Hands-on: gaussian_product

Explore!

- Change the mean value of the prior and the observation.
- Change the standard deviation of the prior.
- What is always true for the mean of the posterior?
- What is always true for the standard deviation of the posterior?

The One-Dimensional Kalman Filter

1. Suppose we have a linear forecast model L
   A. If temperature at time $t_1 = T_{1,}$ then the temperature at $t_2 = t_1 + \Delta t$ is $T_2 = L(T_1)$
   B. Example: $T_2 = T_1 + \Delta t T_1$

The One-Dimensional Kalman Filter

1. Suppose we have a linear forecast model L.

   A. If temperature at time $t_1 = T_{1,}$ then the temperature at $t_2 = t_1 + \Delta t$ is $T_2 = L(T_1)$

   B. Example: $T_2 = T_1 + \Delta t T_1$

2. If posterior estimate at time t_1 is *Normal*$(T_{u,1}, \sigma_{u,1})$ then the prior at t_2 is *Normal*$(T_{p,2}, \sigma_{p,2})$.

   $$T_{p,2} = T_{u,1}, + \Delta t T_{u,1}$$

   $$\sigma_{p,2} = (\Delta t + 1)\, \sigma_{u,1}$$

The One-Dimensional Kalman Filter

1. Suppose we have a linear forecast model L.
   A. If temperature at time $t_1 = T_{1,}$ then the temperature at $t_2 = t_1 + \Delta t$ is $T_2 = L(T_1)$
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3. Given an observation at t_2 with distribution *Normal*(t_0, σ_0) the likelihood is also *Normal*(t_0, σ_0).

The One-Dimensional Kalman Filter

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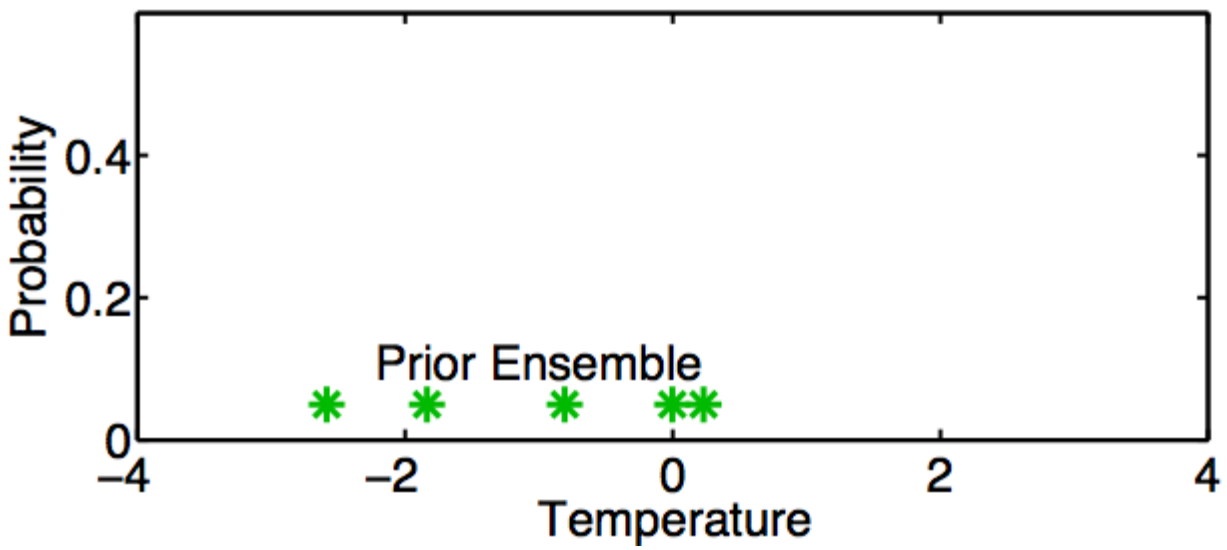
3. Given an observation at t_2 with distribution *Normal*(t_0, σ_0) the likelihood is also *Normal*(t_0, σ_0).
4. The posterior at t_2 is *Normal*$(T_{u,2}, \sigma_{u,2})$ where $T_{u,2}$ and $\sigma_{u,2}$ come from page 19.

A One-Dimensional Ensemble Kalman Filter

Represent a prior pdf by a sample (ensemble) of N values:



A One-Dimensional Ensemble Kalman Filter

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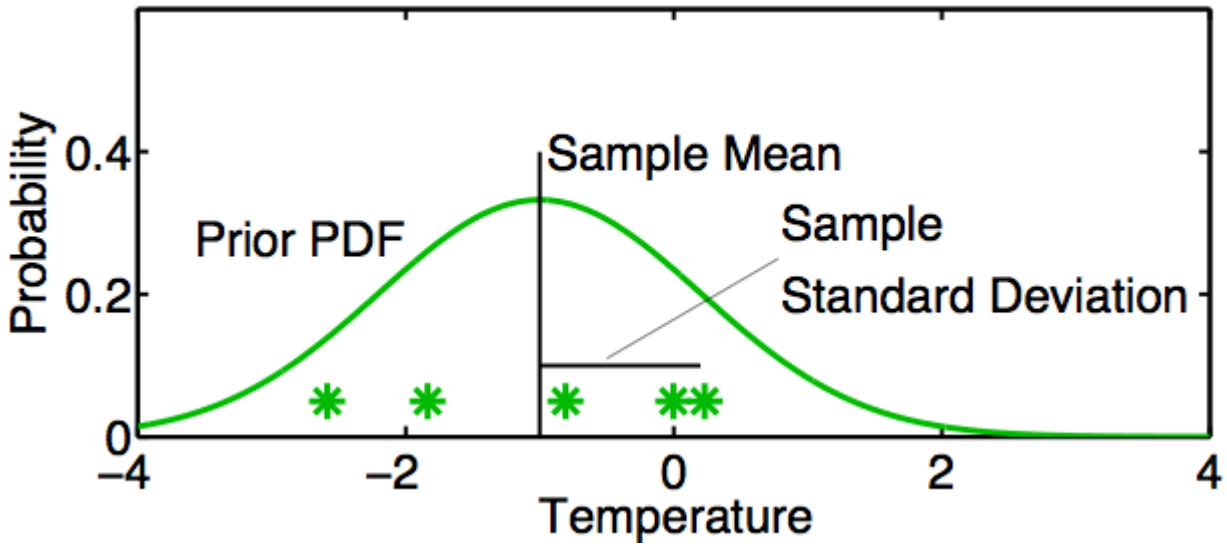


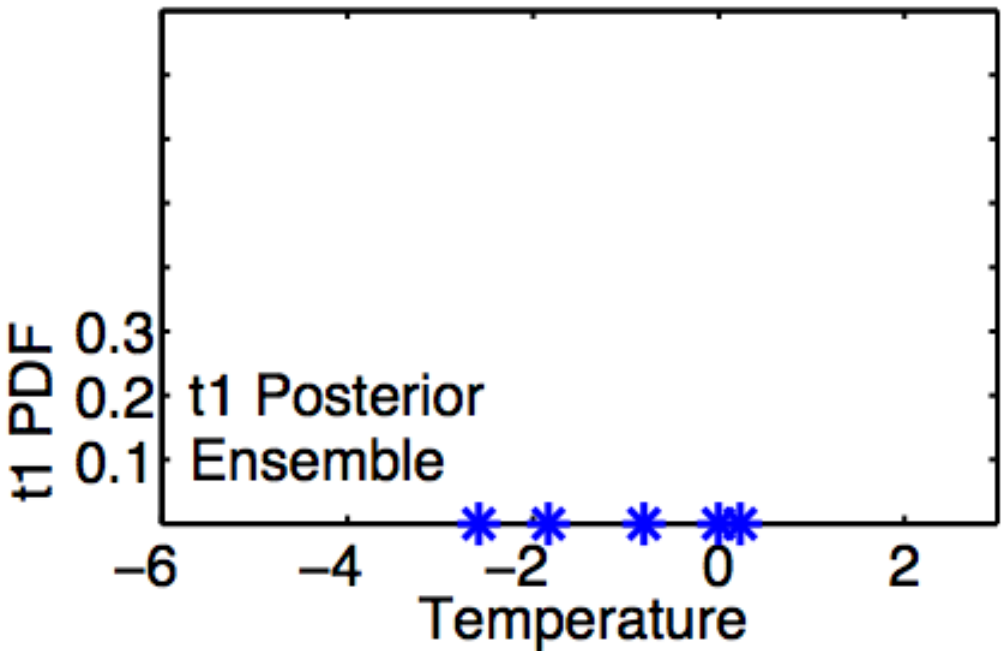
Use sample mean $\overline{T} = \sum_{n=1}^{N} T_n / N$

and sample standard deviation $\sigma_T = \sqrt{\sum_{n=1}^{N} \left(T_n - \overline{T}\right)^2 / \left(N - 1\right)}$

to determine a corresponding continuous distribution $Normal\left(\overline{T}, \sigma_T\right)$

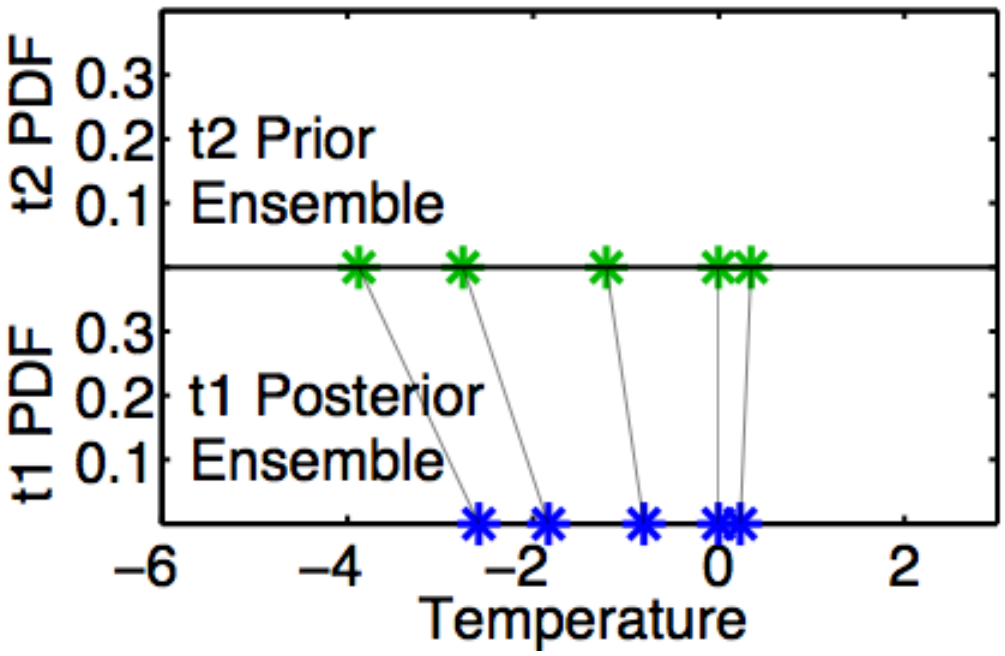
A One-Dimensional Ensemble Kalman Filter: Model Advance

If posterior ensemble at time t_1 is $T_{1,n}$, $n = 1, \ldots, N$



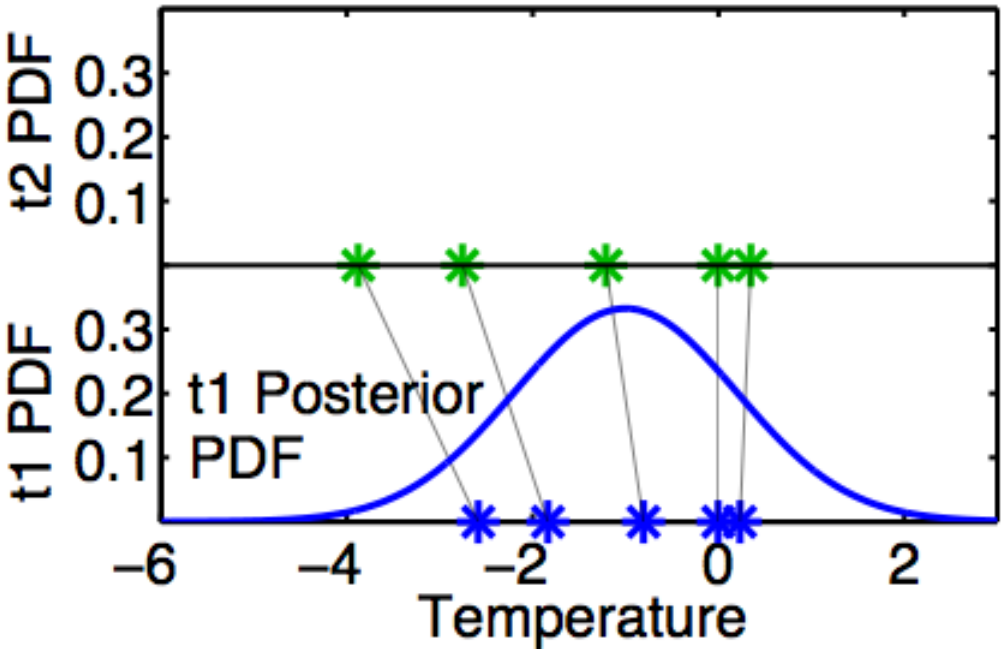
A One-Dimensional Ensemble Kalman Filter: Model Advance

If posterior ensemble at time t_1 is $T_{1,n}$, $n = 1, \ldots, N$
advance each member to time t_2 with model, $T_{2,n} = L(T_{1,n})$ $n = 1, \ldots, N$.



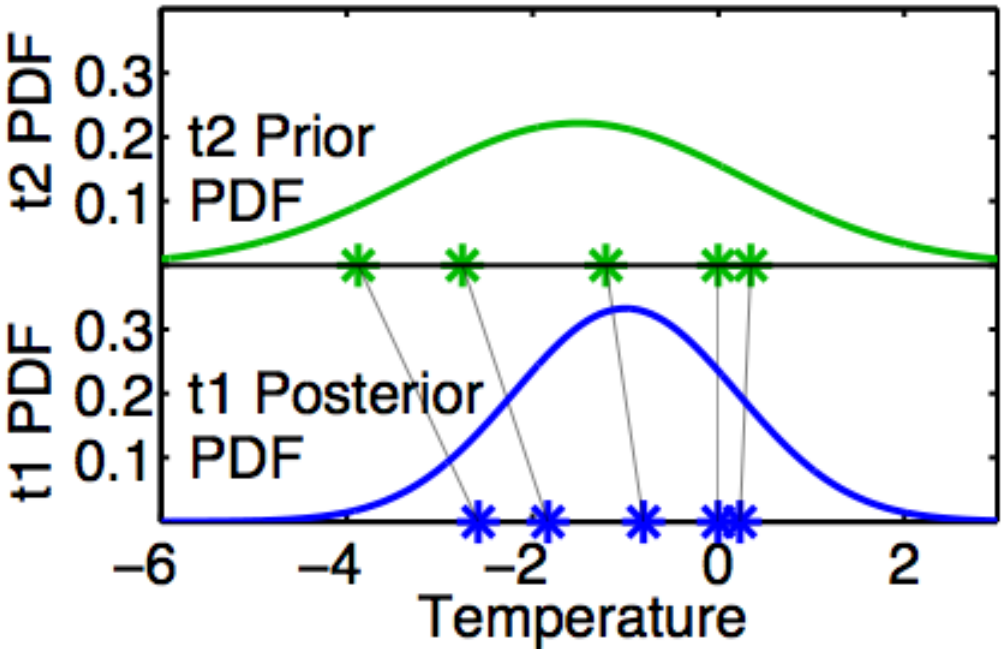
A One-Dimensional Ensemble Kalman Filter: Model Advance

Same as advancing continuous pdf at time t_1 …

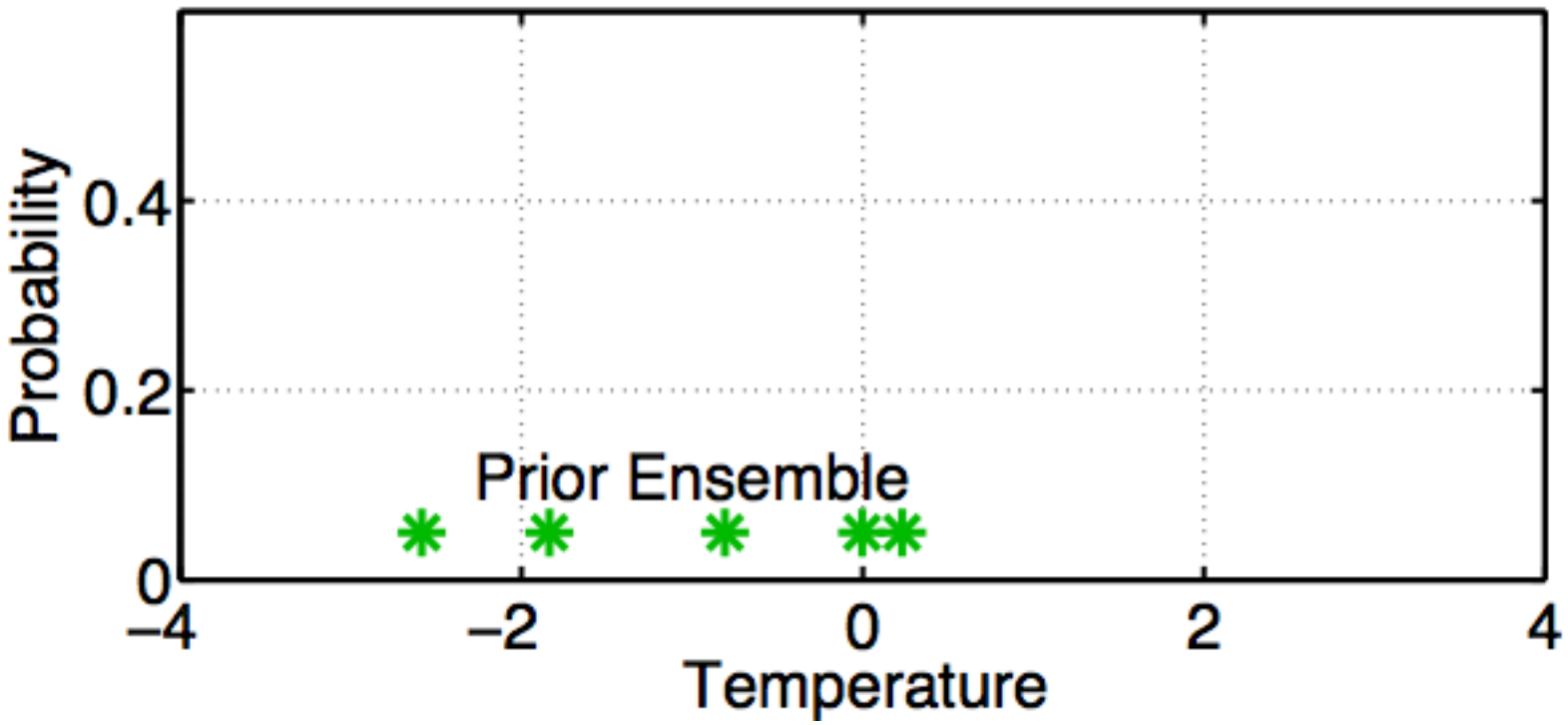


A One-Dimensional Ensemble Kalman Filter: Model Advance

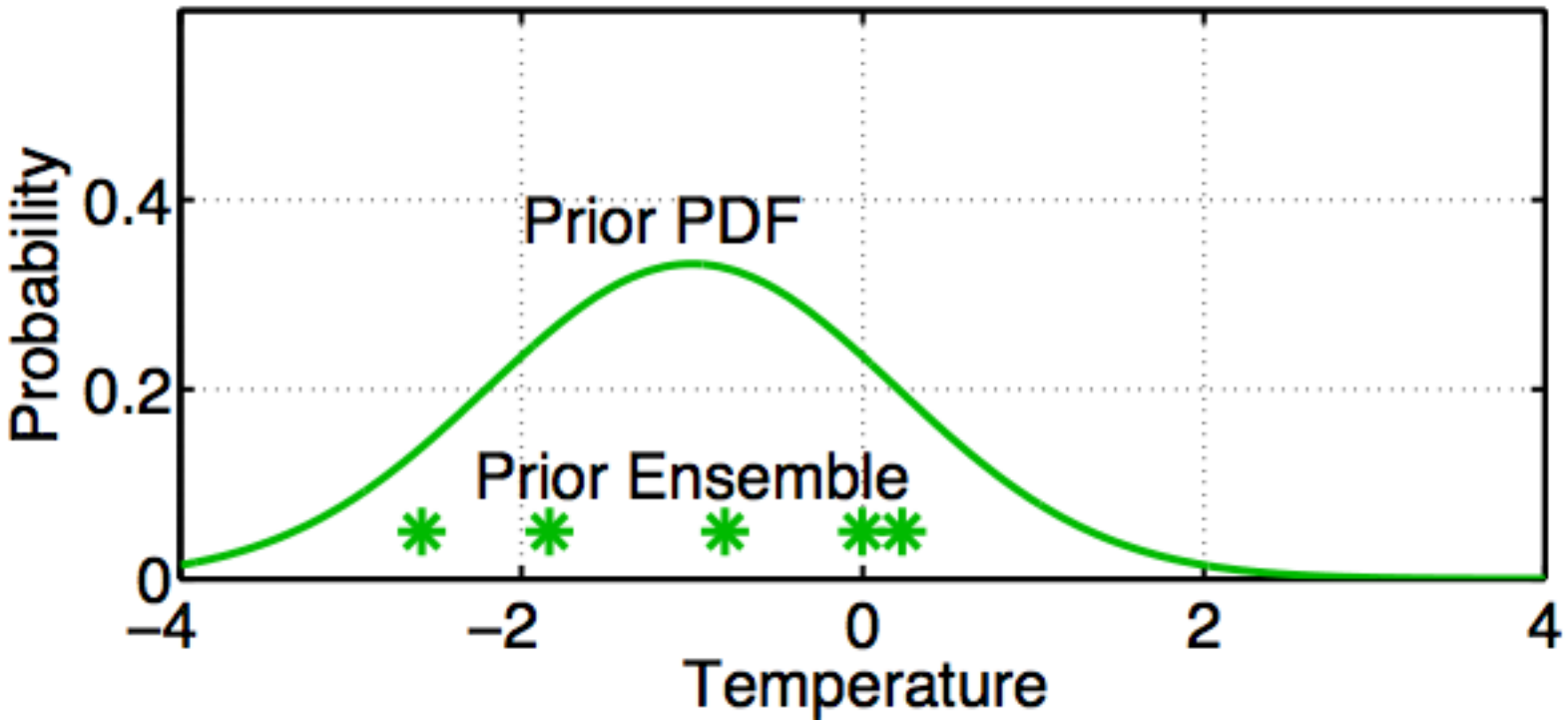
Same as advancing continuous pdf at time t_1 …
to time t_2 with model L.



A One-Dimensional Ensemble Kalman Filter: Assimilating an Observation

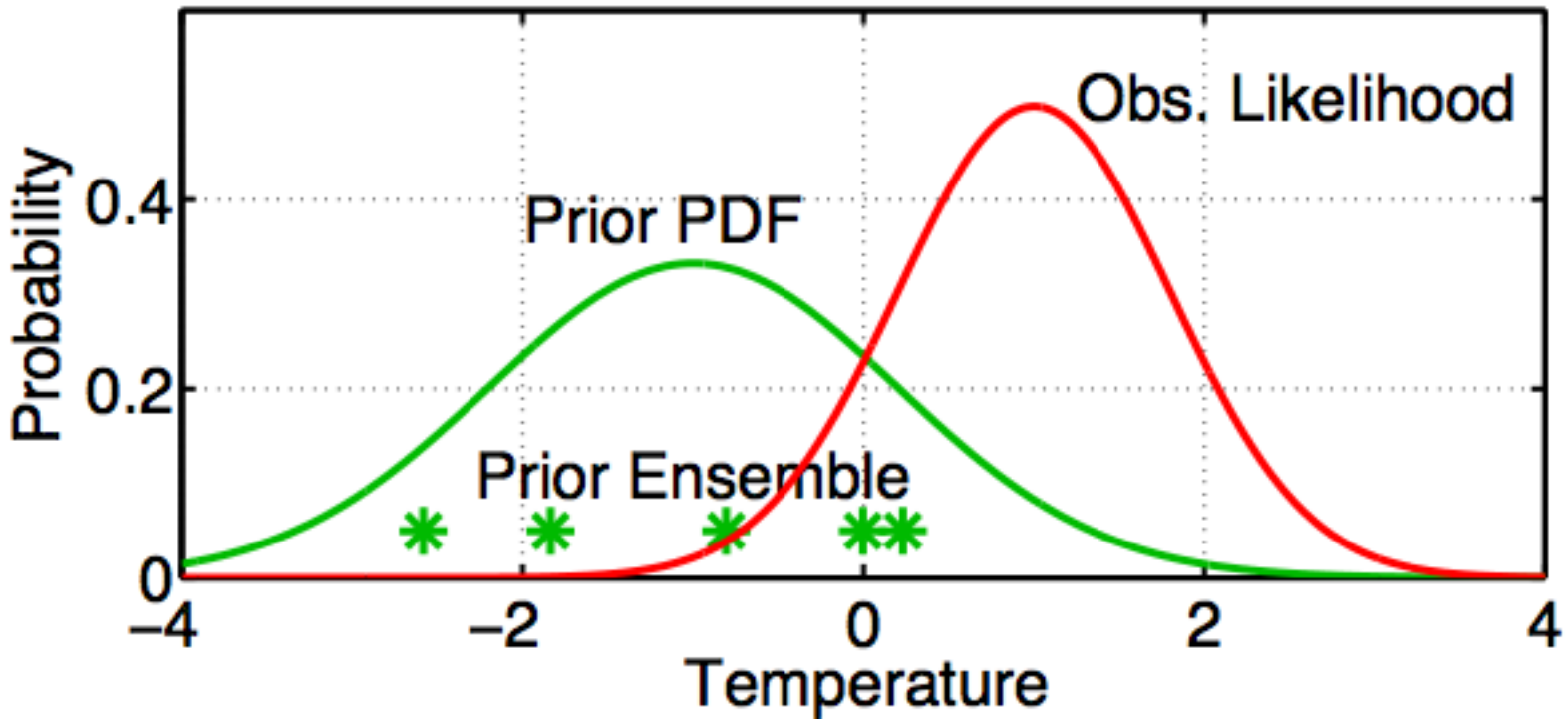


A One-Dimensional Ensemble Kalman Filter: Assimilating an Observation



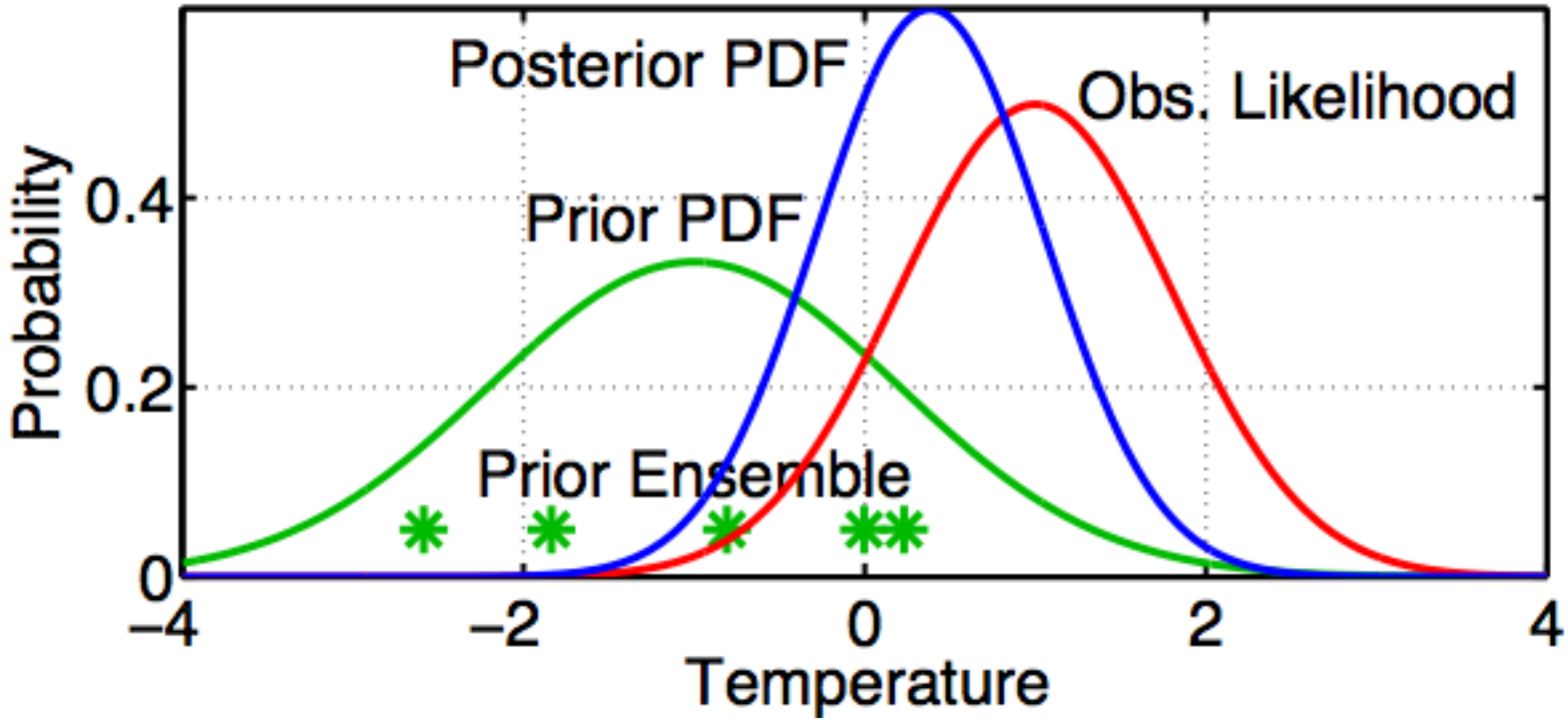
Fit a Gaussian to the sample.

A One-Dimensional Ensemble Kalman Filter: Assimilating an Observation



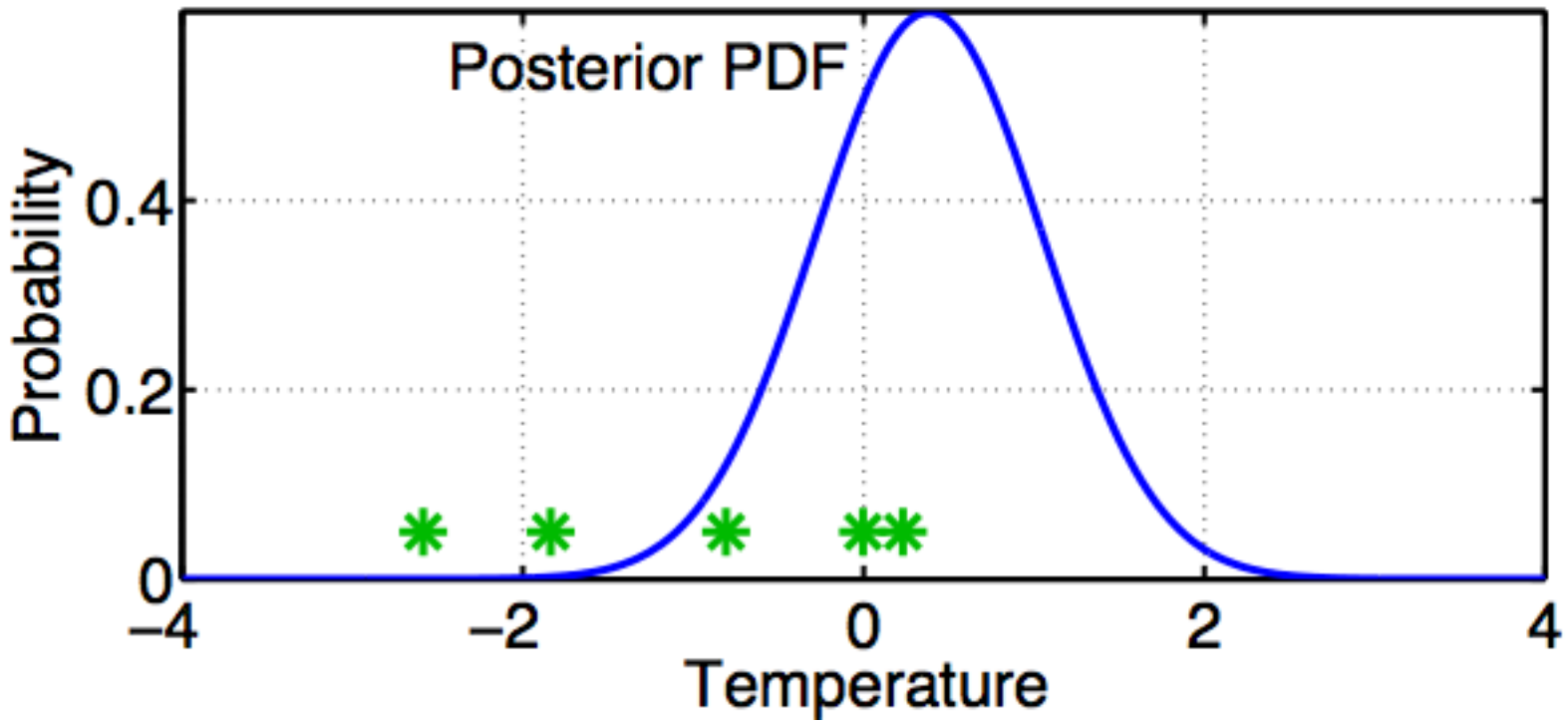
Get the observation likelihood.

A One-Dimensional Ensemble Kalman Filter: Assimilating an Observation



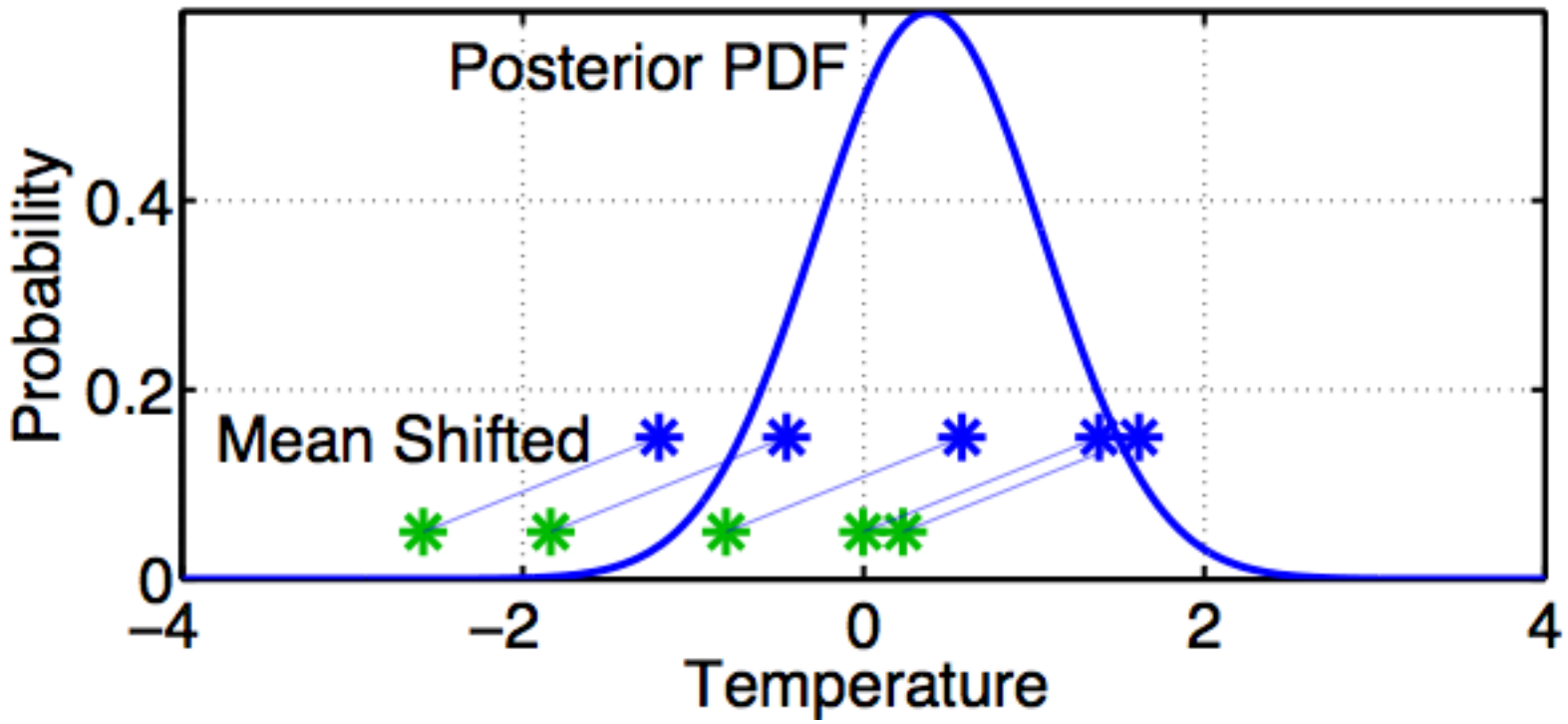
Compute the continuous posterior PDF.

A One-Dimensional Ensemble Kalman Filter: Assimilating an Observation



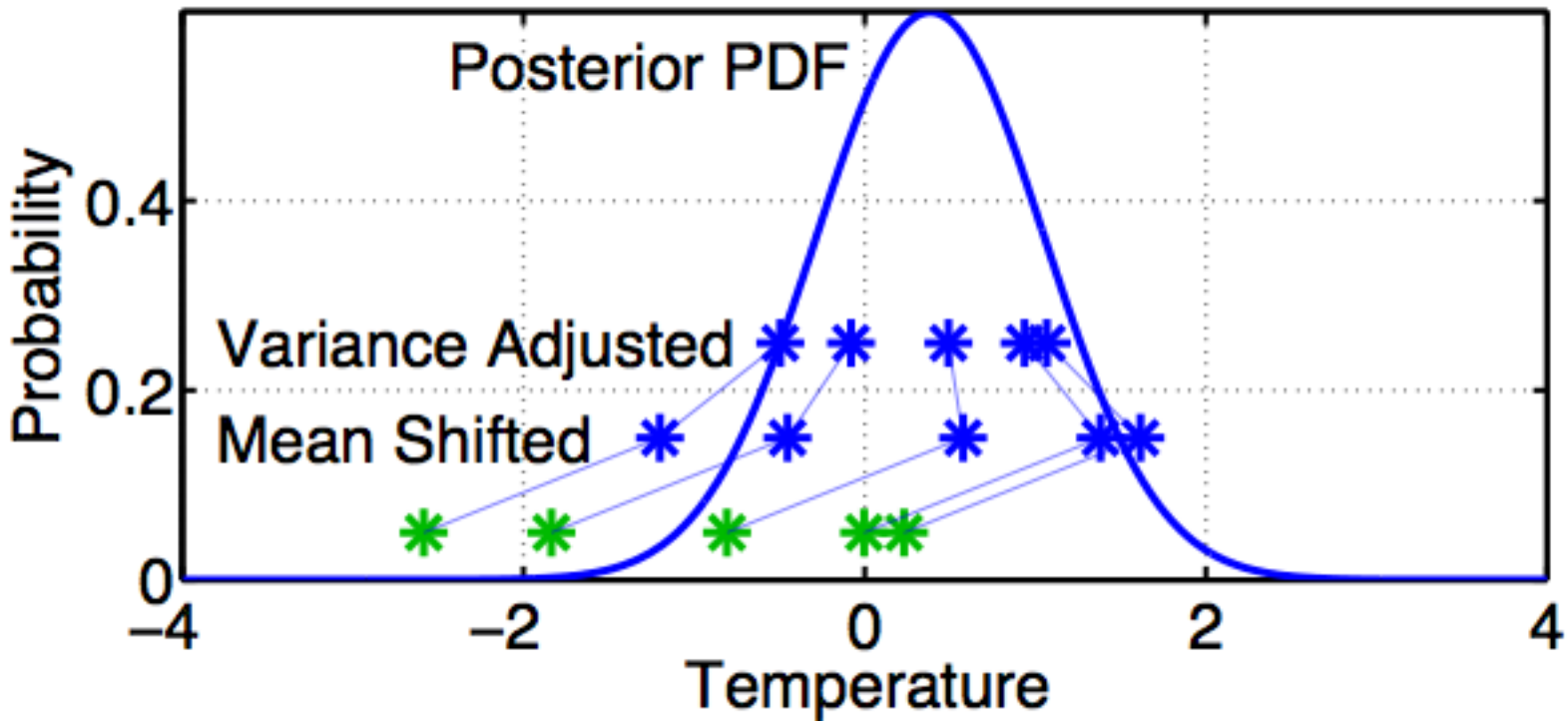
Use a deterministic algorithm to ‘adjust’ the ensemble.

A One-Dimensional Ensemble Kalman Filter: Assimilating an Observation



First, ‘shift’ the ensemble to have the exact mean of the posterior.

A One-Dimensional Ensemble Kalman Filter: Assimilating an Observation

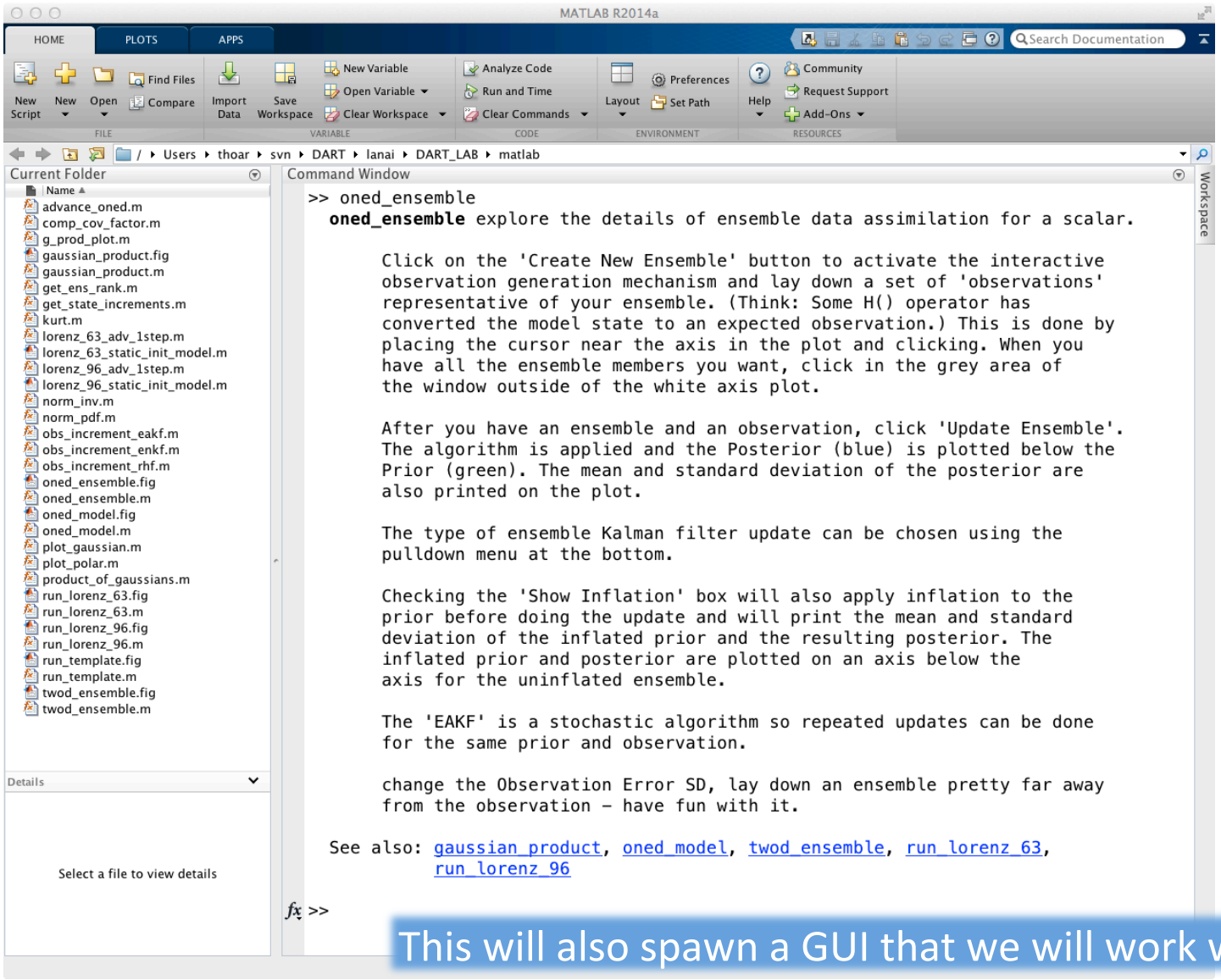


First, 'shift' the ensemble to have the exact mean of the posterior.

Second, linearly contract to have the exact variance of the posterior.

Sample statistics are identical to Kalman filter.

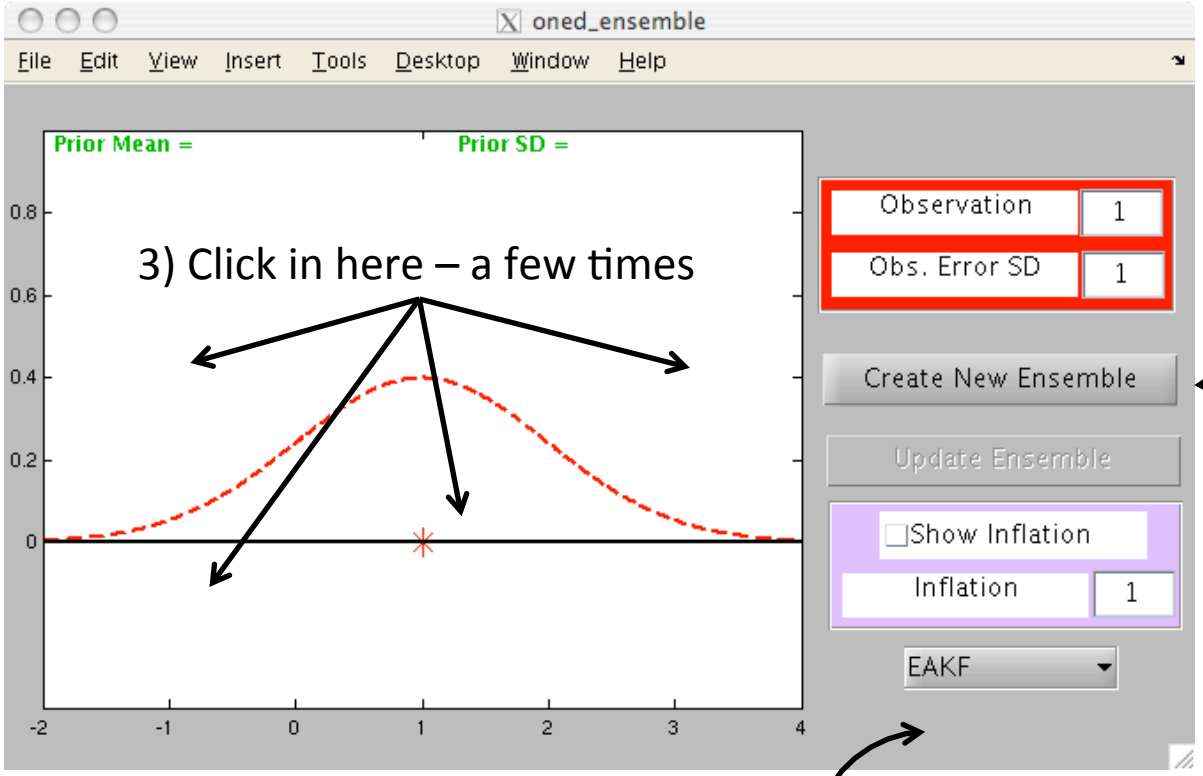
Matlab Hands-On: **oned_ensemble**



This will also spawn a GUI that we will work with.

Matlab Hands-On: oned_ensemble

Purpose: Explore how ensemble filters update a prior ensemble.



1) change these if you want to.

2) Click on Create New Ensemble

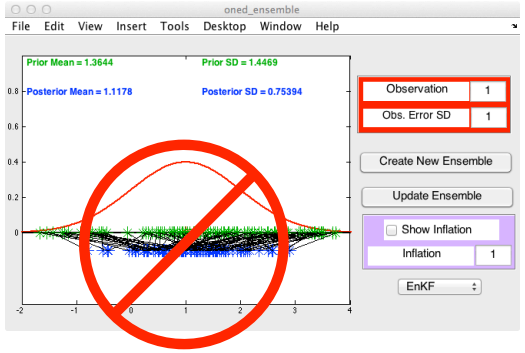
5) Click on Update Ensemble

4) Click outside the axis on the gray (anywhere) to finish defining the ensemble.

Ignore the Inflation and EAKF menus for now.

Matlab Hands-On: oned_ensemble

Explorations:



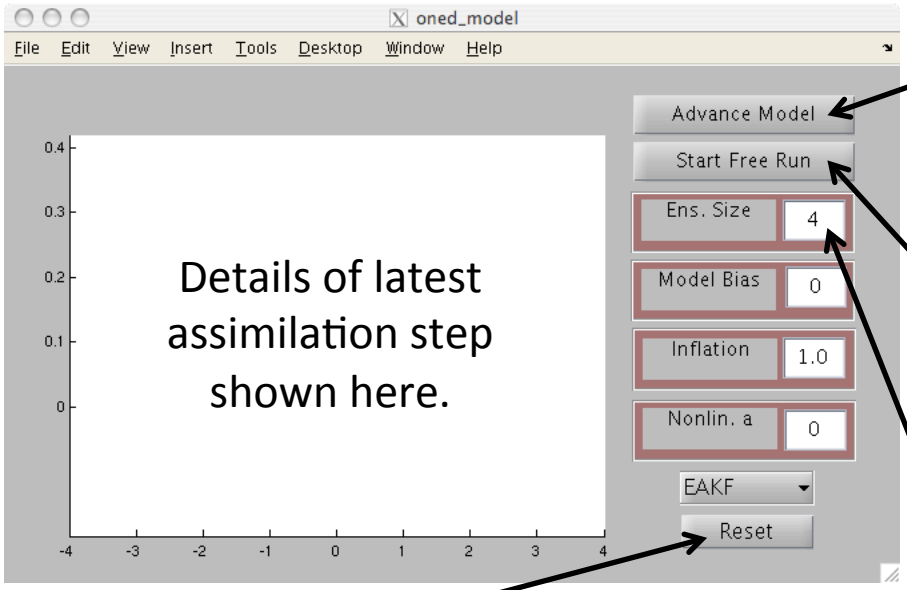
Too Many!

1. Keep your ensembles small, less than 10, for easy viewing.
2. Create a nearly uniformly spaced ensemble. Examine the update.
3. What happens with an ensemble that is confined to one side of the likelihood?
4. What happens with a bimodal ensemble (two clusters of members on either side)?
5. What happens with a single outlier in the ensemble?

Matlab Hands-On: **oned_model**

Purpose:

- Explore behavior of a complete 1-D ensemble filter for a linear system.
- Look at the behavior of different ensemble sizes.



(1)

Advance Model

Assimilate Obs

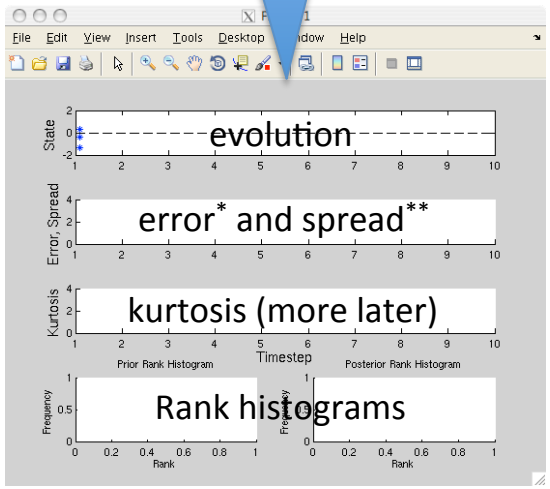
(2)

Start Free Run

Stop Free Run

ensemble size can be changed

This window also pops up. More later.



Reset restarts the exercise.

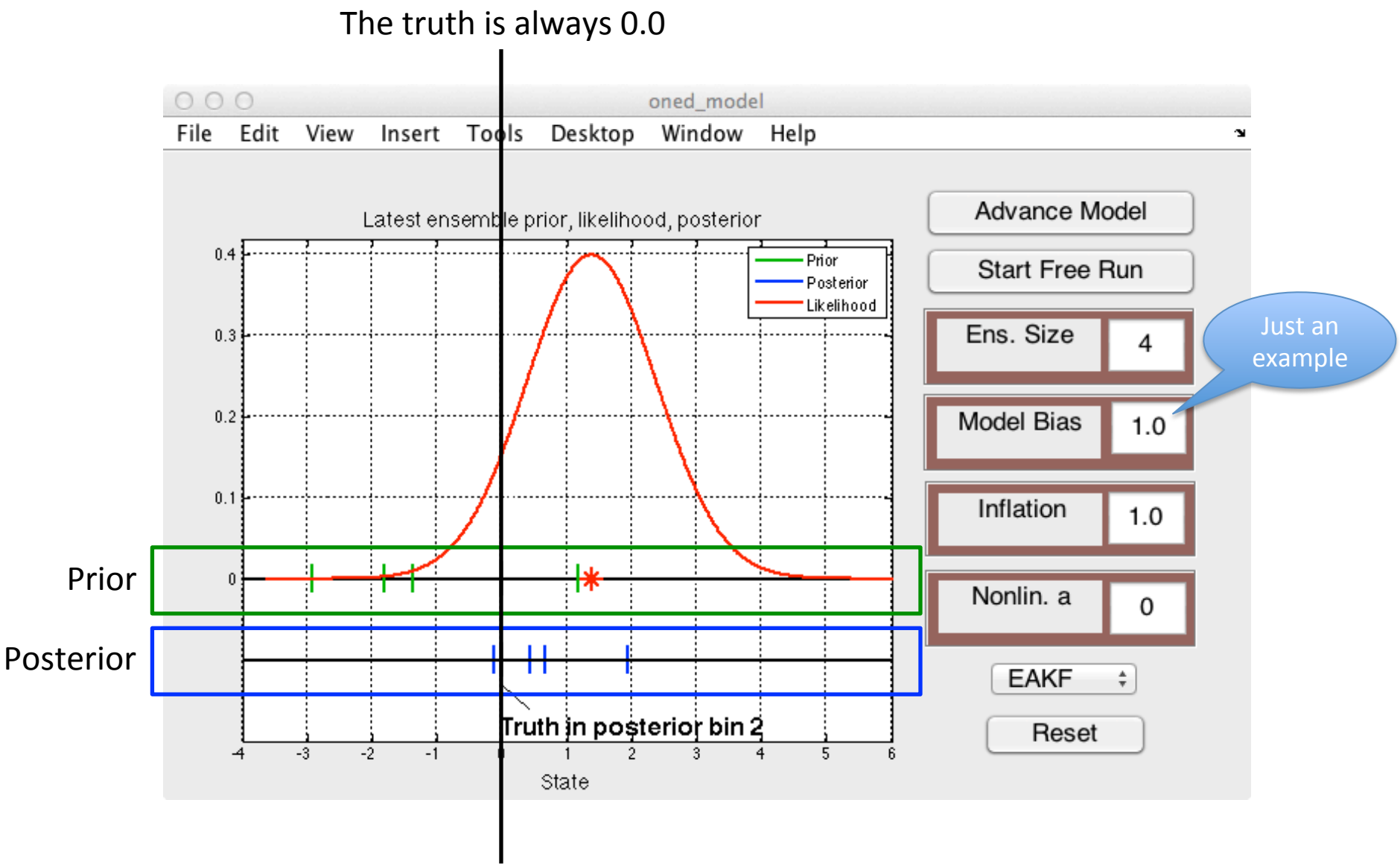
Note:

This uses the simple linear model $dx/dt = x$.

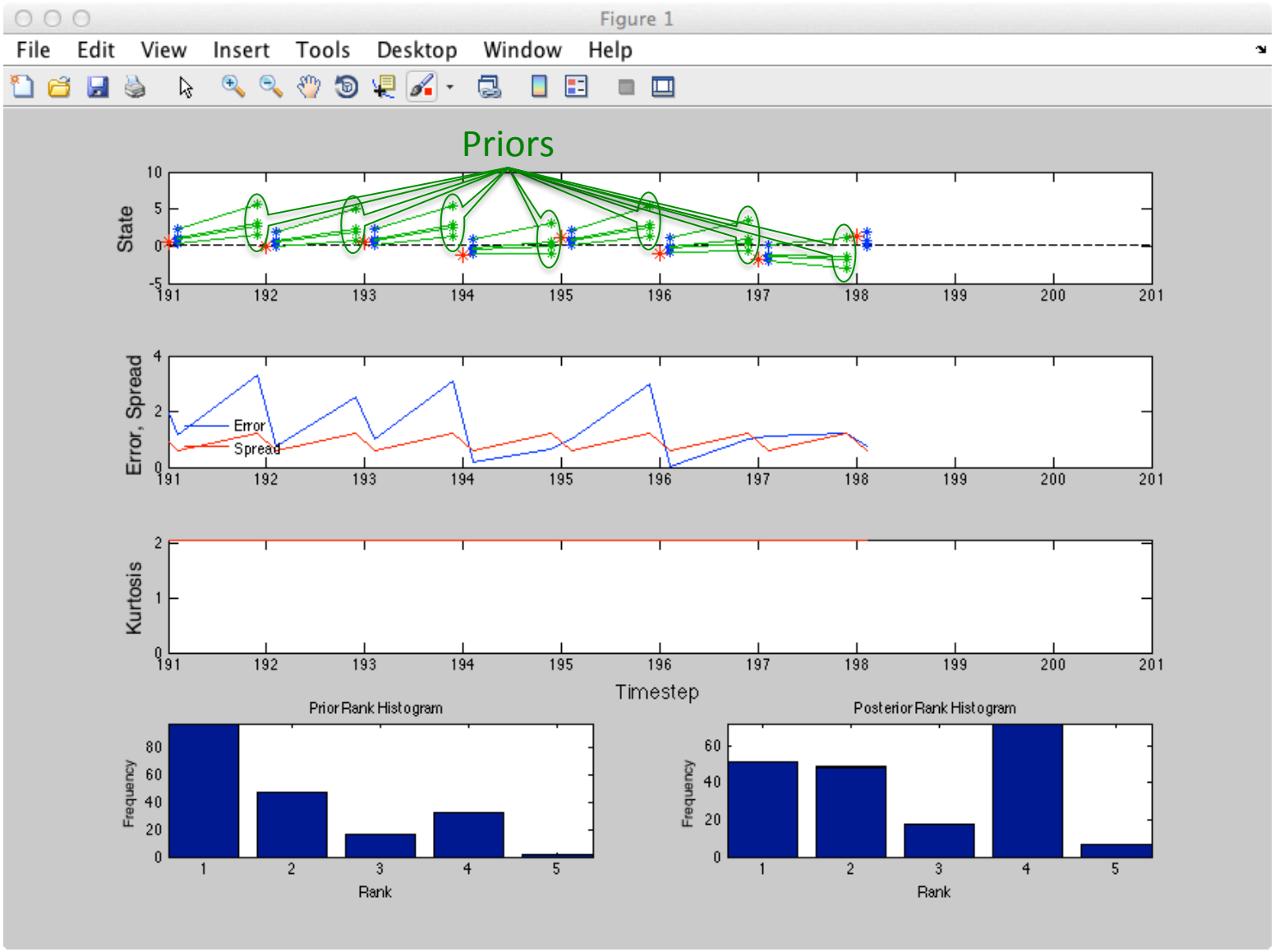
The 'truth' is always 0.

Observation noise is a draw from $N(0,1)$.

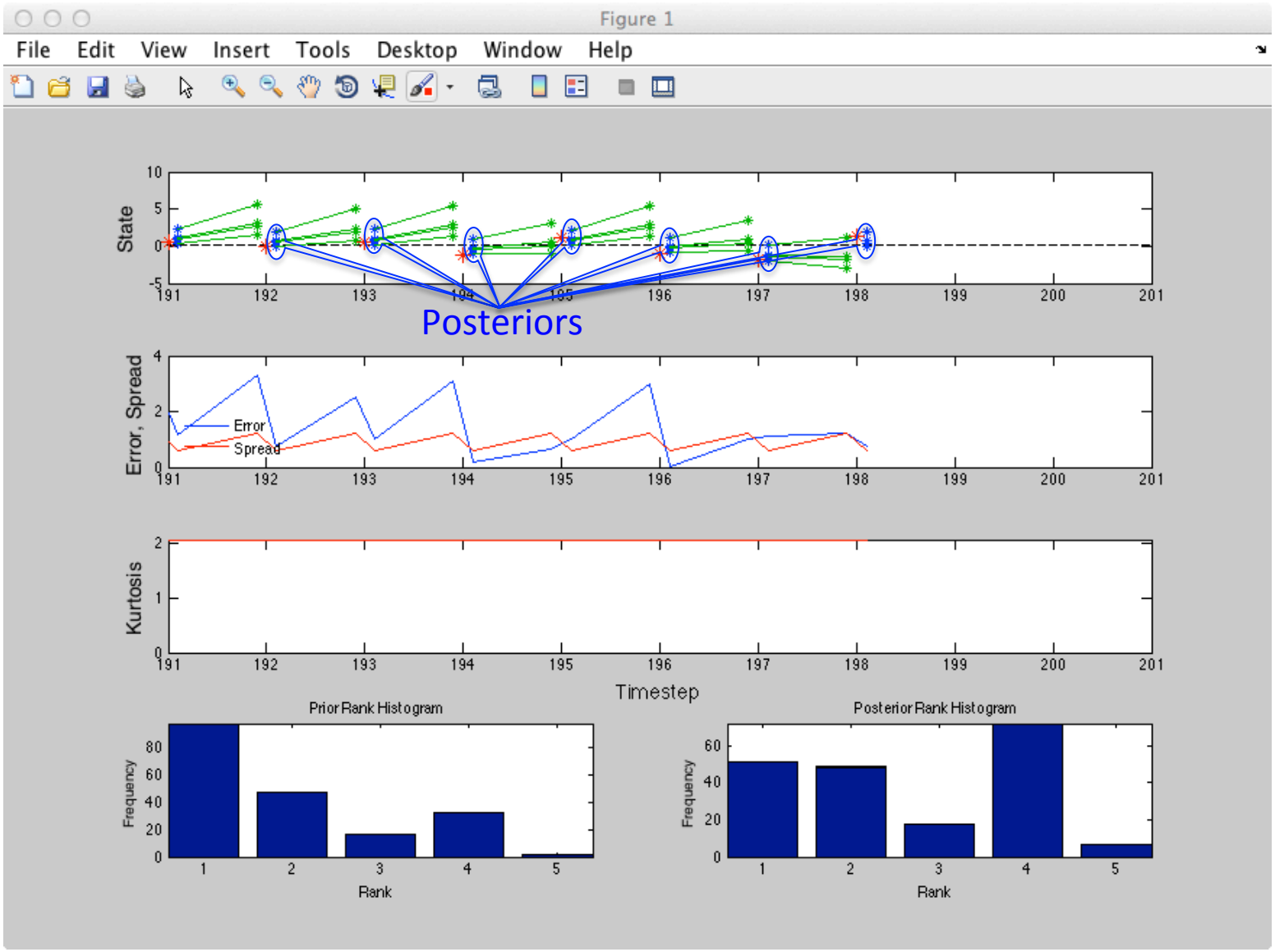
Matlab Hands-On: oned_model



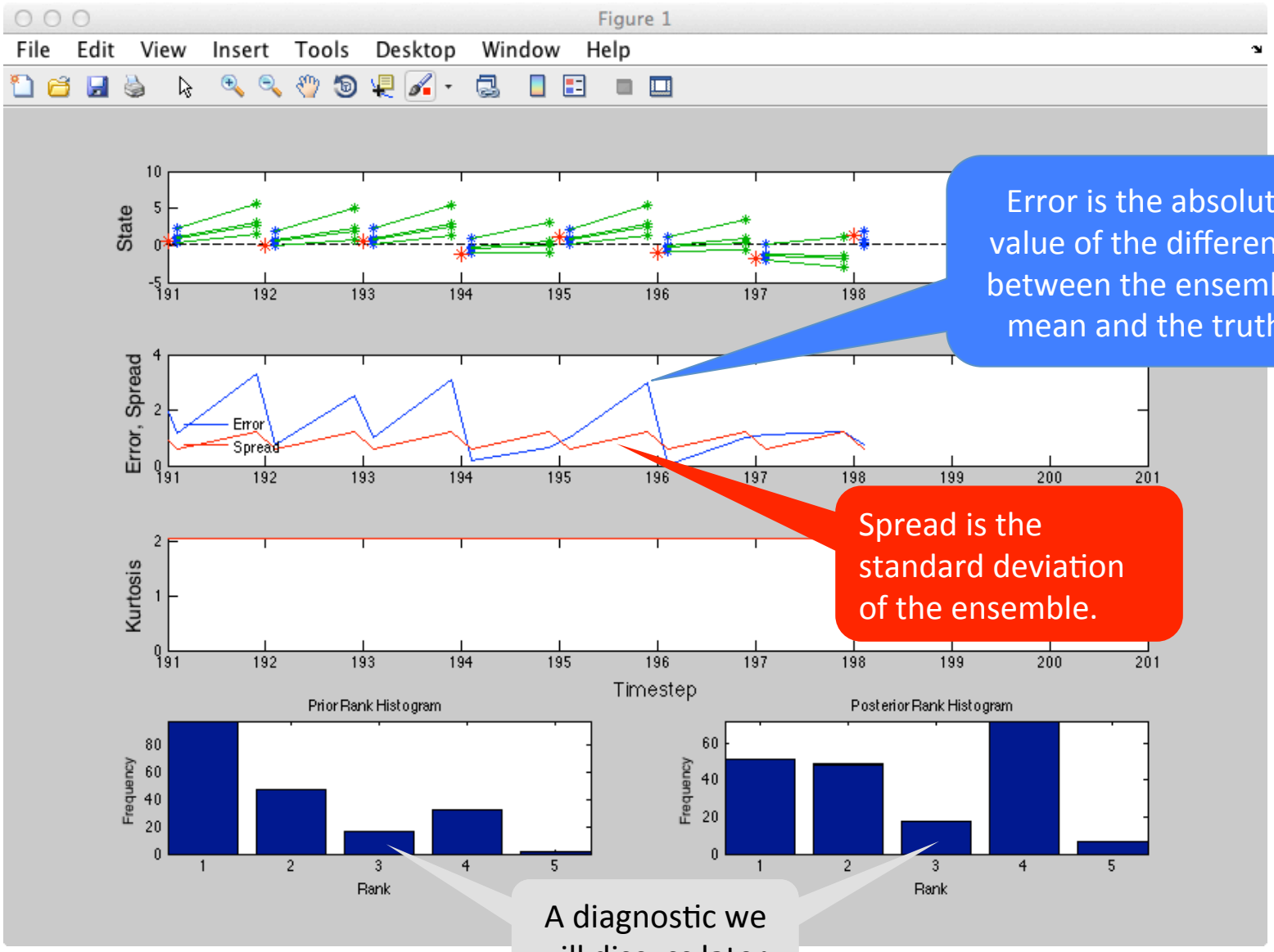
Matlab Hands-On: oned_model



Matlab Hands-On: oned_model



Matlab Hands-On: oned_model



Matlab Hands-On: oned_model

Explorations:

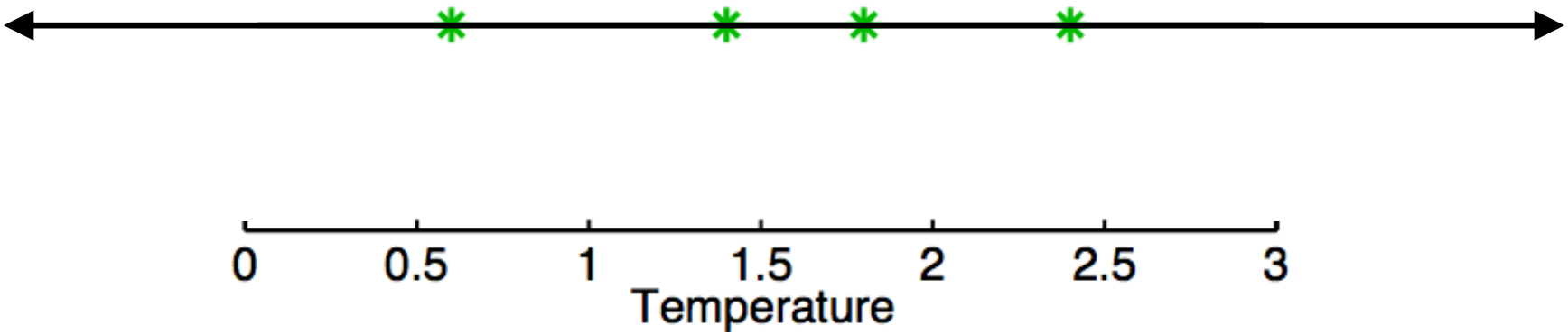
1. Step through a sequence of advances and assimilations with the top button. Watch the evolution of the ensemble, the error and spread.
2. How does a larger ensemble size (< 10 is easiest to see) act?

- Compare the error and spread for different ensemble sizes.
- Note the time behavior of the error and spread.

3. Let the model run freely using the second button.

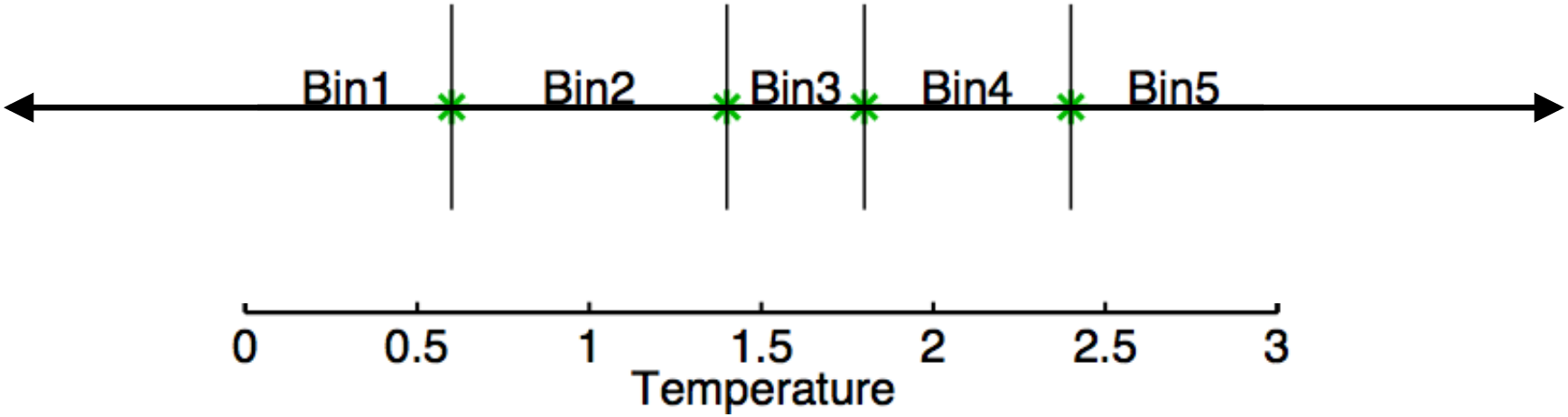
The Rank Histogram: Evaluating Filter Performance

Draw 5 values from a real-valued distribution.
Call the first 4 'ensemble members'.



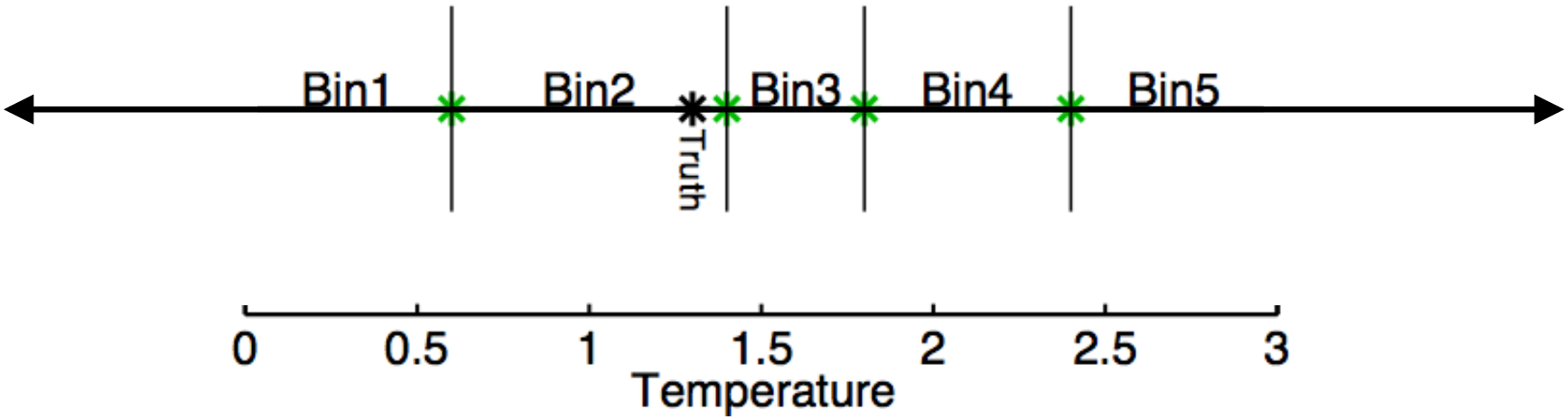
The Rank Histogram: Evaluating Filter Performance

These 4 ‘ensemble members’ partition the real line into 5 bins.



The Rank Histogram: Evaluating Filter Performance

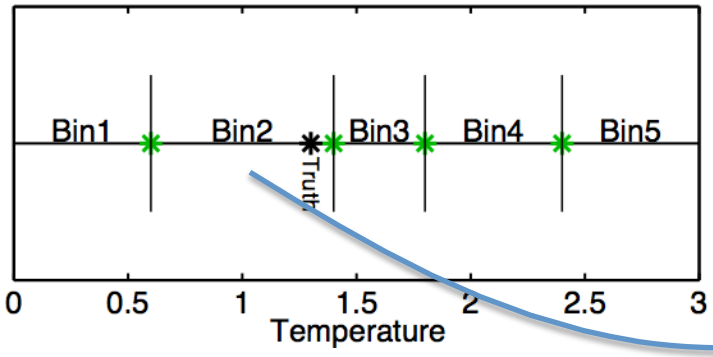
Call the 5th draw the 'truth'.
1/5 chance that this is in any given bin.



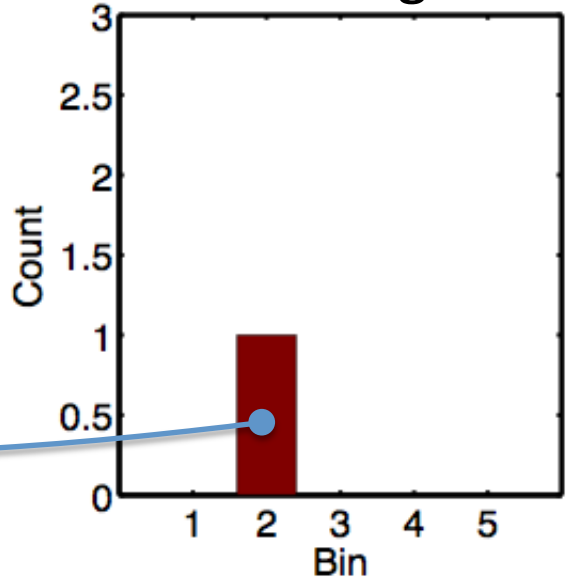
The Rank Histogram: Evaluating Filter Performance

Rank histogram shows the frequency of the truth in each bin over many assimilations.

Same figure as previous slide.

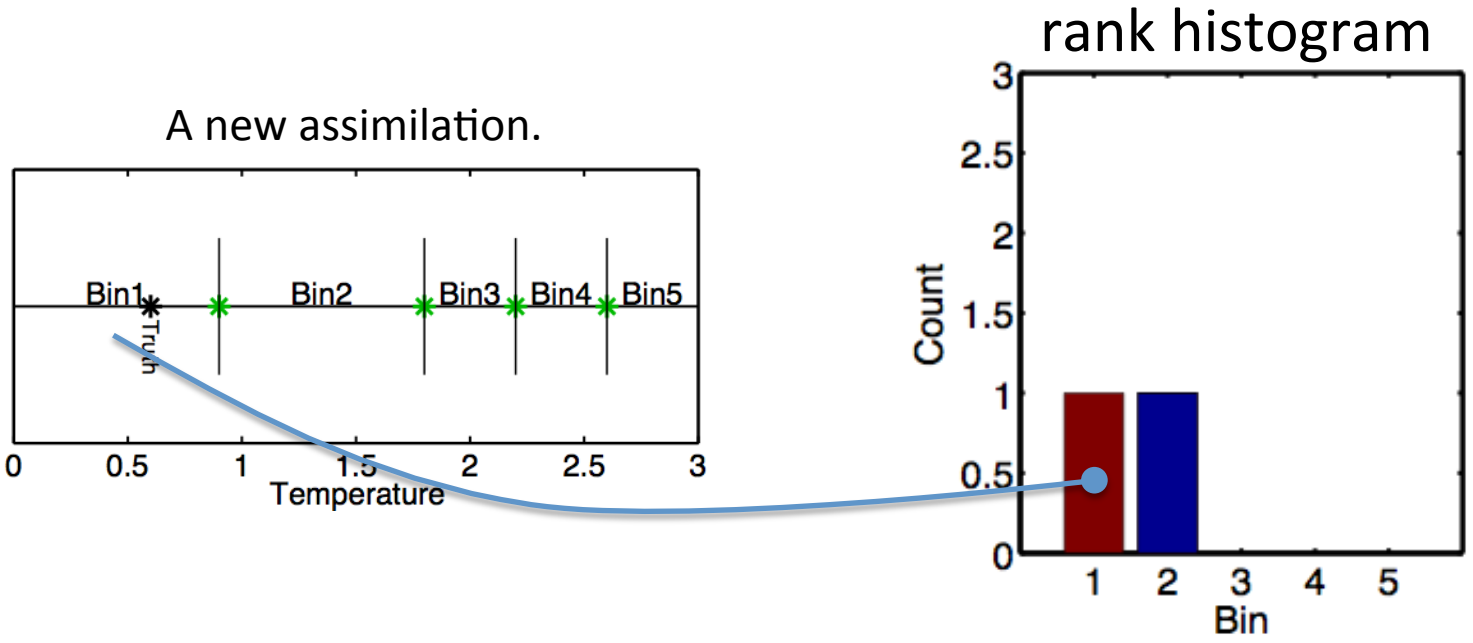


rank histogram



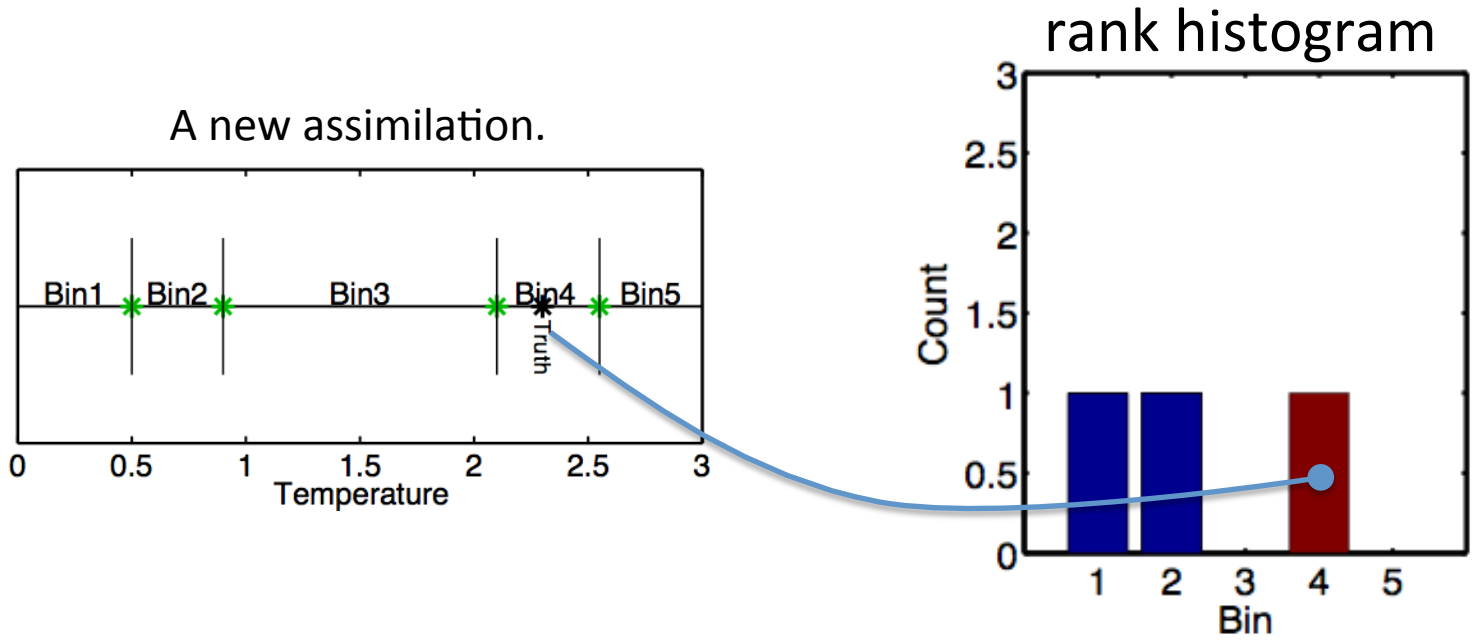
The Rank Histogram: Evaluating Filter Performance

Rank histogram shows the frequency of the truth in each bin over many assimilations.



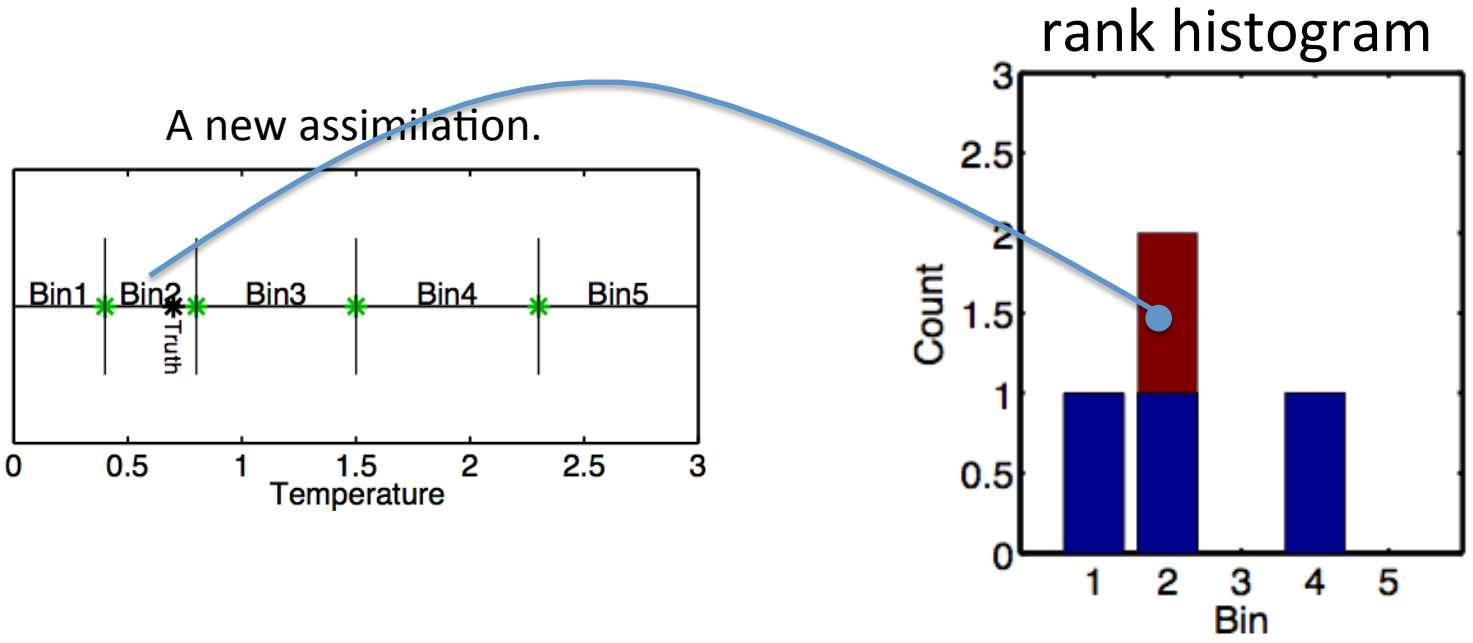
The Rank Histogram: Evaluating Filter Performance

Rank histogram shows the frequency of the truth in each bin over many assimilations.



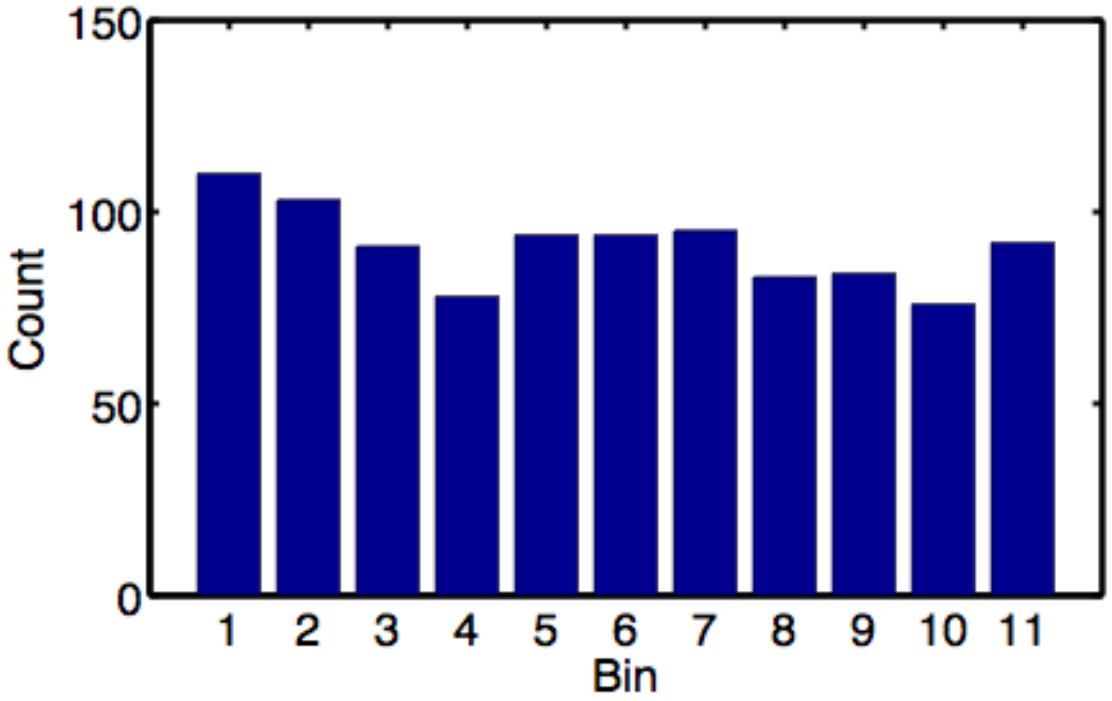
The Rank Histogram: Evaluating Filter Performance

Rank histogram shows the frequency of the truth in each bin over many assimilations.



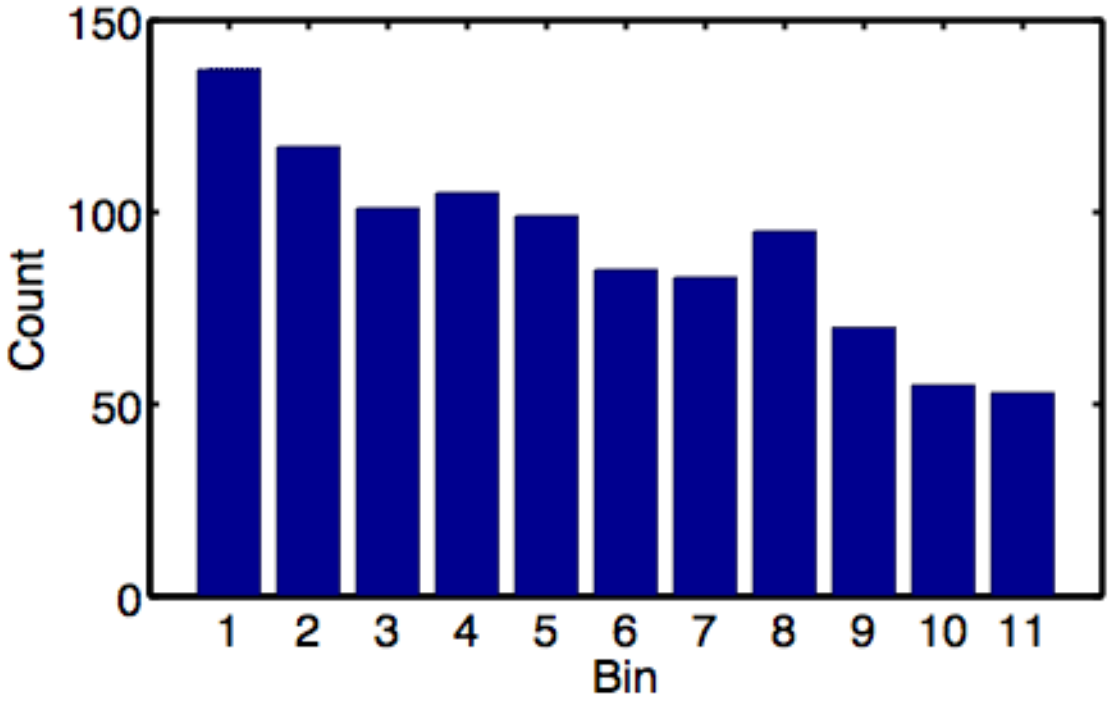
The Rank Histogram: Evaluating Filter Performance

Rank histograms for good ensembles should be uniform (caveat sampling noise).
Want truth to look like random draw from ensemble.



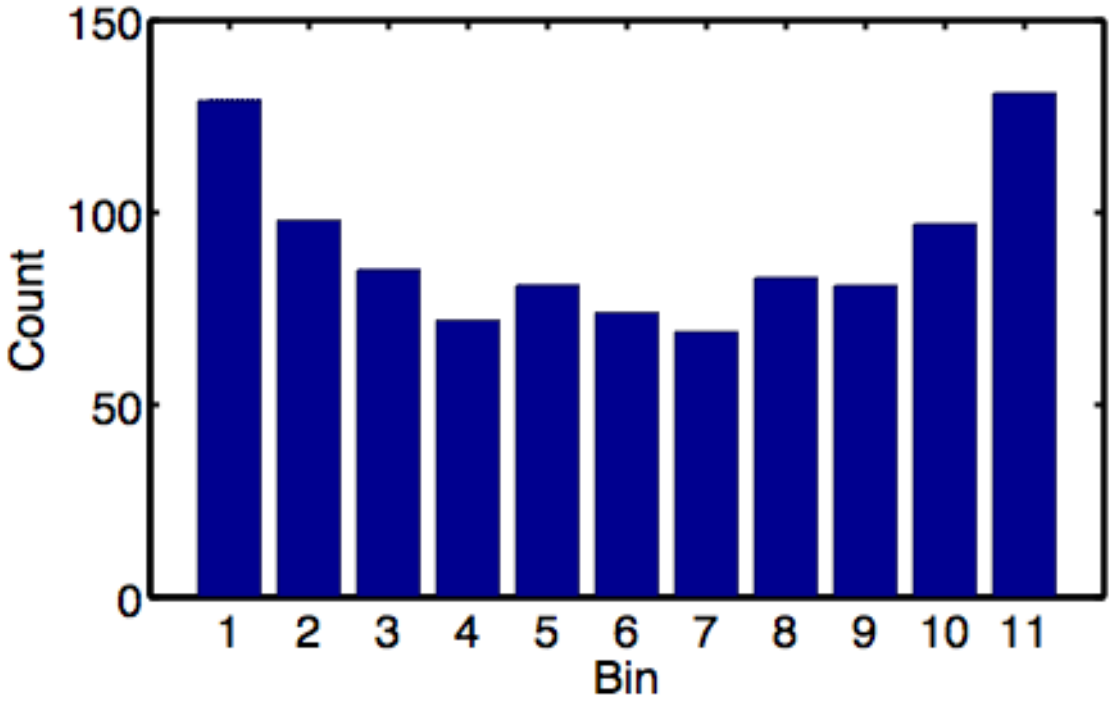
The Rank Histogram: Evaluating Filter Performance

A biased ensemble leads to skewed histograms.



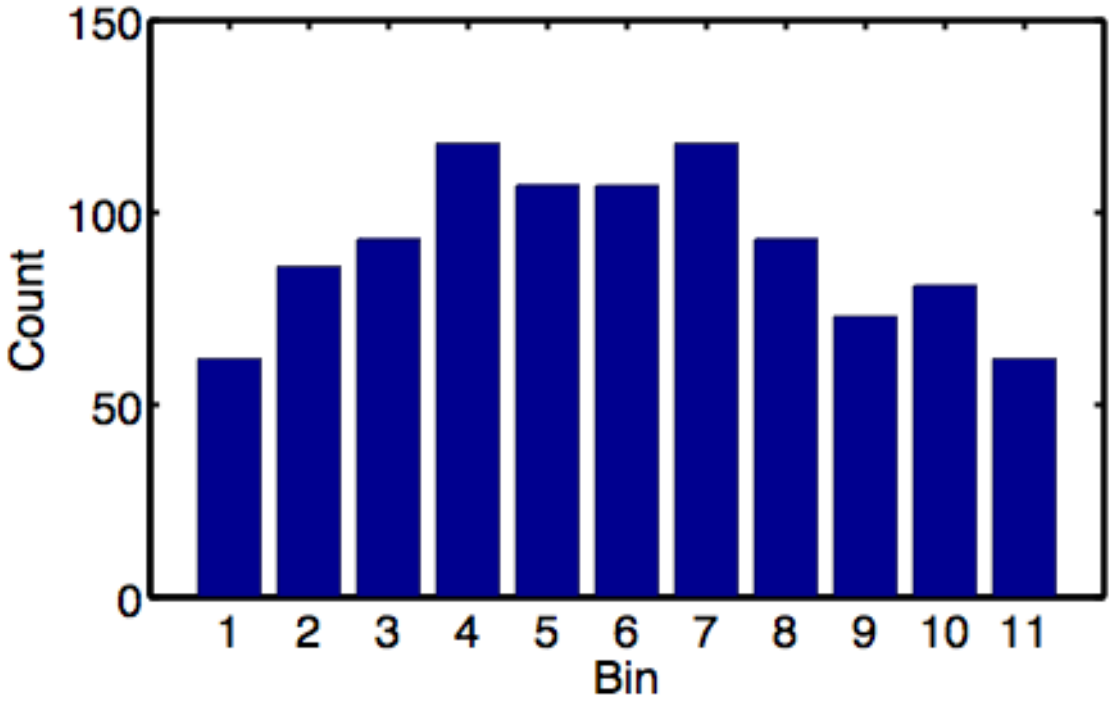
The Rank Histogram: Evaluating Filter Performance

An ensemble with too little spread gives a u-shape.
This is the most common behavior for geophysics.



The Rank Histogram: Evaluating Filter Performance

An ensemble with too much spread is peaked in the center.



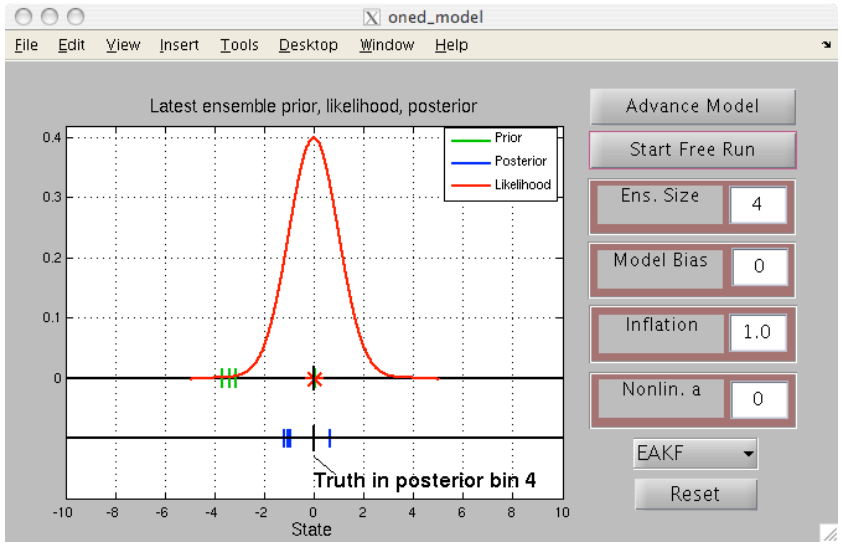
Matlab Hands-On: **oned_model** Understanding the Rank Histogram

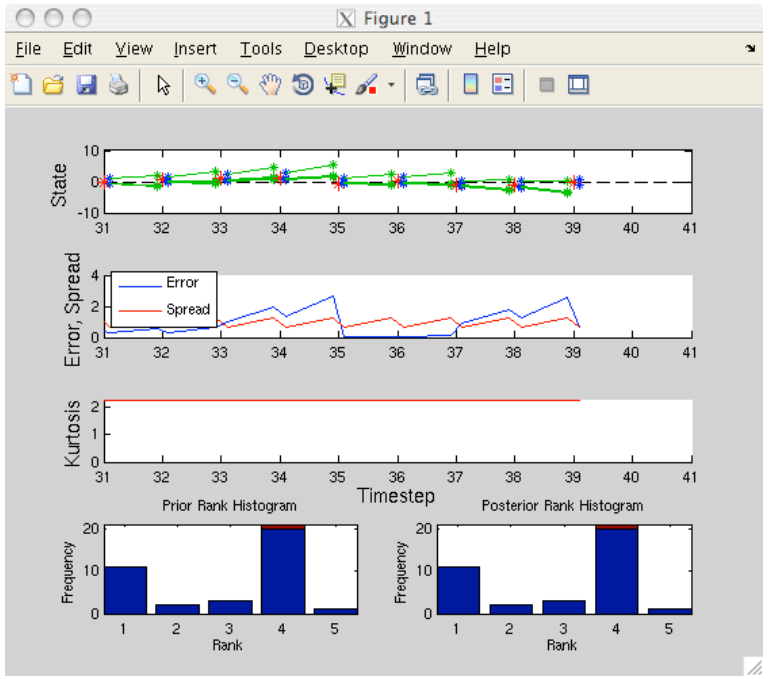
Procedure:

1. The ensemble size can be changed with a dialog box.
2. A model bias can be set with a dialog box.
3. An additional non-linear term can be added to the model with a dialog box.

Note: Model bias is $dx/dt = x + bias$.

- Non-linear model is $dx/dt = x + a\ x\ |x|$. Super-exponential growth.
- Truth is still always 0.
- See matlab script *advance_oned.m* for details.





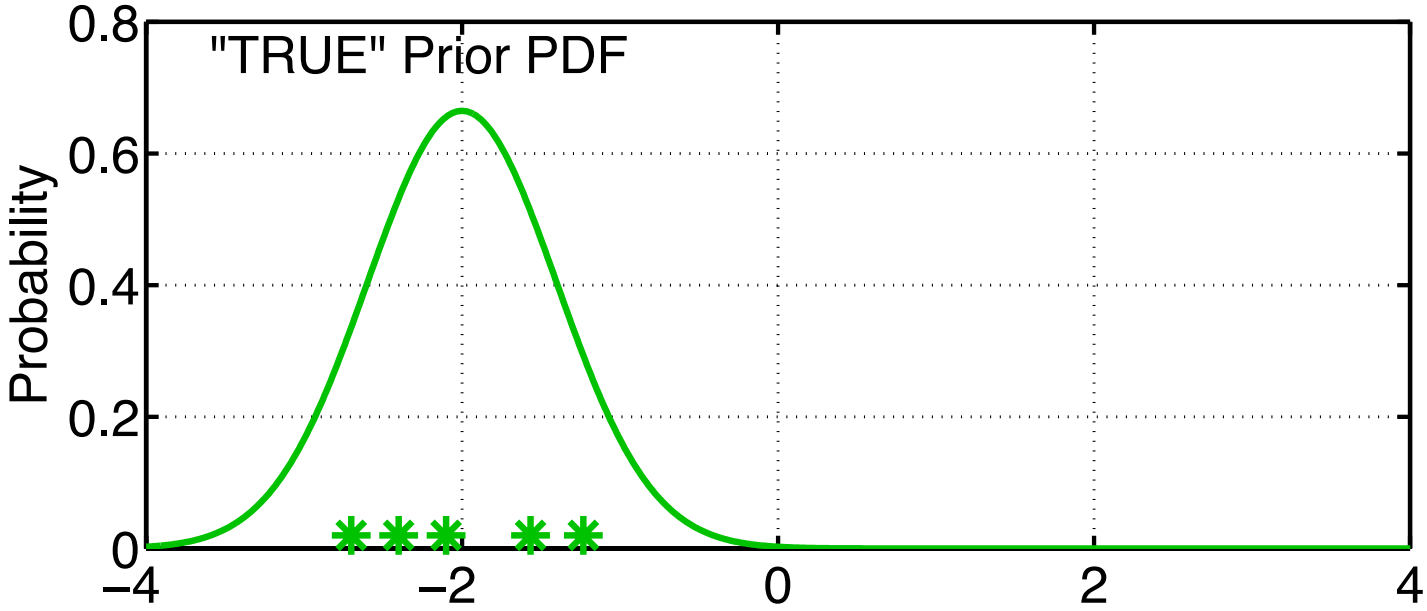
Matlab Hands-On: oned_model Understanding the Rank Histogram

Explorations:

1. Step through a sequence of advances and assimilations with the top button. Watch the evolution of the rank histogram bins.
2. Add some model bias (less than 1 to start) and see how the filter responds.
3. Add some nonlinearity (< 1) to the model. How do the different filters respond?
4. Can you break the filter (find setting so that the ensemble moves away from zero) with the options explored so far?

Dealing with systematic error: Variance Inflation

Observations + physical system → 'true' distribution.

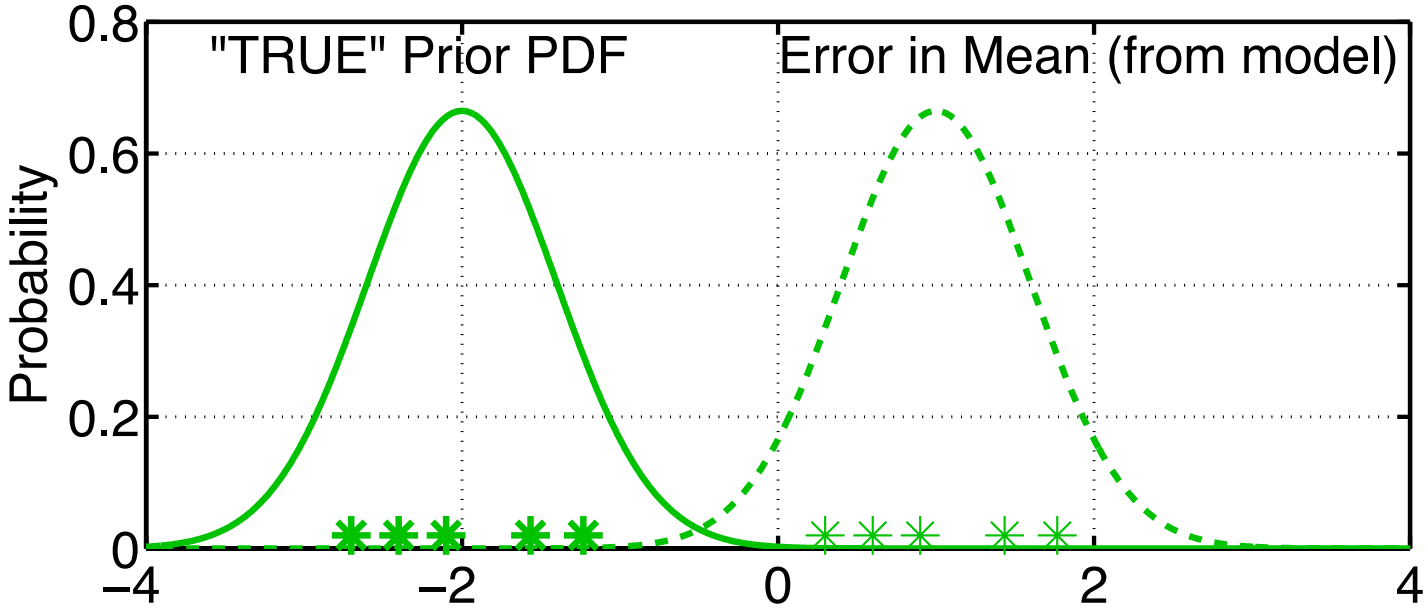


Dealing with systematic error: Variance Inflation

Observations + physical system → 'true' distribution.

Model bias (and other errors) can shift actual prior.

Prior ensemble is too certain (needs more spread).

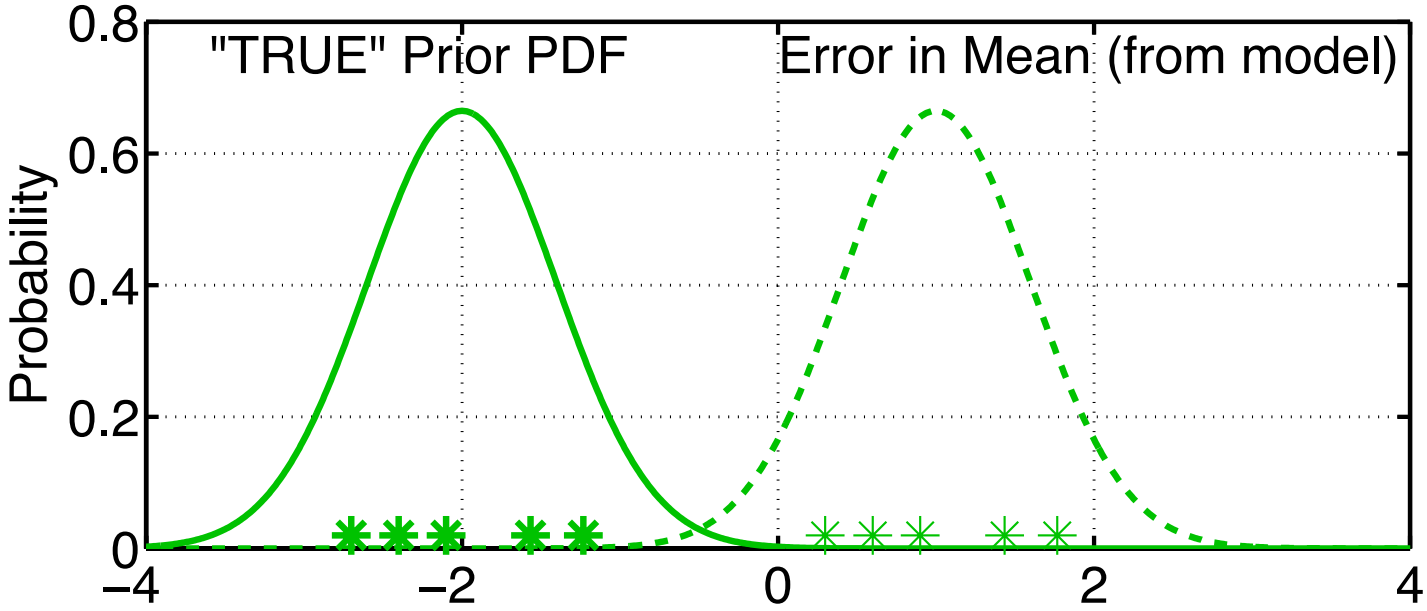


Dealing with systematic error: Variance Inflation

Could correct error if we knew what it was.

With large models, can't know error precisely.

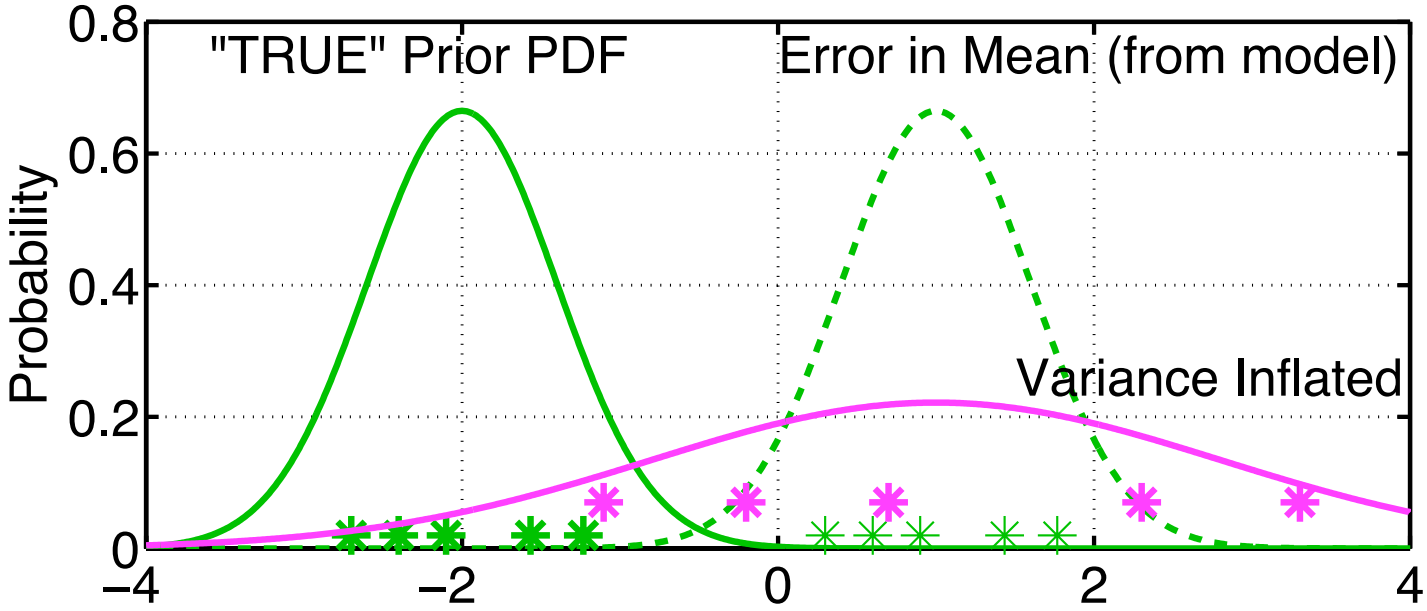
Taking no action can cause observations to be ignored.



Dealing with systematic error: Variance Inflation

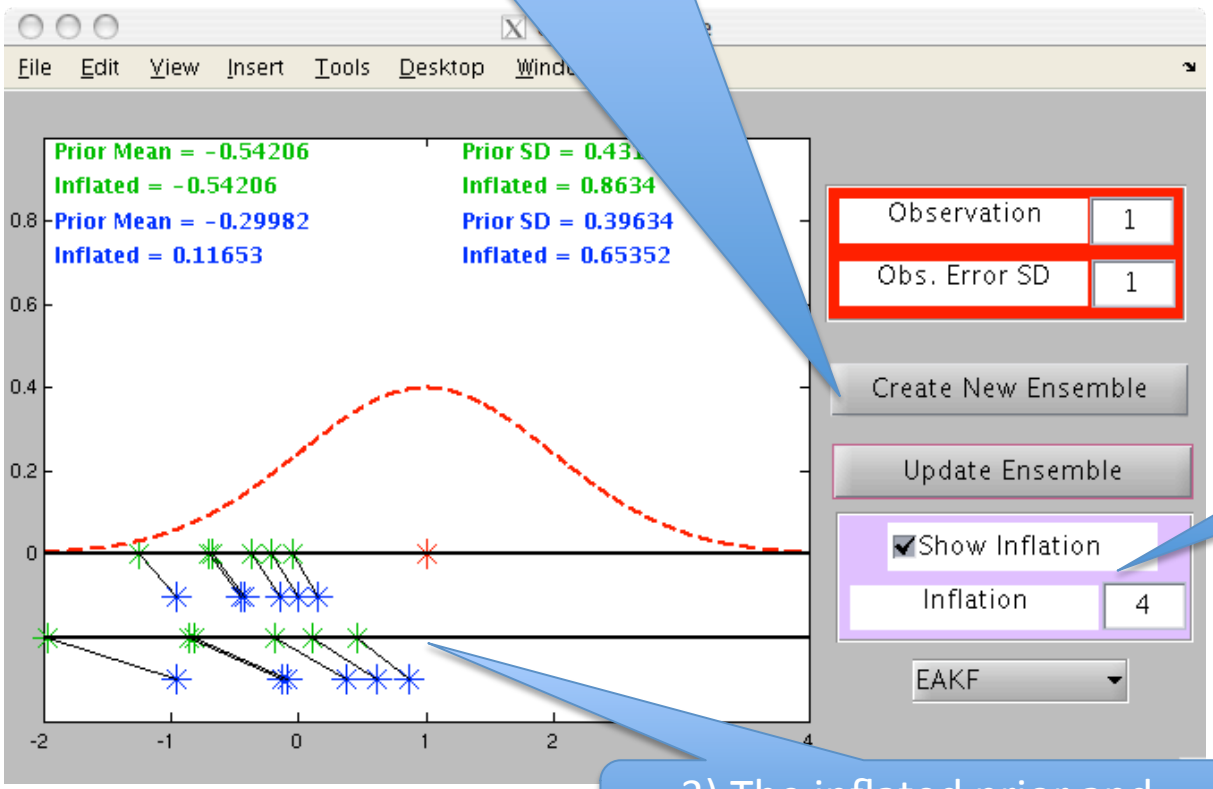
Naïve solution: increase the spread in the prior.

Give more weight to the observation, less to the prior.



Matlab Hands-On: **oned_ensemble** exploring prior inflation

1) Create a new ensemble.



2) Set an inflation value and turn on.

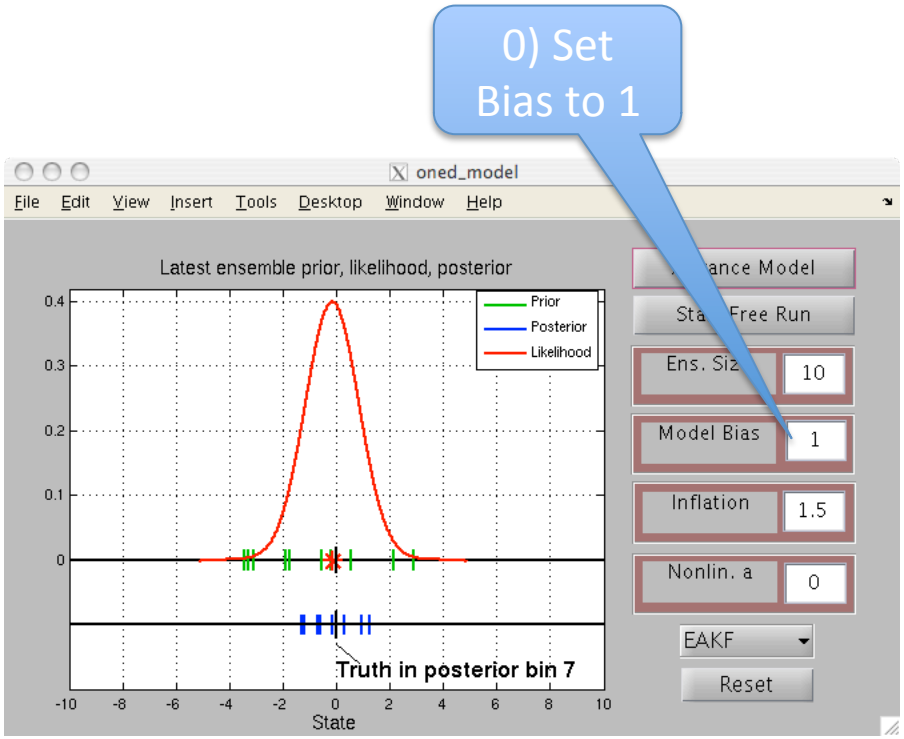
3) The inflated prior and assimilation shows up here.

Matlab Hands-On: exploring prior inflation with oned_ensemble

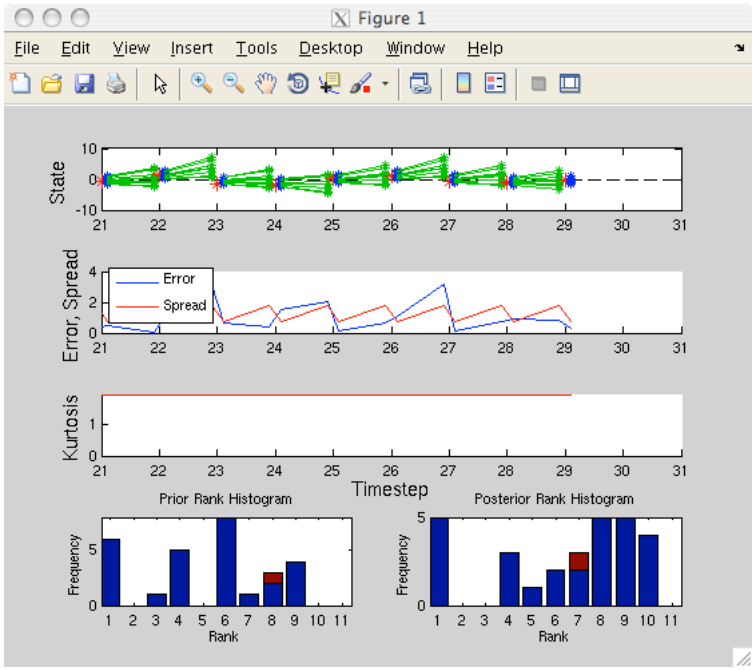
Explorations:

- See how increasing inflation (> 1) changes the posterior mean and standard deviation.
- Look at priors that are not shifted but have small spread compared to the observation error distribution.
- Look at priors that are shifted from the observation.

Matlab Hands-On: **oned_model** using inflation to deal with systematic error



1. Run an assimilation and observe the error, spread, and rank histograms.
2. Add some inflation (try starting with 1.5) and observe how behavior changes.
3. What happens with too much inflation?



Remember:

This uses the simple linear model $dx/dt = x$.

The 'truth' is always 0.

Observation noise is a draw from a $N(0,1)$.

The spread is increased by the square root of the inflation.

Matlab Hands-On: oned_model using inflation to deal with systematic error

Explorations:

- Try a variety of model bias and inflation settings.
- Try using inflation with a nonlinear model.