

Joe Crandall's PHYS 3164 Thermal and Statistical Physics    Used Heavily    Topic    SubTopic    KNOWTHISMATH    Definition/Constant/Operator    Key Observation    Question    Proof    break

**canonical ensemble**  $z = \sum_s e^{-\epsilon_s/\tau}$   $P(s) = \frac{1}{z} e^{-\epsilon_s/\tau}$   $U = \langle \epsilon \rangle = \sum_s \epsilon_s P(s) = \frac{1}{z} \sum_s \epsilon_s e^{-\epsilon_s/\tau} = \tau^2 \frac{\partial \ln z}{\partial \tau}$  characteristic function:  $\sigma = \ln \langle \Gamma \rangle$  **w5 d1 The law of increase of energy** Thermal contact entropy before  $\sigma_I = \ln(g_1(U_1, 0) g_2(U_2, 0))$ ,  $\sigma_F = \ln\left(\sum_{U_1} g_1(U_1) g_2(U_2 - U_1)\right) = \ln(g_1, g_2)_{max}$  when you remove a constraint you expect the entropy to increase **1** Law of increase in entropy is a statistical statement, Law of Thermodynamics, **0th** If system a is in equilibrium with system b and b is in equilibrium with system c, then a is in equilibrium with system c, **1st** Energy conservation, heat is a form of energy **2nd** Law of increase of entropy, when you remove a constraint you expect the entropy increase **3rd** The entropy approaches a constant as  $\tau \rightarrow 0$  and  $\sigma(\tau = 0) = \ln g(\tau = 0)$  the multiplicity of the ground state of the system **Thermal contact** Q. which direction is the energy flow? upon thermal contact.  $\tau_1 > \tau_2$   $\sigma = \ln g$   $\tau = \frac{\partial U}{\partial \sigma}$   $\partial \sigma = \partial \sigma_1 + \partial \sigma_2 = \frac{\partial \sigma_1}{\partial U_1} \partial U_1 + \frac{\partial \sigma_2}{\partial U_2} \partial U_2 = \frac{\partial \sigma_1}{\partial U_1} \partial U - U + \frac{\partial \sigma_2}{\partial U_2} \partial U = (\frac{1}{\tau_1} - \frac{1}{\tau_2}) \partial U \geq 0$  if  $\tau_1 > \tau_2$  then  $\partial U > 0$  if  $\tau_1 < \tau_2$  then  $\partial U < 0$   $\partial U = (\frac{\partial U}{\partial \sigma}) \partial \sigma + (\frac{\partial U}{\partial V}) \partial V$   $\partial U = \tau \partial \sigma - P \partial V$  (first law of thermo dynamics) **pressure** Quantum state  $V \rightarrow V - V_0$   $\delta U = - \frac{\partial \sigma}{\partial V} \cdot \delta V$   $\delta U = P \delta V$   $P_s = -(\frac{\partial \sigma}{\partial V})$  (pressure from a single quantum mechanical state) **Pressure**  $P = \langle P_s \rangle = -(\frac{\partial U}{\partial V})$  ( $P_s$  pressure for a particular quantum state)  $= \sum_s P(s) P_s$  ( $P(s)$  probability of quantum state)  $= \sum_s p(s) - (\frac{\partial \sigma}{\partial V})$   $= -(\frac{\partial \sigma}{\partial V})$  **w5 d2 chapter 3** Boltzman Distribution and Helmholtz free energy, a system in equilibrium with the environment, temperature is fixed **Q**. pdf for Microstates **S**: system,  $\epsilon_s = (s_x, s_p)$  **R**: reservoir  $U_0 - \epsilon_s$ ,  $P(s_x, s_R) = \text{constant}$   $R > s$ ,  $R + s$ : closed system  $U_0$ ,  $P(s_x) = \sum_{s_R} P(s_x, s_R)$   $\frac{P(1)}{P(2)} = \frac{g_R(u_0 - \epsilon_1)}{g_R(u_0 - \epsilon_2)} = e^{\sigma_R(u_0 - \epsilon_1) - \sigma_R(u_0 - \epsilon_2)}$   $\epsilon_1 < U_0$ ,  $\epsilon_2 < U_0$   $\epsilon_R(U_0 - \epsilon) \approx \sigma_R(U_0) - \epsilon (\frac{\partial \sigma_R}{\partial U_0})$   $\frac{P(1)}{P(2)} = e^{-(\epsilon_1 - \epsilon_2)/\tau}$  **for canonical ensemble**  $z = \sum_s e^{-\epsilon_s/\tau}$  Taylor expansion  $P(s) \propto e^{-\epsilon_s/\tau}$  **probability of microstates in thermal equilibrium with its environment**  $P(s) = \frac{1}{z} e^{-\epsilon_s/\tau}$   $U = \langle \epsilon \rangle = \sum_s \epsilon_s P(s) = \frac{1}{z} \sum_s \epsilon_s e^{-\epsilon_s/\tau} = \tau^2 \frac{\partial \ln z}{\partial \tau}$  at  $\tau = 0$ , the system settles in the ground state.  $\sum_s p(s) = 1 \rightarrow z = \sum_s e^{\epsilon_s/\tau}$  **Z**: partition function  $e^{-\epsilon_s/\tau}$ : Boltzman factor  $e^{-\epsilon_s/\tau}$   $z = \sum_s e^{-\epsilon_s/(\Delta + \tau)}$   $= e^{-\Delta/\tau} \sum_s e^{-\epsilon_s/\tau}$  **P**(s)  $\rightarrow \frac{e^{-\epsilon_s/(\Delta + \tau)}}{1 + e^{-\epsilon_s/(\Delta + \tau)}} = \frac{e^{-\epsilon_s/\tau}}{1 + e^{-\epsilon_s/\tau}}$  **Z**( $\tau$ ): high temp limit  $T \rightarrow \infty$ ,  $P(s) \rightarrow \text{uniform} = \frac{1}{z}$ , **Z**: number of microstates in the system, low temp limit  $\tau \rightarrow 0$   $P(s) = 0$  if  $\epsilon_s \neq 0$  and  $\frac{1}{z} = 1$  if  $\epsilon_s = 0$  **ex: one particle, only two states** state 1 energy 0, and state 2 energy  $\epsilon$   $z = 1 + e^{-\epsilon/\tau}$   $U = \frac{1}{1 + e^{-\epsilon/\tau}} (\epsilon e^{-\epsilon/\tau})$  1. microstates and energetics 1.5 decide the ensemble 2. z get the partition function 3. U...  $\leftarrow z$  (from partition function)  $z = e^{-0/\tau} + e^{-\epsilon/\tau} = 1 + e^{-\epsilon/\tau}$   $U = \langle \epsilon \rangle = \frac{\epsilon e^{-\epsilon/\tau}}{1 + e^{-\epsilon/\tau}} = \frac{\tau \partial \ln z}{\partial \tau}$  review  $P(s) = \frac{e^{-\epsilon_s/\tau}}{z}$  and  $z = \sum_s e^{-\epsilon_s/\tau}$  **w6 d1 how do we relate free energy F to microscopic quantities**  $F = -\tau \ln z(B, V, N)$  **demonstrate**  $-\tau \ln(z) = U - \tau \sigma$  connection with ThermoDynamics: free energy [Helmholtz free energy]  $F \equiv U - \tau \sigma$   $\partial U = \tau \partial \sigma - P \partial V$   $\partial F \equiv \partial U - \partial(\tau \sigma) = \tau \partial \sigma - P \partial V - (\tau \sigma + \sigma \partial \tau) = -\sigma \partial \tau - P \partial V$  **Proof** if we identify  $\sigma = -\frac{\partial F}{\partial \tau}$  **main framework for canonical ensemble from T.D.**  $\partial F = -\sigma \partial \tau - P \partial V$   $\sigma = -\frac{\partial F}{\partial \tau}$   $P = -\frac{\partial F}{\partial V}$   $F = U - \tau \sigma$  **another expression of  $\sigma$**  for an isolated system  $\sigma = \log g$  for a system with a given temperature  $\tau$   $\sigma = -(\frac{\partial F}{\partial \tau}) = -\sum_s p(s) \ln p(s)$  for isolated system  $p(s) = \frac{1}{g}$   $\sigma = -\ln \frac{1}{g} = \ln g$  review  $F = -\tau \ln z(\tau)$   $\sigma = -\frac{\partial F}{\partial \tau} \rightarrow \sigma = -\sum_s p(s) \ln p(s)$   $P = -(\frac{\partial F}{\partial V})$  **w6 d2 dU and ideal gas law**  $\partial U = \partial(\sum_s p(s) \epsilon_s) = \sum_s \epsilon_s \partial p(s) + \sum_s p(s) \partial \epsilon_s$  1st term, change in distribution on Microstates, 2. energy spectrum changes, no change in distribution **1a** lets say change  $V \rightarrow V + \delta V$   $\sum_s p(s) \partial \epsilon_s = \sum_s p(s) \frac{\partial \epsilon_s}{\partial V} \delta V = -p \delta V$  proof for a quantum system  $P = -\frac{\partial \sigma}{\partial V}$  for thermal system, carry out ensemble average  $P = -\sum_s p(s) \frac{\partial \epsilon_s}{\partial V} = -\sum_s \frac{e^{-\epsilon_s/\tau}}{z} \frac{\partial \epsilon_s}{\partial V} = -\sum_s \frac{1}{z} \frac{\partial \epsilon_s}{\partial V} e^{-\epsilon_s/\tau} = \frac{1}{z} \frac{\partial}{\partial V} \sum_s e^{-\epsilon_s/\tau} = \frac{1}{z} \frac{\partial}{\partial V} (e^{-\tau \frac{\partial U}{\partial V}}) = P$  **using canonical ensemble Proof**  $\sum_s \epsilon_s \partial p(s) = \tau \partial \sigma$  **ideal gas** System of particles where they do not interact with each other, point particles **one atom in a box**  $V = L^3$   $\epsilon = \frac{\hbar^2}{2M} (\frac{\pi}{L})^2 (n_x^2 + n_y^2 + n_z^2)$  calculate partition function  $Z_1 = \sum_{n_x} \sum_{n_y} \sum_{n_z} e^{-\frac{\hbar^2 \pi^2}{2M L^2} (n_x^2 + n_y^2 + n_z^2)} = \sum_{n_x} e^{-\frac{\hbar^2 \pi^2}{2M L^2} n_x^2} \dots$   $N \rightarrow \infty$  and  $V \rightarrow \infty$  where  $\frac{N}{V} = \text{constant}$  can approximate  $\sum$  with  $\int$   $\sum_n e^{-\alpha^2 n^2}$  where  $\alpha^2 = \frac{\hbar^2 \pi^2}{2m L^2 \tau} = \int_0^\infty \partial n e^{-\alpha^2 n^2} = \frac{1}{\alpha} \int_0^\infty \partial x e^{-x^2} \chi \approx \alpha N$   $\frac{1}{\alpha} \frac{1}{2} \pi^{1/2}$   $Z_1 = \frac{\pi^{3/2}}{8 \alpha^3} = \frac{V}{(2 \pi \hbar^2 / M \tau)^{3/2}}$  rewrite as a ratio of two densities  $n = \frac{1}{V}$   $n_\alpha = (\frac{M \tau}{2 \pi \hbar^2})^{3/2} = \frac{1}{\lambda^3}$  quantum concentration  $\lambda = (\frac{2 \pi \hbar^2}{M \tau})^{1/2}$  where  $\lambda$  is the thermal wavelength  $F = -\tau \ln z = -\tau \ln \frac{n}{n_\alpha}$   $U = \tau^2 \frac{\partial \ln z}{\partial \tau} = \frac{3}{2} \tau U = \frac{3}{2} N \tau$  internal energy of a system with N particles **N atoms in a box** if particles do not interact  $Z_2 = \sum_{n_{21}, n_{22}, n_{23}} \sum_{n_{24}, n_{25}, n_{26}} e^{-(\epsilon_{21} + \epsilon_{22})/\tau} = z_1 \cdot z_1$  **N identical particles in a box.**  $Z_N = \frac{1}{N!} Z_1^N = \frac{1}{N!} (n_\alpha V)^N$   $U = \frac{3}{2} = N \tau$  equation of state  $pV = n \tau$  **w8 d1 Free energy as a force towards equilibrium** Helmholtz free energy F, F is a minimum for a system in thermal contact with a reservoir R, if the volume of the system is kept constant. Consider the total energy of the system plus the reservoir temperature.  $\partial \sigma_{total} = \partial \sigma + \partial \sigma_R$   $\partial \sigma_R = \frac{1}{\tau} \partial U_R + \frac{P}{\tau} \partial V_R$   $P = \text{pressure and given no volume change}$   $\partial \sigma_{total} = \partial \sigma + \frac{1}{\tau} \partial U_R$   $\partial U + \partial U_R = 0$  (Total energy conserved)  $\partial \sigma_{total} = \partial \sigma + \frac{1}{\tau} \partial U_R$   $\partial U + \partial U_R = 0$   $\partial \sigma_{total} = \partial \sigma - \frac{1}{\tau} \partial U$   $\partial \sigma_{total} = -\frac{1}{\tau} \partial F$   $F = U - \tau \sigma$   $\partial F = \partial U - \tau \partial \sigma - \sigma \partial \tau$  under the condition of fixed  $\tau$  and  $V$   $\partial F \leq 0 \rightarrow F$  is a minimum in equilibrium **example using a minimization principle** binary model system  $N_{\uparrow\uparrow}, N = N_\uparrow + N_\downarrow$ ,  $N_{\uparrow\downarrow}, 2s = N_\uparrow - N_\downarrow$   $U(s) = -2smB$  **U** usually (internal energy) thermal average of the systems energy  $Z = \sum_s e^{-\epsilon_s/\tau}$   $U = \sum_s \frac{\epsilon_s e^{-\epsilon_s/\tau}}{z}$  when given  $\tau, B$  can you find  $\langle s \rangle$  two approaches, use canonical ensemble approach, energetics table, partition function, or minimize  $F$  with respect to  $s$   $F = U - \tau \sigma$   $\sigma(s) \approx \ln \binom{N}{N_\uparrow} = -(\frac{N}{2} + s) \ln \left(\frac{N}{2} + \frac{s}{2}\right) - (\frac{N}{2} - s) \ln \left(\frac{N}{2} - \frac{s}{2}\right)$   $U(s) \approx -2smB$  Free energy  $F(s) = U(s) - \tau \sigma(s) = -2smB + -(\frac{N}{2} + s) \tau \ln \left(\frac{N}{2} + \frac{s}{N}\right) + (\frac{N}{2} - s) \tau \ln \left(\frac{N}{2} - \frac{s}{N}\right)$  At equilibrium  $F(s) \rightarrow \min.$   $\frac{\partial F}{\partial s} = 0$  **remember**  $\langle s \rangle$  is the average of the spin axis  $\therefore -smB + \tau \ln \binom{n+2(s)}{N-2(s)} = 0$  and  $\langle s \rangle = \frac{N}{2} \arctan \frac{mB}{\tau}$  **w8 d2 chapter 4 Planck distribution/black body radiation** emission spectrum of the sun, the em spectrum in thermal equilibrium within a cavity Strategy, canonical ensemble, identify microstates, in a cavity the eigenstates of EM waves are modes, lets first look at a single mode  $\epsilon_s = \hbar \omega$ ,  $s = 0, 1, 2, \dots$   $Z = \sum_{s=0}^\infty e^{-s \hbar \omega / \tau} = \frac{1}{1 - e^{-\hbar \omega / \tau}}$   $\langle 1 - x \rangle (1 + x + x^2 + \dots) = 1$  and  $\therefore \langle 1 + x + x^2 + \dots \rangle = \frac{1}{(1-x)}$  Average number of photons  $\langle s \rangle = \frac{1}{z} \sum_s s e^{-s \hbar \omega / \tau}$  plank distribution for any number of photons in a single mode of frequency  $\omega$  Average energy:  $\langle \epsilon \rangle = \hbar \omega \langle s \rangle = \frac{\hbar \omega}{e^{\hbar \omega / \tau} - 1}$  if  $\tau \gg \hbar \omega$  then  $\tau \ll \hbar \omega$  then  $e^{\hbar \omega / \tau} \approx \frac{\hbar \omega}{\tau}$  Now: in a cavity, we have multiple modes Q/ what are the modes, A/ From EM: For a perfectly conducting cavity of  $L^3$   $E_x = E_{01} \sin \omega t \cos \frac{\pi x}{L} \sin \frac{\pi y}{L} \sin \frac{\pi z}{L}$   $E_y = E_{02} \dots$   $E_z = E_{03} \dots$  first sin is time dependence, the rest is spatial dependence  $\nabla \cdot \vec{E} = 0$   $E_{20} n_y + E_{02} n_y + E_{20} n_z = \vec{E}_0 \cdot \vec{n} = 0$  field vector and n vector are perpendicular For a given  $\vec{N}$  we can choose two independent  $\vec{e}_0$ ,  $\rightarrow$  for each  $\vec{n}$  two distinct modes Q/ how is  $\omega$  related to  $(n_x, n_y, n_z)$  sub. eigenmodes into wave equation  $C^2 (\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2}) E_z = \frac{\partial^2 E_z}{\partial t^2}$   $\omega_{n_x, n_y, n_z}$  Now consider all modes, the modes are independent  $U = 2 \sum_{n_x, n_y, n_z} \langle \epsilon_{n_x, n_y, n_z} \rangle = \sum_{n_x, n_y, n_z} \approx \frac{1}{8} \int_{-\infty}^\infty dn_x dn_y dn_z$  Integrand depends only on  $n = \sqrt{(n_x^2 + n_y^2 + n_z^2)} = \frac{1}{z} \int_0^\infty 4 \pi n^2 \partial n$   $U = \frac{\pi^2 \hbar^3 L^3}{15 \epsilon_0 \tau^4} \int_0^\infty dn \frac{n^3}{e^{\frac{\hbar n}{\tau}} - 1}$  **change of variable**  $x = \frac{\pi \hbar n}{L \tau}$   $U = \frac{\pi^2 \hbar^3 L^3}{15 \pi \hbar^4 \tau^4} \int_0^\infty dx \frac{x^3}{e^{x^4} - 1}$   $U = \frac{\pi^2 \hbar^3 L^3}{15 \pi \hbar^4 \tau^4} \frac{4}{15} \int_0^\infty dx \frac{x^3}{e^{x^4} - 1}$   $U = V \cdot \frac{15 \pi^2 \hbar^3 \tau^4}{15 \pi^2 \hbar^4 \tau^4} \tau^4 \propto \tau^4 \rightarrow$  stephan-Boltzmann law of radiation  $U = v \int \partial \omega U \omega$   $U_\omega = \frac{n}{\pi^2 c^3} \frac{\omega^3}{e^{\hbar \omega / \tau} - 1}$  (plank radiation law)  $\tau = 2.73K$  Q/  $\sigma$  A/ at constant volume  $\partial \sigma = \frac{\partial U}{\tau}$  where  $\partial U = \tau \partial \sigma - P \partial V$   $\sigma(\tau) = \int_0^\tau \frac{\partial \sigma}{\partial \tau} = \frac{4 \sigma^4 V}{15 \tau^4} (\frac{\hbar}{\tau c})^3 \propto \tau^3$  **w11 d1 Phonons in solids** quantum harmonic oscillator, partition function  $Z = \frac{1}{1 - e^{-\hbar \omega / \tau}}$  photon gas numbers of em modes is infinite, phonon gas of number of elastic modes is finite photon two possible polarizations, phonon longitudinal 1 and transverse 2  $U = \frac{3 \pi}{2} \int_0^\infty d n n^2 \frac{\hbar \omega_n}{e^{\hbar \omega_n / \tau} - 1}$   $\omega_n = \frac{n v}{L}$   $V = \text{speed of sound in the system}$  Make the integral dimensionless  $X = \frac{\pi v V n}{L \tau}$  therefor  $U = (\frac{3 \pi^2 \hbar^3 V}{2 L^3 \tau^4}) (\frac{\tau^4}{\pi v}) \int_0^\infty dx \frac{x^3}{e^{x^4} - 1}$   $X_D = \frac{\pi \hbar v n_D}{L \tau} = \frac{\theta}{\tau}$  Debye temp  $\theta = \hbar v (\frac{6 \pi^2}{V})^{1/3}$  of particular interest: low  $\tau \ll \theta$  where  $x_D \rightarrow \infty$   $U \approx \frac{3 \pi^2 \hbar^3 V}{2 L^3 \tau^4} \propto \tau^4$   $C_v = (\frac{\partial U}{\partial \tau})$   $\propto \tau^3$  chemical potential and gibbs distribution, when two systems are in diffuse contact and can exchange particles, which way would the particles flow. chemical potential  $\mu$  flow form high  $\mu$  to low  $\mu$ , diffusive contact - particles exchange, thermal contact - temperature exchange consider  $s_1, N_1, V_1 + V_2, s_2, N_2 + N_1, V_1 + V_2$ ,  $F = F_1 + F_2 = U_1 + U_2 - \tau(\sigma_1 + \sigma_2)$  Now consider a small variation  $\partial N_1 = -\partial N_2$  At equilibrium:  $\partial F = (\frac{\partial F_1}{\partial N_1}) \partial N_1 + (\frac{\partial F_2}{\partial N_2}) \partial N_2 = 0$  equilibrium condition  $(\frac{\partial F_1}{\partial N_1}) \tau_1 V_1 = (\frac{\partial F_2}{\partial N_2}) \tau_2 V_2$  Define:  $\mu(T, V, N) \equiv (\frac{\partial F}{\partial N})_{T, V}$   $p$ : pressure  $\partial \sigma = \frac{\partial \sigma}{\partial N} + \frac{\partial \sigma}{\partial V} \partial V + \frac{\partial \sigma}{\partial \tau} \partial \tau = \tau \partial \sigma + p \partial V + \mu \partial N$  Another expression for chemical potential  $\mu = -\tau (\frac{\partial \sigma}{\partial N})_{T, V}$   $\partial \sigma = (\frac{\partial \sigma}{\partial N}) \partial N + (\frac{\partial \sigma}{\partial V}) \partial V + (\frac{\partial \sigma}{\partial \tau}) \partial \tau = \frac{1}{\tau} (\frac{\partial U}{\partial N})_{T, V} + (\frac{\partial \sigma}{\partial V})_{T, V} + \frac{\partial \sigma}{\partial \tau}$   $\mu = (\frac{\partial U}{\partial N})_{T, V} - \tau (\frac{\partial \sigma}{\partial V})_{T, V} = -\tau (\frac{\partial \sigma}{\partial V})_{T, V}$  next topic  $U_0 = U + U_R$  and  $N = N + N_R$ ,  $S + R$  is an isolated system **w11 d2 when system constant temperature and volume**  $U_0 = U_R + U$   $N_0 = N_R + N \propto g(R, S) = g_R \cdot g_s = g_R \cdot 1$   $P(s_\alpha) \propto g_s (N_0 - N, U_0 - \epsilon_{s\alpha})$   $\frac{P(s_1, N_1)}{P(s_2, N_2)} = \frac{g_s (N_0 - N_1, U_0 - \epsilon_{s1})}{g_s (N_0 - N_2, U_0 - \epsilon_{s2})}$   $\sigma_R = \ln g_R$   $\frac{P(s_1, N_1)}{P(s_2, N_2)} = \frac{\exp[\sigma_R (\sigma_R (N_0 - N_1, U_0 - \epsilon_{s1}))]}{\exp[\sigma_R (\sigma_R (N_0 - N_2, U_0 - \epsilon_{s2}))]} = \exp[\Delta \sigma_R]$   $\Delta \sigma_R \approx \sigma_R (\mu_0 - \mu_1, \mu_0 - \epsilon_{s1}) - \sigma_R (\mu_0 - \mu_2, \mu_0 - \epsilon_{s2})$   $\Delta \sigma_R = -(N_1 - N_2) (\frac{\partial \sigma}{\partial N_0}) U_0 - (\epsilon_{s1} - \epsilon_{s2}) (\frac{\partial \sigma}{\partial U_0}) N_0 = \frac{\Delta N}{\tau} \mu - \frac{\Delta \epsilon}{\tau}$   $\frac{P(s_1, N_1)}{P(s_2, N_2)} = \frac{\exp[\frac{N (\mu - \epsilon_s) / \tau]}{\exp[\frac{N (\mu - \epsilon_s) / \tau]}} P(s_N) = \frac{1}{z_s} = \exp[(N \mu - \epsilon_s) / \tau]$   $Z_g = \sum_{N=0}^\infty \sum_{S(N)} \exp[(N \mu - \epsilon_s) / \tau]$   $\langle N \rangle = \frac{1}{z_s} = \sum_N \sum_{S(N)} N P(S_N) = \frac{1}{z_s} \frac{\partial}{\partial \mu} \frac{\partial}{\partial \mu} = \frac{\tau \partial \ln Z_g}{\partial \mu}$  Approach number 1  $\mu = (\frac{\partial F}{\partial N})_{T, V}$   $F = -\tau \ln z$   $\mu = (\frac{\partial F}{\partial N})_{T, V} = -\tau \ln z$   $Z_n = \frac{1}{n!} Z_1^n$   $Z_1 = n_Q \cdot v$   $F = -\tau [N \ln z_1 - \ln N!]$   $\frac{\partial F}{\partial N} = [-\tau \ln z_1 + \partial N \ln \left(\frac{n}{n_Q}\right) + 2 \pi n] (\frac{\partial F}{\partial N})_{T, V} = -\tau \ln \frac{n}{n_Q}$  and density  $= \ln \frac{N}{n_Q}$  Approach 2  $\langle N \rangle = \tau \frac{\partial \ln Z_g}{\partial \mu} = z_1 e^{\mu / \tau}$   $\mu = \tau \ln \left(\frac{N}{n_Q}\right) = \tau \ln \frac{n}{n_Q} - \text{density}$  **w12 d1 Chemical Potential**  $\mu_{tot}(h) = Mgh + c \ln \left(\frac{n(n)}{n_Q}\right)$   $\mu_{tot}(0) = \mu_{tot}(h)$   $\mu_{tot}(h) = \mu_{tot}(h = 0)$   $Mgh + \tau \ln \left(\frac{n(h)}{n_Q}\right) = \tau \ln \left(\frac{n(0)}{n_Q}\right)$   $n(h) = \frac{n(0)}{e^{Mgh/\tau}} = n(0) e^{-Mgh/\tau}$   $PV = nRT$   $n_1 = n(0) \frac{e^{3nD/\tau}}{e^{3nD/\tau} + 1}$   $n_4 = \frac{n(0)}{e^{3nD/\tau} + 1}$  **w12 d2 chapter 6: ideal gas**  $\mu_{tot} = \mu_{ext} + \mu_{int}$  ideal gas in gravity  $\mu_{int} = \tau \ln \left(\frac{n(h)}{n_Q}\right)$   $\mu_{ext} = Mgh$  Identical particles: Fermions:  $\frac{1}{2}$  integer spin, 0 or 1 on an orbital Bosons: integer spin, occupation number on an orbital can arbitrarily 0,1,2... **Grand partition function**  $Z_g = \sum_N \sum_{S_N} e^{(\mu N - \epsilon_{S_N})/\tau}$   $\langle N \rangle = \frac{1}{z_g} = \sum_N \sum_{S_N} N P(S_N)$  Bose Einstein Distribution  $f_{BE} = \frac{1}{e^{(\mu - \epsilon)/\tau} - 1}$  Fermi-Dirac Distribution  $f_{FD} = \frac{1}{e^{(\mu - \epsilon)/\tau} + 1}$  **w13 d1** For classical system  $\langle N \rangle = \lambda \sum_{orb} e^{-\epsilon/\tau}$   $\lambda = e^{\mu/\tau}$  In classical limit  $f \approx e^{(\mu - \epsilon)/\tau}$   $\mu = \tau \ln \left(\frac{n}{n_Q}\right)$   $F(N, V, \tau) = \int_0^N dN \cdot \mu(n, V, \tau)$   $P = -(\frac{\partial F}{\partial V})_{T, N} = \frac{n \tau}{V}$  Microstates: orbital + internal  $\epsilon = \epsilon_n + \epsilon_{int}$  One orbital  $Z_{int} = \sum_{int} e^{-\epsilon_{int}/\tau}$  Reversible isothermal expansion  $p_1 V_1 = p_2 V_2 = N \tau$   $\Delta \sigma = \sigma_f - \sigma_i$   $\Delta \sigma = N \ln \frac{V_f}{V_i}$   $\Delta U = 0$   $U = \frac{3 N \tau}{2}$   $U_f = U_i$   $Q = -W = N \tau \ln \frac{V_f}{V_i}$  **w13 d2 reversible expansion at constant entropy**  $\sigma(\tau, V) = N(\ln \tau^{3/2} + \ln v + \text{constant})$   $\tau_f^{3/2} v_f = \tau_i^{3/2} v_i$   $\tau_f = \tau_i (\frac{v_i}{v_f})^{2/3}$   $PV = N \tau$   $\Delta U = \frac{3}{2} N \Delta \tau$  energy going out of the system is via the work Sudden expansion into a vacuum  $W = 0$  and  $Q = 0$  and  $\Delta \tau = 0$  and  $U_f = U_i$  and  $U = \frac{3}{2} N \tau$  and  $\sigma = N[\ln V + \ln \tau^{3/2} + \text{constant}]$  - Reversible process is an approximation  $\Delta \sigma = N \ln \frac{V_f}{V_i}$  **Chapter 7 Fermi and Bose Gas** For classical ideal gas  $\mu = \tau \ln \frac{n}{n_Q}$   $f_{Boltzman} = \frac{n}{n_Q} e^{\epsilon/\tau} < 1$   $\frac{n}{n_Q} < 1$  classical regime  $\geq 1$  quantum In cases where n is fixed and  $\tau$  adjustable  $\tau_0 = (\frac{2 \pi \hbar^2}{M}) n^{2/3}$  Fermi gas  $\langle N \rangle = \sum_{orbital} f_{orbital}(\mu, \tau)$   $\epsilon_F = \frac{\hbar^2 m}{2m} (3 \pi^2 n)^{2/3}$  what is total energy at ground state  $U = 2 \sum_{n < n_F} \epsilon_n = 2 \cdot \frac{1}{8} \cdot 4 \pi \int_{n_F}^{n_Q} d n n^2 \epsilon_n = \frac{5}{2} N \epsilon_F$  Fermi gas How to determine  $\epsilon_F$   $\epsilon_F = \frac{\hbar^2 m}{2m} (3 \pi^2 \frac{N}{V})^{2/3}$  Keep N constant  $V_1$  and  $\epsilon_F \uparrow$   $\frac{U}{N} = \frac{3}{5} \epsilon_F \uparrow$  symmetry -i effective repulsion **w14 day 1** Fermi gas  $\langle X \rangle = \int \epsilon D(\epsilon) f(\epsilon, \tau, \mu) x(\epsilon)$  How to calculate  $D(\epsilon)$  density of states  $N(\epsilon) = (\frac{V}{3 \pi^2}) (\frac{2 m}{\hbar^2})^{3/2} (\epsilon)^{3/2}$  total number of states under  $\epsilon$   $D(\epsilon) = \frac{\partial N}{\partial \epsilon} = \frac{\partial N}{\partial \epsilon} = \frac{V}{(2 \pi^2)} (\frac{2 m}{\hbar^2})^{3/2} \epsilon^{1/2}$  In high  $\tau$  classical  $c_v \approx \frac{3}{2} K_B N$   $c_v \approx N \frac{c}{\tau} \propto \tau$  quantum what is low  $\tau$   $\epsilon_F = \frac{\hbar^2 m}{2m} (3 \pi^2 n)^{2/3}$  for Cu :  $n = 8.7 \times 10^{28} m^{-3}$  for room temperature  $\frac{c_v}{\tau} \approx 260 >> 1$  **w14 day 2 Bose Gas and Bose Einstein**  $N = \int \partial \epsilon f(\epsilon, \tau, \mu) D(\epsilon)$  ground state ( $\tau = 0$ ) all particle occupy the ground orbital  $\rightarrow \epsilon = 0$   $f(\epsilon = 0, \tau) = \frac{1}{e^{-\mu/\tau} - 1} \geq 0$   $\mu \leq 0$  At  $\tau$  small  $f(0, \tau) \approx \frac{\tau}{(1 - \frac{\tau}{\tau_c}) - 1} = \frac{\tau}{\tau_c}$   $\frac{\tau}{\tau_c} \approx N$  close to zero  $\frac{U}{N} \approx \frac{3}{2} \tau + 0 \frac{1}{N}$  take thermodynamic limit at the ver last state occupancy  $N_0 + N_e = N$   $N_0(\tau, \mu) = \frac{V}{e^{-\mu/\tau} - 1} = \frac{1}{\lambda^{-1} - 1}$   $N_e(\tau, \mu) \approx \frac{V}{4 \pi} (\frac{2 m}{\hbar^2})^{3/2} \tau^{3/2} \int_0^\infty dx \frac{x^{1/2}}{e^{x^4} - 1}$   $\frac{N_0}{N} = 1 - \frac{N_e}{N}$  BE condensation  $\tau_E = \frac{2 \pi \hbar^2}{M} (\frac{2 m}{0.012 V})^{2/3}$   $\frac{N_0}{N} = 1 - (\frac{T}{T_E})^{3/2}$  The critical temperature does not scale linearly with N anymore **chapter 8 hear transfer and work are path dependent**  $\partial U = \partial Q + \partial w$   $dQ \equiv \tau \partial \sigma$   $\partial w = -p \partial V$  heat transfer and work are path dependent, work 100 percent to heat, however not hear to work, there is directionality pay price from going from disorder to order **w15 day 1 thermal engine** First Law  $\partial U = Q + W$  Second law: difference between heat and work  $(\Delta \sigma_{total}) = (\Delta \sigma)_{RL} + (\Delta \sigma)_s = \frac{-\Delta w}{\tau} + \frac{-\Delta w}{Q_s} + 0$  Efficiency  $\eta = \frac{\Delta w}{Q_s} = \frac{2 \tau_c - Q_s}{Q_s} = 1 - \frac{Q_s}{\tau_c}$  For a reversible cycle  $\eta = 1 - \frac{\tau_c}{\tau_h}$  upper limit for efficiency of thermal energy  $\eta_c = \text{carnot efficiency} = 1 - \frac{\tau_c}{\tau_h}$   $\eta \leq \eta_c$  Refrigerator/air conditioner  $(\Delta \sigma)_{total entropy} = \frac{\Delta w}{\tau_h} - \frac{\Delta w}{\tau_c} \geq 0$   $\Rightarrow \gamma \leq \gamma_c$   $\gamma_c = \frac{\tau_c - \tau_h}{\tau_h - \tau_c}$  achievable only for a reversible process isothermal expansion, isothermal expansion, realization of carnot engine classification ideal gas -i reversible cycle **w15 day 2** isothermal  $Q_{12} = \tau_h (\sigma_h - \sigma_l)$  and  $Q_{34} = \tau_l (\sigma_h - \sigma_l)$  isentropic  $Q_{23} = 0$  and  $Q_{41} = 0$   $W = |Q_{12}| - |Q_{34}| = (\tau_h - \tau_l) \cdot \Delta \sigma = \text{area}$  isothermal  $PV = N \tau$  isentropic  $PV^\gamma = \text{constant}$  where  $\gamma > 1$   $1 \rightarrow 2$  Isothermal expansion  $\Delta = 0$  and  $PV = \text{constant}$  and  $P = \text{Cont} \cdot \frac{1}{V}$  and  $|Q_{12}| = |W_{12}| = \int p \partial V = N \tau_N \ln \left(\frac{v_2}{v_1}\right)$   $\eta = \frac{W}{Q_h} = \frac{W}{Q_h} = n \tau_h \ln \frac{v_2}{v_1}$   $W = |Q_h| - |Q_l| = N \tau \ln \left(\frac{v_2}{v_1}\right) - N \tau \ln \left(\frac{v_2}{v_4}\right)$   $\eta = \frac{W}{Q_h} = \frac{\tau_h - \tau_c}{\tau_h} = 1 - \frac{\tau_c}{\tau_h}$  carnot cycle has carnot efficiency **Irreversible cycle**  $\Delta \tau = (\tau_R - \tau_D) \approx 0$  reversible cycle  $\Delta \sigma_s = \int \frac{-\partial \sigma}{\tau}$  in reversible there is no net entropy increase  $\Delta \sigma_R = \int \frac{-\partial \sigma}{\tau}$   $\Delta \sigma_{total} = 0$  **w16 day 1 chemical work**  $\partial U = \partial Q + \partial w - p \partial v$  Chemical work: work performed by the transfer of particles  $U(s, V, N) \approx \partial U = \tau \partial \sigma - p \partial v + \mu \partial N$  chemical work  $\partial w = -p \partial v + \mu \partial N$  Assuming no mechanical work for  $\partial w_i = \partial w_{c1} + \partial w_{c2} = M_2 \partial N_1 + M_2 \partial N_2$  - total number of particles ideal gas  $\mu = \tau \ln \left(\frac{n}{n_Q}\right)$   $\mu_2 + \mu_1 = \tau \ln \left(\frac{n_1}{n_1}\right)$   $W_c = \tau \ln \left(\frac{n_1}{n_1}\right)$  Why  $G \partial G = \partial U - \partial F(\sigma)$  and  $(\partial p v)$  What happens for spontaneous charges  $\Delta \sigma_{total} = \Delta \sigma_s - \frac{Q_p}{\tau_c} = Q_{L2} \leq W_1$  does not require reversibility Simple case if  $\omega_1^+ = 0$  the process is reversible,  $\Delta G_s = 0$   $G_s$  reaches minimum when equilibrium  $W, W_s \neq 0$  and  $\partial w_1^+ = \mu \partial N$   $G_F - G_L \leq \omega_1^+$  the maximum amount of work the system can deliver in Gibbs free energy. Now lets  $\partial w^+ = \mu \partial N$  Extensive and Intensive Quantity  $G_{total} = G_{1+2} = G_1 + G_2 = 2G$ ;  $U_{total} = U_1 + U_2 = 2U$ ; Intensive quantities, no dependence on number of particles,  $\tau$  temperature,  $P$  pressure,  $N$  density,  $\mu$  chemical potential extensive quantities, dependence on number of particles,  $\sigma$  entropy,  $F$  free energy,  $V$  volume,  $N$  number of particles Example of simple ideal gas  $\mu = \tau \ln \left(\frac{n}{n_Q}\right) = \tau \ln \left(\frac{p}{p_Q}\right)$   $F(T, V, \mu) = N \tau [\ln \left(\frac{n}{n_Q}\right) - 1]$  Gibbs free energy  $G = F + pV = N \tau \ln \left(\frac{n}{n_Q}\right) = N \mu$