# TTK4130 - Cheat Sheet

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# https://github.com/haakonbaa/TTK4130-cheatsheet

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## 1 Intro

# 1.2 State space methods

## 1.2.1 State space models

State Space Model is on the form

$$\dot{\boldsymbol{x}} = \boldsymbol{f}(\boldsymbol{x}, \boldsymbol{u}, t)$$

Linear time invariant system

$$\dot{x} = Ax + Bu$$
  
 $y = Cx + Du$   
 $y = Ce^{At}x(0) + \int_0^t Ce^{A(t-\tau)}Bu(t) d\tau + Du(t)$ 

# 1.2.2 Second order models of mechanical systems

Second order models on the form

$$M(q)\ddot{q} + f(q, \dot{q}) = u$$

Can be written as

$$\begin{pmatrix} \dot{\boldsymbol{q}} \\ \dot{\boldsymbol{q}} \end{pmatrix} = \begin{pmatrix} \dot{\boldsymbol{q}} \\ \boldsymbol{M}^{-1}(\boldsymbol{q})(-\boldsymbol{f}(\boldsymbol{q},\dot{\boldsymbol{q}}) + \boldsymbol{u}) \end{pmatrix}$$

#### 1.2.3 Linearization of state space models

Linearization of time varying systems

$$\dot{x} = f(x, u, t)$$
  
 $y = h(x, u, t)$ 

Find two functions  $\boldsymbol{x}_0$  and  $\boldsymbol{u}_0$  begin solutions to the sytem

$$\dot{\boldsymbol{x}}_0 = \boldsymbol{f}(\boldsymbol{x}_0(t), \boldsymbol{u}_0(t), t)$$

Define pertrubations

$$egin{aligned} oldsymbol{x}(t) &= oldsymbol{x}_0(t) + oldsymbol{\Delta} oldsymbol{x}(t) \ oldsymbol{u}(t) &= oldsymbol{u}_0(t) + oldsymbol{\Delta} oldsymbol{u}(t) \ oldsymbol{y}(t) &= oldsymbol{y}_0(t) + oldsymbol{\Delta} oldsymbol{y}(t) \end{aligned}$$

Let  $C = \{x_0(t), u_0(t)\}$ . The linearized system is

$$egin{aligned} oldsymbol{\Delta}\dot{oldsymbol{x}}&pproxrac{\partial oldsymbol{f}}{\partialoldsymbol{x}}\Big|_{\mathcal{C}}oldsymbol{\Delta}oldsymbol{x}+rac{\partial oldsymbol{f}}{\partialoldsymbol{u}}\Big|_{\mathcal{C}}oldsymbol{\Delta}oldsymbol{u} \ oldsymbol{\Delta}\dot{oldsymbol{y}}&pproxrac{\partial oldsymbol{h}}{\partialoldsymbol{x}}\Big|_{\mathcal{C}}oldsymbol{\Delta}oldsymbol{x}+rac{\partial oldsymbol{h}}{\partialoldsymbol{u}}\Big|_{\mathcal{C}}oldsymbol{\Delta}oldsymbol{u} \end{aligned}$$

#### 1.5 ODE's

General formulation

$$\varphi(y^{(m)},\cdots,y,u^{(m-1)},\cdots,u)=0$$

Lipschitz continous. A function

$$f: \mathbb{R} \to \mathbb{R}$$

is said to be Lipschitz continous if

$$\exists c > 0 \in \mathbb{R}$$
 s.t. 
$$||f(x) - f(y)|| < c||x - y|| \qquad \forall x, y \in \mathbb{R}$$

Theorem: existence of unique solution. Consider the ODE

$$\dot{x} = f(x)$$

If f is Lipschitz continous then x(t) exists and is unique for all t

**Mean Value Theorem** suppose f is continuous on [a,b] and differentiable on (a,b) then

$$\exists c \in (a, b) \text{ s.t. } f'(c) = \frac{f(b) - f(a)}{b - a}$$

This can be sused to show that if f is continous and differentiable everywhere it is also Lipschitz.

Theorem 2: existence of unique solution. Consider the ODE

$$\dot{x} = f(x)$$

if f is continously differentiable ( $\frac{\partial f}{\partial x}$  exists and is continous), then the solution to the ODE exists and is unique on some time interval.

#### 2 Rotations

### Vectors

The skew-symetric matrix form of the coordinate vector  $\mathbf{u}$  is defined by

$$\mathbf{u}^x = \begin{pmatrix} 0 & -u_3 & u_2 \\ u_3 & 0 & -u_1 \\ -u_2 & u_1 & 0 \end{pmatrix}$$

Notation:  $v_{ab}^c$  means the vector from point a to point b (or often the origo of the reference frames a and b) described in the reference frame c

#### 6.4 The Rotation Matrix

The coordinate transformation from frame b to frame a is given by

$$oldsymbol{v}^a = oldsymbol{R}_{\iota}^a oldsymbol{v}^b$$

Properties of the rotation matrix

$$egin{aligned} oldsymbol{R}_a^b oldsymbol{R}_b^a &= oldsymbol{I} = oldsymbol{R}_b^a oldsymbol{R}_a^b \\ oldsymbol{(R}_a^b)^{-1} &= oldsymbol{(R}_a^b)^T = oldsymbol{R}_b^a \\ oldsymbol{R}_b^a &= oldsymbol{(b}_1^a \quad oldsymbol{b}_2^a \quad oldsymbol{b}_3^a) \ \det oldsymbol{R}_a^b &= 1 \end{aligned}$$

R is a rotation matrix if and only if it is an element of SO(3)

$$SO(3) = \{ \boldsymbol{R} \in \mathbb{R}^{3 \times 3} | \boldsymbol{R}^T \boldsymbol{R} = \boldsymbol{I} \wedge \det \boldsymbol{R} = 1 \}$$

Rotation matrices in three dimentions

$$\mathbf{R}_{x}(\phi) = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \phi & -\sin \phi \\ 0 & \sin \phi & \cos \phi \end{pmatrix}$$
$$\mathbf{R}_{y}(\theta) = \begin{pmatrix} \cos \theta & 0 & \sin \theta \\ 0 & 1 & 0 \\ -\sin \theta & 0 & \cos \theta \end{pmatrix}$$
$$\mathbf{R}_{z}(\psi) = \begin{pmatrix} \cos \psi & -\sin \psi & 0 \\ \sin \psi & \cos \psi & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

Matrix transformations in different refrence frames

$$egin{aligned} oldsymbol{D}^a &= oldsymbol{R}_b^a oldsymbol{D}^b oldsymbol{R}_a^b \ &(oldsymbol{u}^b)^ imes &= oldsymbol{R}_a^b (oldsymbol{u}^a)^ imes oldsymbol{R}_h^a \end{aligned}$$

The transformation of position and orientation from frame b to frame a is

$$egin{aligned} m{T}_b^a &= egin{pmatrix} m{R}_b^a & m{r}_{ab}^a \ m{0}^T & 1 \end{pmatrix} \ m{T}_b^a egin{pmatrix} m{v}^b \ 1 \end{pmatrix}^T &= m{v}^a \ 1 \end{pmatrix} \ &(m{T}_b^a)^{-1} &= m{T}_a^b &= m{K}_a^b & m{r}_{ba}^b \ m{0}^T & 1 \end{pmatrix} \end{aligned}$$

The Special Euclidean group is the set of all transfor- Quaternion product mations from one reference frames to another

$$SE(3) = \left\{ \boldsymbol{T} = \begin{pmatrix} \boldsymbol{R} & \boldsymbol{r} \\ \boldsymbol{0}^T & 1 \end{pmatrix} \in \mathbb{R}^{3 \times 3} \middle| \boldsymbol{R} \in SO(3) \land \boldsymbol{r} \in \mathbb{R}^3 \right\} \qquad \begin{pmatrix} \alpha_1 \\ \beta_1 \end{pmatrix} \otimes \begin{pmatrix} \alpha_2 \\ \beta_2 \end{pmatrix} = \begin{pmatrix} \alpha_1 \alpha_2 - \boldsymbol{\beta}_1^T \boldsymbol{\beta}_2 \\ \alpha_1 \boldsymbol{\beta}_2 + \alpha_2 \boldsymbol{\beta}_1 + \boldsymbol{\beta}_1^\times \boldsymbol{\beta}_2 \end{pmatrix}$$

#### 6.5 **Euler Angles**

### Roll-Pitch-Yaw Euler angles

$$\mathbf{R}_{a}^{b} = \mathbf{R}_{z}(\psi)\mathbf{R}_{y}(\theta)\mathbf{R}_{x}(\phi)$$

Classical Euler angles. The orientation is described by a rotation bout the z axis, then the resulting y axis. And then again the resulting z axis.

$$\mathbf{R}_{a}^{b} = \mathbf{R}_{z}(\psi)\mathbf{R}_{u}(\theta)\mathbf{R}_{z}(\phi)$$

#### Angle Axis Description of rotation 6.6

#### 6.6.5**Rotation Matrix**

Angle-axis parameters All rotation matrices have an eigen vector with eigen value 1. A rotation can be uniquely described by the direction of this vector and an angle  $\theta$  being the rotatoin about this vector.

$$(\theta, \mathbf{k}) \text{ s.t. } ||\mathbf{k}|| = 1$$
  
 $\mathbf{R}_b^a = \cos \theta \mathbf{I} + \sin \theta (\mathbf{k}_a)^{\times} + (1 - \cos \theta) \mathbf{k}_a \mathbf{k}_a^T$   
 $\mathbf{R}_b^a = \exp{\{\mathbf{k}^{\times} \theta\}}$ 

#### 6.7Euler parameters

#### 6.7.1Definition

$$egin{aligned} \eta &= \cos rac{ heta}{2} \ oldsymbol{\epsilon} &= oldsymbol{k} \sin rac{ heta}{2} \ oldsymbol{R}_e(\eta, oldsymbol{\epsilon}) &= oldsymbol{I} + 2\eta oldsymbol{\epsilon}^{ imes} + 2oldsymbol{\epsilon}^{ imes} oldsymbol{\epsilon}^{ imes} \ \eta^2 + oldsymbol{\epsilon}^T oldsymbol{\epsilon} &= 1 \end{aligned}$$

#### Quaternions

The following can be treated as a unit quaternion

$$oldsymbol{p} = egin{pmatrix} \eta \ oldsymbol{\epsilon} \end{pmatrix}$$

A unit quaternion satisfies

$$\boldsymbol{p}^T\boldsymbol{p} = \eta^2 + \boldsymbol{\epsilon}^T\boldsymbol{\epsilon} = 1$$

$$\begin{pmatrix} \alpha_1 \\ \boldsymbol{\beta}_1 \end{pmatrix} \otimes \begin{pmatrix} \alpha_2 \\ \boldsymbol{\beta}_2 \end{pmatrix} = \begin{pmatrix} \alpha_1 \alpha_2 - \boldsymbol{\beta}_1^T \boldsymbol{\beta}_2 \\ \alpha_1 \boldsymbol{\beta}_2 + \alpha_2 \boldsymbol{\beta}_1 + \boldsymbol{\beta}_1^{\times} \boldsymbol{\beta}_2 \end{pmatrix}$$

# 6.7.6 Euler parameters from the rotation matrix

$$\mathbf{R} = (r_{ij})$$

$$\mathbf{z} = \begin{pmatrix} z_0 & z_1 & z_2 & z_3 \end{pmatrix}^T := 2 \begin{pmatrix} \eta & \epsilon_1 & \epsilon_2 & \epsilon_3 \end{pmatrix}^T$$

$$\mathbf{T} := r_{00} := \text{Trace } \mathbf{R}$$

The algorithm from Shepperd (1978) goes like this:

- Let  $i = \arg \max_i \{r_{ii}\}$
- Compute  $|z_i| = \sqrt{1 + 2r_{ii} T}$
- Determine sign of  $z_i$
- ullet Determine the rest of z from equations below

$$z_0 z_1 = r_{32} - r_{23}$$
  $z_2 z_3 = r_{32} + r_{23}$   
 $z_0 z_2 = r_{13} - r_{31}$   $z_3 z_1 = r_{13} + r_{31}$   
 $z_0 z_3 = r_{21} - r_{12}$   $z_1 z_2 = r_{21} + r_{12}$ 

## 6.8 Angular Velocity

Let  $R \in SO(3)$ 

$$0 = \frac{d}{dt}(\mathbf{I}) = \frac{d}{dt}(\mathbf{R}\mathbf{R}^T) = \dot{\mathbf{R}}\mathbf{R}^T + \mathbf{R}(\dot{\mathbf{R}})^T$$

$$\Rightarrow \dot{\mathbf{R}}\mathbf{R}^T skew\text{-symmetric}$$

Definition of angluar velocity

$$egin{aligned} (oldsymbol{\omega}_{ab}^a)^{ imes} &= \dot{oldsymbol{R}}_b^a (oldsymbol{R}_b^a)^T \Rightarrow \ \dot{oldsymbol{R}}_b^a &= (oldsymbol{\omega}_{ab}^a)^{ imes} oldsymbol{R}_b^a \ \dot{oldsymbol{R}}_b^a &= oldsymbol{R}_b^a (oldsymbol{\omega}_{ab}^b)^{ imes} \end{aligned}$$

It can be shown that

$$\omega = \dot{\theta} k$$

Where  $\theta$  and k are Angle Axis parameters.

$$egin{aligned} oldsymbol{\omega}_{ad}^a &= oldsymbol{\omega}_{ab}^a + oldsymbol{\omega}_{bc}^a + oldsymbol{\omega}_{cd}^a \ \dot{oldsymbol{u}}^a &= oldsymbol{R}_b^a (\dot{oldsymbol{u}}^b + (oldsymbol{\omega}_{ab}^b)^{ imes} oldsymbol{u}^b) \end{aligned}$$

# 6.9 Kinematic differential equations

#### 6.9.4 Euler Angles

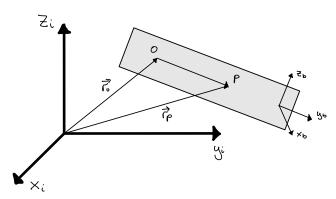
$$\omega_{ad}^{a} = \begin{pmatrix} 0 \\ 0 \\ \dot{\psi} \end{pmatrix} + \mathbf{R}_{z,\psi} \begin{pmatrix} 0 \\ \dot{\theta} \\ 0 \end{pmatrix} + \mathbf{R}_{z,\psi} \mathbf{R}_{y,\theta} \begin{pmatrix} \dot{\phi} \\ 0 \\ 0 \end{pmatrix}$$
$$= \begin{pmatrix} -\sin\psi\dot{\theta} + \cos\psi\cos\theta\dot{\phi} \\ \cos\psi\dot{\theta} + \sin\psi\cos\theta\dot{\phi} \\ \dot{\psi} - \sin\theta\dot{\phi} \end{pmatrix}$$

# 3 Rigid Body Dynamics

# 6.12 Kinematics of a rigid body

 $\vec{\omega}_{io}$  is the angular velocity of the o frame with respect to the i frame.

 $\frac{i}{dt}\vec{r}_o$  is the derivative of  $\vec{r}_o$  in the *i* frame.



Velocity and Acceleration

$$\begin{split} \vec{v}_p &:= \frac{^i d}{dt} \vec{r}_p \\ &= \vec{v}_o + \frac{^b d}{dt} \vec{r} + \vec{\omega}_{ib} \times \vec{r} \\ \vec{a}_p &:= \frac{^i d^2}{dt^2} \vec{r}_p \\ &= \vec{a}_o + \frac{^b d^2}{dt^2} \vec{r} + 2 \vec{\omega}_{ib} \times \frac{^b d}{dt} \vec{r} + \vec{\alpha}_{ib} \times \vec{r} + \vec{\omega}_{ib} \times (\vec{\omega}_{ib} \times \vec{r}) \end{split}$$

The last three terms are, respectively, the coriolis acceleration, Transveral acceleration and Centripetal acceleration. Note that

$$\vec{a}_o = \frac{^i d}{dt} \vec{v}_o = \frac{^b d}{dt} \vec{v}_o + \vec{\omega}_{ib} \times \vec{v}_o$$

## 6.13 The center of mass

The center of mass of a rigid body  $\mathcal{C}$  is defined to be

$$\vec{r}_c := \frac{1}{m} \int_{\mathcal{C}} \vec{r}_p \, dm$$

It can be shown that

$$\vec{v}_c = \frac{1}{m} \int_{\mathcal{C}} \vec{v_p} \, dm$$
  $\vec{a}_c = \frac{1}{m} \int_{\mathcal{C}} \vec{a_p} \, dm$ 

where c denotes center

## 7.2 Forces and torques

**Moment.** The moment about a point P of the set  $S = \{F_j\}_{j \in [1, n_F]}$  for forces is

$$\vec{N}_{S/P} = \sum_{j=1}^{n_F} r_{Pj} \times \vec{F}_j$$

Where  $\vec{r}_{Pj}$  is an arbitrary point along the line of action of  $\vec{F}_i$ 

**Torque** is defined as the moment of the couple C. A couple being a set of forces with  $\mathbf{0}$  resultant force.

# 7.3 Newton-Euler Equations for rigid bodies

**Angular Momentum**. The angular momentum of the body b about the center of mass c is

$$egin{aligned} m{h}_{b/c} &= \int_b m{r} imes m{v} \, dm \ &= m{M}_{b/c} m{\omega}_{ib} \ m{T}_{bc} &= rac{d}{dt} m{h}_{b/c} \end{aligned}$$

Rotational Inertia / The intertia dyadic. The inertia matrix of the body b about the point c is

$$M_{b/c} = -\int_{b} \mathbf{r}^{\times} \mathbf{r}^{\times} dm$$

$$= \int_{b} (\mathbf{r}^{T} \mathbf{r} \mathbb{I} - \mathbf{r} \mathbf{r}^{T}) dm$$

$$= \begin{pmatrix} \mathbf{I}_{xx} & -\mathbf{I}_{xy} & -\mathbf{I}_{xz} \\ -\mathbf{I}_{xy} & \mathbf{I}_{yy} & -\mathbf{I}_{yz} \\ -\mathbf{I}_{xz} & -\mathbf{I}_{yz} & \mathbf{I}_{zz} \end{pmatrix}$$

Where r is the distance vector from the center of mass to the mass element being integrated

$$I_{xx} = \int_b y^2 + z^2 dm$$
  $I_{xy} = \int_b xy dm$   $I_{yy} = \int_b x^2 + z^2 dm$   $I_{xz} = \int_b xz dm$   $I_{zz} = \int_b x^2 + y^2 dm$   $I_{yz} = \int_b yz dm$ 

$$\boldsymbol{M}_{b/c}^i = \boldsymbol{R}_b^i \boldsymbol{M}_{b/c}^b \boldsymbol{R}_i^b$$

Equations of motion. Let b denote body, i an intertial frame, c the center of mass of b,  $\mathbf{F}_{bc}$  a resultant force acting on b with line of action through c and  $\mathbf{T}_{bc}$  the torque about c. Then

$$egin{aligned} F_{bc} &= mm{a}_c \ T_{bc} &= m{M}_{b/c}m{lpha}_{ib} + m{\omega}_{ib} imes (m{M}_{b/c}m{\omega}_{ib}) \end{aligned}$$

On compact matrix form

$$egin{pmatrix} egin{pmatrix} m{m} \mathbb{I} & \mathbf{0} \ \mathbf{0} & m{M}_{b/c}^b \end{pmatrix} m{igg(m{a}_c^b \ m{lpha}_{ib}^b \end{pmatrix}} + m{igg(m{0} \ m{\omega}_{ib}^b)^ imes m{M}_{b/c}^b m{\omega}_{ib}^b \end{pmatrix}} = m{igg(m{F}_{bc}^b \ m{T}_{b/c}^b m{\omega}_{ib}^b m{\omega$$

**Kinetic energy**. The kinetic energy of the body b in an inertial refrence frame i is

$$K = \frac{1}{2}m(\boldsymbol{v}_c^b)^T\boldsymbol{v}_c^b + \frac{1}{2}(\boldsymbol{\omega}_{ib}^b)^T\boldsymbol{M}_{b/c}^b\boldsymbol{\omega}_{ib}^b$$

The parallel axes theorem. The inertia matrix of b about o is related to the inertia matrix of b about c according to

$$egin{aligned} oldsymbol{M}_{b/o}^b &= oldsymbol{M}_{b/c}^b - m(oldsymbol{r}_g^b)^ imes (oldsymbol{r}_g^b)^ imes \ &= oldsymbol{M}_{b/c}^b - m(||oldsymbol{r}_g^b||^2 \mathbb{I} - oldsymbol{r}_g^b (oldsymbol{r}_g^b)^T) \end{aligned}$$

Where  $\boldsymbol{r}_g^b$  is the vector from c to o

# 4 Lagrange Mechanics

## 8.2 Lagrange Mechanics

The lagrangian. Define a set of generalized coordinates  $\mathbf{q}$ . Let  $T(\mathbf{q}, \dot{\mathbf{q}}, t)$  be the kinetic energy and  $U(\mathbf{q}, \dot{\mathbf{q}}, t)$  the potential energy (Sometimes V). Then the lagrangian is defined to be

$$L(\boldsymbol{q}, \dot{\boldsymbol{q}}, t) = T(\boldsymbol{q}, \dot{\boldsymbol{q}}, t) - U(\boldsymbol{q})$$

The equation of motion is

$$\frac{d}{dt} \left( \frac{\partial L}{\partial \dot{q}_i} \right) - \frac{\partial L}{\partial q_i} = \tau_i$$

$$\frac{d}{dt} \left( \nabla_{\dot{q}} L \right) - \nabla_{\dot{q}} L = \tau$$

Where  $\tau_i$  is a generalized actuator force

$$\tau_i = \sum_{k=1}^{N} \frac{\partial \boldsymbol{r_k}}{\partial q_i} \cdot \boldsymbol{F_k}$$

 $r_k(q)$  is the position of the point of application of the force  $F_k$ . In general  $\tau_i$  is a force or a torque.

Constrained Lagrange. Having the constraints

$$c(q) = 0$$

The system can be described by

$$L(\boldsymbol{q}, \dot{\boldsymbol{q}}, t) = T(\boldsymbol{q}, \dot{\boldsymbol{q}}, t) - U(\boldsymbol{q}) - \boldsymbol{z}^T \boldsymbol{c}(\boldsymbol{q})$$
$$\frac{d}{dt} (\nabla_{\dot{\boldsymbol{q}}} L) - \nabla_{\boldsymbol{q}} L = \boldsymbol{\tau}$$
$$\boldsymbol{c}(\boldsymbol{q}) = \boldsymbol{0}$$

Baumgarte stabilization. Instead of imposing

$$\ddot{\boldsymbol{c}}(\boldsymbol{q}) = \mathbf{0}$$

Impose

$$\ddot{\boldsymbol{c}} + 2\alpha\dot{\boldsymbol{c}} + \alpha^2\boldsymbol{c} = \boldsymbol{0}$$

As to reduce drifts in the constraints resulting from

$$\ddot{c} = 0$$

not begin satisfied exactly when doing numeric computations.

#### 5 Differential Algebraic Equa- Theorem: Solvability of DAE. A fully implicit tions

#### 14.2 **Preliminaries**

# Differential Algebraic Equations

**Definition of DEA**. The differential equation defined

$$F(\dot{x}, x, u, t) = 0$$

Es a DAE if

$$\frac{\partial \boldsymbol{F}}{\partial \dot{\boldsymbol{x}}}$$

is rank defficient

## Fully-explicit DAE

$$oldsymbol{F}(\dot{oldsymbol{x}},oldsymbol{x},oldsymbol{z},oldsymbol{z},oldsymbol{u})=\mathbf{0}$$

$$\det \left| \frac{\partial \mathbf{F}}{\partial \dot{\mathbf{x}}} \right| = 0$$

Can be rewritten as

$$\dot{x} = v$$
 $0 = F(v, x, z, u)$ 

#### Semi-explicit DAE

$$\dot{x} = f(x, z, u)$$
 $0 = g(x, z, u)$ 

This can be rewritten as

$$\boldsymbol{F}(\dot{\boldsymbol{x}},\boldsymbol{x},\boldsymbol{z},\boldsymbol{u}) = \begin{pmatrix} \dot{\boldsymbol{x}} - \boldsymbol{f}(\boldsymbol{x},\boldsymbol{z},\boldsymbol{u}) \\ \boldsymbol{g}(\boldsymbol{x},\boldsymbol{z},\boldsymbol{u}) \end{pmatrix} = \boldsymbol{0}$$

Tikhonov Theorem. Consider the ordinary differential equation

$$\dot{m{x}} = m{f}(m{x},m{z})$$

$$\varepsilon \dot{\boldsymbol{z}} = \boldsymbol{g}(\boldsymbol{x}, \boldsymbol{z})$$

If

- dynamics of  $\dot{z} = g(x, z)$  stable  $\forall x$
- $\frac{\partial \mathbf{g}}{\partial \mathbf{z}}$  is full rank

then

$$\lim_{\epsilon \to 0} \boldsymbol{x}(t) = \boldsymbol{x}_0(t)$$

$$\lim_{\epsilon \to 0} \boldsymbol{z}(t) = \boldsymbol{z}_0(t)$$

where  $z_0(t)$  and  $x_0(t)$  is the solution to the ODE above modified to a DAE where  $\varepsilon=0$ 

DAE with smooth

$$F(\dot{x}, x, z, u) = 0$$

Can be readily solved (solved for  $\dot{x}$  and z) if

$$\begin{pmatrix} \frac{\partial \mathbf{F}}{\partial \dot{\mathbf{x}}} & \frac{\partial \mathbf{F}}{\partial \mathbf{z}} \end{pmatrix}$$

is full rank on all trajectories  $\dot{x}$ , z, x and u. Note that all Index 1 DAEs fullfil these requirements. The theorem implies that

$$\dot{\boldsymbol{x}} = \boldsymbol{f}(\boldsymbol{x}, \boldsymbol{z}, \boldsymbol{u})$$

$$\mathbf{0} = g(x, z, u)$$

with smooth  $\boldsymbol{f}$  can be solved if

$$\frac{\partial \boldsymbol{g}}{\partial \boldsymbol{z}}$$

is full rank on all trajectories z, x and u

Definition: Differential index of a DAE is the numer of times the differentiation operator  $\frac{d}{dt}$  must be applied to the equations in order to convert the DAE into an ODE.

# 6 Simulation methods

#### Butcher tableau

$$K_n = f(t_k + \Delta t c_n, x_n + \Delta t(A_n)K)$$
$$x_{k+1} = x_k + \Delta t \mathbf{b}^T K$$

The method is **consistent** if

$$c_n = \Sigma_i A_{ni}$$
$$\Sigma_i \mathbf{b}_i = 1$$

The **stages** of a Runge Kutta method is the number of elements in c. The Butcher tableu defines an explicit integrator if and only if the diagonal elements and the upper-diagonal elements are zero.

#### Stability function For the Butcher Tableu

The stability function is

$$R(z) = \frac{\det |\mathbb{I} - z(\boldsymbol{A} - \boldsymbol{1}\boldsymbol{b}^T)|}{\det |\mathbb{I} - z\boldsymbol{A}|}$$

#### **Explicit Euler**

$$\boldsymbol{x}_{k+1} = \boldsymbol{x}_k + \Delta t \cdot \boldsymbol{f}(\boldsymbol{x}, \boldsymbol{u})$$
  $\frac{0 \mid 0}{\mid 1}$ 

Global error  $||\boldsymbol{x}_N - \boldsymbol{x}(T)|| = \mathcal{O}(\Delta t)$ Stability function R(z) = 1 + z

#### Explicit mid-point rule

$$\begin{array}{c|cccc}
0 & 0 & 0 \\
1/2 & 1/2 & 0 \\
\hline
& 0 & 1/2 \\
\end{array}$$

Global error  $||\boldsymbol{x}_N - \boldsymbol{x}(T)|| = \mathcal{O}(\Delta t^2)$ 

#### Ralston's RK2

$$\begin{array}{c|cccc}
0 & 0 & 0 \\
2/3 & 2/3 & 0 \\
\hline
& 1/4 & 1/3
\end{array}$$

#### Heun's RK2

$$\begin{array}{c|cccc}
0 & 0 & 0 \\
1 & 1 & 0 \\
\hline
& 1/2 & 1/2
\end{array}$$

#### Generic second-order method

$$\begin{array}{c|cccc}
0 & 0 & 0 \\
\alpha & \alpha & 0 \\
\hline
& 1 - \frac{1}{2\alpha} & \frac{1}{2\alpha}
\end{array}$$

#### Generic third-order method

## "The" RK4 method

Global error  $||\boldsymbol{x}_N - \boldsymbol{x}(T)|| = \mathcal{O}(\Delta t^4)$ 

## One-Step error

$$||x_{k+1} - x(t_{k+1})||$$

Taylor expand  $\boldsymbol{x}(t)$  around  $t_k$ . Do the same for  $\boldsymbol{x}_{k+1}$ . Subtract and get the results

**Global error** Get error on timestep T by multiplying local error with  $N = \frac{T}{\Delta t}$ 

stability of integration method: Consdier the ODE

$$\dot{x} = \lambda x \qquad \qquad x(0) = x_0$$

Perform one step of the RK method to find

$$x_{n+1} = R(z)x_n$$
  $z = \Delta t\lambda$ 

The integration method is unstable if

**A-stability** A method is A-stable if the region of stability is the entire left-half plane.

$$|R(\lambda \Delta t)| \le 1 \forall \lambda \in \mathbb{C} \text{ s.t. } \mathcal{R}(\lambda) \le 0$$

**L-stability**. A Runge kutta method is L-stable if it is A-stable and

$$\lim_{\omega \to \pm \infty} |R(i\omega h)| = 0$$

### Implicit methods

- Can chaieve high and systematic orders
- Can be stable regardless of the step size
- Can handle the simulation of DAEs (must be index 1)
- Can give huge speed-up when system is stiff
- Can achieve order o=2s for any stage s. Explicit methods can only achieve o=s and only for  $s\leq 4$

### Implicit Euler

$$\begin{array}{c|c}
1 & 1 \\
\hline
& 1
\end{array}$$

$$k_1 = f(x_n + \Delta t k_1)$$
  

$$x_{n+1} = x_n + \Delta t k_1$$
  

$$= x_n + \Delta t f(x_{n+1})$$

Gauss-Legendre collocation method Generates order o=2s A-stable IRK methods. Let s be the number of stages. Find the roots  $(\tau_i)$  of

$$P_s(\tau) = \frac{1}{s!} \frac{d^s}{d\tau^s} \left( (\tau^2 - \tau)^s \right)$$

Build polynomials

$$\ell_i(\tau) = \prod_{j \neq i} \frac{\tau - \tau_j}{\tau_i - \tau_j}$$

Integrate them

$$L_i(\tau) = \int_0^{\tau} \ell_i(\xi) \, d\xi$$

Calculate A , $\boldsymbol{b}$  and  $\boldsymbol{c}$ 

$$A_{ji} = L_i(\tau_j)$$
  $b_i = L_i(1)$   $c_j = \tau_j$ 

**Adaptive integrator**. Use two methods:  $x_{n+1}$  and  $\hat{x}_{n+1}$ , of different order: p and p+1. The local error is

$$\epsilon_{n+1} := |x_{n+1} - \hat{x}_{n+1}|$$

$$\approx \mathcal{O}(\Delta t^{p+1}) - \mathcal{O}(\Delta t^{p+2})$$

$$\approx C\Delta t^{p+1}$$

$$e_{\text{tol}} = C\Delta t_{\text{new}}^{p+1} \Rightarrow$$

$$\Delta t_{\text{new}} = \Delta t \left(\frac{e_{\text{tol}}}{\epsilon_{n+1}}\right)^{\frac{1}{p+1}}$$

Newtons method

$$egin{aligned} oldsymbol{f}(oldsymbol{x} + \Delta oldsymbol{x}) &pprox oldsymbol{f}(oldsymbol{x}) + rac{\partial oldsymbol{f}}{\partial oldsymbol{x}}(oldsymbol{x}) \Delta oldsymbol{x} &= oldsymbol{0} \left(rac{\partial oldsymbol{f}}{\partial oldsymbol{x}}(oldsymbol{x})
ight)^{-1} oldsymbol{f}(oldsymbol{x}) \ oldsymbol{x}_{n+1} &= oldsymbol{x}_n + lpha \Delta oldsymbol{x} \end{aligned}$$

- May not work if jacobian is singular
- Has quadratic convergence when it converges
- May diverge

# 7 Math

Inverse of  $2 \times 2$  Matrix

$$\begin{pmatrix} a & b \\ c & d \end{pmatrix}^{-1} = \frac{1}{ad - bc} \begin{pmatrix} d & -b \\ -c & a \end{pmatrix}$$

Multivariable derivative rules

$$egin{aligned} rac{\partial}{\partial oldsymbol{x}} oldsymbol{a}^T oldsymbol{x} = oldsymbol{a}^T & 
abla_{oldsymbol{x}} oldsymbol{a}^T oldsymbol{x} = oldsymbol{a}^T & 
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abla_{oldsymbol{x}} oldsymbol{A} oldsymbol{x} - oldsymbol{a}^T oldsymbol{a} = oldsymbol{a}^T & 
abla_{oldsymbol{x}} oldsymbol{A} oldsymbol{x} - oldsymbol{a}^T oldsymbol{a} - old$$

Second order terms

$$\frac{\partial}{\partial x} \frac{1}{2} x^T A x = \frac{1}{2} x^T (A^T + A)$$
$$\nabla_x \frac{1}{2} x^T A x = \frac{1}{2} (A^T + A) x$$

Multivariable Chain rule

$$\begin{split} \frac{\partial f(g(x))}{\partial x} &= \frac{\partial f}{\partial g} \frac{\partial g}{\partial x} \\ \frac{\partial f(g(x), h(x))}{\partial x} &= \frac{\partial f}{\partial g} \frac{\partial g}{\partial x} + \frac{\partial f}{\partial h} \frac{\partial h}{\partial x} \end{split}$$

Some derivatives

$$\frac{d}{dt}\sinh(t) = \cosh(t)$$

$$\frac{d}{dt}\cosh(t) = \sinh(t)$$

$$\frac{d}{dt}\tanh(t) = \frac{d}{dt}\frac{\sinh(t)}{\cosh(t)} = \frac{1}{\cosh^2 t} = 1 - \tanh^2(x)$$

$$\frac{d}{dt}\arctan(t) = \frac{1}{1+t^2}$$

### Taylor's theorem

Let  $k \geq 1$  be an integer and let the function  $f : \mathbb{R} \to \mathbb{R}$  be k+1 times differentiable at the point  $a \in \mathbb{R}$ . Then

$$f(x) = \sum_{n=0}^{k} \frac{1}{n!} f^{(n)}(a)(x-a)^n + \frac{1}{(1+k)!} f^{(k+1)}(\xi)(x-a)^{k+1}$$

for some  $\xi \in [x, a]$ 

Common Taylor expansions

$$e^{x} = \sum_{n=0}^{\infty} \frac{x^{n}}{n!}$$

$$\sin(x) = \sum_{n=1}^{\infty} (-1)^{n} \frac{x^{2n+1}}{(2n+1)!}$$

$$\cos(x) = \sum_{n=0}^{\infty} (-1)^{n} \frac{x^{2n}}{(2n)!}$$