

# Any title, doesnt really matter

Hunter Gabbard\* and Chris Messenger  
*SUPA, School of Physics and Astronomy,  
University of Glasgow,  
Glasgow G12 8QQ, United Kingdom*

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This may be an abstract one day.

## *Introduction.—*

With the recent detections of [some number] binary black holes [?] and binary neutron merger GW170817 [?] by the LIGO-Virgo Scientific Collaboration (LVC), the era of gravitational-wave astronomy has begun. The LVC is composed of three detectors in Hanford, Washington State, Livingston, Louisiana, and Pisa, Italy. Once the detectors will have reached design sensitivity, they will be able probe a search distance on the order of 200 Mpc. As the detectors become more sensitive, we will be limited by the sheer number of signals to process. As such, it is of paramount importance that we are able to get full parameter estimate posteriors on signals in a timely and efficient manner.

A variant of machine learning, generative adversarial networks (GANs), has gained some traction in recent years [?]. Some successful applications of GANs can be seen in image retrieval of historical archives [?], text to image synthesis [?], simulation of high energy particle showers [?], along with many others.

GANs are composed of two neural networks, a discriminator and a generator. The generator network attempts to map random latent variables to some distribution which the user wants the GAN to approximate, whereas the discriminator network will attempt to distinguish between samples from the real distribution and fake samples made by the generator network. You need only to feed into the generator as input a vector of latent random variables (can be uniform, Gaussian, etc.). The variables are latent in that they are not directly related to the distribution that the generator is trying to emulate, however in our case they do have an underlying Gaussian distribution. An additional motivation for using GANs is that they are also semi-supervised, so few training samples are needed, and you are left with by default two tools (one for classification and one for sample generation).

We use a variant of a GAN called deep convolutional generative adversarial networks (DCGAN) [?]. DCGANs are composed of two networks which act as the generator and discriminator, however the difference between DCGANs and their more standard GAN brethren is that both networks are entirely made of convolutional layers. We follow the architecture guidelines laid out in Radford et al. [?] which are as follows: re-

place pooling (downsampling) layers with strided convolutions (alternative to pooling), use batch normalization in both the generator and the discriminator, remove all fully connected hidden layers, use ReLU activation functions [?] in the generator except for the output, and use LeakyReLU activation functions [?] in the discriminator for all layers.

To-do:

- GWs.
- Bayesian Parameter Estimation.

## *Methods—*

To-do:

- How do you incorporate priors.
- Convergence.
- Volume of training data.
- Modifications from standard GAN (most of section).

## *Results—*

To-do:

GW150914. Show PE estimates on mass, spins, etc. Small section on waveform reconstruction. Plot of this as well.

- Which priors were used. Same as GW150914 analysis.

## *Conclusions—*

To-do:

Precision that we get, speed (sell this hard),

- Sell speed, with the right caveats.
- Waveform reconstruction.
- We are not model independent.
- Non-Gaussian noise.

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