

ASTC02 - LECTURE 1 - PROF. HANNO REIN

COORDINATE SYSTEMS

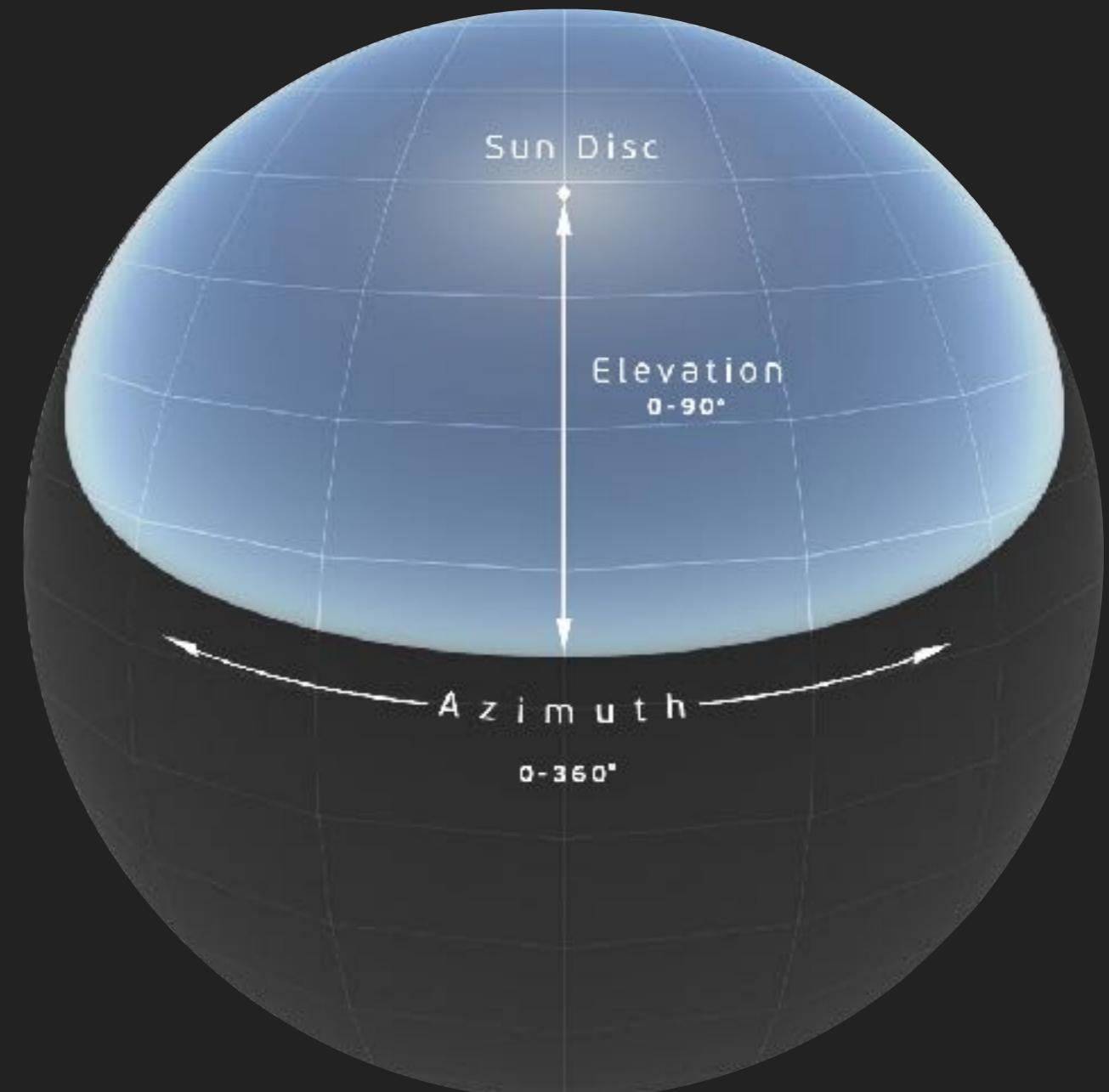
CELESTIAL COORDINATE SYSTEMS

- ▶ Need a way to specify the location of celestial objects
- ▶ Can be in 3D or in 2D
- ▶ Different coordinate systems exists for different purposes
- ▶ Spherical / cartesian, different origins, different orientation
- ▶ Can convert between them



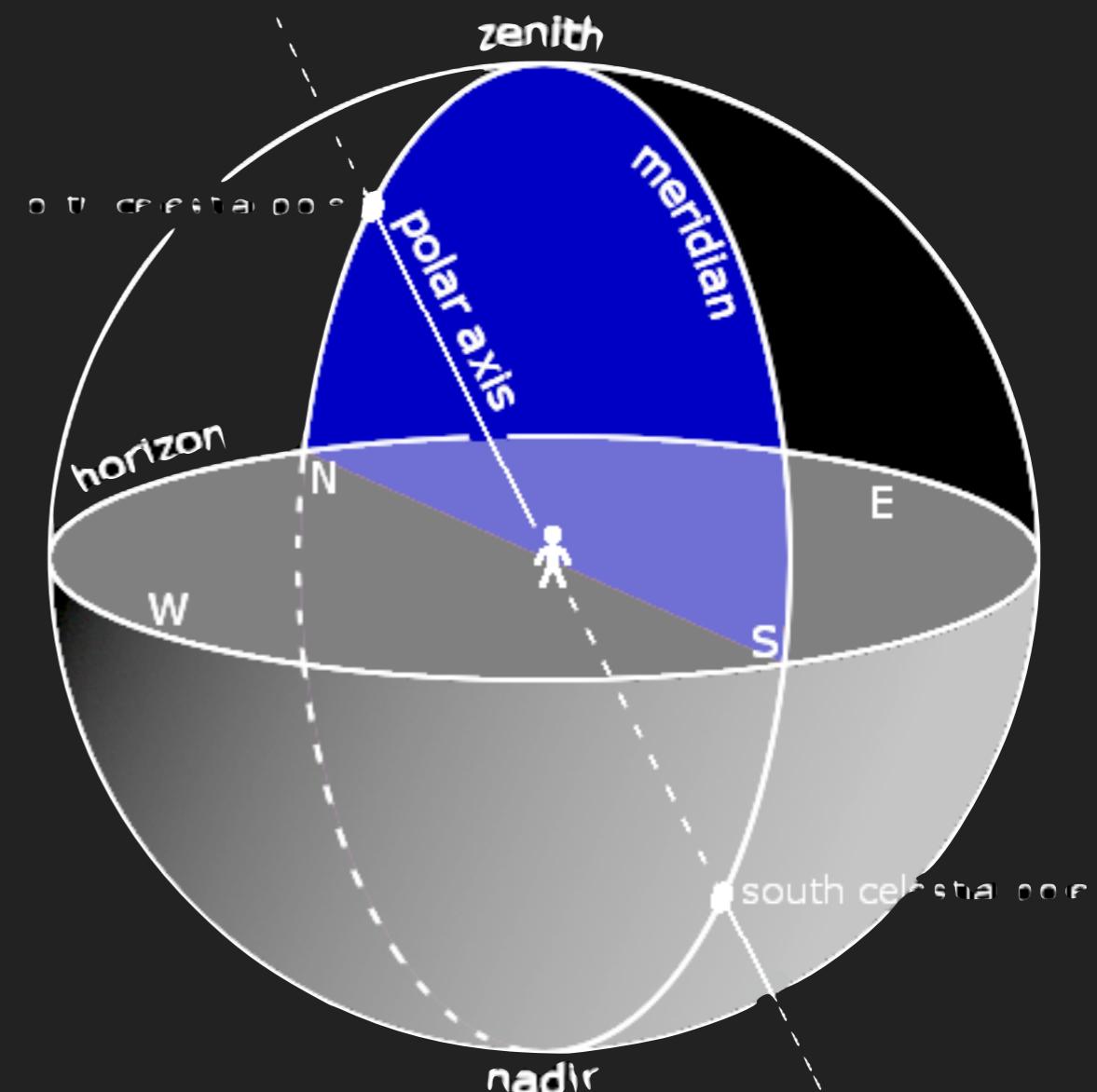
HORIZONTAL COORDINATE SYSTEM

- ▶ Local observer's horizon is the fundamental plane
- ▶ Altitude (alt) / Azimuth (az)
- ▶ Azimuth measured from north, increasing towards east
- ▶ Altitude from horizon upwards



HORIZONTAL COORDINATE SYSTEM

- ▶ Meridian is the line from North to the Zenith to South
- ▶ Azimuth 0 and 180



HORIZONTAL COORDINATE SYSTEM PROS

- ▶ Know exactly where to look

HORIZONTAL COORDINATE SYSTEM CONS

- ▶ Depends on time and location

EQUATORIAL COORDINATE SYSTEM

- ▶ Fundamental plane is the Earth's equator
- ▶ primary direction towards the vernal equinox
- ▶ Declination (dec) / Right Ascension (ra)



EQUATORIAL COORDINATE SYSTEM PROS

- ▶ Fixed stars have fixed coordinates
- ▶ Coordinates do not depend on time or date

EQUATORIAL COORDINATE SYSTEM CONS

- ▶ Harder to find objects

ANGLES IN ASTRONOMY

- ▶ Both coordinate systems use angles
- ▶ Multiple way to specify angles:
 - ▶ Degrees 0° - 360°
 - ▶ Radians 0 - 2π
 - ▶ Hours $0h$ - $24h$

DEGREES

- ▶ 1 full circle = 360°
- ▶ 1 degree = 60 arc minute = $60'$
- ▶ $1'$ = 60 arc seconds = $60''$
- ▶ $1''$ = 1000 milli arc seconds = 1000 mas
- ▶ 1 mas = 1000 micro arc seconds = $1000 \mu\text{as}$

Venus

Type: planet

Magnitude: -4.03 (extincted to: -3.76)

Absolute Magnitude: 27.33

RA/Dec (J2000.0): 5h28m23.11s/+21°20'31.4"

RA/Dec (J2017.6): 5h29m26.19s/+21°21'18.9"

Hour angle/DE: 19h21m41.41s/+21°22'26.5" (apparent)

Az/Alt: +87°25'20.8"/+29°07'24.5" (apparent)

Ecliptic longitude/latitude (J2000.0): +82°38'14.4"/-1°53'43.7"

Ecliptic longitude/latitude (J2017.6): +82°52'57.5"/-1°53'35.5"

Galactic longitude/latitude: -175°38'54.7"/-7°20'58.3"

Obliquity (of date, for Earth): +23°26'13.2"

Distance: 1.101AU (164.681 Mio km)

Apparent diameter: +0°00'15.2"

Sidereal period: 224.70 days (0.615 a)

Sidereal day: 5832h28m47.1s

Mean solar day: 2802h0m52.2s

Phase Angle: +63°45'25"

Elongation: +39°49'08"

Phase: 0.72

Illuminated: 72.1%



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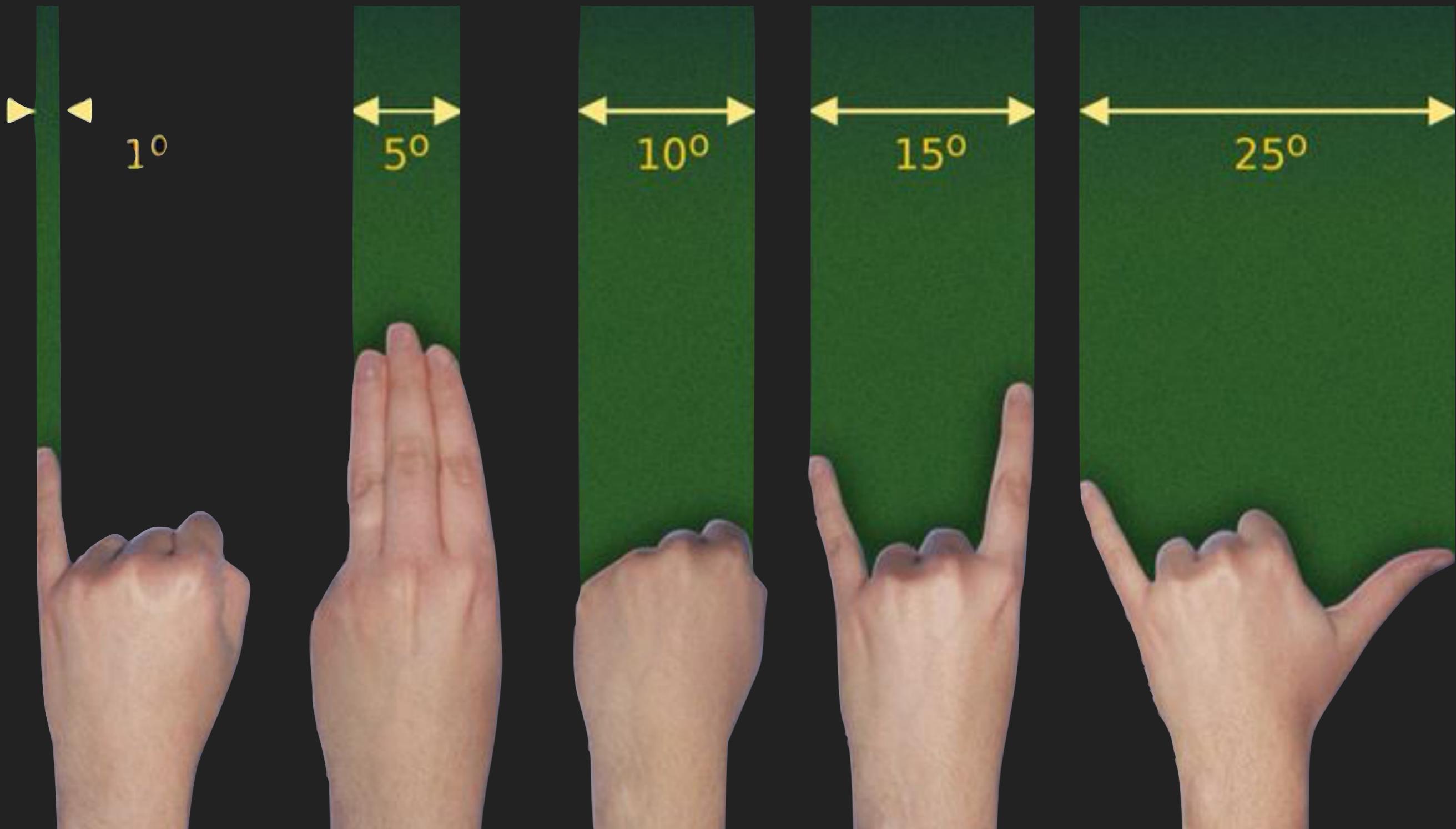
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DEGREES, MEASURED BY HAND



HOURS

- ▶ 1 full circle = 24h
- ▶ 1h = 60 minutes = 60 m
- ▶ 1m = 60 seconds = 60 s

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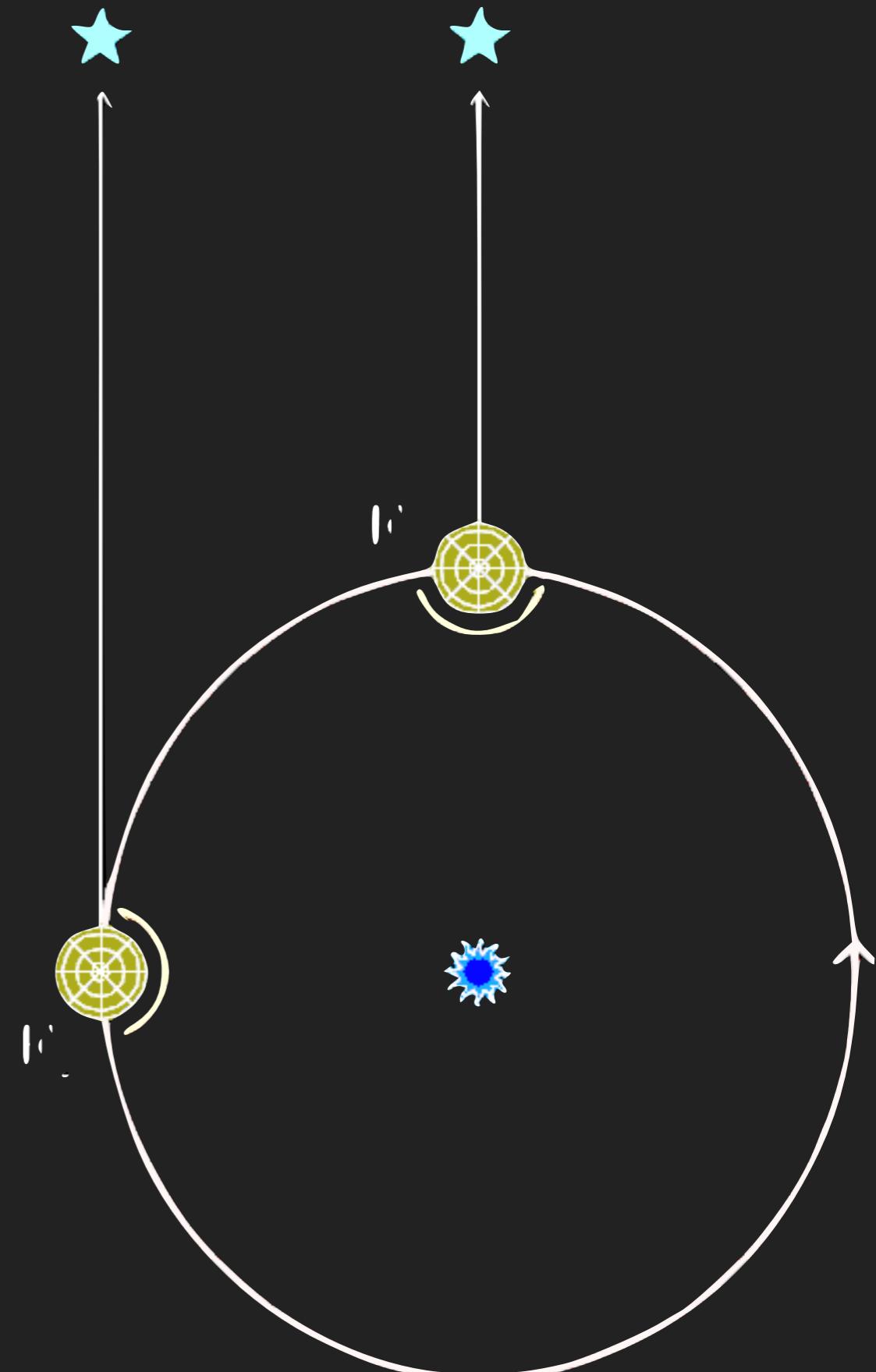
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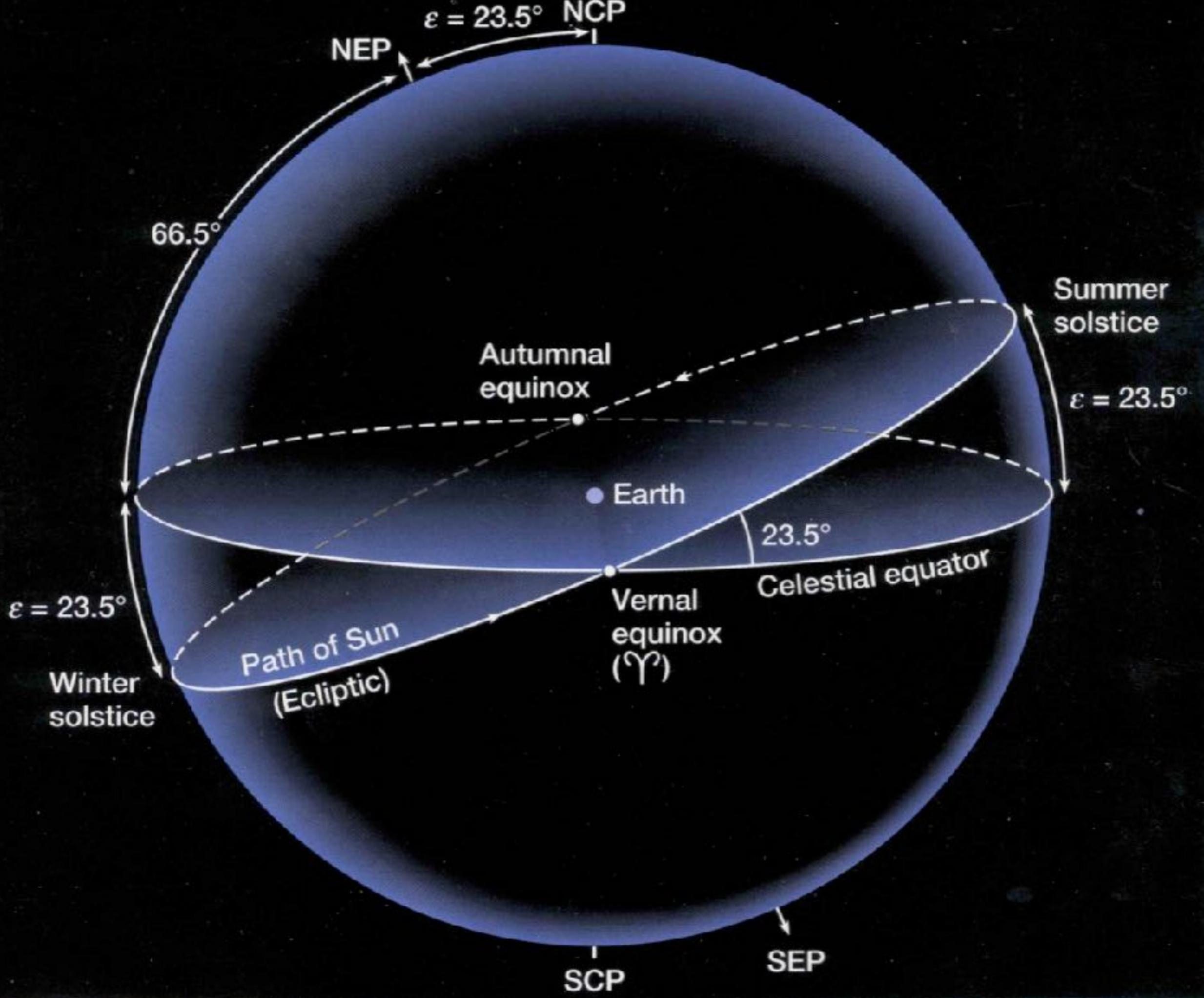
CONVERSION BETWEEN COORDINATE SYSTEMS

- ▶ Not difficult, just two rotations
- ▶ Do not remember formula, but do look at it and try to understand it
- ▶ To go between RA/DEC and AZ/ALT one also needs
 - ▶ Time
 - ▶ Location
 - ▶ How to specify time? Sidereal time

SIDEREAL TIME

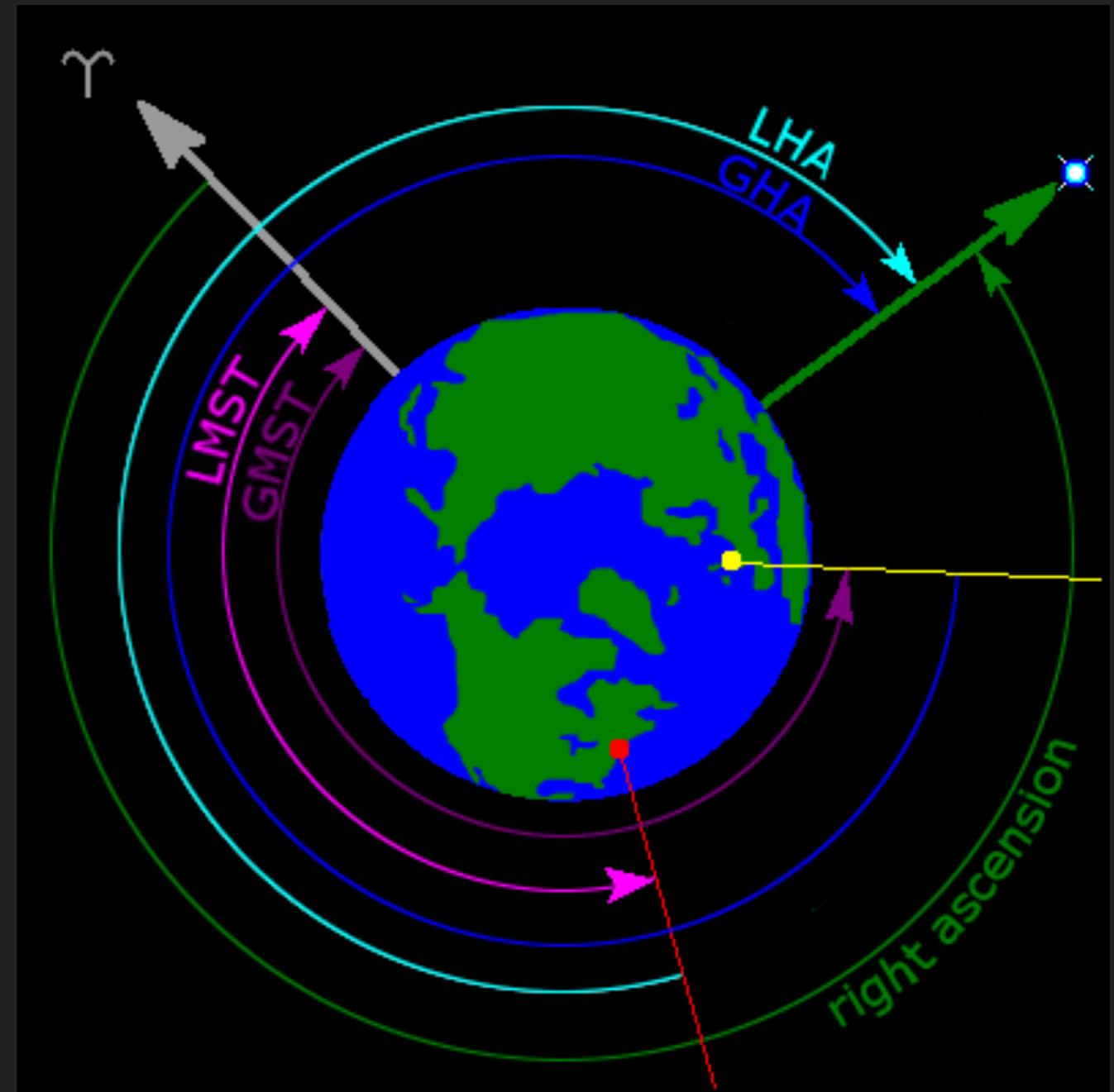
- ▶ Which star is on our local meridian?
- ▶ Depends on time and date
- ▶ Our normal clocks use solar time
- ▶ Astronomers are interested in sidereal time
- ▶ Local Sidereal Time (LST) is 0 hours when the vernal equinox (RA=0) is on local meridian





SIDEREAL TIME

- ▶ Hour angle
 $HA = LST - RA$
- ▶ Tells you where your object is with respect to the meridian.
- ▶ $|HA| > 6$ hours hard to observe (but depends on declination)



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COORDINATE SYSTEMS

EXAMPLE (WITHOUT THE CELESTIAL SPHERE)

At midnight on 1998 February 4th, LST
at St. Andrews was 8h45m.

St. Andrews has longitude $2^{\circ}48'W$.

What was the Local Hour Angle of
Betelgeuse ($RA = 5h55m$) at midnight?

At what time was Betelgeuse on the
meridian at St. Andrews?

At what time was Betelgeuse on the
meridian at Greenwich?

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SOLUTIONS

2h 50m

21h 10m

20h 59m

PROBLEMS WITH THE EQUATORIAL SYSTEM

- ▶ Equatorial coordinates change slowly
- ▶ Timescale 25770 years
- ▶ This is because Earth's rotation axis precesses around the orbital plane
- ▶ Must also specify Epoch, the standard nowadays is J2000

GALACTIC COORDINATE SYSTEM

- ▶ Earth at centre
- ▶ Latitude and longitude
- ▶ 0 towards galactic centre

