

Higher dim Osc

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1 Introduction

We make the metric Ansatz:

$$ds^2 = \alpha^2 dt^2 + a^2 dr^2 + r^2 d\Omega_3 \quad (1)$$

where $A(t, r) = a^2$ and $C(t, r) = (a/\alpha)^2$.

The Einstein equation t,t component in 3+1D

$$\begin{aligned} A^{(0,1)}(t, r) = & \frac{1}{2}rA(t, r)\phi^{(1,0)}(t, r)^2C(t, r) + \frac{1}{2}rA(t, r)\phi^{(0,1)}(t, r)^2 + \frac{1}{2}rA(t, r)^2\phi(t, r)^2 \\ & - \frac{1.A(t, r)^2}{r} + \frac{1.A(t, r)}{r} \end{aligned} \quad (2)$$

The Einstein equation t,t component in 4+1D

$$\begin{aligned} A^{(0,1)}(t, r) = & \frac{1}{3}rA(t, r)\phi^{(1,0)}(t, r)^2C(t, r) + \frac{1}{3}rA(t, r)\phi^{(0,1)}(t, r)^2 + \frac{1}{3}rA(t, r)^2\phi(t, r)^2 \\ & - \frac{2.A(t, r)^2}{r} + \frac{2.A(t, r)}{r} \end{aligned} \quad (3)$$

The Einstein equation t,r component in 3+1D

$$A^{(1,0)}(t, r) = rA(t, r)\phi^{(0,1)}(t, r)\phi^{(1,0)}(t, r) \quad (4)$$

The Einstein equation t,r component in 4+1D

$$A^{(1,0)}(t, r) = \frac{2}{3}rA(t, r)\phi^{(0,1)}(t, r)\phi^{(1,0)}(t, r) \quad (5)$$

The Einstein equation r,r component in 3+1D

$$C^{(0,1)}(t, r) = rA(t, r)\phi(t, r)^2C(t, r) - \frac{2.A(t, r)C(t, r)}{r} + \frac{2.C(t, r)}{r} \quad (6)$$

The Einstein equation r,r component in 4+1D

$$C^{(0,1)}(t, r) = \frac{2}{3}rA(t, r)\phi(t, r)^2C(t, r) - \frac{4.A(t, r)C(t, r)}{r} + \frac{4.C(t, r)}{r} \quad (7)$$

The Klein Gordon equation in 3+1D in as similar form as the paper

$$\begin{aligned}\phi^{(2,0)}(t,r)C(t,r) = & -\frac{C^{(0,1)}(t,r)\phi^{(0,1)}(t,r)}{2C(t,r)} - A(t,r)\phi(t,r) \\ & - \frac{1}{2}C^{(1,0)}(t,r)\phi^{(1,0)}(t,r) + \frac{2\phi^{(0,1)}(t,r)}{r} + \phi^{(0,2)}(t,r)\end{aligned}\quad (8)$$

The Klein Gordon equation in 4+1D in as similar form as the paper

$$\begin{aligned}\phi^{(2,0)}(t,r)C(t,r) = & -\frac{C^{(0,1)}(t,r)\phi^{(0,1)}(t,r)}{2C(t,r)} - A(t,r)\phi(t,r) \\ & - \frac{1}{2}C^{(1,0)}(t,r)\phi^{(1,0)}(t,r) + \frac{3\phi^{(0,1)}(t,r)}{r} + \phi^{(0,2)}(t,r)\end{aligned}\quad (9)$$

The Klein Gordon equation in 3+1D in as similar form as the Code

$$\begin{aligned}\phi^{(2,0)}(t,r)C(t,r) = & -0.5rA(t,r)\phi^{(0,1)}(t,r)\phi(t,r)^2 \\ & + \frac{1.A(t,r)\phi^{(0,1)}(t,r)}{r} - A(t,r)\phi(t,r) - \frac{1}{2}C^{(1,0)}(t,r)\phi^{(1,0)}(t,r) \\ & + \frac{1.\phi^{(0,1)}(t,r)}{r} + \phi^{(0,2)}(t,r) + 0.\end{aligned}\quad (10)$$

The Klein Gordon equation in 4+1D in as similar form as the Code

$$\begin{aligned}\phi^{(2,0)}(t,r)C(t,r) = & -\frac{1}{3}rA(t,r)\phi^{(0,1)}(t,r)\phi(t,r)^2 \\ & + \frac{2.A(t,r)\phi^{(0,1)}(t,r)}{r} - A(t,r)\phi(t,r) - \frac{1}{2}C^{(1,0)}(t,r)\phi^{(1,0)}(t,r) \\ & + \frac{1.\phi^{(0,1)}(t,r)}{r} + \phi^{(0,2)}(t,r) + 0.\end{aligned}\quad (11)$$

2 Axion equation

The Einstein equation t,t component in 3+1D

$$\begin{aligned}A^{(0,1)}(t,r) = & \frac{1}{2}rA(t,r)\phi^{(1,0)}(t,r)^2C(t,r) + \frac{1}{2}rA(t,r)\phi^{(0,1)}(t,r)^2 + rA(t,r)^2V(\phi(t,r)) \\ & - \frac{A(t,r)^2}{r} + \frac{A(t,r)}{r}\end{aligned}\quad (12)$$

The Einstein equation t,r component in 3+1D

$$A^{(1,0)}(t,r) = rA(t,r)\phi^{(0,1)}(t,r)\phi^{(1,0)}(t,r)\quad (13)$$

The Einstein equation r,r component in 3+1D

$$C^{(0,1)}(t, r) = 2rA(t, r)V(\phi(t, r))C(t, r) - \frac{2A(t, r)C(t, r)}{r} + \frac{2C(t, r)}{r} \quad (14)$$

The Klein Gordon equation in 3+1D as in Paper

$$\begin{aligned} \phi^{(2,0)}(t, r)C(t, r) = & -\frac{C^{(0,1)}(t, r)\phi^{(0,1)}(t, r)}{2C(t, r)} - A(t, r)V'(\phi(t, r)) \\ & - \frac{1}{2}C^{(1,0)}(t, r)\phi^{(1,0)}(t, r) + \frac{2\phi^{(0,1)}(t, r)}{r} + \phi^{(0,2)}(t, r) \end{aligned} \quad (15)$$

The Klein Gordon equation in 3+1D as in Code

$$\begin{aligned} \phi^{(2,0)}(t, r)C(t, r) = & -rA(t, r)\phi^{(0,1)}(t, r)V(\phi(t, r)) \\ & + \frac{A(t, r)\phi^{(0,1)}(t, r)}{r} - A(t, r)V'(\phi(t, r)) - \frac{1}{2}C^{(1,0)}(t, r)\phi^{(1,0)}(t, r) \\ & + \frac{\phi^{(0,1)}(t, r)}{r} + \phi^{(0,2)}(t, r) \end{aligned} \quad (16)$$

3 Fourier Space

$$\phi(t, x) = \sum_{j=0}^{\infty} \hat{\phi}_j(x) \cos(j\omega t) \quad (17)$$

$$A(t, x) = \sum_{j=0}^{\infty} \hat{A}_j(x) \cos(j\omega t) \quad (18)$$

$$C(t, x) = \sum_{j=0}^{\infty} \hat{C}_j(x) \cos(j\omega t) \quad (19)$$

$$\begin{aligned} \phi(t, x)A(t, x) &= \sum_{j=0}^{\infty} \sum_{i=0}^{\infty} \hat{\phi}_j(x)\hat{A}_i(x) \cos(i\omega t) \cos(j\omega t) \\ &= \sum_{n=0}^{\infty} \sum_{i=0}^{\infty} \hat{\phi}_{n-i}(x)\hat{A}_i(x) \frac{1}{2}(\cos((j+i)\omega t) + \cos((j-i)\omega t)) \\ &\stackrel{\underbrace{\hspace{1cm}}}{=} \sum_{n=i+j}^{\infty} \sum_{n=0}^{\infty} \sum_{i=0}^{\infty} \hat{\phi}_{n-i}(x)\hat{A}_i(x) \frac{1}{2}(\cos(n\omega t) + \cos((j-i)\omega t)) \\ &= \sum_{n=0}^{\infty} \sum_{i=0}^{\infty} \hat{\phi}_{n-i}(x)\hat{A}_i(x) \frac{1}{2}(\cos(n\omega t)) + \sum_{n=0}^{\infty} \sum_{i=0}^{-\infty} \hat{\phi}_{n+i}(x)\hat{A}_{-i}(x) \frac{1}{2}(\cos(n\omega t)) \\ &= \sum_{n=0}^{\infty} \sum_{i=-\infty}^{\infty} \hat{\phi}_{n-|i|}(x)\hat{A}_{|i|}(x) \cos(n\omega t) \end{aligned} \quad (20)$$

will be written as

$$A(t, x) = \sum_{j=0}^{\infty} \hat{A}_j(x) \partial_t \cos(j\omega t) = \sum_{j=0}^{\infty} \hat{A}_j(x) j\omega \cos(j\omega t) \quad (21)$$

$$\begin{aligned} A^{(0,1)}(t, r) &= \frac{1}{2} r A(t, r) \phi^{(1,0)}(t, r)^2 C(t, r) + \frac{1}{2} r A(t, r) \phi^{(0,1)}(t, r)^2 + \frac{1}{2} r A(t, r)^2 \phi(t, r)^2 \\ &\quad - \frac{1 \cdot A(t, r)^2}{r} + \frac{1 \cdot A(t, r)}{r} \end{aligned} \quad (22)$$

In Fourier space

$$\begin{aligned} \hat{A}_j^{(0,1)}(t) &= \frac{1}{2} r j^2 \omega^2 (\hat{A} * \hat{\phi} * \hat{\phi} * \hat{C})_j + \frac{1}{2} r (\hat{A} * \hat{\phi}^{(0,1)} * \hat{\phi}^{(0,1)})_j + \frac{1}{2} r (\hat{A} * \hat{A} * \hat{\phi} * \hat{\phi})_j \\ &\quad - \frac{(\hat{A} * \hat{A})_j}{r} + \frac{A_j}{r} \end{aligned} \quad (23)$$

4 ADM mass

Asymptotic behaviour of Metric

$$ds^2 = - \left(A - \frac{2M}{r^{d-3}} \right) dt^2 + \left(C - \frac{2M}{r^{d-3}} \right)^{-1} + r^2 d\Omega_{d-2}^2 \quad (24)$$

where $C, B \in R$ we can find the asymptomatic mass by :

$$\begin{aligned} A &= \left(C - \frac{2M}{r^{d-3}} \right)^{-1} \\ M &= \frac{1}{2} \left(C - \frac{1}{A} \right) r^{-3+d} \end{aligned} \quad (25)$$

5 Non-relativistic limit

$$i\hbar \partial_t \psi = - \frac{\hbar^2}{2m} \nabla^2 \psi + mU\psi \quad (26)$$

$$\nabla^2 U = 4\pi G |\psi|^2 \quad (27)$$

we Fourier transform $\nabla = 1/L$ where L is the characteristic length of the Osc.

$$i\hbar \partial_t \psi = - \frac{\hbar^2}{2m} \nabla^2 \psi + mU\psi \quad (28)$$

$$\nabla^2 U = 4\pi G |\psi|^2 \quad (29)$$

6 Conclusion

3+1D \neq 4+1D

7 AdS Solitons

7.1 Ansatz

We make the metric Ansatz:

$$ds^2 = \left(\alpha^2 - \frac{r^2 \Lambda}{6} \right) dt^2 + a^2 dr^2 + r^2 d\Omega_3 \quad (30)$$

where $A(t, r) = a^2$ and $C(t, r) = (a/\alpha)^2$.

compared to the ansatz in BRITO PROCA Star

$$ds^2 = \sigma^2(r) F(r) dt^2 + \frac{1}{F(r)} dr^2 + r^2 d\Omega \quad (31)$$

where $F(r) = 1 - \frac{2m(r)}{r^2} - \Lambda r^2/6$

7.2 Stability

In this section we discuss solution for higher-dimensional scalar and vector Boson stars [3]. One finds that the maximum mass for Boson Stars in 5 dimension diverges for $\phi_0(0) \rightarrow 0$. The binding energy seems to be negative for all values, indicating that such solution are unstable (analogous for Boson Stars: [2]). In [3] they hypothesise that for $\Lambda < 0$ there are stable solutions. In [1] the construction of AdS boson stars is described.

References

- [1] Dumitru Astefanesei and Eugen Radu. Boson stars with negative cosmological constant. *Nucl. Phys.*, B665:594–622, 2003.
- [2] Yves Brihaye and Betti Hartmann. Minimal boson stars in 5 dimensions: classical instability and existence of ergoregions. *Class. Quant. Grav.*, 33(6):065002, 2016.
- [3] Miguel Duarte and Richard Brito. Asymptotically anti-de Sitter Proca Stars. *Phys. Rev.*, D94(6):064055, 2016.

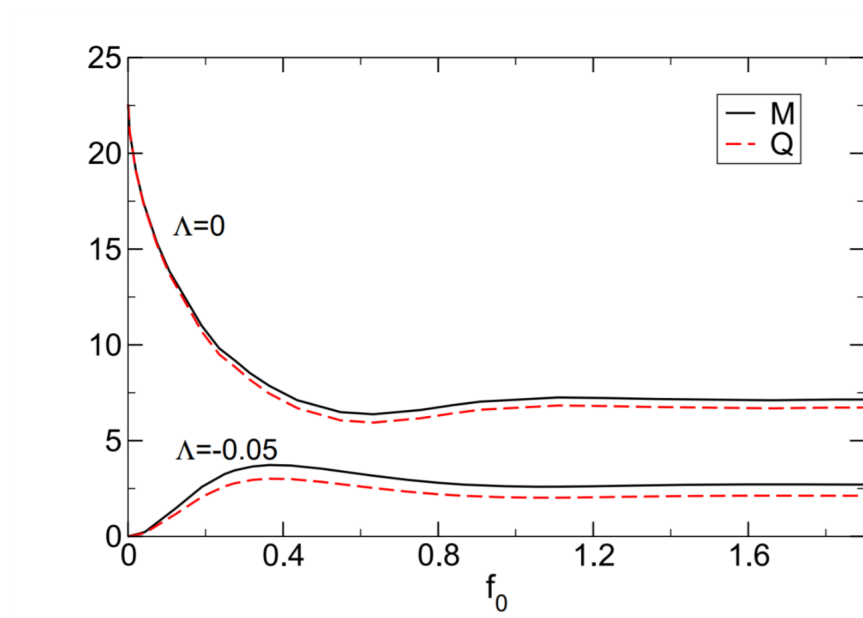


Figure 1: **Stability of vector Boson stars in 5 dimension:** Where f_0 is the parameter defining Proca stars. Adding a negative cosmological constant gives the upper mass a bound.