

Liquid Argon optical properties for Geant4 and Opticks Simulations

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Abstract

In Geant4 and Opticks optical properties like e.g. the refractive index, absorption length, Rayleigh scattering length etc. as well as surface properties are inputs that have to be provided. In this paper we collect the optical properties relevant for simulating the optical response of liquid Argon TPC's.

Contents

1	Introduction	3
2	Light production	6
2.1	Scintillation Properties of liquid Argon	6
2.2	Cerenkov spectrum and Yield	7
3	Light propagation	9
3.1	Refraction Index and group velocity	9
3.2	Absorption length	9
3.3	Rayleigh Scattering length	10
4	Boundary processes	14
4.1	Refraction and total internal Reflection	14
4.2	Reflection	14
5	Photon Detection	15
5.1	Quantum efficiency and absorption length of the tetraphenyl butadiene (TPB) wave length shifter	15
6	Software	17
6.1	physics configuration	17
6.2	Sensitive Detector PhotonSD and PhotonHit	18
7	Conclusions and Outlook	18

1 Introduction

In Geant4, the optical physics has an exceptional position among the physics processes, as it adds special particles (optical photons) and new properties for materials and optical surfaces. Being the only particle that can be reflected or refracted at optical surfaces, as well as being only created in optical processes like scintillation, Cerenkov radiation, and wavelength-shifting (WLS) makes the G4OpticalPhoton differing from the “usual” high-energy particle-physics photon (G4Gamma) in Geant4. Optical properties need to be assigned to the materials whenever optical physics processes are to be considered in the simulation. Every material needs at least a refractive index spectrum (which corresponds to the dispersion relation) and an attenuation length spectrum, though the attenuation length is by default set to infinity if it is not defined. Special optical materials, i.e. scintillating and WLS materials, additionally require the specification of the emission spectra as well as of the rise and decay times. More properties can be assigned to optical surfaces between volumes, e.g. the reflectivity of the surface. An article describing the peculiarities (and pitfalls) of Geant4 optical simulation can be found here [5]

In Geant4 [1] and Opticks [2] optical properties like e.g. the materials refractive index, Rayleigh scattering length or absorption length are inputs that have to be provided when the detector is constructed. The properties are either bulk properties (e. g. the refractive index) or boundary properties (G4OpticalSurfaces e.g. reflectivity). Geant4 only traces optical photons for materials for which at least the refractive index is defined other wise the optical photons are killed. High-precision modeling of light production, transport and detection in liquid Argon experiments requires the use of the best available values to describe the properties of liquid Argon.

In this article we briefly describe the physical processes relevant to the production, transport and detection of optical photons in liquid Argon. We collect the values and parameterizations of optical properties relevant for liquid Argon TPC’s. We provide scripts that plot this quantities and that convert this values into a g4mdl description that can be directly used in the Geant4 Detector description. All values are summarized in the file material.xml which can be found in the github repository [8]. Usually quantities are given as a function of photon wavelength but Geant4 requires the photon energy.

A nice overview about the properties of Scintillation Light in Liquid Argon can be found in [12]. The motion of the charged particles liberates charge from the surrounding argon (ionization) and produces light (scintillation)

The specific properties will be covered in detail in the next sections. Here we want to give an overview of the basic properties.

- Liquid Argon has a high scintillation yield in order of 20,000–40,000 photons/MeV depending on purity. The Scintillation yield is E-field and particle dependent. For a proton the yield is listed in Table 1.
- Tracking detectors can be constructed by applying an electric field across the bulk and collecting electrons freed by ionization. Ionization and Scin-

electric Field [V/cm]	Scintillation yield [photons/MeV]
0	40000
500	24000

Table 1: Scintillation yield in liquid Argon for a proton for various electric fields.

tillation are competing processes and the number of electrons and scintillation photons is correlated (see e.g [16]).

- The wavelength of the scintillation photons are in the vacuum ultraviolet. The emission spectrum is modeled as a Gaussian:

$$G(\lambda) = ae^{-\frac{(\lambda-b)^2}{2\sigma^2}}$$

where:

a = height of the curve's peak

b = the position of the center of the peak (128 nm)

σ = the standard deviation (10 nm)

e = Euler's number

With both a fast (6nsec $3\Sigma_u$ excimer) and slow component (1590 nsec via $1\Sigma_u$ excimer state) Liquid argon produces scintillation light via two distinct scintillation mechanisms, each of which has a different characteristic time constant The fast scintillation path

- Relatively complicated including various excimer states can be induced by ionization but not by scintillation photons. For that reason reemission of scintillation light doesn't play a role in liquid Argon.
- Liquid Argon is highly transparent to its own scintillation light with absorption length in the order of several meters depending on purity.
- Rayleigh scattering length is 90 (55) cm at the scintillation wavelength of pure liquid argon (128 nm). This significantly diffuses the scintillation light for large detectors required for neutrino physics.
- To match the photon wavelength to the quantum efficiency (QE) of photo detectors wavelength shifters are used. A typical wavelength shifting material used to coat the photo detectors in liquid Argon TPCs is TPB with a reemission spectrum peaking at 420 nm.

The Geant4 keywords used in this article refer to Geant4 version > 11.0 released in December 2021. ¹

Note while in the literature properties like e.g. the refraction index are give as a function of the wavelength λ while Geant4 requires this properties to be

¹Note: the latest version introduced changes to the Geant4 API with regards to optical material properties.

expressed as a function of photon energy. The Conversion between wavelength (in nm) and photon energy (eV) is given in Equation 1

$$E_{\gamma}(eV) = \frac{hc}{\lambda_{\gamma} \times 10^{-9}} \quad (1)$$

with:

speed of light: $c = 299792458m/sec$

Planck constant: $h = 4.13566743 \times 10^{-15}eV/sec$

2 Light production

There are two relevant primary sources of light production when a charged particle passes through a medium. One is Scintillation light the other is Cerenkov radiation. The two sources have very different characteristics and yield. Scintillation light produced when a charged particle ionizes the material. The light is emitted isotropic from the point where it is produced. Cerenkov radiation ([14],[15]) is electromagnetic radiation emitted when a charged particle passes through a dielectric medium at a speed greater than the phase velocity (speed of propagation of a wavefront in a medium) of light in that medium. Cere Cerenkov radiation directional emitted as a conical wave front around the direction of the charge particle with the emission angle given by 4. Note Geant4 provides the transition radiation process as another primary source of photons but this doesn't play a role in liquid Argon TPCs. The wave length shifting (WLS) process where a WLS material absorbs optical photons and then re-emits them with a longer wavelength is a source of secondary photons.

2.1 Scintillation Properties of liquid Argon

Efficient scintillator with typical Light yields in the order of a few 10,000's of photons per MeV deposited (depends on E field, particle type and purity) (SCINTILLATIONYIELD: 50000/ MeV when no electric field present)

Liquid argon produces scintillation light via two distinct scintillation mechanisms, each of which has a different characteristic time constant. The emission spectra are passed to Geant4 as a 2 column matrix where the first column is the photon energy and the second is the value. result into excited, Ar_2^* , or ionized, Ar_2^+ , argon dimers. Ar_2^+ recombines with a thermalized electron to form Ar_2^* which in turn decays non-radiatively to the first singlet and triplet excited states $1\Sigma_u^+$ and $3\Sigma_u^+$. These two states, whose dis-excitation leads to the emission of the scintillation photons, have approximately the same energy with respect to the dissociative ground state, while the lifetimes are very different: in the nanosecond range for $1\Sigma_u^+$ and in the microsecond range for $3\Sigma_u^+$.

[17]

Property	Geant4 keyword	value
light yield	SCINTILLATIONYIELD	50000 $\gamma'/s/MeV$ (no electric field)
Wavelength of emission	SCINTILLATIONCOMPONENT1	128nm ($FWHM = 10nm$) see 1
Wavelength of emission	SCINTILLATIONCOMPONENT2	128nm ($FWHM = 10nm$) see 1
fast component	SCINTILLATIONTIMECONSTANT1	6ns
fast fraction	SCINTILLATIONYIELD1	0.75
slow component	SCINTILLATIONTIMECONSTANT2	1500ns
slow fraction	SCINTILLATIONYIELD2	0.25
	RESOLUTIONSCALE	1

Table 2: Scintillation Properties of liquid Argon.

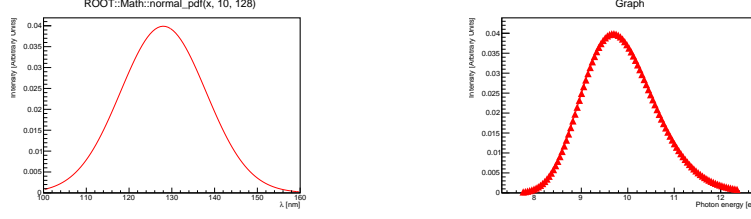


Figure 1: Scintillation emission spectrum.

Scintillation Quenching, Birks law below:

$$\frac{dL}{dx} = S \frac{\frac{dE}{dx}}{1 + kB \frac{dE}{dx}}. \quad (2)$$

where L is the light yield, S is the scintillation efficiency, dE/dx is the specific energy loss of the particle per path length, kB is the Birks coefficient. Its value depends on the scintillating material and particle type.

2.2 Cerenkov spectrum and Yield

A charged particle radiates if its speed is greater than the local phase speed of light v_p . In Geant4 the process is not contributing to energy loss.

the charged particle travels in a medium with speed v_p such that $\frac{c}{n} < v_p < c$. Cerenkov radiation as conical wave front with the emission angle given by

$$\cos \theta = \frac{1}{n\beta} \quad (3)$$

with:

ratio between the speed of the particle v_p and the speed of light as $\beta = \frac{v_p}{c}$.

from [14]

$$\cos(\theta_C) = \frac{1}{n\beta} \quad (4)$$

$$\frac{d^2N}{dE dx} = \frac{2\pi\alpha z^2}{\lambda^2} \left(1 - \frac{1}{(\beta^2 n^2(\lambda))} \right) \quad (5)$$

α is the fine-structure constant.

$$\beta = \frac{v_p}{c}$$

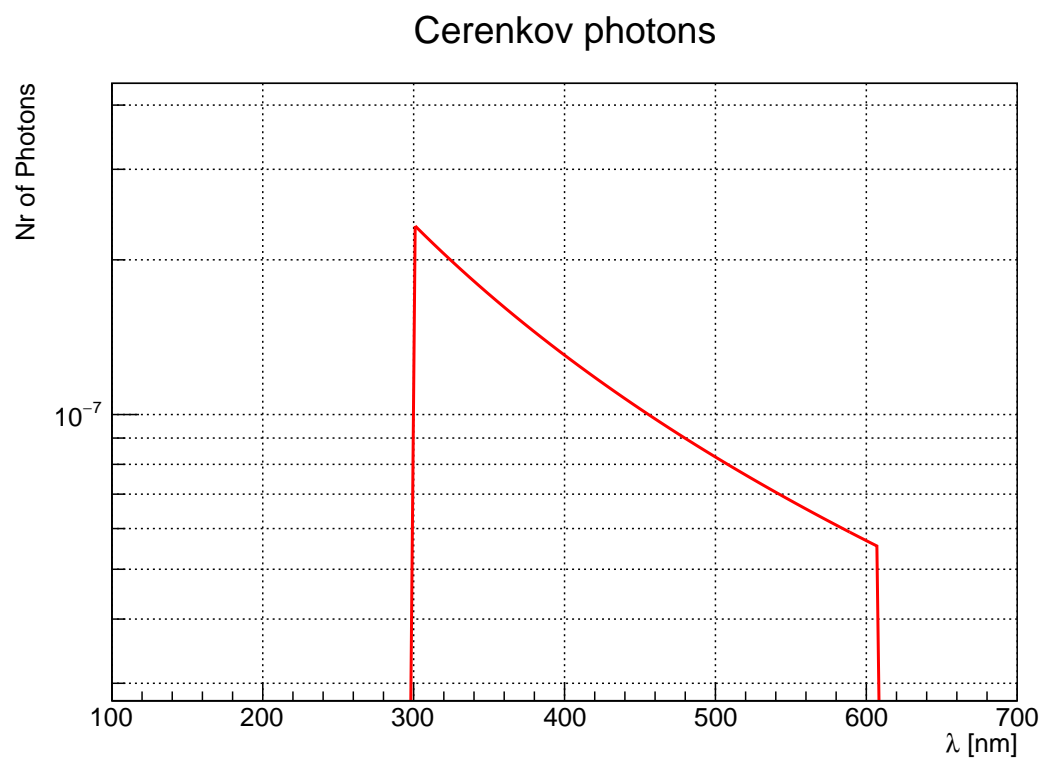


Figure 2: Cerenkov spectrum

3 Light propagation

In this section we discuss the material properties related to light propagation in a medium. The values are passed to Geant4 as a 2 column matrix where the first column is the photon energy and the second is the value.

optical material property	Geant4 Keyword
Refraction index as function of photon energy	RINDEX
Absorption length as function of photon energy	ABSLENGTH
Rayleigh scattering length as function of photon energy	RAYLEIGH

Table 3: Material properties related to light propagation in a medium

3.1 Refraction Index and group velocity

Every material needs at least a refractive index spectrum and an absorption length spectrum. The absorption length is by default set to infinity if it is not defined. If the group velocity v_g is not provided explicitly it is calculated from the refraction index. The speed of optical photons in Geant4 is given by the group velocity v_g . The relation between group velocity and the refraction index is given by:

$$v_g(\lambda) = \frac{c}{n - \lambda \frac{\partial n}{\partial \lambda}} \quad (6)$$

Figure shows the calculated group velocity compared to the phase velocity $1/n$.

In [11] the refraction index and group velocity at $128nm$ are measured at $n = 1.358 \pm 0.003$ and $\frac{1}{v_g} = 7.46 \pm 0.08ns/m$. (compared to $n = 1.45 \pm 0.07$ [10])

$\frac{1}{v_g} = 7.46 \pm 0.08ns/m$ corresponds to $0.134m/nsec$ which is approximately $c_0/2240$ where c_0 is the speed of light in vacuum. (reading it from the gdml dump calculated by Geant4 one gets $c_0/2600$)

The Sellmeier equation 7 below is an empirical relationship between refractive index and wavelength for a particular transparent medium.

$$n^2 = a_0 + \frac{a_{UV}\lambda^2}{\lambda^2 - \lambda_{UV}^2} + \frac{a_{IR}\lambda^2}{\lambda^2 - \lambda_{IR}^2}. \quad (7)$$

where the parameters a_0 , a_{UV} and a_{IR} known as Sellmeier coefficients have to be determined experimentally.

3.2 Absorption length

Argon is highly transparent to its own scintillation light. (ABSLENGTH) > $1.1m$ (ArXiv:1511.07725)

Scintillation λ (nm)	UV Resonance λ_{UV} (nm)	IR Resonance λ_{IR} (nm)
128	106.6	908.3

Table 4: blabla bla.

$T(K)$	a_0	a_{UV}	a_{IR}
83.81	1.24 ± 0.09	0.27 ± 0.09	0.00047 ± 0.007
90	1.26 ± 0.09	0.23 ± 0.09	0.0023 ± 0.007

Table 5: Sellmeier coefficients

3.3 Rayleigh Scattering length

Rayleigh scattering length (RAYLEIGH). In the literature one can find the following calculated values at $128nm$: 90 cm [11] and $55 \pm 5cm$ [10]. The range of values for the Rayleigh scattering length l is due to the different refraction indices n at $128nm$ input to equation 8 below:

$$l^{-1} = \frac{16\pi^3}{6\lambda^4} \left[kT\rho^2 k_T \left(\frac{(n^2 - 1)(n^2 + 2)}{3} \right)^2 \right] \quad (8)$$

with

l : the scattering length

λ : the wavelength of light

n : the index of refraction corresponding the wavelength of light

T : the temperature

ρ : the density

k_T : the isothermal compressibility

k : the Boltzman constant

Figure 5 shows the Rayleigh scattering length as a function of λ calculated using formula 8.

T=83.81 K

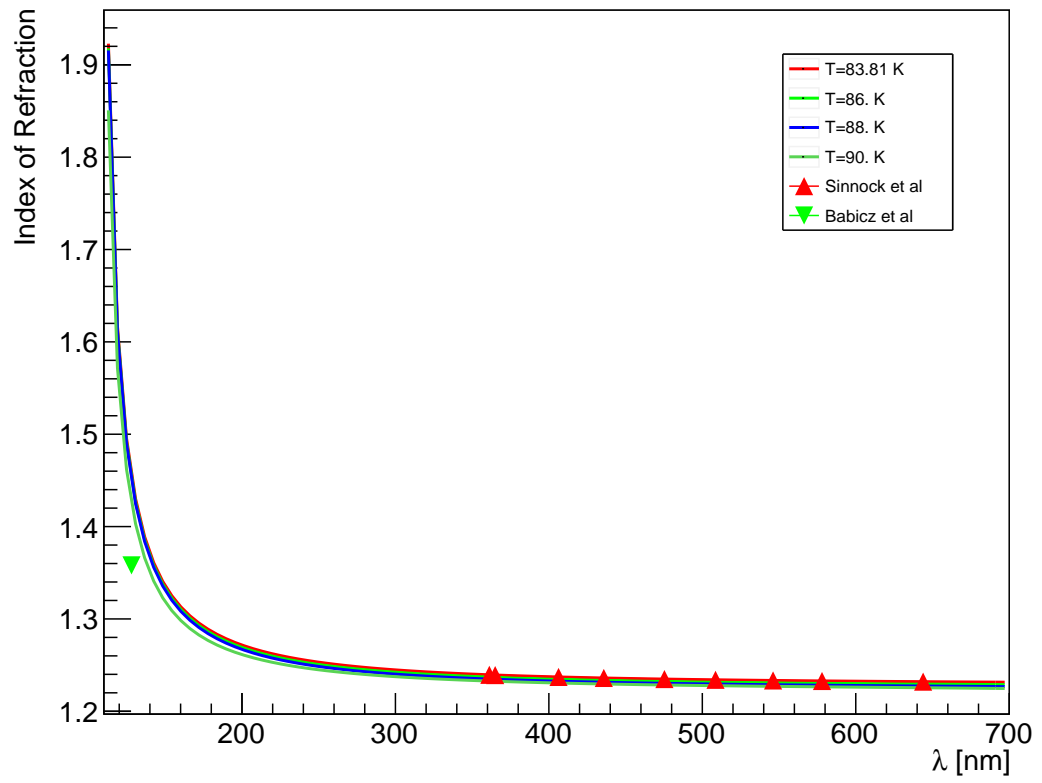


Figure 3: Refraction index as a function of λ for various temperatures. The experimental data is from [6] and [11].

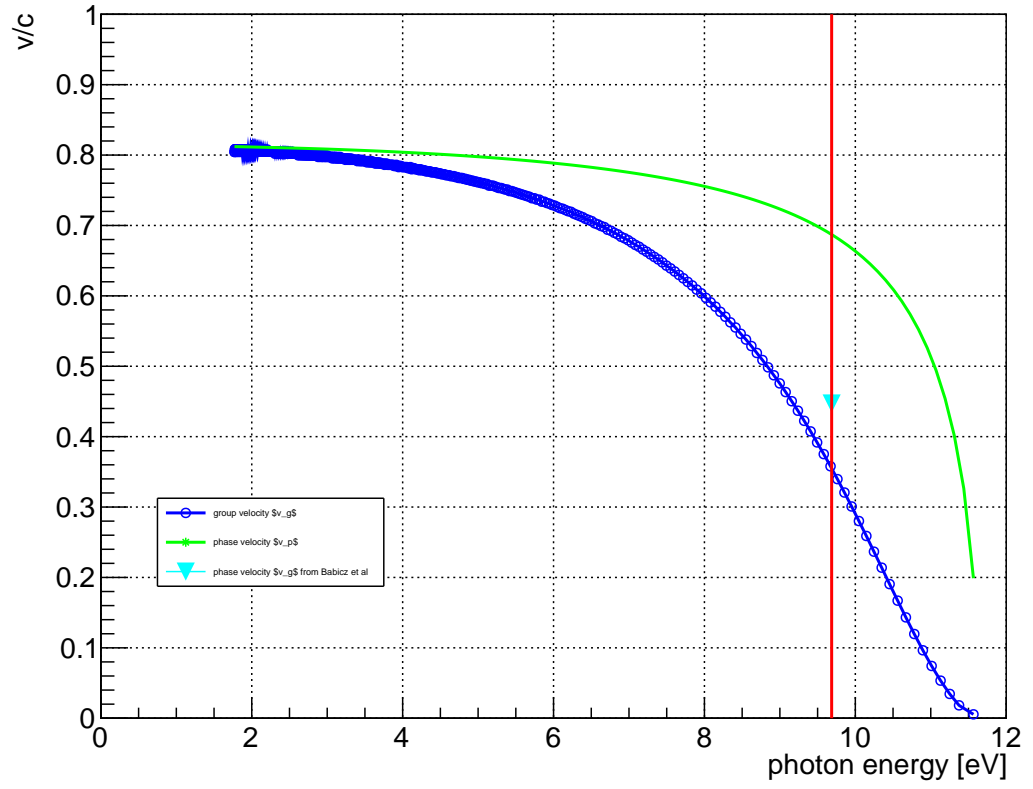


Figure 4: Group velocity v_g and phase velocity v_p compared to speed of light in a vacuum c_0 as function of photon energy. The red line represents the position of the scintillation peak (128 nm). Experimental data point for v_g at 128 nm from [11].

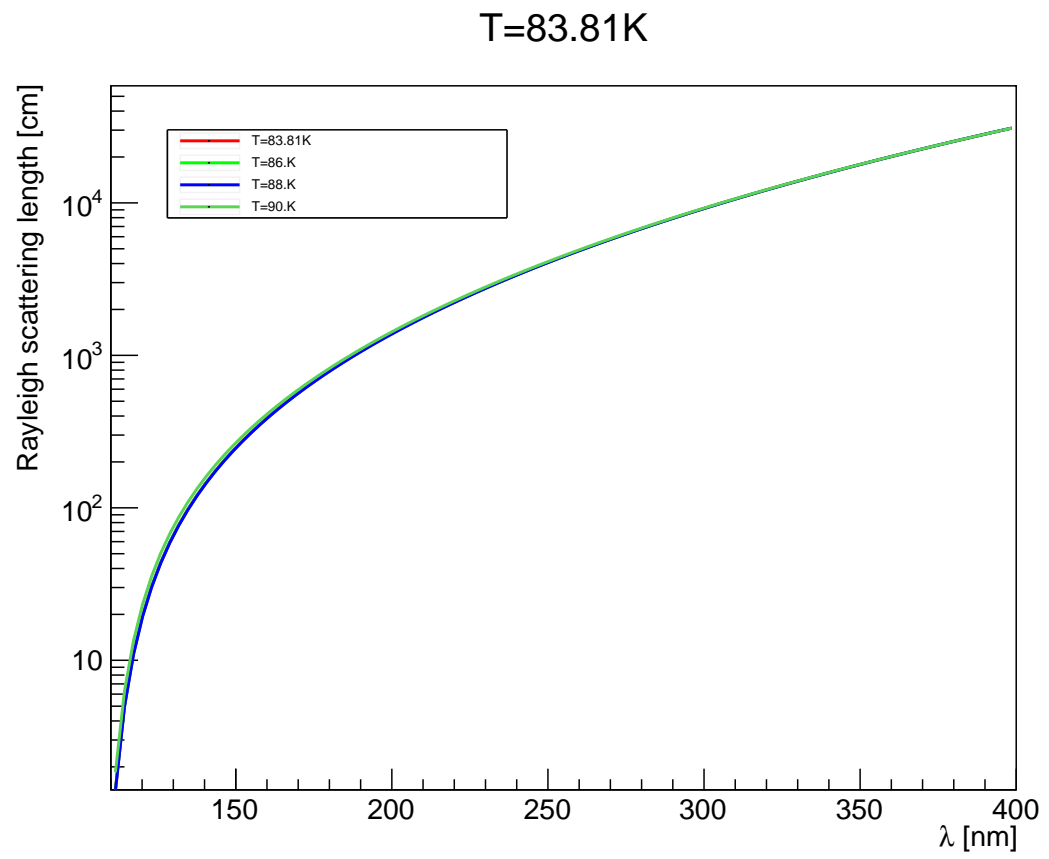


Figure 5: Rayleigh scattering length as a function of λ .

4 Boundary processes

4.1 Refraction and total internal Reflection

Refraction is the change in direction of a wave passing from one medium to another or from a gradual change in the medium. Total internal reflection is the optical phenomenon in which waves arriving at the interface (boundary) from one medium to another (e.g., from water to air) are not refracted into the second ("external") medium, but completely reflected back into the first ("internal") medium. It occurs when the second medium has a higher wave speed (lower refractive index) than the first, and the waves are incident at a sufficiently oblique angle on the interface. For light, refraction follows Snell's law, which states that, for a given pair of media, the ratio of the sines of the angle of incidence θ_1 and angle of refraction θ_2 is equal to the ratio of phase velocities (v_1 / v_2) in the two media, or equivalently, to the indices of refraction (n_2 / n_1) of the two media.

$$\frac{\sin \theta_1}{\sin \theta_2} = \frac{v_1}{v_2} = \frac{n_2}{n_1} \quad (9)$$

in which waves arriving at the interface (boundary) from one medium to another. The Fresnel equations (or Fresnel coefficients) describe the reflection and transmission of light (or electromagnetic radiation in general) when incident on an interface between different optical media. (boundary between liquid Ar and wls, wls and photo-detector.)

4.2 Reflection

Specular reflection, or regular reflection, is the mirror-like reflection of waves, such as light, from a surface. The law of reflection states that a reflected ray of light emerges from the reflecting surface at the same angle to the surface normal as the incident ray, but on the opposing side of the surface normal in the plane formed by the incident and reflected rays. (boundary between liquid Argon and metal walls of cryogenic vessel)

Total Reflectivity Fraction specular / diffuse

5 Photon Detection

5.1 Quantum efficiency and absorption length of the tetraphenyl butadiene (TPB) wave length shifter

WLS materials absorb optical photons and then re-emit them with a longer wavelength. This effect is referred to as photoluminescence. The main properties characterising a WLS material are the absorption and emission spectrum, the rise and decay time, and the multiplication factor.

In addition to the ABSLENGTH variable, which defines the usual absorption of optical photons, the WLSABSLENGTH variable corresponds to absorption which triggers the WLS process.

The extracted VUV absorption lengths at 128 nm was measured to be ≈ 400 nm [9]

The decay time spectrum of the WLS process (also essential for the accurate simulation of the signal shape/timing) is by default a δ -function rather than an exponential spectrum. If an exponential spectrum is to be used, this has to be activated when registering the optical physics process in the Physics List. thicknesses between $0.5 \mu m$ and $3.7 \mu m$

propagate to the TPB surface boundary where they pass into bulk TPB or are reflected, according to Snell's law. The TPB/vacuum interface is modeled as a rough surface using the GLISUR surface model in GEANT4 with a polish value of 0.01 ± 0.09 . No dependence of the reemission spectrum of TPB on incident wavelength was observed

optical material property	Geant4 Keyword	value
Emission spectrum	WLSCOMPONENT	see Figure 6 (from [9])
Absorption length	WLSABSLENGTH	400nm at 128nm
emission time constant	WLSTIMECONSTANT	0.5ns
Emission spectrum	WLSCOMPONENT2	NA
Absorption length	WLSABSLENGTH2	NA
emission time constant	WLSTIMECONSTANT2	NA

Table 6: Properties of the TPB wavelength shifter (values from [9]).

$$f(\lambda|A, \alpha, \sigma_1, \mu_1, \sigma_2, \mu_2, C) = C \times \left(A \frac{\alpha}{2} e^{\frac{\alpha}{2}(2\mu_1 + \alpha\sigma_1^2 - 2\lambda)} \times \text{erfc} \left(\frac{\mu_1 + \alpha\sigma_1^2 - \lambda}{\sqrt{2}\sigma_1} \right) + \frac{(1-A)}{\sqrt{2\sigma_2^2\pi}} e^{\frac{-(\lambda - \mu_2)^2}{2\sigma_2^2}} \right) \quad (10)$$

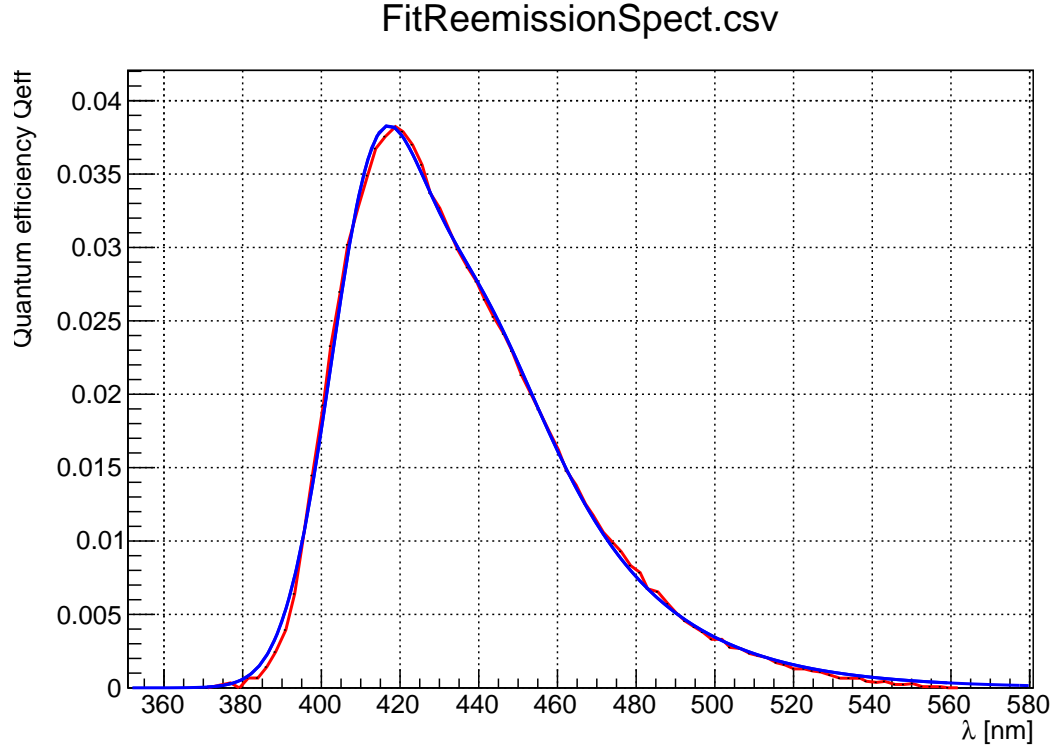


Figure 6: TPB wave length spectrum extracted from [9](red) with fit superimposed (blue).

Parameter	Value	Uncertainty (\pm)
A	0.696747	0.0212897
α	0.0393111	0.000930845
σ_1	15.1679 nm	0.55542 nm
μ_1	421.863 nm	1.11517 nm
σ_2	11.2098 nm	0.203709 nm
μ_2	411.922 nm	0.299189 nm
C	2.27282	0.0092999

Table 7: Results of fit to the TPB wavelength shifter spectrum.

6 Software

6.1 physics configuration

"*FTFP_BERT + OPTICAL + STEPLIMIT*" refeence physics list optical
physics constructor steplimit constructor neutron killer

```
G4OpticalParameters::Instance()->SetProcessActivation("Cerenkov", true);
G4OpticalParameters::Instance()->SetProcessActivation("Scintillation", true);
G4OpticalParameters::Instance()->SetScintFiniteRiseTime(false);
G4OpticalParameters::Instance()->SetProcessActivation("OpAbsorption", true);
G4OpticalParameters::Instance()->SetProcessActivation("OpRayleigh", true);
G4OpticalParameters::Instance()->SetProcessActivation("OpMieHG", false);
G4OpticalParameters::Instance()->SetProcessActivation("OpWLS", true);
G4OpticalParameters::Instance()->SetProcessActivation("OpWLS2", false);
G4OpticalParameters::Instance()->SetCerenkovStackPhotons(false);
G4OpticalParameters::Instance()->SetScintStackPhotons(false);
G4OpticalParameters::Instance()->SetScintTrackSecondariesFirst(
    true); // only relevant if we actually stack and trace the optical photons
G4OpticalParameters::Instance()->SetCerenkovTrackSecondariesFirst(
    true); // only relevant if we actually stack and trace the optical photons
G4OpticalParameters::Instance()->SetCerenkovMaxPhotonsPerStep(100);
G4OpticalParameters::Instance()->SetCerenkovMaxBetaChange(10.0);
G4OpticalParameters::Instance()->SetWLSTimeProfile("exponential");
G4OpticalParameters::Instance()->SetWLS2TimeProfile("exponential");
```

6.2 Sensitive Detector PhotonSD and PhotonHit

The sensitive detector for optical photons is PhotonSD which produces PhotonHits. It registers the properties of every photon that reaches the sensitive volume and then kills the optical photon. The photon properties that are collected are listed in Table 6.2. Note that properties like quantum efficiency, geometrical fill factor etc. are regarded as detector response and are not included in the simulation.

name	meaning
unsigned int fid	Detector ID of the Photodetector
unsigned int fpid	ID of Process that produced the photon (Scintillation, Cerenkov, WLS)
G4double fwavelength	wavelength of photon in nm
G4double ftime	arrival time at photon at detector surface
G4ThreeVector fposition	global position where photon hits the detector
G4ThreeVector fdirection	direction of photon
G4ThreeVector fpolarization	polarization of photon

Table 8: Data members of the PhotonHit class.

7 Conclusions and Outlook

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