

Homework 3

1. Jupiter is the largest planet in the solar system. It is surrounded by a large, rotating disc of *plasma* (ionized gas) which comes originally from the volcanoes on the surface of the moon, Io. Disc material situated at different distances r from the centre of Jupiter rotates at different values of angular velocity ω . The variation of angular velocity of the disc material as a function of distance is approximately described by the following differential equation:

$$\frac{1}{2}r \frac{d\omega}{dr} + \omega = \frac{(1-\omega)\kappa_0}{\dot{M}r^4}$$

The constant $\kappa_0 = 10^9$ represents the efficiency of the interaction between the rotating planet and the plasma disc.

The constant \dot{M} is the rate at which material is added to the disc, in units of kg / s. The units for ω are chosen so that a value $\omega=1$ indicates perfect ~~corotation~~ with the planet. The units for r are chosen so that a value $r=1$ represents one planet radius.

Use `NDSolve` to obtain and plot the five solutions of the above differential equation for the following five values of \dot{M} : 500 kg/s, 1000 kg/s, 2000 kg/s, 4000 kg/s and the hypothetical case $\dot{M}=\infty$ (infinite mass loading). Use the boundary condition: $\omega=1$ at $r=6$ (Io orbit), and obtain the solutions over the range $r=6$ to 40.

Plot your five solutions on the *same* graphics panel, label the axes appropriately, and clearly indicate which plot corresponds to which value of the mass loading rate, \dot{M} . For each of the five solutions, determine the value of r for which $\omega=0.7$ - if such a point exists.

For the $\dot{M}=\infty$ solution only, make a separate plot of the quantity $l = \omega r^2$ (specific angular momentum) as a function of r . Label the axes appropriately, and comment on the appearance of the plot.

[Q1: 19 marks]

2. The differential equation in part 1 can be expressed as a *finite-difference* equation. If we replace the exact derivative with a centred-difference 'derivative', and rearrange, this finite-difference version of the equation is equivalent to the set of equations:

$$\frac{1}{2}r_j(\omega_{j+1} - \omega_{j-1}) + 2\Delta r \left(\frac{\kappa_0}{\dot{M}r_j^4} + 1 \right) \omega_j = \frac{2\kappa_0\Delta r}{\dot{M}r_j^4}$$

Where these equations are defined for the interior points ($j=2, 3, \dots, (n-1)$) of a uniformly spaced grid in radial distance, given by

$r_k = (6 + (k-1)\Delta r)$ ($k=1, 2, \dots, n$), with grid spacing $\Delta r = (40-6)/(n-1)$, and ω_j is the value of the solution at radial distance r_j .

This set of $(n-2)$ equations is thus equivalent to the matrix equation:

$$\mathbf{M} \begin{pmatrix} \omega_2 \\ \vdots \\ \omega_{n-1} \end{pmatrix} = \mathbf{b},$$

where \mathbf{M} is a $(n-2) \times (n-2)$ tridiagonal matrix, whose elements are related to the constant coefficients of the angular velocity values. \mathbf{b} is a vector of constant values, some of which involve the exterior values $\omega = \omega_1 = 1$ at $r = r_1 = 6$ (~~corotation~~ with the planet at the closest distance) and $\omega = \omega_n$ at $r = r_n = 40$.

Using a grid of $n=23$ points, $\dot{M}=500$ kg/s, and ω_n equal to the corresponding value from the `NDSolve` solution (obtained in question 1), compute all the elements of the matrix \mathbf{M} and the vector \mathbf{b} . Then solve the matrix equation to obtain the vector of angular velocity values at the interior points of the grid.

Plot the set of n points given by $\{r_k, \omega_k\}$ ($k=1, 2, \dots, n$) which represent the matrix solution - label the axes appropriately. On the same graphics panel, plot a continuous curve illustrating the `NDSolve` solution you obtained for $\dot{M}=500$

kg/s in question 1. Calculate the maximum value of the absolute relative difference $|(\omega_k - \omega(r_k))/\omega(r_k)|$ between the matrix solution ω_k and the NDSolve solution evaluated at the same points, denoted here by $\omega(r_k)$. Is the matrix solution in agreement with NDSolve to better than 0.5 percent?

[Q2: 19 marks]

Solutions are due by 5pm on Monday March 12th. Make a copy of your final solutions with the output deleted (Cell|Delete All Output), name that file yourname_hw2 with an appropriate substitution for your-name, upload that file to Moodle and submit it. The file will be evaluated using Evaluation|Evaluate Notebook, so make sure the final file you submit will produce the results you expect when that is done. If you submit a file without deleting the output, the output will be deleted and regenerated for purposes of assessment.