

***TITLE:- Simulate Orbital Eclipse & Flux Deficiency Stress Test***

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***Scientific Summary:- This Simulation models a Dyson Node Entering Orbital Shadow Phase, such as***

- ★ *Lunar Eclipse*
- ★ *Planetary Occlusions*
- ★ *Artificial Obstruction (from other Dyson nodes)*

***During this time, Solar Flux drops Significantly, potentially to Zero. The Node must Survive off internal Reserves or Slow its Systems – but must not fallback unless Conditions become Catastrophic (i.e., Dangerously Cold).***

***Objectives:-***

- ★ *Test Thermal Behavior of Node in Low-to-no flux environments.*
- ★ *Log how Internal temperature drops due to flux loss.*
- ★ *Simulate Buffered Operation.*
- ★ *Confirm no Premature fallback,*

***Physics Behind It:- When a Spaceborne System enters a planetary Shadow;***

- ★ *It receives Negligible Radiative Energy.*
- ★ *Without Energy, Internal temperature begins to fall due to Heat Loss.*
- ★ *Satellites in Earth's Shadow Cool by ~200°C within 60- 90 minutes.*

***In Our Aim;***

★ We Assume  $1361 \text{ W/m}^2 \gg 200 - 400 \text{ W/m}^2$  during Eclipse.

★ Radiative Loss Simulated by slow temperature Decreases.

★ Node must maintain Stability until Sun Returns.

*System Response System:-*

(i) Shadow flux modeling

$\gg F(t) = 300 + \sin(i) * 50$ , (simulates low- flux turbulence during partial eclipse).

(ii) Cooling Equation

$\gg T_{i+1} = T_i - \gamma + \delta * (i \bmod 5)$ ,  $\gamma = 0.002$  (passive Radiation to space) and  $\delta = 0.0001$  (Fluctuations for Instability).

(iii) Fallback Trigger (Low temperature Condition)

$\gg T < 19.5^\circ\text{C} \Rightarrow \text{fallback} = \text{true}$ .

*Expected Results:-*

★ Gradual Temperature Decreases from  $24^\circ\text{C}$  towards  $\sim 20^\circ\text{C}$ .

★ Node Survives  $\sim 3000+$  cycles before Any Risk.

★ Final Log Shows “Eclipse Phase Complete” if fallback not triggered.

★ If temperature drops too fast  $\gg$  Node Activates passive heating + logs fallback.