2.10. Geometrical optics: Optical instruments



TEASER?

• ANSWER FRAGMENT

... dummy ...

geometrical optics: optical instruments

- using reflection and refraction to study lenses and optical instruments
- covers simple devices like magnifying glasses to complex ones like telescopes and the human eye
- introduces key concepts such as image formation, aberrations, and optical limitations

primer on lenses & the lensmaker's equation

- lenses have two surfaces which can be planar, convex or concave
- convex lenses are converging while concave lenses are diverging
- when a ray enters air-to-lens and then lens-toair, it is refracted according to snell's law
- focal length is the same on both sides (assuming a thin lens)

lensmaker's equation

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frac1f = (n-1) \ left( \ frac1R_1 + \ frac1R_2 \ right)
```

- *n* is the index of refraction
- R_1 and R_2 are the radii of curvature (positive for convex, negative for concave)

thin lenses model

- approximates the lens as having negligible thickness
- ullet assumes paraxial rays (small angle approximations: sin(theta)

 $theta) \ approx \ theta$

1

```
tan(
theta)
approx
theta
cos(
theta)
approx1
```

- uses a single principal plane instead of two separate refraction planes
- distinguishes between real and virtual focal points

image formation at thin lenses via ray tracing

- three principal rays determine the image location:
 - parallel ray: travels parallel to the principal axis
 - focal point ray: passes through the focal point
 - central (optical center) ray: goes
 through the lens center with little

deviation

- for convex lenses:
 - parallel ray refracts through the focal point on the opposite side
- for concave lenses:
 - refracted rays appear to diverge from the focal point on the same side as the object

thin lens equation & lateral magnification

- ullet object distance d_o and image distance d_i are measured along the principal axis
- for convex lenses, the thin lens equation is:

$$frac1d_{o} + \ frac1d_{i} = \ frac1f$$

• for concave lenses, the equation becomes:

$$egin{aligned} frac1d_o - \ frac1d_i = \ frac1f \end{aligned}$$

lateral magnification is defined as:

$$m = \ frach_i h_o = - \ fracd_i d_o$$

- sign conventions:
 - h_o always positive
 - h_i positive if upright, negative if inverted
 - d_o positive for real objects
 - d_i positive for real images and negative for virtual images
 - f positive for convex and negative for concave lenses

combining lenses

- analyze the first lens using the thin lens equation:
 - lacktriangledown calculate d_{i1} and magnification $m_1 = \ frac d_{i1} d_{o1}$
- treat the image from the first lens as the object for the second lens

- determine the second object distance
 d_{o2} based on the lens separation and image location
- apply the thin lens equation for the second lens:
 - ullet calculate d_{i2} and magnification $m_2 = \ frac d_{i2} d_{o2}$
- total magnification is:

$$m_{eff}=m_1 \ times m_2$$

careful use of sign conventions is crucial

pinhole cameras

- a lensless camera where a tiny pinhole focuses light onto a light-sensitive surface
- image formation is based on similar triangles:

$$frach_i d_i = \ frach_o d_o$$

optimal pinhole diameter is given by:

$d_{pin} \ approx$

sqrt2.44 lambdaf

- lambda is the wavelength of light
- f is the distance from the pinhole to the image plane
- advantages:
 - simple, inexpensive, infinite depth of field
- disadvantages:
 - low light sensitivity and sharpness limited by diffraction

the human eye

- functions similarly to a camera with additional biological sophistication
- key components:
 - cornea (
 n
 approx1.376
): primary site of light bending

- iris and pupil: control light entry
- lens (n approx 1.385 text to 1.405): adjusts shape (accommodation) to focus light
- retina and fovea: convert light into neural signals; fovea is the sharpest point
- the eye has a fixed image distance
- corrective lenses are used to adjust the focus for nearsightedness, farsightedness, and astigmatism

corrective lenses

- nearsightedness (myopia):
 - image forms in front of the retina

- corrected using diverging (negative)
- farsightedness (hyperopia):
 - image forms behind the retina
 - corrected using converging (positive)lenses
- astigmatism:
 - irregular cornea or lens shape causes blurred images
 - corrected with cylindrical lenses

magnifying glass

- a simple magnifier that uses a single converging lens to enlarge the apparent size of objects
- ullet the apparent size is based on the angular magnification M = fractheta'
- with a relaxed eye (image at infinity):
 - object is placed at the focal point f
 - angular magnification is:

$$M = fracNf$$

where

$$N=25$$

, textcm is the near point

- when the eye is focused at the near point:
 - ullet the virtual image is formed at $d_i = -N$
 - using the thin lens equation:

$$egin{aligned} frac1f = \ frac1d_o - \ frac1N \end{aligned}$$

• the object distance becomes:

$$d_o = \ frac N f N + f$$

angular magnification is then:

$$M=1+ \\ fracNf$$

 a larger subtended angle leads to a larger apparent size

Telescopes

introduction

- telescopes are optical instruments designed to magnify distant objects
- they allow observation of celestial bodies and faraway phenomena in greater detail

components

- objective lens/mirror collects light from the object and forms a real image
- eyepiece lens magnifies the image formed by the objective

angular magnification

ullet the total angular magnification M of a telescope is defined as the ratio of the angle subtended by the image at the eye to the angle subtended by the object as seen by the unaided eye

ullet for an objective with focal length f_o and an eyepiece with focal length f_e , the magnification is given by

$$M=rac{ heta'}{ heta}=rac{h/f_e}{h/f_o}=rac{f_o}{f_e}$$

ullet with the small angle approximation $heta'pprox h/f_e$ and $hetapprox h/f_o$

image formation

- the image of the first lens (objective) becomes the object for the second lens (eyepiece)
- objective lens forms a real, inverted image near its focal point
- eyepiece lens acts as a magnifying glass producing a virtual, inverted final image types of telescopes

REFRACTING TELESCOPES (KEPLERIAN)

- design uses two converging lenses
- objective lens has a long focal length to collect light and form a real, inverted, diminished image at its focal plane

- eyepiece lens has a short focal length and magnifies the real image
- advantages include simple design and good image quality for small apertures
- disadvantages include heavy, expensive objective lenses and chromatic aberration

REFLECTING TELESCOPES (NEWTONIAN)

- design uses a concave mirror as the objective to collect and focus light
- light reflects off the primary mirror forming a real, inverted image at the focal point
- a secondary flat mirror reflects the image to the side for viewing through an eyepiece
- advantages include absence of chromatic aberration and easier manufacture of large mirrors
- disadvantages include diffraction effects from the secondary mirror and the need for periodic realignment

Aberration

introduction

- aberration refers to imperfections in image formation due to the breakdown of the small angle approximation
- real lenses have thickness and imperfections that can lead to blurring or distortion types of aberrations
 - monochromatic aberrations occur with a single wavelength and include spherical aberration, coma, astigmatism, curvature of field, and distortion
 - chromatic aberration arises from dispersion where different wavelengths are refracted by different amounts

spherical aberration

- occurs when rays far from the optical axis focus at different points than rays close to the axis
- causes a blurred image or a circle of least confusion

 correction methods include aspherical lenses, aperture stops, or compound lenses
 coma

- causes off-axis points to be imaged as asymmetrical, comet-shaped blurs
- results from off-axis rays being refracted by different amounts
- can be minimized using aspherical or carefully designed compound lenses

astigmatism

- happens when a lens or mirror has different focal lengths in two perpendicular planes
- a point object is imaged as two short perpendicular line segments
- correction involves using a combination of lenses or cylindrical lenses

curvature of field

- the image of a flat object is formed on a curved surface instead of a flat plane
- results in a sharp center with blurred edges or vice versa

 field-flattening lenses or specific lens designs can correct this aberration

distortion

- causes the shape of the image to differ from the shape of the object
- barrel distortion makes straight lines bow outwards and pincushion distortion makes them bow inwards
- symmetric compound lenses or software correction can reduce distortion

chromatic aberration

- arises because the refractive index of a lens varies with the wavelength of light
- results in colored fringes especially at high contrast edges
- achromatic doublets or apochromatic lenses are used for correction

aberrations in mirrors

 spherical aberration in mirrors occurs for the same geometric reasons as in lenses

- parabolic mirrors can correct spherical aberration for objects at infinity
- mirrors do not exhibit chromatic aberration because reflection is independent of wavelength

summary of aberrations

- spherical aberration: caused by the spherical shape of the lens/mirror; corrected by aspherical lenses, aperture stops, or compound lenses
- coma: off-axis rays cause a comet-shaped blur;
 minimized with aspherical or compound lenses
- astigmatism: different focal lengths in perpendicular planes; corrected by lens combinations or cylindrical lenses
- curvature of field: focal points on a curved surface lead to uneven focus; corrected with field-flattening lenses
- distortion: non-uniform magnification leads to curved lines; minimized with symmetric compound lenses or digital correction

 chromatic aberration: dispersion causes colored fringes; corrected with achromatic doublets or apochromatic lenses