

# Contracted Bianchi identities – Field equations

Hendrik

December 3, 2013

## 1 Teleparallel Gravity

As discussed in the document `cartan_geo_ext.pdf` the geometric setting underlying Teleparallel Gravity is that of a Riemann-Cartan geometry with vanishing curvature. The Bianchi identities for a generic Riemann-Cartan\* geometry are given by

$$\begin{aligned}dR + [\omega, R] &\equiv 0 , \\dT + [\omega, T] + [e, R] &\equiv 0 ,\end{aligned}$$

which in the case of Teleparallel Gravity reduce to

$$\begin{aligned}0 &\equiv 0 , \\dT + [\omega, T] &\equiv 0 .\end{aligned}$$

Apart from these Bianchi identities for the torsion  $T$ , there are two identities at hand for the corresponding contortion  $K$ . The latter is defined through the splitting  $\omega = \hat{\omega} + K$ , where  $\hat{\omega}$  is the unique Levi-Civita spin connection, i.e., the spin connection without torsion. Introducing the  $\mathfrak{so}(1, 3)$ -valued two form  $Q := dK + [\omega, K] - \frac{1}{2}[K, K]$ , these identities are of the form (for a generic RC geometry)

$$\begin{aligned}dQ + [\omega, Q] - [K, Q] + [K, R] &\equiv 0 , \\[e, Q] - [e, R] &\equiv 0 .\end{aligned}$$

Specializing for Teleparallel Gravity, one finds that

$$dQ + [\omega, Q] - [K, Q] \equiv 0 , \tag{1.1a}$$

$$[e, Q] \equiv 0 . \tag{1.1b}$$

Of course, the two sets of Bianchi identities for  $\omega$  and  $K$  are related by considering the identities for the Levi-Civita connection  $\hat{\omega}$ .

---

\*RC geometry.

Introducing the notation  $\mathring{D} := d + \mathring{\omega}$ , the first identity (1.1a) can be expanded as  $\mathring{D}_{[\rho} Q^{ab}_{\mu\nu]} \equiv 0$ . Contracting this equation twice with the vielbein results in

$$\mathring{D}_\rho \mathcal{Q} - 2\mathring{D}_\rho e_a^\mu Q i_a^\mu + 2e_a^\mu \mathring{D}_\mu Q i_a^\rho - 2e_a^\mu \mathring{D}_\mu e_b^\nu Q^{ab}_{\nu\rho} \equiv 0 ,$$

where the notation  $\mathcal{Q} := Q^{ab}_{\mu\nu} e_a^\mu e_b^\nu$  and  $Q i_a^\mu := Q^{ab}_{\mu\nu} e_b^\nu$  has been used.\* The vielbein postulate  $D_\rho e_a^\mu = -\Gamma^\mu_{\nu\rho} e_a^\nu$  allows us to eliminate the covariant derivatives on the tetrads, which renders

$$\begin{aligned} 0 &\equiv e_a^\mu \mathring{D}_\mu Q i_a^\rho - e_a^\mu \mathring{\Gamma}^\sigma_{\mu\rho} Q i_a^\sigma - \frac{1}{2} \partial_\rho \mathcal{Q} \\ &\equiv \mathring{\nabla}_\mu (Q i_a^\mu{}_\rho - \frac{1}{2} \delta^\mu_\rho \mathcal{Q}) . \end{aligned}$$

This equation can be rewritten as follows:

$$\begin{aligned} 0 &\equiv \mathring{\nabla}_\mu Q i_a^\mu e_a^\rho + Q i_a^\mu \mathring{\nabla}_\mu e_a^\rho - \frac{1}{2} \mathring{\nabla}_\mu e_a^\rho e_a^\mu \mathcal{Q} - \frac{1}{2} e_a^\rho \mathring{\nabla}_\mu e_a^\mu \mathcal{Q} - \frac{1}{2} e_a^\rho e_a^\mu \partial_\mu \mathcal{Q} \\ &= \partial_\mu Q i_a^\mu e_a^\rho + \partial_\mu \ln e Q i_a^\mu e_a^\rho - Q i_a^\mu \mathring{\omega}^a_{b\mu} e_b^\rho + \frac{1}{2} \mathring{\omega}^a_{b\mu} e_b^\rho e_a^\mu \mathcal{Q} \\ &\quad - \frac{1}{2} e_a^\rho \partial_\mu e_a^\mu \mathcal{Q} - \frac{1}{2} e_a^\rho \partial_\mu \ln e e_a^\mu \mathcal{Q} - \frac{1}{2} e_a^\rho e_a^\mu \partial_\mu \mathcal{Q} \\ &= e_a^\rho (\partial_\mu \ln e Q i_a^\mu + \mathring{D}_\mu Q i_a^\mu - \frac{1}{2} \partial_\mu \ln e e_a^\mu \mathcal{Q} - \frac{1}{2} \mathring{D}_\mu e_a^\mu \mathcal{Q} - \frac{1}{2} e_a^\mu \partial_\mu \mathcal{Q}) , \end{aligned}$$

which, given that the vierbein be invertible, is true if and only if

$$\mathring{D}_\mu (e Q i_a^\mu - \frac{1}{2} e e_a^\mu \mathcal{Q}) \equiv 0 . \quad (1.2)$$

To conclude what conditions the contracted Bianchi identity (1.2) puts on the field equations of Teleparallel Gravity, we further work out the tensors  $Q i_a^\mu$  and  $\mathcal{Q}$ . These were defined as contractions of the Lorentz algebra valued two-form

$$Q^{ab}_{\mu\nu} = D_\mu K^{ab}_{\nu} - D_\nu K^{ab}_{\mu} - K^a_{c\mu} K^{cb}_{\nu} + K^a_{c\nu} K^{cb}_{\mu} .$$

Let us first calculate the scalar function  $\mathcal{Q} = Q^{ab}_{\mu\nu} e_a^\mu e_b^\nu$ :

$$\begin{aligned} \mathcal{Q} &= \partial_\mu K^{\mu\nu}_{\nu} - D_\mu e_a^\mu K^{a\nu}_{\nu} - D_\mu e_b^\nu K^{\mu b}_{\nu} - \partial_\nu K^{\mu\nu}_{\mu} + D_\nu e_a^\mu K^{a\nu}_{\mu} \\ &\quad + D_\nu e_b^\nu K^{\mu b}_{\mu} - K^\mu_{\rho\mu} K^{\rho\nu}_{\nu} + K^\mu_{\rho\nu} K^{\rho\nu}_{\mu} \\ &= 2\partial_\mu K^{\mu\nu}_{\nu} + \Gamma^\mu_{\rho\mu} e_a^\rho K^{a\nu}_{\nu} + \Gamma^\nu_{\rho\mu} e_b^\rho K^{\mu b}_{\nu} - \Gamma^\mu_{\rho\nu} e_a^\rho K^{a\nu}_{\mu} \\ &\quad - \Gamma^\nu_{\rho\nu} e_b^\rho K^{\mu b}_{\mu} - K^\mu_{\rho\mu} K^{\rho\nu}_{\nu} + K^\mu_{\rho\nu} K^{\rho\nu}_{\mu} \\ &= 2\partial_\mu K^{\mu\nu}_{\nu} + 2\mathring{\Gamma}^\mu_{\rho\mu} K^{\rho\nu}_{\nu} + 2K^\mu_{\rho\mu} K^{\rho\nu}_{\nu} + 2\mathring{\Gamma}^\nu_{\rho\mu} K^{\mu\rho}_{\nu} + 2K^\nu_{\rho\mu} K^{\mu\rho}_{\nu} \\ &\quad - K^\mu_{\rho\mu} K^{\rho\nu}_{\nu} + K^\mu_{\rho\nu} K^{\rho\nu}_{\mu} \\ &= \frac{2}{e} \partial_\mu (e K^{\mu\nu}_{\nu}) + K^\nu_{\rho\mu} K^{\mu\rho}_{\nu} - K^\mu_{\rho\mu} K^{\nu\rho}_{\nu} . \end{aligned}$$

---

\*Let us note, *en passant*, that due to the identity (1.1b),  $Q i_{\mu\nu}$  is a symmetric tensor.

Subsequently we take a better look at  $Qi^\mu_a = Q^{cb}{}_{\rho\nu} e_c^\mu e_a^\rho e_b^\nu$ :

$$\begin{aligned}
Qi^\mu_a &= D_\rho(e_a^\rho K^{\mu\nu}{}_\nu) - D_\rho e_c^\mu e_a^\rho K^{c\nu}{}_\nu - D_\rho e_a^\rho K^{\mu\nu}{}_\nu - e_a^\rho D_\rho e_b^\nu K^{\mu b}{}_\nu - D_\nu K^{\mu\nu}{}_a \\
&\quad + D_\nu e_c^\mu K^{c\nu}{}_a + D_\nu e_a^\rho K^{\mu\nu}{}_\rho + D_\nu e_b^\nu K^{\mu b}{}_a - K^\mu{}_{\rho a} K^{\rho\nu}{}_\nu + K^\mu{}_{\rho\nu} K^{\rho\nu}{}_a \\
&= D_\nu K^{\nu\mu}{}_a - \Gamma^\nu{}_{\rho\nu} e_b^\rho K^{\mu b}{}_a + D_\rho(e_a^\rho K^{\mu\nu}{}_\nu) + \Gamma^\rho{}_{\sigma\rho} e_a^\sigma K^{\mu\nu}{}_\nu + \Gamma^\mu{}_{\sigma\rho} e_c^\sigma e_a^\rho K^{c\nu}{}_\nu \\
&\quad + e_a^\rho \Gamma^\nu{}_{\sigma\rho} e_b^\sigma K^{\mu b}{}_\nu - \Gamma^\mu{}_{\rho\nu} e_c^\rho K^{c\nu}{}_a - \Gamma^\rho{}_{\sigma\nu} e_a^\sigma K^{\mu\nu}{}_\rho - K^\mu{}_{\rho a} K^{\rho\nu}{}_\nu \\
&\quad + K^\mu{}_{\rho\nu} K^{\rho\nu}{}_a \\
&= e^{-1} D_\nu(e K^{\nu\mu}{}_a) + K^\nu{}_{\rho\nu} K^{\rho\mu}{}_a + e^{-1} D_\nu(e e_a^\nu K^{\mu\rho}{}_\rho) + K^\rho{}_{a\rho} K^{\mu\nu}{}_\nu \\
&\quad + e_a^\rho \Gamma^\mu{}_{\sigma\rho} K^{\sigma\nu}{}_\nu + e_a^\rho (\Gamma^\nu{}_{\sigma\rho} - \Gamma^\nu{}_{\rho\sigma}) K^{\mu\sigma}{}_\nu - K^\mu{}_{\rho\nu} K^{\rho\nu}{}_a - K^\mu{}_{\rho a} K^{\rho\nu}{}_\nu \\
&\quad + K^\mu{}_{\rho\nu} K^{\rho\nu}{}_a \\
&= e^{-1} D_\nu(e K^{\nu\mu}{}_a + e e_a^\nu K^{\mu\rho}{}_\rho) + T^\nu{}_{\sigma a} K^{\sigma\mu}{}_\nu + e_a^\rho \Gamma^\mu{}_{\sigma\rho} K^{\sigma\nu}{}_\nu + K^\rho{}_{a\rho} K^{\mu\nu}{}_\nu .
\end{aligned}$$

One may consider the difference

$$\begin{aligned}
Qi^\mu_a - \frac{1}{2} e_a^\mu \mathcal{Q} &= e^{-1} D_\nu(e K^{\nu\mu}{}_a + e e_a^\nu K^{\mu\rho}{}_\rho) + T^\nu{}_{\sigma a} K^{\sigma\mu}{}_\nu + e_a^\rho \Gamma^\mu{}_{\sigma\rho} K^{\sigma\nu}{}_\nu + K^\rho{}_{a\rho} K^{\mu\nu}{}_\nu \\
&\quad - e^{-1} \partial_\nu(e e_a^\mu K^{\nu\rho}{}_\rho) + \partial_\nu e_a^\mu K^{\nu\rho}{}_\rho - \frac{1}{2} e_a^\mu K^\nu{}_{\rho\sigma} K^{\sigma\rho}{}_\nu + \frac{1}{2} e_a^\mu K^\sigma{}_{\rho\sigma} K^{\nu\rho}{}_\nu \\
&= e^{-1} D_\nu(e K^{\nu\mu}{}_a + e e_a^\nu K^{\mu\rho}{}_\rho - e e_a^\mu K^{\nu\rho}{}_\rho) + T^b{}_{\sigma a} K^{\sigma\mu}{}_b + K^\rho{}_{a\rho} K^{\mu\nu}{}_\nu \\
&\quad + e_a^\rho (\Gamma^\mu{}_{\sigma\rho} - \Gamma^\mu{}_{\rho\sigma}) K^{\sigma\nu}{}_\nu - \frac{1}{2} e_a^\mu \mathcal{L}_{\text{tg}} \\
&= \frac{1}{2} e^{-1} D_\nu(e W_a{}^{\nu\mu}) + T^b{}_{\sigma a} K^{\sigma\mu}{}_b - T^\mu{}_{\sigma a} K^{\sigma\nu}{}_\nu + K^\rho{}_{a\rho} K^{\mu\nu}{}_\nu - \frac{1}{2} e_a^\mu \mathcal{L}_{\text{tg}} ,
\end{aligned}$$

where the notation  $W_a{}^{\nu\mu} = 2(K^{\nu\mu}{}_a + e_a^\nu K^{\mu\rho}{}_\rho - e_a^\mu K^{\nu\rho}{}_\rho)$  is used in the last line. Since  $K^{a\nu}{}_\nu = T^{\nu a}{}_\nu$ , one further concludes that

$$\begin{aligned}
Qi^\mu_a - \frac{1}{2} e_a^\mu \mathcal{Q} &= \frac{1}{2} e^{-1} D_\nu(e W_a{}^{\nu\mu}) + T^b{}_{\nu a} K^{\nu\mu}{}_b - T^b{}_{\nu a} e_b^\mu K^{\nu\rho}{}_\rho + T^\rho{}_{\rho a} K^{\mu\nu}{}_\nu - \frac{1}{2} e_a^\mu \mathcal{L}_{\text{tg}} \\
&= \frac{1}{2} e^{-1} D_\nu(e W_a{}^{\nu\mu}) + T^b{}_{\nu a} (K^{\nu\mu}{}_b + e_b^\nu K^{\mu\rho}{}_\rho - e_b^\mu K^{\nu\rho}{}_\rho) - \frac{1}{2} e_a^\mu \mathcal{L}_{\text{tg}} \\
&= \frac{1}{2} [e^{-1} D_\nu(e W_a{}^{\nu\mu}) + T^b{}_{\nu a} W_b{}^{\nu\mu} - e_a^\mu \mathcal{L}_{\text{tg}}] .
\end{aligned}$$

Finally, the contracted Bianchi identities (1.2) imply that

$$\mathring{D}_\mu [D_\nu(e W_a{}^{\nu\mu}) + e T^b{}_{\nu a} W_b{}^{\nu\mu} - e e_a^\mu \mathcal{L}_{\text{tg}}] \equiv 0 . \quad (1.3)$$