# **Spatial Statistics with Image Analysis**

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# Outline

### Spatial Statistics with Image Analysis

Bayesian statistics

Hierarchical modelling

**Estimation Procedures** 

### Spatial Statistics

Stochastic Fields

Gaussian Markov Random Fields

Image Reconstruction

Environmental Data

### Non-Gaussian Data

Examples

### Corrupted Pixels

NDVI

Learn more

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# Spatial Statistics

Many things vary in space and obsevations may depend on what happens at nearby locations. To modell and analys such data we need spatial dependence.

### Spatial Interpolation

Given observations at some locations (pixels),

$$y(\mathbf{u}_i), i = 1 \dots n$$

we want to make statements about the value at unobserved location(s),  $x(u_0)$ .

Stationary Stochastic Processes (FMSF10) in 2+ dimensions!

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# Bayesian modelling

We assume that there is some unknown truth, that we would like to find out about. This "reality" can be measured, usually with measurement variation, and often only partially.

### Bayesian modelling

A Bayesian model consists of

- ▶ A **prior**, "a priori", model for reality, x, given by the probability density  $\pi(x)$ .
- A conditional model for data, y, given reality, with density p(y|x).

The **prior** can be expanded into several **layers** creating a **Bayesian hierarchical model**.

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# Bayes' Formula

How should the **prior** and **data model** be combined to make statements about the reality **x**, given observations of **y**?

### Bayes' Formula

$$p(\mathbf{x}|\mathbf{y}) = \frac{p(\mathbf{y}|\mathbf{x})\pi(\mathbf{x})}{p(\mathbf{y})} = \frac{p(\mathbf{y}|\mathbf{x})\pi(\mathbf{x})}{\int_{\mathbf{x}' \in \Omega} p(\mathbf{y}|\mathbf{x}')\pi(\mathbf{x}') d\mathbf{x}'}$$

p(x|y) is called the **posterior**, or "a posteriori", distribution.

Often, only the proportionality relation

$$p(x|y) \propto p(x,y) = p(y|x)\pi(x)$$

is needed, when seen as a function of x.

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Bayesian statistics Hierarchical Estimatio

### Hierarchical Models

- ▶ We often have some **prior knowledge** of the reality.
- ► Given knowledge of the true reality, what can we say about images and other data?
- Construct a model for observations given that we know the truth.
- ► Given data, what can we say about the unknown reality?

This is the inverse problem.

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# Bayesian hierarchical modelling (BHM)

A hierarchical model is constructed by systematically considering components/features of the data, and how/why these features arise.

### Bayesian hierarchical modelling

A Bayesian hierarchical model typically consists of (at least)

Data model, p(y|x): Describing how observations arise assuming known latent variables x.

Latent model,  $p(x|\vartheta)$ : Describing how the latent variables (reality) behaves, assuming known parameters.

Parameters,  $\pi(\vartheta)$ : Describing our, sometimes vauge, **prior knowledge** of the parameters.

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### **Estimation Procedures**

Maximum A Posteriori (MAP): Maximise the posterior distribution p(x|y) with respect to x.

- Standard optimisation methods
- Specialised procedures, using the model structure

Simulation: Simulate samples from the posterior distribution p(x|y). Estimate statistical properties from these samples. The samples can be seen as representative "possible realities", given the available data.

- Markov chain Monte Carlo (MCMC)
- Gibbs sampling

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Fields GMRF Reconstruction Environmental

# Image Reconstruction

### Spatial Interpolation

Given observations at some locations (pixels),  $y(u_i)$ , i = 1 ... n

we want to make statements about the value at unobserved location(s),  $x(\mathbf{u}_0)$ .

The typical model consists of a latent Gaussian field

$$x \in N(\mu, \Sigma)$$
,

observed at locations  $u_i$ , i = 1, ..., n, with additive Gaussian noise (nugget or small scale variability)

$$y_i = x(\mathbf{u}_i) + \varepsilon_i$$

$$arepsilon_{i}\in\mathsf{N}\left(0,\sigma_{arepsilon}^{2}
ight).$$

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### Stochastic Fields

To perform the reconstruction (interpolation) we need a model for the **spatial dependence** between locations (pixels).

1. Assume a latent Gaussian field

$$extbf{ extit{x}} \in \mathsf{N}\left(\mu, \Sigma
ight).$$

- 2. Assume a regresion model for  $\mu = B\beta$ .
- 3. Assume a parametric (stationary) model for the dependence (covariance)

$$\Sigma_{i,j} = \mathsf{C}(x(\mathbf{u}_i), x(\mathbf{u}_i)) = r(\mathbf{u}_i, \mathbf{u}_i; \vartheta) = r(\|\mathbf{u}_i - \mathbf{u}_i\|; \vartheta).$$

 $r(\mathbf{u}_i, \mathbf{u}_i; \vartheta)$  is called the **covariance function**.

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# Matérn covariances functions

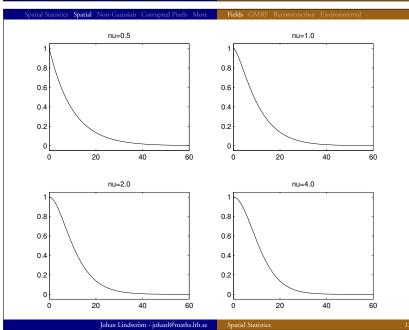
- ▶ One of the most common families of covariance functions is named after Bertil Matérn, who worked for Statens Skogsforskningsinstitut (Forest Research Institute of Sweden).
- ▶ Variance  $\sigma^2 > 0$ , scale parameter x > 0 and shape parameter y > 0

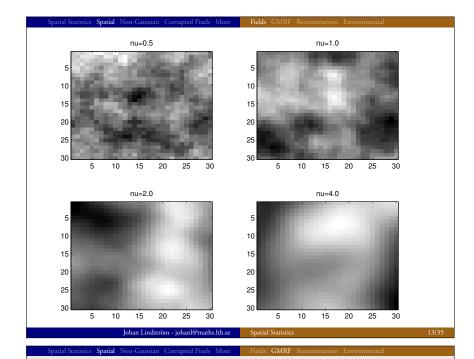
$$r_{M}(\boldsymbol{h}) = \frac{\sigma^{2}}{\Gamma(\nu) 2^{\nu-1}} (\boldsymbol{x} \| \boldsymbol{h} \|)^{\nu} K_{\nu}(\boldsymbol{x} \| \boldsymbol{h} \|), \quad \boldsymbol{h} \in \mathbb{R}^{d},$$

• A measure of the range is given by  $\rho = \sqrt{8\nu/\kappa}$ .

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# A local model

Instead of specifying the covariance function we could consider the **local behaviour** of pixels. A popular model is the **conditional autoregressive**, **CAR(1)** model.

$$\begin{split} x_{ij} &= \frac{1}{4 + \varkappa^2} \left( x_{i-1,j} + x_{i+1,j} + x_{i,j-1} + x_{i,j+1} \right) + \varepsilon, \\ \varepsilon &\in N\left(0, \frac{1}{\tau^2}\right). \end{split}$$

This corresponds to a model for x where

$$extbf{\textit{x}} \in N\left(0, extbf{\textit{Q}}^{-1}
ight),$$

where **Q** is called the precision matrix

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### Matérn covariances

### The Matérn covariance family

The covariance between two points at distance  $\|\boldsymbol{h}\|$  is

$$r_{M}(\boldsymbol{h}) = \frac{\sigma^{2}}{\Gamma(\nu) 2^{\nu-1}} (\boldsymbol{x} \| \boldsymbol{h} \|)^{\nu} K_{\nu} (\boldsymbol{x} \| \boldsymbol{h} \|)$$

Fields with Matérn covariances are solutions to a **Stochastic Partial Differential Equation (SPDE)** (Whittle, 1954),

$$\left(x^2 - \Delta\right)^{\alpha/2} x(\boldsymbol{u}) = \mathcal{W}(\boldsymbol{u}).$$

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# Lattice on $\mathbb{R}^2$

Order  $\alpha = 1$  ( $\nu = 0$ ):

$$\chi^{2} \left[\begin{array}{c} 1 \\ 1 \end{array}\right] + \left[\begin{array}{ccc} -1 \\ -1 & 4 & -1 \\ -1 & \end{array}\right] \approx -\Delta \left(G\right)$$

Order  $\alpha = 2$  ( $\nu = 1$ ):

$$\chi^{4} \left[\begin{array}{c} 1 \\ 1 \\ \end{array}\right] + 2\chi^{2} \left[\begin{array}{cccc} -1 \\ -1 & 4 & -1 \\ -1 \end{array}\right] + \left[\begin{array}{ccccc} 1 \\ 2 & -8 & 2 \\ 1 & -8 & 20 & -8 & 1 \\ 2 & -8 & 2 \\ & 1 \end{array}\right] \\
\approx \Delta^{2} \left(G_{2} = GC^{-1}G\right)$$

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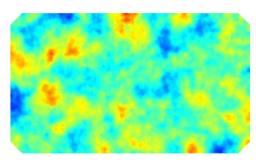
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Spatial models for data

GMRF representations of SPDEs can be constructed for oscillating, anisotropic, non-stationary, non-separable spatio-temporal, and multivariate fields on manifolds.

$$(\varkappa^2 - \Delta)(\tau \mathbf{x}(\mathbf{u})) = \mathcal{W}(\mathbf{u}), \quad \mathbf{u} \in \mathbb{R}^d$$



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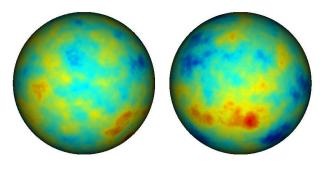
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# Spatial models for data

GMRF representations of SPDEs can be constructed for oscillating, anisotropic, non-stationary, non-separable spatio-temporal, and multivariate fields on manifolds.

$$(\varkappa^2 - \Delta)(\tau \mathbf{x}(\mathbf{u})) = \mathcal{W}(\mathbf{u}), \quad \mathbf{u} \in \Omega$$



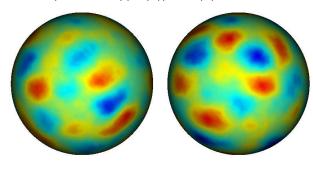
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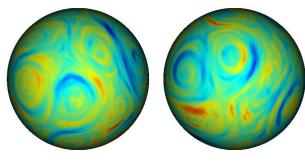
$$(\varkappa^2 e^{i\pi\vartheta} - \Delta)(\tau x(\boldsymbol{u})) = \mathcal{W}(\boldsymbol{u}), \quad \boldsymbol{u} \in \Omega$$



### Spatial models for data

GMRF representations of SPDEs can be constructed for oscillating, anisotropic, non-stationary, non-separable spatio-temporal, and multivariate fields on manifolds.

$$(\mathbf{x}_{\mathbf{u}}^2 + \nabla \cdot \mathbf{m}_{\mathbf{u}} - \nabla \cdot \mathbf{M}_{\mathbf{u}} \nabla)(\tau_{\mathbf{u}} \mathbf{x}(\mathbf{u})) = \mathcal{W}(\mathbf{u}), \quad \mathbf{u} \in \Omega$$



# Image Reconstruction II

Model with observations, y, and latent field, x,

$$\textbf{\textit{y}}|\textbf{\textit{x}}\in N\left(\textbf{\textit{A}}\textbf{\textit{x}},\sigma^{2}\textbf{\textit{I}}\right) \hspace{1cm} \textbf{\textit{x}}\in N\left(\mu,\textbf{\textit{Q}}^{-1}\right).$$

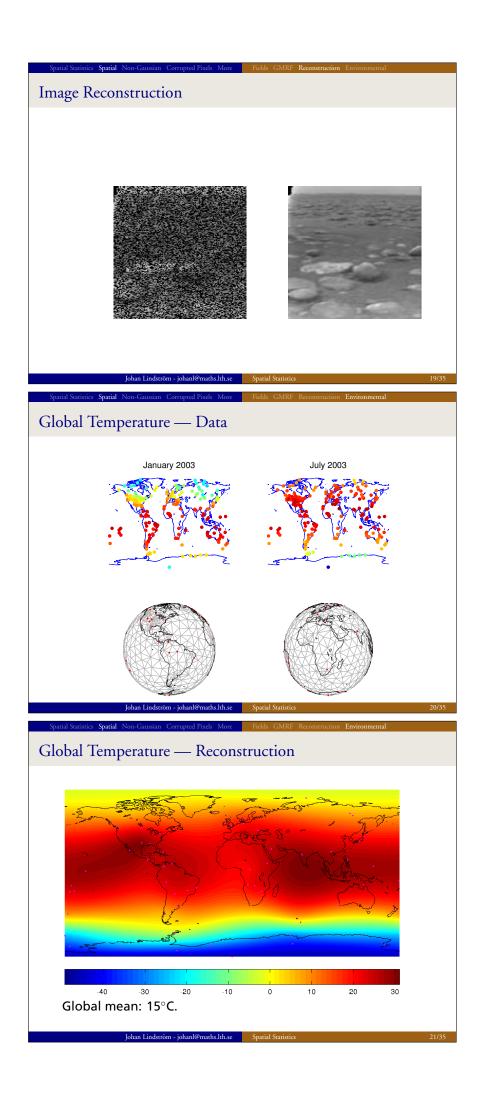
and 
$$Q = \chi^2 C + G$$
 or  $Q = \chi^4 C + 2\chi^2 G + GC^{-1}G$ .

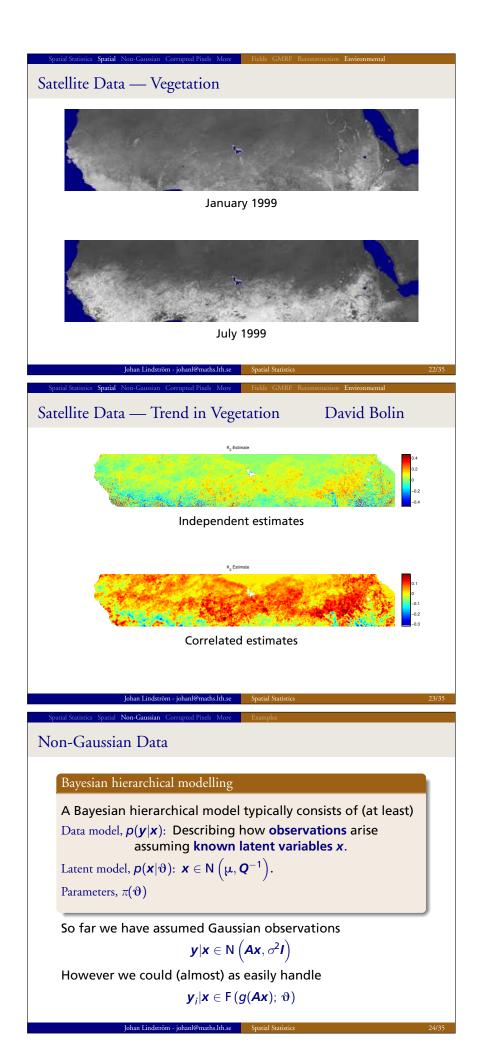
### Interpolation using a GMRF

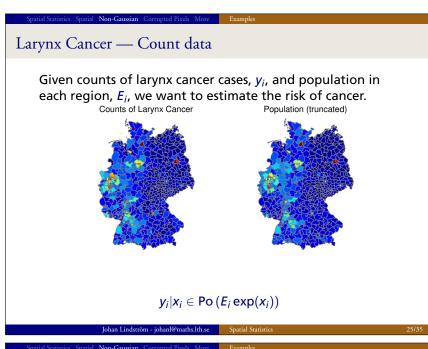
$$\mathsf{E}\left(\boldsymbol{x}|\boldsymbol{y}\right) = \mu + \frac{1}{\sigma^2} \boldsymbol{Q}_{\boldsymbol{x}|\boldsymbol{y}}^{-1} \boldsymbol{A}^{\top} \left(\boldsymbol{y} - \boldsymbol{A}\mu\right)$$

$$V(\boldsymbol{x}|\boldsymbol{y}) = \boldsymbol{Q}_{\boldsymbol{x}|\boldsymbol{y}}^{-1} = \left(\boldsymbol{Q} + \frac{1}{\sigma^2} \boldsymbol{A}^{\top} \boldsymbol{A}\right)^{-1}$$

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Insurance Claims — Count data

Oscar Tufvesson

Given the number of insurance claims,  $y_i$ , we want to estimate the risk of an accident.





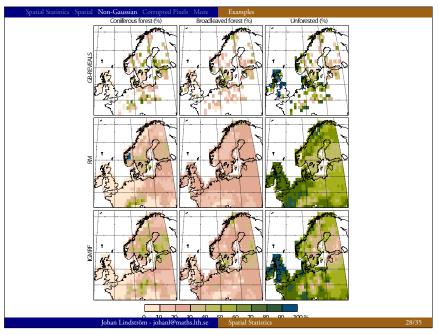
$$y_i | \eta_i \in Po\left(E_i \exp(\eta_i)\right)$$
  
 $oldsymbol{\eta} = oldsymbol{\mathcal{B}}eta + oldsymbol{x}$ 

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# Parana Rainfall — Positive data January Precip. January log(Precip.) $y_i|x_i \in \Gamma\left(b, \frac{e^{x_i}}{b}\right)$ Examples Examples



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# Image Reconstruction — Corrupted Pixels

- ▶ Typically we don't know which pixels are bad.
- ► A better model is then
  - Assume an underlying image, x.
  - ► Assume an indicator image for **bad** pixels, **z**.
  - Given the indicator we either observe the correct pixel value from x or noise.
- ► Use Bayes' formula to compute the distribution for the unknown image (and indicator) given observations and parameters.

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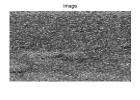
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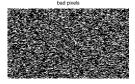
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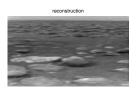
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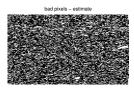
NDVI

# Image Reconstruction — Corrupted pixels









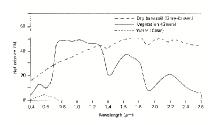
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ND

# Normalized difference vegetation index (NDVI)



$$\mathsf{NDVI} = \frac{R_{\scriptscriptstyle \mathrm{NIR}} - R_{\scriptscriptstyle \mathrm{RED}}}{R_{\scriptscriptstyle \mathrm{NIR}} + R_{\scriptscriptstyle \mathrm{RED}}}$$

- $R_{\rm RED}$  is the amount of reflected red light (0.58 0.68  $\mu m$ )
- ►  $R_{\rm NIR}$  is the amount of reflected near-infrared light (0.72 1.00  $\mu m$ )

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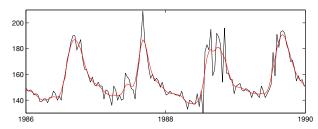
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### Smoothed version of the NDVI Data

Smooth the data to fill in missing values and remove noise due to cloud cover, etc.



Important ecological questions:

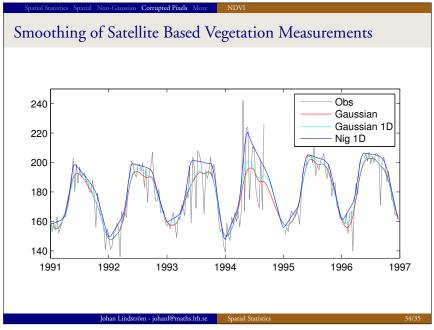
- ▶ Plant phenology (start and end of season)
- Plant productivity (integral)

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# Spatial Statistics Spatial Non-Gaussian Corrupted Pixels More NDVI Smoothing of Satellite Based Vegetation Measurements NDVI May 1993 Smooth 240 220 200 180 160 140 1991 1992 1993 1994 1995 1996



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### Learn more!

### What?

Spatial statistics with image analysis, FMSN20

### When?

HT2-2018, October-December

### Where?

Information and Matlab files will be available at www.maths.lth.se/matstat/kurser/fmsn20masm25/ (currently the 2017 webpage, updated soon)

# Who?

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