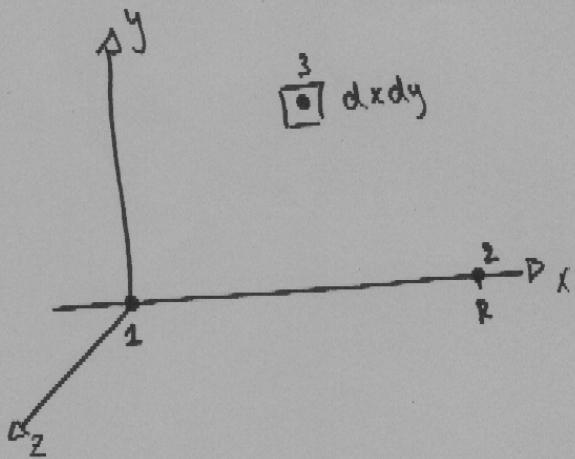


\* In order to solve the Kirkwood Equation numerically, it is convenient to use the symmetries in the system and change coordinates such that the 3-dimensional integral over  $\vec{r}_3$  becomes a 2-dimensional integral.

Due to the translational symmetry you can choose the position of the origin and due to the spherical rotational symmetry you can choose the orientation of the coordinate system.

$\Rightarrow$  Place the origin at  $\vec{r}_1$ , and align the x-axis along  $\vec{r}_1 - \vec{r}_2$ , such that

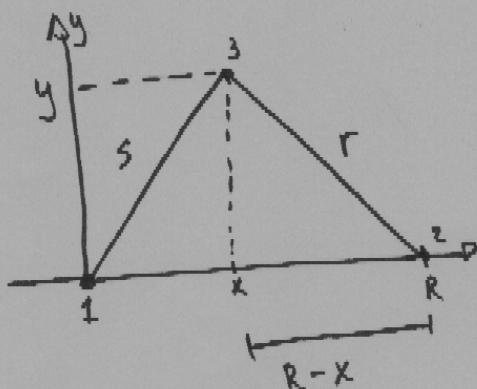
$$\vec{r}_1 = \begin{bmatrix} 0 \\ 0 \\ 0 \end{bmatrix} \quad \vec{r}_2 = \begin{bmatrix} 0 \\ 0 \\ R \end{bmatrix} \quad \text{where } R \in |\vec{r}_1 - \vec{r}_2|$$



\* The integral only depends on  $|\vec{r}_1 - \vec{r}_3|$  and  $|\vec{r}_2 - \vec{r}_3|$ , so it is invariant under rotation about the x-axis. Rotating about rotating the area  $dx dy$  about the x-axis creates a volume  $2\pi y dx dy$ , so the volume element  $d\vec{r}_3$  becomes  $2\pi y dx dy$ , and you only need to integrate over  $-\infty \leq x \leq \infty \quad 0 \leq y \leq \infty$ .

\* Make a coordinate transformation from  $x, y, z$  to

$$R = |\vec{r}_1 - \vec{r}_2| \quad \text{or} \quad r = |\vec{r}_2 - \vec{r}_3| \quad s = |\vec{r}_1 - \vec{r}_3|$$



$$\text{so } s = \sqrt{x^2 + y^2} \quad r = \sqrt{y^2 + (R-x)^2}$$

\* Next, calculate the Jacobian of the transformation

$$J = \det \begin{Bmatrix} \frac{\partial s}{\partial x} & \frac{\partial s}{\partial y} \\ \frac{\partial r}{\partial x} & \frac{\partial r}{\partial y} \end{Bmatrix} \quad \begin{aligned} \frac{\partial s}{\partial x} &= \frac{1}{2s} 2x = \frac{x}{s} & \frac{\partial s}{\partial y} &= \frac{y}{s} \\ \frac{\partial r}{\partial x} &= -\frac{2(R-x)}{2r} = -\frac{(R-x)}{r} & & \\ \frac{\partial r}{\partial y} &= \frac{y}{r} & & \end{aligned}$$

$$J = \frac{x}{s} \frac{y}{r} + \frac{(R-x)}{r} \frac{y}{s} = \frac{yR}{rs}$$

$$\text{so } ds dr = \frac{yR}{rs} dx dy$$

$$dx dy = \frac{rs}{yR} ds dr$$

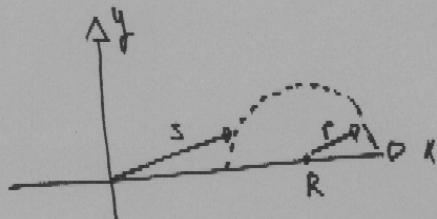
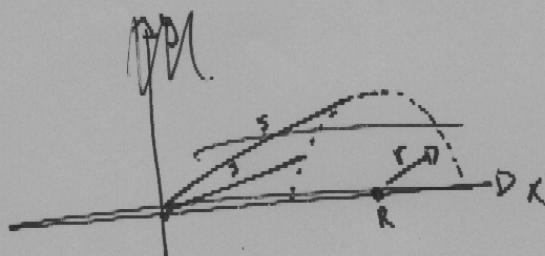
$$d\vec{r}_3 = 2\pi y dx dy = 2\pi \frac{rs}{R} ds dr$$

\* The domain of integration should cover the upper half of the  $x-y$  plane. There are more than one possibility for the limits of  $r$  and  $s$ .  
 One possibility is

$$0 \leq r \leq \infty$$

$$|R-r| \leq s \leq R+r$$

For a fixed  $r$ , the integration over  $s$  will be over a semicircle.



$r$  fixes the radius of the dashed semicircle. The integral over  $s$  for a fixed  $r$  or over the dashed semicircle.

Putting this all together gives

$$-kT \ln g(R; \xi) = \xi U(R) + \frac{2\pi P}{R} \int_0^\infty dr r [g(r)-1] \left\{ \int_0^\xi ds s u(s) g(s; \xi') \right\}_{|R-r|}$$

$$-kT \ln(g(R; \xi)) = \xi U(R) + \frac{2\pi p}{R} \int_0^\infty dr r [g(r)-1] \int_0^\xi ds' \int_{|r-R|}^{R+r} ds s U(s) g(s'; \xi')$$

$$R \ln(g(R; \xi) e^{\frac{\xi}{kT} U(R)}) = - \frac{2\pi p}{kT} \int_0^\infty dr r [g(r)-1] \int_0^\xi ds' \int_{|r-R|}^{R+r} ds s U(s) g(s'; \xi')$$

define  $Z(R; \xi) \equiv R \ln(g(R; \xi) e^{\frac{\xi}{kT} U(R)})$

from the  $ds$  integral of the RHS you get  $Z(0; \xi) = 0$

$$Z'(R; \xi) = - \frac{2\pi p}{kT} \int_0^\infty dr r [g(r)-1] \int_0^\xi ds' \left[ (R+r) U(R+r) g(R+r; \xi') \right. \\ \left. - |R-r| U(|R-r|) g(|R-r|; \xi') \partial_R |R-r| \right]$$

$$= - \frac{2\pi p}{kT} \int_0^\infty dr r [g(r)-1] \int_0^\xi ds' \left[ (R+r) U(R+r) g(R+r; \xi') \right. \\ \left. + (R-r) U(|R-r|) g(|R-r|; \xi') \right]$$

$$\Rightarrow Z(R; \xi) = \int_0^R dR' Z'(R'; \xi)$$

$$g(R; \xi) = \exp \left\{ \frac{1}{R} Z(R; \xi) - \frac{\xi}{kT} U(R) \right\}$$