GNG1106

Fundamentals of Engineering Computation

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University of Ottawa

Fall 2023 ~

Outline

1 Teaching Evaluation

2 Numerical Method II: Trapezoidal Rule For Integration

3 Numerical Method III: Euler's Method For Solving DE

Outline

1 Teaching Evaluation

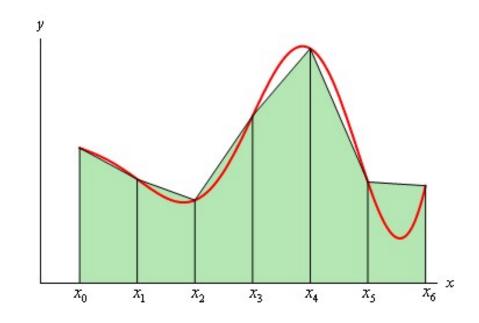
2 Numerical Method II: Trapezoidal Rule For Integration

3 Numerical Method III: Euler's Method For Solving DE

- Computing the integral of a function, say f(x) is often needed.
- For many functions f(x), there is no analytic solution.
- We now consider numerical computation of the integral

$$\int_{a}^{b} f(x)dx$$

• Idea: Cut the range [a, b] into n equal subintervals and approximate the integral as the sum of the areas of n trapezoids.



Highlight

The area of a trapezoid is equal to:

$$\frac{\text{Base1} + \text{Base2}}{2} \times \text{Height}$$

- The smaller are the subintervals, the more accurate is the approximation.
- Larger range [a, b] results in decreased approximation accuracy.

Algorithm

- Input: integration limits a and b, subinterval size Δ
- Output: integral value
- Let $n = \text{Round}(\frac{b-a}{\Delta})$. Cut [a, b] into n subintervals $(a_0, a_1), (a_1, a_2), \ldots, (a_{n-1}, a_n)$, where $a_i a_{i-1} = \Delta$. Then

Intergal
$$\approx \frac{f(a_0) + f(a_1)}{2} \Delta + \frac{f(a_1) + f(a_2)}{2} \Delta + \dots + \frac{f(a_{n-1}) + f(a_n)}{2} \Delta$$

$$= \Delta \cdot \left(\frac{f(a_0)}{2} + \sum_{i=1}^{n-1} f(a_i) + \frac{f(a_n)}{2} \right)$$

• Algorithm/Code: Do it yourself.

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Differential Equations (DE)

- Engineering applications often need to solve differential equations.
- We consider the first-order DE of the following form.

$$\frac{dy}{dt} = f(y,t) \tag{1}$$

$$y(t_0) = y_0 \tag{2}$$

$$y(t_0) = y_0 (2)$$

where

- y, or written as y(t), is an unknown function of time t.
- f is a given function of y and t.
- t_0 is a given initial time and y_0 is the given initial value of y at $t=t_0$.
- Solving this DE means finding function y(t) for $t > t_0$.

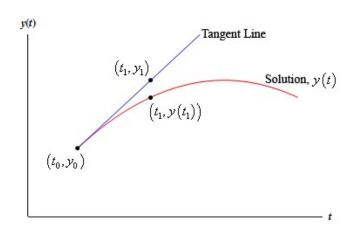
• For example, when an object falls with air friction/resistance, we have

$$\frac{dv}{dt} = \frac{mg - cv}{m}$$
$$v(t_0) = v_0$$

where m is object mass, g is gravitational acceleration, c is some constant, and v is the object falling speed.

- In some cases, analytic solutions are attainable using calculus.
- In other cases, analytic solutions are difficult to obtain or do not exist.
- We would like to solve such a DE numerically, i.e, instead of obtaining an analytic expression of the y(t), we aim at obtaining a numerical solution: for a list of t values with $t > t_0$, we find their corresponding y(t) values.

Euler's Method for Solving DE



$$\frac{dy}{dt} = \lim_{\Delta \to 0} \frac{y(t + \Delta) - y(t)}{\Delta}$$

$$\approx \frac{y(t + \Delta) - y(t)}{\Delta} \text{ (for small } \Delta)$$

$$f(y, t) \approx \frac{y(t + \Delta) - y(t)}{\Delta} \text{ (for small } \Delta)$$

Highlight

If we select a set of equally space time points $\{t_1, t_2, \ldots, t_n\}$ with $t_i - t_{i-1} = \Delta$, we have

$$y(t_i) \approx y(t_{i-1}) + f(y(t_{i-1}), t_{i-1})\Delta$$

The Algorithm

- Input: t_0, y_0, Δ, n
- Output: (t_1, t_2, \ldots, t_n) and (y_1, y_2, \ldots, y_n) $(y_i \text{ is the short-hand notation for } y(t_i)).$
- Algorithm (not the code):

```
t = t_0;

y = y_0;

for (i=0; i<n; i++)

{

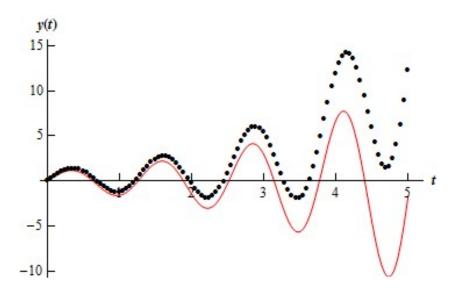
y = y + f(y, t)\Delta;

t = t + \Delta;

print (t, y, "\n");

}
```

$$\frac{dy}{dt} = y - \frac{1}{2}e^{t/2}\sin(5t) + 5e^{t/2}\cos(5t)$$



- The smaller the step size Δ , the more accurate is the approximation.
- As t increases, the approximation usually gets less accurate.

In-Class Exercise: