Project 2 in FYS4411: Computational Physics 2

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May 31, 2017

Abstract

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Contents

1	Intr	oduction	2		
2	The	Theory			
	2.1	Variational Monte Carlo	3		
	2.2	The Metropolis Algorithm	4		
		2.2.1 Uniform sampling			
		2.2.2 Importance sampling			
		2.2.3 Steepest descent			
	2.3	Electron wave function			
		2.3.1 Slater determinants			
		2.3.2 The Jastrow factor			
	2.4	Two electron case			
		2.4.1 Unperturbed expectation value			
		2.4.2 Local energy			
		2.4.3 Quantum force			
		2.4.4 α and β derivatives in the steepest descent algorithm .			
	2.5	N electron case			
		2.5.1 Local energy			
		2.5.2 Quantum force			
		2.5.3 α and β derivatives in the steepest descent algorithm.			
3	Imp	ementation	17		
	3.1	TESTS	17		
	3.2	Slater determinants	17		

4	Results	18
5	Discussion and conclusion	18
6	ppendix 1	
	6.1 Hermite polynomials	18
	6.2 Useful formulas	18

1 Introduction

The goal of this project is to study closed shell systems of electrons confined in a harmonic oscillator potential - a quantum dot. Within this scope we are investigating the ground state energies, exception values kinetic and potential energies, single particle densities and one-body densities. The system we are interested in is a two dimensional system of N electrons, and since we have closed shell systems we will look at N=2,6 and 12 electrons.

2 Theory

As tradition demands we begin by looking at the Hamiltonian of the system we are to solve,

$$\hat{H} = \sum_{i=1}^{N} \left(-\frac{1}{2} \nabla_i^2 + \frac{1}{2} \omega^2 r_i^2 \right) + \sum_{i < j}^{N} \frac{1}{r_{ij}}$$
 (1)

In order to accommodate a modern notational benefits and simplifications, we use natural units ($\hbar = c = e = m_e = 1$). We can also observe that N is the number of particles we are using, and the ω is the oscillator frequency for the harmonic oscillator part. The first part, we recognize as the unperturbed part of the Hamiltonian,

$$\hat{H}_0 = \sum_{i=1}^N \left(-\frac{1}{2} \nabla_i^2 + \frac{1}{2} \omega^2 r_i^2 \right),\tag{2}$$

and the last part is the perturbation to our system,

$$\hat{H}_1 = \sum_{i < j}^N \frac{1}{r_{ij}} \tag{3}$$

such that $\hat{H} = \hat{H}_0 + \hat{H}_1$. The distance between two electrons is defined as following,

$$r_{ij} = |\mathbf{r}_i - \mathbf{r}_j| = \sqrt{(x_i - x_j)^2 + (y_i - y_j)^2}$$
 (4)

2.1 Variational Monte Carlo

In order to go through the basic steps of variational Monte Carlo, we begin by deriving the basic formulas needed. Let ψ_T be our trial wave function, which we can expand into an energy basis which we assume is normalized,

$$\psi_T = \sum_{i=0} c_i |\psi_i\rangle$$

The expectation value of the energy given by the Hamiltonian \hat{H} (from now one omitting the hat) is,

$$\langle E \rangle = \langle \Psi_T | H | \psi_T \rangle$$

$$= \sum_{i=0}^{\infty} \sum_{j=0}^{\infty} c_i^* c_j \langle \psi_i | H | \psi_j \rangle$$

$$= \sum_{i=0}^{\infty} \sum_{j=0}^{\infty} c_i^* c_j E_j \langle \psi_i \psi_j \rangle$$

$$= \sum_{i=0}^{\infty} |c_i|^2 E_i$$

From the variational principle in quantum mechanics, we have that the energy of the ground state can be approximated by a trial wave function,

$$E_0 \ge E[\psi_T] = \frac{\langle \psi_T | H | \psi_T \rangle}{\langle \psi_T | \psi_T \rangle} \tag{5}$$

Writing this on integral form, we get

$$E[H] = \langle H \rangle = \frac{\int d\mathbf{r} \psi_T^*(\mathbf{r}) H(\mathbf{r}) \psi_T(\mathbf{r})}{\int d\mathbf{r} \psi_T^*(\mathbf{r}) \psi_T(\mathbf{r})}$$

$$= \frac{\sum_{ij} c_i^* c_j \int d\mathbf{r} \psi_i^*(\mathbf{r}) H(\mathbf{r}) \psi_j(\mathbf{r})}{\sum_{ij} c_i^* c_j \int d\mathbf{r} \psi_i^*(\mathbf{r}) \psi_j(\mathbf{r})}$$

$$= \frac{\sum_i |c_i|^2 E_i}{\sum_i |c_i|^2}$$

$$\geq E_0$$

In order to make any progress with variational Monte Carlo, we need to get ourselves a wave function. This wave function will have to take N particles $\mathbf{r} = (\mathbf{r}_1, \mathbf{r}_2, ..., \mathbf{r}_N)$, and have can be varied by a set of variational parameters $\alpha = {\alpha_i}$. The integral we wish to solve is now,

$$E[H] = \frac{\int d\mathbf{r} \psi_T^*(\mathbf{r}, \alpha) H(\mathbf{r}) \psi_T(\mathbf{r}, \alpha)}{\int d\mathbf{r} \psi_T^*(\mathbf{r}, \alpha) \psi_T(\mathbf{r}, \alpha)}$$
(6)

As with Monte Carlo methods, we need to define ourselves a probability distribution function(PDF) which will be based off our trial wave function $\psi_T(\mathbf{r})$,

$$P(\mathbf{r}) = \frac{|\psi_T(\mathbf{r})|^2}{\int |\psi_T(\mathbf{r})|^2 d\mathbf{r}}$$
 (7)

We can now define E_L as our local energy,

$$E_L = \frac{1}{\psi_T} H \psi_T \tag{8}$$

We can use this and the PDF, we can rewrite our E[H] which now also depends on α ,

$$E[H(\alpha)] = \int P(\mathbf{r}) E_L(\mathbf{r}) d\mathbf{r} \approx \frac{1}{N_{MC}} \sum_{i=1}^{N_{MC}} P(\mathbf{r}_i, \alpha) E_L(\mathbf{r}_i, \alpha)$$

where N_{MC} is the number of Monte Carlo cycles which we run for. We now have an estimate for the ground state energy from the average of our Monte Carlo samples. The algorithm for the VMC(Variational Monte Carlo) method will be based of the Metropolis algorithm CITE METROPOLIS!!, such that a quick review of this is in order.

2.2 The Metropolis Algorithm

The Metropolis algorithm SITER is a method for obtaining a sequence of samples when direct sampling is problematic. Given a PDF $P_i^{(n)}$, where n is the time step, and i is the current state, the algorithm is such that a transitioning probability to a new state j is given by $T_{i\to j}$. Then, the probability of accepting this state is given by $A_{i\to j}$. If rejected, no move is performed and we remain at state i. We will require that A and T is such

that $P_i^{(n\to\infty)}\to p_i$. This gives rise to the relation

$$\begin{split} P_{i}^{(n)} &= \sum_{j} \left[P_{j}^{(n-1)} T_{j \to i} A_{j \to i} + P_{i}^{(n-1)} T_{i \to j} (1 - A_{i \to j}) \right] \\ &= \sum_{j} \left[P_{j}^{(n-1)} T_{j \to i} A_{j \to i} - P_{i}^{(n-1)} T_{i \to j} A_{i \to j} + P_{i}^{(n-1)} T_{i \to j}) \right] \\ &= \sum_{j} \left[P_{j}^{(n-1)} T_{j \to i} A_{j \to i} - P_{i}^{(n-1)} T_{i \to j} A_{i \to j} \right] + P_{i}^{(n-1)} \sum_{j} T_{i \to j}) \\ &= \sum_{j} \left[P_{j}^{(n-1)} T_{j \to i} A_{j \to i} - P_{i}^{(n-1)} T_{i \to j} A_{i \to j} \right] + P_{i}^{(n-1)} \cdot 1 \end{split}$$

Since we require our system to make some transition into a new state j, $\sum_{j} T_{i \to j} = 1$. Now using the requirement $P_i^{(n \to \infty)} \to p_i$, we get

$$p_{i} = \sum_{j} [p_{j}T_{j\rightarrow i}A_{j\rightarrow i} - p_{i}T_{i\rightarrow j}A_{i\rightarrow j}] + p_{i}$$

$$\downarrow$$

$$0 = \sum_{j} [p_{j}T_{j\rightarrow i}A_{j\rightarrow i} - p_{i}T_{i\rightarrow j}A_{i\rightarrow j}]$$

The requirement applies for the system as a whole when time tends to infinity. This is a rather weak requirement, so we seek to enforce a similar condition to each state. Namely,

$$p_i T_{i \to i} A_{i \to i} = p_i T_{i \to j} A_{i \to j}$$

which we observe implies that transitioning and accepting a state $i \to j$ is equal to that of $j \to i$. By rearranging we get,

$$\frac{A_{j\to i}}{A_{i\to j}} = \frac{p_j T_{j\to i}}{p_i T_{i\to j}} \tag{9}$$

This is a much stronger requirement, and is called *detailed balance*. The Metropolis algorithm is now all about maximizing A, such that

$$A_{i \to j} = \min\left(1, \frac{p_i T_{i \to j}}{p_j T_{j \to i}}\right)$$

The ratio which we uses to accept or reject a step is defined as

$$q(\mathbf{r}^{\text{new}}, \mathbf{r}^{\text{old}}) = \frac{p_i T_{i \to j}}{p_j T_{j \to i}}$$
(10)

Let us then recapitulate what the Metropolis algorithm becomes as applied to our VMC problem.

Algorithm 1 Metropolis algorithm for Variational Monte Carlo.

- 1: Set up initial conditions for our system.
- 2: Sample a new state j with transition probability $T_{i\to j}$ by the desired sampling method.
- 3: Accept new state j with acceptance probability $A_{i \to j}$.
- 4: If rejected, we go back to state i.
- 5: Sample the system.
- 6: Repeat step 2-5 N_{MC} times.

2.2.1 Uniform sampling

For sampling a new state(by which we mean a new position), we use

$$\mathbf{r}^{\text{new}} = \mathbf{r}^{\text{old}} + \xi \Delta \mathbf{r}^{\text{old}} \tag{11}$$

where ξ is a random number chosen from a uniform distribution, and $\Delta \mathbf{r}$ is the step length which we update for. For this kind of uniform sampling, we seek to tune $\Delta \mathbf{r}$ to such that we accept roughly 50% all suggested moves.

The acceptance ratio (10) as applied to our VMC calculation, is (dubbing it as R instead of $q(\mathbf{r}^{\text{new}}, \mathbf{r}^{\text{old}})$),

$$R \equiv \frac{\psi_T^{\text{new}}}{\psi_T^{\text{old}}} \tag{12}$$

2.2.2 Importance sampling

If we wish to improve our VMC Metropolis algorithm, we can use something called importance sampling. It is based upon the Fokker-Planck equation, which characterizes a move through coordinate space. For one dimension, it reads

$$\frac{\partial P}{\partial t} = D \frac{\partial}{\partial x} \left(\frac{\partial}{\partial x} - F \right) P(x, t) \tag{13}$$

The D is a referred to as the diffusion constant since it is derived from the Fokker-Planck equation, and the F is called the drift term. A new position in coordinate space is chosen from solving the Langevin equation,

$$\frac{\partial x(t)}{\partial t} = DF(x(t)) + \eta \tag{14}$$

 η is a random variable. This is solved using Euler's method. For our problem, the solution and sampling of a new position is defined by

$$\mathbf{r}^{\text{new}} = \mathbf{r}^{\text{old}} + D\mathbf{F}(\mathbf{r}^{\text{old}})\Delta t + \xi \sqrt{\Delta t}$$
(15)

Where D = 0.5 in atomic units, and **F** is the quantum force as defined by,

$$\mathbf{F} = \frac{2}{\psi_T} \nabla \psi_T \tag{16}$$

We typically choose $\Delta t \in [0.001, 0.01]$. The random variable ξ is chosen from a Gaussian distribution around 0 with a standard deviation of 1. Further, the acceptance ratio $q(\mathbf{r}^{\text{new}}, \mathbf{r}^{\text{old}})$ now becomes,

$$q(\mathbf{r}^{\text{new}}, \mathbf{r}^{\text{old}}) = \frac{G(\mathbf{r}^{\text{old}}, \mathbf{r}^{\text{new}}) |\psi_T(\mathbf{r}^{\text{new}})|^2}{G(\mathbf{r}^{\text{new}}, \mathbf{r}^{\text{old}}) |\psi_T(\mathbf{r}^{\text{old}})|^2}$$
(17)

where the $G(\mathbf{r}^{\text{new}}, \mathbf{r}^{\text{old}})$ is a Green's function which is a solution to the Fokker-Planck equation. This is a transition probability, and is defined by

$$G(\mathbf{r}^{\text{new}}, \mathbf{r}^{\text{old}}) = \frac{1}{(4\pi D\Delta t)^{3N/2}} \exp\left(-(\mathbf{r}^{\text{new}} - \mathbf{r}^{\text{old}} - D\delta t \mathbf{F}(\mathbf{r}^{\text{old}}))^2 / 4D\Delta t\right)$$
(18)

A more detailed derivation of importance sampling can be found in SITER MORTEN KOMPENDIUM HER. Is we can already observe, it is possible to simplify the Greens ratio (17) due to the exponential. We will begin by using that the squaring of the exponential implies a dot product for vectors. For simplicity, we dub $\mathbf{x} = \mathbf{r}^{\text{old}}$ and $\mathbf{y} = \mathbf{r}^{\text{new}}$. Through a lot of patience, we

get REMOVE THIS???

$$\frac{G(\mathbf{x}, \mathbf{y})}{G(\mathbf{y}, \mathbf{x})} = \exp \left[\frac{1}{4D\Delta t} \left\{ -\left(x_x^2 - x_x y_x - x_x D\Delta t F_x(\mathbf{y}) + y_x^2 - y_x x_x + y_x D\Delta t F_x(\mathbf{y}) - x_x D\Delta t F_x(\mathbf{y}) + y_x D\Delta t F_x(\mathbf{y}) + D^2 \Delta t^2 F_x^2(\mathbf{y}) \right) + \left(y_x^2 - y_x x_x - y_x D\Delta t F_x(\mathbf{x}) + x_x^2 - x_x y_x + x_x D\Delta t F_x(\mathbf{x}) + D^2 \Delta t^2 F_x^2(\mathbf{x}) - y_x D\Delta t F_x(\mathbf{x}) + x_x D\Delta t F_x(\mathbf{x}) \right) - \left(x_y^2 - x_y y_y - x_y D\Delta t F_y(\mathbf{y}) - x_y y_y + y_y^2 + y_y D\Delta t F_y(\mathbf{y}) - x_y D\Delta t F_y(\mathbf{y}) + y_y D\Delta t F_y(\mathbf{y}) + D^2 \Delta t^2 F_y^2(\mathbf{y}) \right) + \left(y_y^2 - y_y x_y - y_y D\Delta t F_y(\mathbf{x}) + x_y^2 - x_y y_y + x_y D\Delta t F_y(\mathbf{x}) + D^2 \Delta t^2 F_y^2(\mathbf{x}) - y_y D\Delta t F_y(\mathbf{x}) + x_x D\Delta t F_x(\mathbf{x}) + D^2 \Delta t^2 F_x^2(\mathbf{x}) - y_x D\Delta t F_x(\mathbf{x}) + x_x D\Delta t F_x(\mathbf{x}) + x_x D\Delta t F_x(\mathbf{y}) - y_x D\Delta t F_x(\mathbf{y}) + x_x D\Delta t F_x(\mathbf{y}) - y_x D\Delta t F_x(\mathbf{y}) - D^2 \Delta t^2 F_x^2(\mathbf{y}) - y_y D\Delta t F_y(\mathbf{y}) + x_y D\Delta t F_y(\mathbf{y}) - y_y D\Delta t F_y(\mathbf{y}) + x_y D\Delta t F_y(\mathbf{y}) - y_y D\Delta t F_y(\mathbf{y}) - D^2 \Delta t^2 F_y^2(\mathbf{y}) \right] \\ = \exp \left[\frac{1}{4D\Delta t} \left\{ -y_x D\Delta t \left[F_x(\mathbf{x}) + F_x(\mathbf{x}) + F_x(\mathbf{y}) + F_x(\mathbf{y}) \right] - D^2 \Delta t^2 \left[F_x^2(\mathbf{y}) - F_x^2(\mathbf{x}) \right] + 2x_x D\Delta t \left[F_y(\mathbf{x}) + F_y(\mathbf{y}) \right] - D^2 \Delta t^2 \left[F_y^2(\mathbf{y}) - F_y^2(\mathbf{x}) \right] \right\} \\ = \exp \left[-\frac{y_x}{2} \left(F_x(\mathbf{x}) + F_x(\mathbf{y}) \right) - \frac{D\Delta t}{4} \left(F_x^2(\mathbf{y}) - F_x^2(\mathbf{x}) \right) + \frac{x_x}{2} \left(F_x(\mathbf{x}) + F_x(\mathbf{y}) \right) - \frac{D\Delta t}{4} \left(F_y^2(\mathbf{y}) - F_y^2(\mathbf{x}) \right) \right] \\ - \frac{D\Delta t}{4} \left(F_y^2(\mathbf{y}) - F_y^2(\mathbf{x}) \right) + \frac{x_y}{4} \left(F_y(\mathbf{x}) + F_y(\mathbf{y}) \right) \right]$$

Through the magic of simplification (but mostly patience), we arrive at

$$\frac{G(\mathbf{x}, \mathbf{y})}{G(\mathbf{y}, \mathbf{x})} = \exp\left[\frac{1}{2}(x_x - y_x) \left(F_x(\mathbf{x}) + F_x(\mathbf{y})\right) + \frac{1}{2}(x_y - y_y) \left(F_y(\mathbf{x}) + F_y(\mathbf{y})\right) - \frac{D\Delta t}{4} \left(F_x^2(\mathbf{y}) - F_x^2(\mathbf{x}) + F_y^2(\mathbf{y}) - F_y^2(\mathbf{x})\right)\right]$$

Summing over dimensions, and we get

$$\frac{G(\mathbf{x}, \mathbf{y})}{G(\mathbf{y}, \mathbf{x})} = \exp \left[\sum_{i=1}^{N_{\text{dim}}} \frac{1}{2} \left(F_i(\mathbf{x}) + F_i(\mathbf{y}) \right) \left(\frac{D\Delta t}{2} \left(F_i(\mathbf{x}) - F_i(\mathbf{y}) \right) - y_i + x_i \right) \right]$$
(19)

2.2.3 Steepest descent

The steepest descent method is a simple iterative method for finding a local minimum in parameters. It is used when we are tasked to solve systems of linear equations on the form of

$$A\mathbf{x} = \mathbf{b}$$

where we through an iterative process seeks to minimize the following,

$$r = b - Ax$$

where the exact solution implies that $\mathbf{r} = 0$. The process for solving this through steepest descent can be set up as,

$$\mathbf{x}_{i+1} = \mathbf{x}_i - \nabla f(\mathbf{x}_i)\delta t \tag{20}$$

Applied to our VMC problem, where we will have one or two variational parameters, α and β , the equations will take on the form of (20),

$$\begin{pmatrix} \alpha_{i+1} \\ \beta_{i+1} \end{pmatrix} = \begin{pmatrix} \alpha_i \\ \beta_i \end{pmatrix} - \delta t \begin{pmatrix} \frac{d\langle E_L[\alpha] \rangle}{d\alpha} \\ \frac{d\langle E_L[\beta] \rangle}{d\beta} \end{pmatrix}$$
(21)

Where the $\langle E_L \rangle$ is the expectation value of the local energy and δt is the step size of the method. The expectation value of the local energy is found by

$$\frac{d \langle E_L[\alpha_i] \rangle}{d\alpha_i} = 2 \left(\left\langle \frac{\bar{\psi}_{\alpha_i}}{\psi_T[\alpha_i]} E_L[\alpha_i] \right\rangle - \left\langle \frac{\bar{\psi}_{\alpha_i}}{\psi[\alpha_i]} \right\rangle \langle E_L[\alpha_i] \rangle \right) \tag{22}$$

with the derivative of ψ with respect to the variational parameter is defined as

$$\bar{\psi_{\alpha_i}} = \frac{\psi[\alpha_i]}{d\alpha_i}$$

As is evident, this method may oscillate around the solution, since the step size is fixed. Other more optimal methods such as the conjugate gradient method existreference?, but are not implemented in the project.

2.3 Electron wave function

Our wave function will consist of two parts: one that comes from the harmonic oscillator potential and is built up based on the fermionic nature of the system, and one that proved a correlation between two particles - the so-called Jastrow factor. The wave function we construct, will be called our trial wave function.

$$\psi_T(\mathbf{r}) = \psi_J(\mathbf{r})\psi_{SD}(\mathbf{r}) \tag{23}$$

2.3.1 Slater determinants

The wave function for an electron in a two dimensional harmonic oscillator potential can be written as a Hermite polynomial,

$$\phi_{n_x,n_y}(x,y) = AH_{n_x}(\sqrt{\omega\alpha}x)H_{n_y}(\sqrt{\omega\alpha}y)\exp\left(-\frac{\omega\alpha}{2}\left(x^2 + y^2\right)\right)$$
 (24)

The details surrounding the mysterious variational parameter α will be explained further on. For now, we will place these ϕ 's into a Slater determinant. The Slater determinant is a creature that describes the wave function of a fermionic system, while also enforcing anti-symmetrization and thus the Pauli principle. Our Slater determinant will take the following form, when we describe the specific state of a system by n_x, n_y and a specific particle $\mathbf{r}_i = x_i \mathbf{i} + y_i \mathbf{j}$,

$$Det(\Phi(\mathbf{r})) \equiv \frac{1}{\sqrt{N!}} \begin{vmatrix} \phi_1(\mathbf{r}_1) & \phi_2(\mathbf{r}_1) & \dots & \phi_N(\mathbf{r}_1) \\ \phi_1(\mathbf{r}_2) & \phi_2(\mathbf{r}_2) & \dots & \phi_N(\mathbf{r}_2) \\ \vdots & \vdots & \ddots & \vdots \\ \phi_1(\mathbf{r}_N) & \phi_2(\mathbf{r}_N) & \dots & \phi_N(\mathbf{r}_N) \end{vmatrix}$$
(25)

From how we have described the position, we introduce following, useful notation,

$$r_i = \sqrt{x_i^2 + y_i^2} \tag{26}$$

Note that we have defined $\mathbf{r} \equiv (\mathbf{r_1}, \mathbf{r_2}, \dots, \mathbf{r_N})$ in the Slater determinant (25). To pull this definition back to our trial wave function (23), we get

$$\psi_{SD}(\mathbf{r}) = Det(\Phi(\mathbf{r})) \tag{27}$$

Where the OB stands for one body, as in one body wave function. We now need to look into the part that accounts for many body effects, the Jastrow factor.

2.3.2 The Jastrow factor

The correlation term is called a *Jastrow factor* is as mentioned here to take into account many-body effects of our system. The general shape of it is

$$\psi_J(\mathbf{r}) = \prod_{i < j}^N \exp\left(\frac{ar_{ij}}{1 + \beta r_{ij}}\right) = \prod_{i=1}^N \prod_{j=i+1}^N \exp\left(\frac{ar_{ij}}{1 + \beta r_{ij}}\right)$$
(28)

where the C stands for correlation. The a is a parameter that is 1 for parallel spin, and 1/3 for anti-parallel spin. The β is another variational parameter and the r_{ij} is defined by the equation (4) as the distance between two electrons. For now, we shall begin by looking closer at the two-electron case.

2.4 Two electron case

For the two electron case the Hamiltonian takes on familiar form,

$$H = -\frac{1}{2} \left(\frac{\partial^2}{\partial x_1^2} + \frac{\partial^2}{\partial y_1^2} + \frac{\partial^2}{\partial x_2^2} + \frac{\partial^2}{\partial y_2^2} \right) + \frac{1}{2} \omega^2 (x_1^2 + y_1^2 + x_2^2 + y_2^2) + \frac{1}{r_{12}}$$
(29)

We will assume the spin is anti-parallel and the total spin is 0 in the ground state for the 2 electron case due to Pauli's exclusion principle.

2.4.1 Unperturbed expectation value

We begin by looking at the unperturbed expectation value, in which we will try to convince ourselves that the energy comes out to be 2ω in atomic units. Given a wave function,

$$\Phi_0(\mathbf{r}_1, \mathbf{r}_2) = C \exp\left(-\frac{\omega}{2}(r_1^2 + r_2^2)\right)$$
(30)

with r_1 and r_2 being defined by the notation introduced in the Slater determinant (26).

$$\langle E_0 \rangle = \langle \Psi_0 | H_0 | \Psi_0 \rangle$$

$$= \int C \exp\left(-\frac{\omega}{2}(r_1^2 + r_2^2)\right) \left(-\frac{1}{2}\left(\nabla_1^2 + \nabla_2^2\right) + \frac{1}{2}\omega^2\left(r_1^2 + r_2^2\right)\right)$$

$$\times C \exp\left(-\frac{\omega}{2}(r_1^2 + r_2^2)\right) dx_1 dy_1 dx_2 dy_2$$

Since the derivatives are similar, we can look a single derivative, and then solve for that one.

$$\frac{\partial^2}{\partial x^2} \exp\left(-\frac{\omega}{2}x^2\right) = \frac{\partial}{\partial x} \left(-\omega x\right) \exp\left(-\frac{\omega}{2}x^2\right)$$
$$= \left(\omega^2 x^2 - \omega\right) \exp\left(-\frac{\omega}{2}x^2\right)$$

Inserting this, into our integral, we get,

$$I_{x_1} = -\frac{1}{2} \int \exp\left(-\frac{\omega}{2} \left(x_1^2 + y_1^2 + x_2^2 + y_2^2\right)\right) \frac{\partial^2}{\partial x_1^2}$$

$$\times \exp\left(-\frac{\omega}{2} \left(x_1^2 + y_1^2 + x_2^2 + y_2^2\right)\right) dx_1 dy_1 dx_2 dy_2$$

$$= -\frac{1}{2} \int \left(\omega^2 x_1^2 - \omega\right) \exp\left(-\omega \left(x_1^2 + y_1^2 + x_2^2 + y_2^2\right)\right) dx_1 dy_1 dx_2 dy_2$$

$$= -\frac{\omega}{2} \left(\omega \int x^2 \exp\left(-\omega \left(x_1^2 + y_1^2 + x_2^2 + y_2^2\right)\right) dx_1 dy_1 dx_2 dy_2$$

$$- \int \exp\left(-\omega \left(x_1^2 + y_1^2 + x_2^2 + y_2^2\right)\right) dx_1 dy_1 dx_2 dy_2$$

We can now use the integral formulas seen in the appendix, integral (44) and (45). Noting that there are actually four integrations happening, we get

$$-\frac{\omega}{2} \left(\omega \int x^2 \exp\left(-\omega \left(x_1^2 + y_1^2 + x_2^2 + y_2^2\right)\right) dx_1 dy_1 dx_2 dy_2 \right)$$

$$-\int \exp\left(-\omega \left(x_1^2 + y_1^2 + x_2^2 + y_2^2\right)\right) dx_1 dy_1 dx_2 dy_2$$

$$= -\frac{\omega}{2} \left(\left(\frac{\pi^{\frac{1}{2}}}{2}\omega^{-\frac{3}{2}}\right) \left(\pi^{\frac{1}{2}}\omega^{-\frac{1}{2}}\right)^3 - \left(\pi^{\frac{1}{2}}\omega^{-\frac{1}{2}}\right)^4 \right)$$

$$= -\frac{\omega}{2} \left(\frac{\pi^2}{2}\omega^{-3} + \pi^2\omega^{-2}\right)$$

$$= -\frac{\pi^2}{4\omega} + \frac{\pi^2}{2\omega}$$

$$= \frac{\pi^2}{4\omega}$$

Putting this together with the other Gaussian symmetric integral coming from the harmonic oscillator, we get

$$\langle E_0 \rangle = C^2 \left(4 \frac{\pi^2}{4\omega} + 4 \frac{\pi^2}{4\omega} \right)$$
$$= C^2 \frac{2\pi^2}{\omega}$$

We now need to find the constant C.

$$\langle \Phi_0 || \Phi_0 \rangle = C^2 \int \exp\left(-\omega \left(x_1^2 + y_1^2 + x_2^2 + y_2^2\right)\right) dx_1 dy_1 dx_2 dy_2$$

= $C^2 \left(\sqrt{\frac{\pi}{\omega}}\right)^4$

Since our wave function is required to be normalized C becomes,

$$C = \frac{\omega}{\pi}$$

Inserting this, and we get our expectation value

$$\langle E_0 \rangle = \frac{\omega^2}{\pi^2} \frac{2\pi^2}{\omega} = 2\omega$$

Which, is the answer we expected. This will provide a useful test for us later on.

2.4.2 Local energy

We now seek to find an expression for the local energy as given by equation (8), for the two electron case.

$$E_L = \frac{1}{\psi_T} H \psi_T \tag{31}$$

The trial wave function is given as

$$\psi_T = \psi_{SD}\psi_J = C \exp\left(-\frac{\omega\alpha}{2}\left(r_1^2 + r_2^2\right)\right) \exp\left(\frac{ar_{12}}{1 + \beta r_{12}}\right)$$
(32)

When the Hamiltonian is given as in equation (29), the main problem is to find the laplacian,

$$\nabla^2 = \nabla_1^2 + \nabla_2^2 = \frac{\partial^2}{\partial x_1^2} + \frac{\partial^2}{\partial y_1^2} + \frac{\partial^2}{\partial y_2^2} + \frac{\partial^2}{\partial y_2^2}$$

and it is thus where we shall begin. For simplicity, we write

$$\frac{\nabla^2 \psi_T}{\psi_T} = \frac{\nabla^2 \left(C \exp\left(-\frac{\omega \alpha}{2} \left(r_1^2 + r_2^2\right) + \frac{a r_{12}}{1 + \beta r_{12}}\right) \right)}{\psi_T}$$

We start by doing a single derivative.

$$\frac{\partial}{\partial x_1} \psi_T = \psi_T \left(-\omega \alpha x_1 + \frac{a(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2} \right)$$
 (33)

For the derivative with respect to x_2 we get an additional minus sign in the last term. For the second derivative, we get

$$\begin{split} \frac{\partial^2}{\partial x_1^2} \psi_T &= \frac{\partial}{\partial x_1} \left(\psi_T \left(-\omega \alpha x_1 + \frac{a(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2} \right) \right) \\ &= \psi_T \left(\frac{\partial}{\partial x_1} \left(-\omega \alpha x_1 \right) + \frac{\partial}{\partial x_1} \left(\frac{a(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2} \right) \right) \\ &+ \left(-\omega \alpha x_1 + \frac{a(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2} \right) \frac{\partial}{\partial x_1} \left(\psi_T \right) \end{split}$$

We now do the second term by itself,

$$\frac{\partial}{\partial x_1} \left(\frac{a(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2} \right) = \frac{a}{r_{12}(1 + \beta r_{12})^2} - \frac{a(x_1 - x_2)^2}{r_{12}^3(1 + \beta r_{12}^2)} - \frac{2a\beta(x_1 - x_2)^2}{r_{12}^2(1 + \beta r_{12})^3}$$

An overall minus sign appears when taking the derivative with respect to x_2 . We now look at the last term in our second derivative.

$$\left(-\omega\alpha x_1 + \frac{a(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2}\right) \frac{\partial}{\partial x_1} (\psi_T) = \left(\omega^2 \alpha^2 x_1^2 - \frac{\omega\alpha a x_1(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2} - \frac{\omega\alpha a x_1(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2} + \frac{a^2(x_1 - x_2)^2}{r_{12}^2(1 + \beta r_{12})^4}\right) \psi_T$$

$$= \left(\omega^2 \alpha^2 x_1^2 - \frac{2\omega\alpha a x_1(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2} + \frac{a^2(x_1 - x_2)^2}{r_{12}^2(1 + \beta r_{12})^4}\right) \psi_T$$

Putting all this together, the second derivative becomes

$$\frac{\partial^{2}}{\partial x_{1}^{2}}\psi_{T} = \left(-\omega\alpha + \omega^{2}\alpha^{2}x_{1}^{2} - \frac{2\omega\alpha ax_{1}(x_{1} - x_{2})}{r_{12}(1 + \beta r_{12})^{2}} + \frac{a^{2}(x_{1} - x_{2})^{2}}{r_{12}^{2}(1 + \beta r_{12})^{4}}\right)$$

$$+ \frac{a}{r_{12}(1 + \beta r_{12})^{2}} - \frac{a(x_{1} - x_{2})^{2}}{r_{12}^{3}(1 + \beta r_{12}^{2})} - \frac{2a\beta(x_{1} - x_{2})^{2}}{r_{12}^{2}(1 + \beta r_{12})^{3}}\psi_{T}$$

$$= \left(-\omega\alpha + \omega^{2}\alpha^{2}x_{1}^{2} - \frac{2\omega\alpha ax_{1}(x_{1} - x_{2})}{r_{12}(1 + \beta r_{12})^{2}} + \frac{2a}{(1 + \beta r_{12})^{2}}\right)$$

$$\times \left(\frac{1}{2r_{12}} - \frac{(x_{1} - x_{2})^{2}}{2r_{12}^{3}} - \frac{\beta(x_{1}x_{2})^{2}}{r_{12}^{2}(1 + \beta r_{12})} + \frac{a(x_{1} - x_{2})^{2}}{2r_{12}^{2}(1 + \beta r_{12})^{2}}\right)\psi_{T}$$

We can now pull the derivative with respect to x_1 and x_2 together.

$$\left(\frac{\partial^2}{\partial x_1^2} + \frac{\partial^2}{\partial x_1^2}\right)\psi_T = \left(-2\omega\alpha + \omega^2\alpha^2\left(x_1^2 + x_2^2\right) - \frac{2\omega\alpha a(x_1 - x_2)^2}{r_{12}(1 + \beta r_{12})^2} + \frac{2a}{(1 + \beta r_{12})^2}\left(\frac{1}{r_{12}} - \frac{(x_1 - x_2)^2}{r_{12}^3} - \frac{2\beta(x_1 - x_2)^2}{r_{12}^2(1 + \beta r_{12})} + \frac{a(x_1 - x_2)^2}{r_{12}^2(1 + \beta r_{12})^2}\right)\right)\psi_T$$

Putting this together with the derivatives with respect to y_1 and y_2 and using the relation given in the beginning (4), we get

$$\frac{\nabla^2 \psi_T}{\psi_T} = -4\omega\alpha + \omega^2 \alpha^2 (r_1^2 + r_2^2) - \frac{2\omega\alpha a r_{12}}{(1 + \beta r_{12})^2} + \frac{2a}{(1 + \beta r_{12})^2} \left(\frac{2}{r_{12}} - \frac{r_{12}^2}{r_{12}^3} - \frac{2\beta r_{12}^2}{r_{12}^2 (1 + \beta r_{12})} + \frac{a r_{12}^2}{r_{12}^2 (1 + \beta r_{12})^2} \right)$$

which becomes our final expression for the Laplacian,

$$\frac{\nabla^2 \psi_T}{\psi_T} = \omega^2 \alpha^2 (r_1^2 + r_2^2) - 4\omega\alpha - \frac{2\omega\alpha a r_{12}}{(1 + \beta r_{12})^2}$$
 (34)

$$+\frac{2a}{(1+\beta r_{12})^2} \left(\frac{a}{(1+\beta r_{12})^2} + \frac{1}{r_{12}} - \frac{2\beta}{(1+\beta r_{12})} \right)$$
(35)

Which gives the total local energy by adding the harmonic oscillator term and coulomb interaction.

$$E_L = -\frac{1}{2} \left((\alpha^2 - 1)\omega^2 (r_1^2 + r_2^2) - 4\omega\alpha - \frac{2\omega\alpha a r_{12}}{(1 + \beta r_{12})^2} \right)$$
 (36)

$$+\frac{2a}{(1+\beta r_{12})^2} \left(\frac{a}{(1+\beta r_{12})^2} + \frac{1}{r_{12}} - \frac{2\beta}{(1+\beta r_{12})} \right) + \frac{1}{r_{12}}$$
(37)

2.4.3 Quantum force

We also have to find the quantum force given by equation (16), which involves taking the gradient of the wave function. As we already know the first derivatives from taking the gradient (33), the process is trivial.

$$\mathbf{F} = \left(-\omega \alpha x_1 + \frac{a(x_1 - x_2)}{r_{12}(1 + \beta r_{12})^2}\right)\mathbf{i} + \left(-\omega \alpha y_1 + \frac{a(y_1 - y_2)}{r_{12}(1 + \beta r_{12})^2}\right)\mathbf{j}$$

For r_2 do we get a minus in front of the last term.

2.4.4 α and β derivatives in the steepest descent algorithm

For the steepest descent method (20), we have to find derivatives of the trial wave function (32) with respect to α and β .

From the equation for the energy expectation value (22), we see that we have to find the derivative of the wave function (30) with respect to the variational parameters α and β .

$$\frac{d\psi_T[\alpha]}{d\alpha} = -\frac{\omega}{2}(r_1^2 + r_2^2)\psi_T \tag{38}$$

For β we get,

$$\frac{d\psi_T[\beta]}{d\beta} = -\frac{ar_{12}^2}{(1+\beta r_{12})^2}\psi_T \tag{39}$$

Which is what is need in order to apply steepest descent to the two-body Jastrow factor case.

2.5 N electron case

For the N-electron case, we apply the Hamiltonian as states in equation (1) at the beginning of the theory section on the trial wave function as stated in (23). The N electron energy is once more given as the local energy (8).

From looking at the general Hamiltonian (1), we see that a major task will be to once more find the laplacian of our trial wave function (23).

$$\nabla_i^2 \psi_T = \frac{\nabla_i^2 \psi_{SD}}{\psi_{SD}} + \frac{\nabla_i^2 \psi_J}{\psi_J} + 2 \frac{\nabla_i \psi_{SD}}{\psi_{SD}} \cdot \frac{\nabla_i \psi_J}{\psi_J} \tag{40}$$

We begin by finding the Jastrow terms. The full derivation can be found in the lecture notes for Computational Physics 2 CITE KOMPENDIEM . k is here used instead of ias in the equation to the Hamiltonian (1). HA MED GENERELL FORM OG VIS LITT MER FRAMGANGSMETODE?

$$\frac{1}{\psi_J} \nabla_k \psi_J = \sum_{i \neq k}^N \frac{\mathbf{r}_{ik}}{r_{ik}} \frac{a_{ik}}{(1 + \beta r_{ik})^2}$$

$$\tag{41}$$

The laplacian of the Jastrow factor is given as

$$\frac{1}{\psi_J} \nabla_k^2 \psi_k = \left(\frac{\nabla_k \psi_J}{\psi_J}\right)^2 + \sum_{i \neq k}^N \frac{a_{ik} (1 - \beta r_{ik})}{r_{ik} (1 + \beta r_{ik})^3} \tag{42}$$

We can now look at the Slater determinant. As shown in http://onlinelibrary.wiley.com/doi/10.1002 we can split the spin parts into two Slater determinants, one for spin up and one for spin down since our Hamiltonian is spin independent. If we do not do this the determinant will be zero.

- 2.5.1 Local energy
- 2.5.2 Quantum force
- 2.5.3 α and β derivatives in the steepest descent algorithm

Kanskje ikke ha med noe her?

- 3 Implementation
- 3.1 TESTS
- 3.2 Slater determinants

An efficient way of updating the inverse of the Slater matrix can be found by SITER .

$$D_{kj}^{-1}(\mathbf{x}^{\text{new}}) = \begin{cases} D_{kj}^{-1}(\mathbf{x}^{old}) - \frac{D_{ki}^{-1}(\mathbf{x}^{old})}{R} \sum_{l=1}^{N} D_{il}(\mathbf{x}^{new}) D_{lj}^{-1}(\mathbf{x}^{old}) & \text{if } j \neq i \\ \frac{D_{ki}^{-1}(\mathbf{x}^{old})}{R} \sum_{l=1}^{N} D_{il}(\mathbf{x}^{old}) D_{lj}^{-1}(\mathbf{x}^{old}) & \text{if } j = i \end{cases}$$

$$(43)$$

- 4 Results
- 5 Discussion and conclusion
- 6 Appendix
- 6.1 Hermite polynomials

Ha med polynomene som brukes

6.2 Useful formulas

Solution to a Gaussian integral is given as,

$$\int \exp(-\omega x^2) dx = \sqrt{\frac{\pi}{\omega}} = \sqrt{\pi}\omega^{-\frac{1}{2}}$$
(44)

Another symmetric Gaussian integral is given as,

$$\int x^2 \exp\left(-\omega x^2\right) dx = \frac{\sqrt{\pi}}{2} \omega^{-\frac{3}{2}} \tag{45}$$

REFERANSER!!