# Note on Reconstruction beyond the Zel'dovich approximation

Hong-Ming Zhu,<sup>1,2</sup> Yu Yu,<sup>3</sup> Ue-Li Pen,<sup>4,5,6,7</sup> and Xuelei Chen<sup>1,2,8</sup>

<sup>1</sup>Key Laboratory for Computational Astrophysics, National Astronomical Observatories,
Chinese Academy of Sciences, 20A Datun Road, Beijing 100012, China

<sup>2</sup>University of Chinese Academy of Sciences, Beijing 100049, China

<sup>3</sup>Key Laboratory for Research in Galaxies and Cosmology, Shanghai Astronomical Observatory,
Chinese Academy of Sciences, 80 Nandan Road, Shanghai 200030, China

<sup>4</sup>Canadian Institute for Theoretical Astrophysics, University of Toronto,
60 St. George Street, Toronto, Ontario M5S 3H8, Canada

<sup>5</sup>Dunlap Institute for Astronomy and Astrophysics, University of Toronto,
50 St. George Street, Toronto, Ontario M5S 3H4, Canada

<sup>6</sup>Canadian Institute for Advanced Research, CIFAR Program in
Gravitation and Cosmology, Toronto, Ontario M5G 1Z8, Canada

<sup>7</sup>Perimeter Institute for Theoretical Physics, 31 Caroline Street North, Waterloo, Ontario, N2L 2Y5, Canada

<sup>8</sup>Center of High Energy Physics, Peking University, Beijing 100871, China
(Dated: September 22, 2016)

In this paper we introduce a new way to reconstruct BAO peaks in real space.

PACS numbers:

#### I. INTRODUCTION

The imprint of baryon acoustic oscillations (BAO) in the large-scale structure provides a standard ruler to measure the expansion rate of the Universe. Precise measurements of the BAO feature are crucial for probing the dynamics of dark energy. Nonlinear evolution in the density field damps the oscillations in the linear power specturm, reducing the BAO signal can be extracted from observations. The lost linear BAO information can be partially recovered by the reconstruction technique []. Precision modelling of the reconstructed density field and further improvements of the reconstruction method are of great importance for measureing the BAO peak at the subpercent level in the current and future dark energy experiments.

In the standard BAO reconstruction algorithm, the negative Zel'dovich (linear) displacement is used to reverse the large-scale bulk flows. The small-scale inhomogeneities in the nonlinear density field is ususally suppressed using a Gaussian smoothing of scale R so that the smoothed density field provides a reliable estimation on the linear displacement. The smoothing scale R should be comparable to the nonlinear scale  $\sim 10 \text{ Mpc/}h$ , where linear approximation breaks down. Thus, the nonlinear information is not used in the reconstruction. However, using smaller smoothing scales in order to include the small-scale modes would make the Zel'dovich approximation unvalid and induce nonlinearities which are difficult to model. The gain would be huge if these small-scale modes can be exploited, providing the information about the higher-order displacement field. The estimated negative Zel'dovich displacement is evaluated at the Eulerian position x rather than the Lagrangian position q. The difference between these two positions is often neglected in the reconstruction, which is a reasonable first order approximation (See Section 3.1 in Ref. [] for a discussion about this). Recently it is suggested that a new correction term due to this should be included in modelling the displacement (See Appendix A in Ref. for more detailed discussions). In this Letter, we point out that the nonlinear displecement can be solved from the nonlinear density field and present a new method to reconstruct the linear BAO information from the nonlinear density field.

The basic idea is to build a curvilinear coordinate system, where the mass per volume element is constant [3,4]. Consider a numerical grid of coordinates  $\boldsymbol{\xi} \equiv (\xi_1, \ \xi_2, \ \xi_3)$ . In order to determine the physical position of each lattice point, one needs to specify the Cartesian coordinate  $\boldsymbol{x}(\boldsymbol{\xi}, t)$  of each curvilinear coordinate. In the Cartesian coordinate system, the metric is just the Kronecker delta function  $\delta_{ij}$ , while the curvilinear metric is

$$g_{\alpha\beta} = \frac{\partial x^i}{\partial \xi^{\alpha}} \frac{\partial x^j}{\partial \xi^{\beta}} \delta_{ij}, \tag{1}$$

where Latin indices denote Cartesian coordinate labels  $x^i$ , while Greek indices imply curvilinear coordinates  $\xi^{\alpha}$ . As in Refs. [3,4], we define a coordinate transformation that is a pure gradient,

$$x^i = \xi^\mu \delta^i_\mu + \Delta x^i, \tag{2}$$

where

$$\Delta x^i \equiv \frac{\partial \phi}{\partial \mathcal{E}^{\nu}} \delta^{i\nu}.\tag{3}$$

The lattice displacement  $\Delta x$  is completely determined by the potential  $\phi$ . The mass density at the curvilinear coordinate  $\boldsymbol{\xi}$  is

$$\rho(\boldsymbol{\xi}) = \sqrt{g}\rho(\boldsymbol{x}),\tag{4}$$

where  $\sqrt{g} \equiv \det(\partial x^i/\partial \xi^{\alpha})$  is the volume element. The potential field  $\phi(\boldsymbol{\xi})$  which gives  $\rho(\boldsymbol{\xi}) = 0$  can be solved

using the multigrid alogirthm described in Ref. []. Since we only have the density field, which is a scalar field, it only allows us to reconstruct the scaler potential  $\phi$  rather than the vector. The reconstructed density field  $\delta_r(\boldsymbol{\xi})$  is given by

$$\delta_r(\boldsymbol{\xi}) = \nabla_{\boldsymbol{\xi}} \cdot \Delta \boldsymbol{x}(\boldsymbol{\xi}). \tag{5}$$

In the Lagrangian approach to the nonlinear structure formation, the displacement field  $\Psi(q,\tau)$  fully describes the motion of the mass elements. The Eulerian position x of a mass element is

$$x = q + \Psi(q, \tau), \tag{6}$$

where q is the initial Lagrangian position of this mass element. The displacement field  $\Psi$  can be decomposed into an irrotational part and a curl part,

$$\Psi = \Psi_E + \Psi_B, \tag{7}$$

where  $\nabla \times \Psi_E = 0$  and  $\nabla \cdot \Psi_B = 0$ . In the case that particles describe an irrotational potential flow,  $\Psi_E$  fully describes the motion of the mass elements. In reality, shell crossing happens in the nonlinear regime, and the emergence of vorticity leads to the growth of  $\Psi_B$ . The reconstructed lattice displacement  $\Delta x$  describes the nonlinear motion beyond the Zel'dovich displacement.

The real displacement field is available in N-body simulations. It can be decomposed and compared with the reconstructed lattice displacement to see how much information is reconstructed.

## II. DISPLACEMENT DECOMPOSITION

The displacement field  $\Psi$  can be decomposed into an irrotational part and a curl part,

$$\Psi = \Psi_E + \Psi_B, \tag{8}$$

where  $\nabla \times \Psi_E = 0$  and  $\nabla \cdot \Psi_B = 0$ . We further decompose  $\Psi_E$  as

$$\Psi_E = \Psi_E^{(l)} + s_E^{(nl)}. \tag{9}$$

The density field is given by

$$\delta_E^{(l)} = -\nabla \cdot \mathbf{\Psi}_E^{(l)} \tag{10}$$

The density field is given by

$$\delta_E^{(nl)} = -\nabla \cdot \mathbf{\Psi}_E^{(nl)} \tag{11}$$

The displacement can be written as

$$\mathbf{\Psi} = \mathbf{\Psi}_E^{(l)} + \mathbf{\Psi}_E^{(nl)} + \mathbf{\Psi}_B, \tag{12}$$

## A. Power spectra of displacement fields

Show the power spectra of different displacement componenets  $\dots$ 

#### B. PDFs

Show the pdf of the displacement fields, maybe the density field?

### C. Cross correlations

Show the cross correlation coefficients of the nonlinear density field and the linear density field ...

## III. RECONSTRUCTION ALGORITHM

Discuss how to solve the deformation potential.

The Lagrangian displacement can be solved easily in 1+1 dimensions by ordering of the mass elements.

- (1) Solve the displacement  $\Psi(q)$  field by ordering of the mass elements.
- (2) Take the differential derivative of  $\Psi(q)$  to get the reconstructed density field  $\delta_r(\mathbf{q}) = -\nabla_{\mathbf{q}}\Psi(\mathbf{q})$ .  $\delta_r(\mathbf{q}) = -\partial\Psi(q)/\partial q$ .

## IV. SIMULATIONS AND PERFORMANCE

We adopt the 1D N-body simulations in Ref. [1] and use outputs at z=0. The simulation box is  $10^8$  Mpc with  $3\times10^8$  grids and  $3\times10^8$  PM elements. We scale the intial density field by the linear growth factor to get the linear density field  $\delta_L$  at z=0. Note that  $\delta_L$  has been rescaled using the linear growth function to the same redshift as  $\delta$ .

#### A. Deformation potential

Show the visual grid deformation.

#### B. Statistics

Show resonctruction from linear density fields and z = 0 density field. Different configurations.

To convenienty quantify the linear information  $\delta_L$  in the nonlinear density field  $\delta$ , we decompose the nonlinear density field  $\delta$  as

$$\delta(\mathbf{k}) = b(\mathbf{k})\delta_L(\mathbf{k}) + \delta_N(\mathbf{k}). \tag{13}$$

Here,  $b\delta_L$  is completely correlated with the linear density field  $\delta_L$  and  $b = P_{\delta\delta_L}/P_{\delta_L}$ . Nonlinear evolution drives b to drop from unity, reducing the linear signal.  $\delta_N$  is generated in the nonlinear evolution and thus uncorrelated with the linear density field  $\delta_L$ , further reducing  $b\delta_L$  with respect to  $\delta$ . This part induces noises in the measurement of BAO. Hence we denote it with a subscript "N". Such

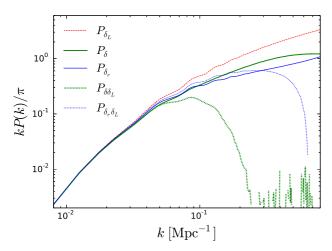


FIG. 1: The power spectra of the linear (dashed line), non-linear (thick solid line), and reconstructed (thin solid line) fields. We also plot the nonlinear-linear (thick dotted line) and reconstructed-linear (thin dotted line) cross-correlation power spectra.

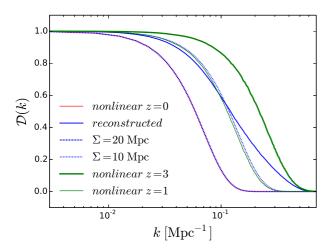


FIG. 2: The damping factors for the nonlinear (thin solid line) and reconstructed (thick solid line) fields. The Gaussian BAO damping models with  $\Sigma=20$  Mpc (thick dotted line) and  $\Sigma=10$  Mpc (thin dotted line).

decomposition helps to write the nonlinear power spectrum as

$$P_{\delta}(k) = \mathcal{D}(k)P_{\delta_L}(k) + P_{\delta_N}(k), \tag{14}$$

where  $\mathcal{D}=b^2$  describes the damping of the linear power spectrum. The reconstructed power spectrum  $P_{\delta_r}$  can be describe in the same way. Here,  $b(\mathbf{k})$  is often referred as the "propagator" and  $n(\mathbf{k})$  is usually called the modecoupling term [2–4]. still to be modified

Figure 2 shows the damping functions for the raw and reconstructed fields. The nonlinear damping of the linear power spectrum is significantly reduced after reconstruction. We also overplot the best-fitting Gaussian BAO

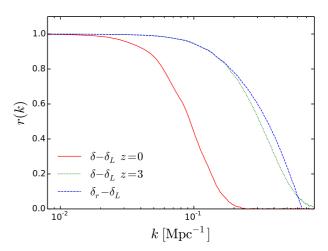


FIG. 3: The  $\delta - \delta_L$  correlation coefficients at z = 0 (solid line) and z = 3 (dotted line), as well as the  $\delta_r - \delta_L$  correlation coefficient (dashed line).

damping model,

$$\mathcal{D}(k) = e^{-k^2 \Sigma^2 / 2},\tag{15}$$

with  $\Sigma$  =? Mpc and ? Mpc for the nonlinear and reconstructed fields. The new BAO reconstruction algorithm reduces the the nonlinear damping scale  $\Sigma$  by ?? per cent, i.e., a of ??. The damping factor is above 0.9 for  $k \lesssim$ ? Mpc<sup>-1</sup> indicating (almost) perfect reconstruction. how to best quantify the reduction of damping? However, the 100 per cent reconstruction, cancelling any nonlinear effects, is still unachievable, as some information has been irreversibly lost. (more discissions)

Reconstruction also reduces the noise term  $P_{\delta_N}$ . To demonstrate this, in Fig. 3 we plot the cross-correlation coefficient

$$r(k) = \frac{P_{\delta\delta_L}(k)}{\sqrt{P_{\delta}(k)P_{\delta_L}(k)}} = \frac{1}{\sqrt{1+\eta(k)}},$$
 (16)

where  $\eta = P_n/(\mathcal{D}P_{\delta_L})$  quantifies the relative amplitude of  $\delta_N$  with respect to  $b\delta_L$ . The correlation of  $\delta_r$  with  $\delta_L$  is as good as that of  $\delta$  at z=3. how to quantify this better?

Distribution functions

#### V. DISCUSSIONS AND CONCLUSIONS

The new method significantly improves the expansion rate measurement from BAO. (more discussions?)

This method can be generalized to the 3D case. We leave this to future work.

Comparision with and Implications for the standard BAO reconstruction: exact Lagrangian displacement, nonlinear displacement, which is easier to model.

If use the displacement solved in this paper for the standard BAO rec, we expect the performance will become much better but still not as good as our results.

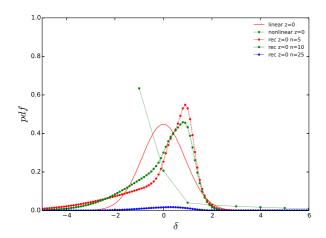


FIG. 4: The distribution functions.

## VI. ACKNOWLEDGEMENT

We acknowledge the support of the Chinese MoST 863 program under Grant No. 2012AA121701, the CAS Science Strategic Priority Research Program XDB090000000, the NSFC under Grant No. 11373030, IAS at Tsinghua University, and NSERC. The Dunlap Institute is funded through an endowment established by the David Dunlap family and the University of Toronto. Research at the Perimeter Institute is supported by the Government of Canada through Industry Canada and by the Province of Ontario through the Ministry of Research & Innovation.

M. McQuinn and M. White, J. Cosmology Astropart. Phys. 1, 043 (2016), 1502.07389.

<sup>[2]</sup> M. Crocce and R. Scoccimarro, Phys. Rev. D 73, 063520 (2006), astro-ph/0509419.

<sup>[3]</sup> M. Crocce and R. Scoccimarro, Phys. Rev. D 77, 023533 (2008), 0704.2783.

<sup>[4]</sup> T. Matsubara, Phys. Rev. D **77**, 063530 (2008), 0711.2521.