Bayesian Learning (732A73) Lab 1 Hoda Fakharzadehjahromy (hodfa840), Ravinder Atla(ravat601) **Assisgnment 1 Daniel Bernoulli** Bernoulli ...again. Let  $y_1,\ldots,y_n| heta\sim \mathrm{Bern}( heta)$  and assume that you have obtained a sample with s=8 successes in n=24trials. Assume a  $\mathrm{Beta}(lpha_0,eta_0)$  prior for heta and let  $lpha_0=eta_0=3$ (a )Draw random numbers from the posterior  $heta|y\sim ext{Beta}(lpha_0+s,eta_0+f)$ , where  $y=(y_1,\dots,y_n)$ ,; and verify graphically that the posterior mean and standard deviation converges to the true values as the number of random draws grows large. (b) Use simulation (nDraws = 10000) to compute the posterior probability  $\Pr( heta < 0.4|y)$  and compare with the exact value. (c) Compute the posterior distribution of the log-odds  $\phi = \log rac{ heta}{1- heta}$  (nDraws = 10000) (a) True mean for  $Beta(\alpha, \beta)$  is:  $E[ heta] = rac{lpha}{lpha + eta}$ True  $\sigma^2$  for  $Beta(\alpha, \beta)$  is:  $\sigma^2(\theta) = \frac{\alpha\beta}{(\alpha+\beta)^2(\alpha+\beta+1)}$ The prior function for this problem is:  $p(\theta) \propto \theta^{\alpha_0-1} (1-\theta^{\beta_0-1})$ The likelihood function is:  $p(y_1,\ldots,y_n| heta) \propto \prod_{i=1}^n p(y_i| heta) = \prod_{i=1}^n heta^{y_i} (1- heta)^{1-y_i} = heta^s (1- heta)^f$ posterior is: posterior  $p(\theta|y) \propto Beta(\alpha_0 + s, \beta_0 + f) = Beta(11, 19)$ In [1]:  $alpha_0 = beta_0 = 3$ In [2]: NDraws = 10000 res = matrix(,NDraws,3) colnames(res) = c("iter", "mean", "sd") alpha\_n=alpha\_0+s beta  $n = beta_0+f$ TrueMean = (alpha\_n)/(alpha\_n+beta\_n) TrueVar = sqrt((alpha\_n\*beta\_n)/((alpha\_n+beta\_n)^2\*(alpha\_n+beta\_n+1))) TrueMean TrueVar 0.366666666666667 0.0865507910219387 In [3]: PostB <- function(alpha= alpha n,beta =beta n ,n=NDraws) {</pre> for(i in 1:n) { postB = rbeta(i,alpha,beta) res[i,] = c(i, mean(postB), sd(postB))return(list(post=postB, mean sd=res)) Posterior = PostB() In [4]: Posterior\$mean\_sd[1,3]=0 head(Posterior\$mean\_sd) A matrix:  $6 \times 3$  of type dbl iter 1 0.3629400 0.00000000 2 0.2669733 0.07396499 3 0.3443858 0.06782267 4 0.3124737 0.09665964 0.3704016 0.11454616 In [5]:  $x \leftarrow seq(from = -5, to = 10, by = 0.001)$ hist(Posterior\$post,col='lightyellow',freq = FALSE,breaks=50,xlab='posterior',main='Halse,breaks curve(dbeta(x,alpha\_n,beta n),add=TRUE,col='blue', lwd =2,lty=3) **Histogram of Posterior** 0.1 0.2 0.3 0.4 0.5 0.6 0.7 posterior In [6]: plot(x=1:NDraws,y=Posterior\$mean\_sd[,2],col='lightblue',lwd =1,type='l',xlab="Number of the color of the abline(h=TrueMean,col='deeppink3',lwd=3) 0.40 2000 4000 6000 8000 10000 Number of Draws In [7]: plot(x=1:NDraws,y=Posterior\$mean sd[,3],col='lightblue',lwd =1,type='l',xlab="Number abline(h=TrueVar,col='deeppink3',lwd=3) 0.14 0.12 0.10 90.0 0.04 0.02 0.00 2000 4000 6000 8000 10000 Number of Draws As we can see from the two plots above, as the number of draws increases, both  $\mu$  and  $\sigma$  converge to the true values. (b) we can see that our estimation of  $Pr(\theta > 0.4|y)$  is very close to true theoratical value. In [8]: exact prb =1- pbeta(0.4,alpha n,beta n) pos prob = length(Posterior\$post[Posterior\$post > 0.4])/length(Posterior\$post) prb = data.frame(exact prb,pos prob) colnames(prb) = c('excact prob', "simulated value") excact\_prob simulated\_value <dbl> <dbl> 0.3426654 0.3402 (C) Histogram and kernel density of of the data simulated from the posterior distribution of the  $\phi = log(\theta (1 - \theta))$  with 10000 draws. In [9]: phi = log(Posterior\$post/(1- Posterior\$post)) hist(phi,breaks=100,probability = TRUE,col='cyan',xlab=expression(phi)) lines(density(phi), col='deeppink3', lty=2, lwd=3) Histogram of phi 1.0 9.0 Density 0.4 0.2 -2.0 -1.5 -1.0 -0.5 0.5 -2.5 **Assignment 2** Log-normal distribution and the Gini coefficient. Assume that you have asked 10 randomly selected persons about their monthly income (in thousands Swedish Krona) and obtained the following ten observations: 14, 25, 45, 25, 30, 33, 19, 50, 34 and 67. A common model for non-negative continuous variables is the log-normal distribution. The log-normal distribution  $\log \mathcal{N}(\mu, \sigma^2)$ has density function  $p(y|\mu,\sigma^2) = rac{1}{y\cdot\sqrt{2\pi\sigma^2}} ext{exp}(-rac{1}{2\sigma^2}(\log y - \mu)^2)$ where y>0 ,  $\mu>0$  and  $\sigma^2>0$ The log-normal distribution is related to the normal distribution as follows:  $y \sim \log \mathcal{N}(\mu, \sigma^2)$  then  $\log y \sim \mathcal{N}(\mu,\sigma^2)$  Let  $y_1,\ldots,y_n|\mu,\sigma^2 \stackrel{iid}{\sim} \log \mathcal{N}(\mu,\sigma^2)$  , where  $\mu=3.8$ is assumed to be known but  $\sigma^2$  is unknown with non-informative prior  $p(\sigma^2) \propto 1/\sigma^2$  The posterior for  $\sigma^2$ is the  $Inv-\chi^2(n, au^2)$ distribution, where  $au^2=rac{\sum_{i=1}^n(\log y_i-\mu)^2}{n}.$ (a) Simulate 10, 000 draws from the posterior of  $\sigma$ 2 (assuming  $\mu$  = 3.8) and com- pare it with the theoretical Inv –  $\chi$ 2(n,  $\tau$  2) posterior distribution. In [10]: num draws = 10000 #sample size mu = 3.8observ = c(38,20, 49, 58, 31, 70, 18, 56, 25,78)n obser = length(observ) In [11]: tau2 = function(data,mu,n\_obser) { sum((log(data) -mu)^2)/n\_obser In [12]: tau2(observ,mu,n\_obser) 0.261043665681099 In [13]: rinvchisq <- function(num\_draws, n\_obser, tau\_sq) {</pre> set.seed(1234) x <- rchisq(num draws, df = n obser-1)</pre>  $x inv \leftarrow ((n obser-1)*tau sq)/x$ return(x\_inv) In [14]: post sigma2 <- function(m) {</pre> set.seed(12345) rinvchisq(num draws = num draws, n obser = n obser, tau sq = tau2(observ, mu, n obser= In [15]: dinvchisq <- function(x,n\_obser,tau\_sq){</pre> res <- (((tau\_sq\*(n\_obser-1))/2)^(n\_obser-1)/2 \* exp(-(tau\_sq\*(n\_obser-1))/(2\*x)))/ return (res) In [16]:  $x \leftarrow seq(from = 0, to = 10, by = 0.001)$ hist(post sigma2(m = num draws), probability = TRUE, col='pink', breaks = 100, main=expression(paste('Simulated and theoratical posterior of ', sigma^2) ), ,xlab = expression(paste(sigma ^2))) curve(dinvchisq(x ,n\_obser= n\_obser,tau\_sq = tau2(observ,mu,n\_obser)),add=TRUE,cd Simulated and theoratical posterior of  $\sigma^2$ 0 2 (b) The most common measure of income inequality is the Gini coeffcient, G, where 0 < G < 1. G=0 means a completely equal income distribution, whereas G=1 means complete income inequality (see Wikipedia for more information about the Gini coeffcient). It can be shown that  $G=2\Phi\left(\,\sigma/\sqrt{2}\,
ight)-1$  when incomes follow a  $\log \mathcal{N}(\mu, \sigma^2)$  distribution. $\Phi(z)$  is the cumulative distribution function (CDF) for the standard normal distribution with mean zero and unit variance. Use the posterior draws in a) to compute the posterior distribution of the Gini coeffcient G for the current data set In [17]: sigma2 = post\_sigma2(m) g sigma = sqrt(sigma2)/sqrt(2)  $gpdf = 2* pnorm(q = g_sigma, mean = 0, sd = 1)-1$ hist(gpdf,probability = TRUE,col='yellow',breaks=50,main='the posterior distribution of the Gini coefficient', xlab=expression(paste('Gini ', sigma))) lines(density(gpdf),col='darkblue',lty=3,lwd=2) the posterior distribution of the Gini coefficient 9 2 0.2 0.3 0.4 0.5 0.6 0.7 0.8 Gini  $\sigma$ (c) Use the posterior draws from b) to compute a 90% equal tail credible interval for G. A 90% equal tail interval (a; b) cuts off 5% percent of the posterior probability mass to the left of a, and 5% to the right of b. Also, do a kernel density estimate of the posterior of G using the density function in R with default settings, and use that kernel density estimate to compute a 90% Highest Posterior Density Interval (HPDI) for G. Compare the two intervals. In [18]:  $alpha_conf = 0.1$ q\_lower = quantile(gpdf,alpha\_conf/2) q\_upper = quantile(gpdf,1-alpha\_conf/2) c(q\_lower, q\_upper) true mean = mean(gpdf) In [19]: true mean 0.305486649437337 HPD: shortest possible interval that under the posterior has the siginificance probability (ie. 0.9) In [20]: KerDenG=density(gpdf) xx = KerDenG\$xyy = sort(KerDenG\$y, decreasing = TRUE, index.return=TRUE) we will use the cumulative sum of kernel density values to find the shortest interval in data that contains 0.9 of the data. the reason we can use cumulative to find HPD interval is because our data is unimodal. In [21]: CumSum = cumsum(yy\$x) $HPD_D = CumSum[length(CumSum)]*0.9$ temp = which(CumSum<HPD D)</pre> HPD\_interval = range(xx[which(CumSum<HPD\_D)])</pre> In [22]: df = data.frame('HPD\_interval'=HPD\_interval,'credible\_interval'=c(q\_lower, q\_upper)) In [24]: y temp=density(gpdf, n=10000) plot(x=y temp\$x,y=y temp\$y,type = '1', lwd = 2, col = 'violet', main= 'distribution of ,xlab=expression(paste('Gini ', sigma)),ylab='Density') polygon(x=y\_temp\$x,y=y\_temp\$y, col = 'lightgrey',border = 'darkblue') abline(v = true mean, col="deeppink", lwd=3, lty=2) segments (x0 = q lower, y0 = 0.5, x1 = q upper, y1 = 0.5, col='deepskyblue1', lwd=3)abline(v=q\_lower,col='deepskyblue1' ,lwd=3) abline(v=q upper,col='deepskyblue1',lwd=3) segments(x0=HPD\_interval[1],y0=0.8,x1=HPD\_interval[2],y=0.8,col='darkolivegreen',lwd= abline(v=HPD interval[1], col='darkolivegreen', lwd=3) abline(v=HPD interval[2], col='darkolivegreen', lwd=3)  $arrows(x0 = q_lower, y0 = 0.5, x1 = q_upper, y1 = 0.5, col='deepskyblue1', lwd=3, code = 3)$ arrows(x0=HPD\_interval[1],y0=0.8,x1=HPD\_interval[2],y=0.8,col='darkolivegreen',lwd=3,c legend("topright", legend = c(" Credible interval", "HPD", expression(mu)), col = c('deepskyblue1','darkolivegreen','deeppink'), lty=c(1,1,2), cex=0.8,lwd=3,bg= distribution of the Gini coefficient Credible\_interval HPD 5 3 7 0.2 0.4 0.6 8.0 Gini  $\sigma$ **Assignment 3** Bayesian inference for the concentration parameter in the von Mises distribution distributions for circular data we convert the data into radians. The 10 observations in radians are (-2.44, 2.14, 2.54, 1.83, 2.02, 2.33, -2.79, 2.23, 2.07, 2.02).Assume that these data points are independent observations following the von Mises distribution  $p(y|\mu,\kappa) = rac{\exp[\kappa \cdot \cos(y-\mu)]}{2\pi I_0(\kappa)}, \; -\pi \leq y \leq \pi,$ where  $I_0(\kappa)$  is the modified Bessel function of the first kind of order zero. The parameter  $-\pi \leq \mu \leq \pi$  is the mean direction and  $\kappa > 0$  is called the concentration parameter. Large  $\kappa$  gives a small variance around  $\mu$ , and vice versa. Assume that  $\mu$  is known to be 2.39. Let  $\kappa \sim \mathrm{Exponential}(\lambda=1)$  a priori, where  $\lambda$  is the rate parameter of the exponential distribution (so that the mean is  $1/\lambda$ ). (a) Plot the posterior distribution of  $\kappa$  for the wind direction data over a fine grid of  $\kappa$  values. (b) Find the (approximate) posterior mode of  $\kappa$  from the information in a). (a) Posterior:  $p(\kappa \mid y_1, y_2, \dots, y_n) \propto p(y_1, y_2, \dots, y_n \mid \kappa) \cdot p(\kappa)$  $p(\kappa \mid y_1, y_2, \dots, y_n) \propto \left[rac{1}{I_0(\kappa)}
ight]^n \cdot \exp \left[\sum_{i=1}^n \kappa \cdot cos(y_i - \mu) - \lambda \kappa
ight]$  $p(\kappa \mid y_1, y_2, \dots, y_n) \propto \left[rac{1}{I_0(\kappa)}
ight]^n \cdot \exp\left[\sum_{i=1}^n \kappa \cdot cos(y_i - 2.39) - \kappa
ight]$ In [25]: y\_data =c(-2.44, 2.14, 2.54, 1.83, 2.02, 2.33, -2.79, 2.23, 2.07, 2.02) In [26]: n =length(y\_data) mu = 2.3910 In [27]: # a function to compute prior for k (exponential with lambda = 1) prior\_k <- function(k){</pre> dexp(x = k, rate = 1)In [28]: likelikood\_k <- function(y,mu ,n,k){</pre>  $\exp(k*sum(cos(y-mu)))/((2*pi*besselI(k,nu = 0))^n)$ In [29]: k = seq(from = 0, to = 20, by = 0.01)In [30]: posterior k <- function(y=y\_data,mu=mu ,n=n,k=k) {</pre> likelikood\_k(y,mu ,n,k )\*prior\_k(k) In [31]: post=posterior k(y=y data,mu=mu ,n=n,k=k) In [32]: plot(y=post, x=k, col='blue', type='l', lwd=3) 1e-06 10 15 20 (b) Mode In [33]: k[which.max(post)] 2.12 In [34]: hist(y data,probability = TRUE,breaks = 50,col='brown') Histogram of y\_data 3.0 2.5 2.0 1.5 1.0 0.5 -2 -1 0 1 y\_data from the plot above we can see that  $\kappa=2.12$  is the actual mode of our data. In [35]: y data[which(y data<0)]</pre> -2.44 · -2.79