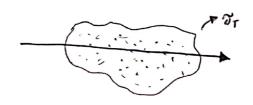
Home work #4

1. The warm contend interstellar medicm (ISM) of our Gabaxy in the solar neighborhood can be approximated by a disk of half-thickness half vpc and declarate we action density. Ne 20.1cm-3 what 15 the optical depth of the ISM to Thompson scattering in the direction normal to the disk?

So, If we consider the problem in this way, we have



Assuming that the Midhess of the deject is given by he and considering that only Thompson scattering is used for the calculations we have that,

$$J = \int_0^5 \alpha(s) ds$$
 or more precisely $J_v(s) = \int_{s_0}^5 \alpha v(s') ds'$

and, furthermore, we have to keep in mind that,

with 11 being the particle desity and on the effective absorbing area. Now, the electron scrittering is frequency independent so,

$$\mathcal{I} = \int_0^h h \sigma v ds = h \sigma v N \Rightarrow \mathcal{I} = h \cdot \frac{\partial \pi}{3} r_0^2 \cdot h$$

2. Consider a medium containing a large number of radialing particles. (For definitioness you way wish to imagine electrons emitting bremsstraling). Each porticle emits a pulse of radiation with an electric field Eo(t) as a function of time. An observer will detect a series of such pulses, all with the some shape but with random arrival times to, to, to, to, the measured electric field will be,

$$E(t) = \sum_{i=1}^{N} E_0(t - \epsilon_i)$$

9) Show that the Farner transform of E(t) is É(w) = Ê(w) & e(wti So, considering that the general form to the Fourier transform is,

$$F(\omega) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} f(t)e^{i\omega t} dt$$

and $f(t) = \sum_{i=1}^{N} E_0(t - \epsilon_i)$, we then obtain,

$$F(\omega) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \left[\sum_{i=1}^{N} E_0(t-t_i) \right] e^{i\omega t} dt$$

Now, considering that if is a linear combination of linearly independent functions we can rearrange the equation such that,

$$F(w) = \frac{1}{\sqrt{2\pi^2}} \sum_{i=1}^{N} \int_{-\infty}^{\infty} E_0(t-ii) e^{i\omega t} dt$$

Continuing, we can make a charge of variable given that to conbe represented as a constant related to $\sum_{i=1}^{N}$, so

$$0 = t - ti$$

$$\frac{dv}{dt} = 1 = > t = v + ti$$

$$dv = dt$$

$$e^{i\omega t} = e^{i\omega v} + i\omega ti$$

$$e^{i\omega v} = e^{i\omega t}$$

$$\Rightarrow F(\omega) = \frac{1}{\sqrt{2\pi'}} \sum_{l=1}^{N} e^{i\omega t_l} \int_{-\infty}^{\omega} E_0(0) e^{i\omega t_l} dt$$

Now, we know that this term

is equal to Eo(w), 50

$$E(\omega) = E_0(\omega) \sum_{i=1}^{N} e^{i\omega t_i}$$

$$\left|\sum_{i=1}^{N} e^{i\omega t_i}\right|^2 = N$$
 when averaged over the random arrival times.

Considering that
$$\sum_{i=1}^{n} \sum_{j=1}^{n} a_{ij} = \sum_{ij=1}^{n} a_{ij}$$
 and considering the random arrival times we then

$$\left|\sum_{i=1}^{N} e^{i\omega t_i}\right|^2 = \sum_{i=1}^{N} \sum_{j=1}^{N} e^{i\omega t_j} e^{-i\omega t_i} \text{ being i related to } (t-t_i) \text{ retaited times.}$$
b) when i = i = i = 1

$$\sum_{j=1}^{N} e^{i\omega t_{j}} e^{-i\omega t_{j}} = \sum_{j=1}^{N} e^{i\omega(t_{j} - t_{i})} = \sum_{j=1}^{N} e^{j\omega(t_{j} - t_{i})} = \sum_{j=1}^{N} e^{j\omega(t_{j} - t_{i})}$$

$$\sum_{j=1}^{N} \sum_{i=1}^{N} e^{i\omega(\ell_j - \ell_i)}$$

We know that the measured spectrum Follow the robotion

$$\frac{dw}{dAdw} = c |\hat{E}(w)|^2$$
, now, from this $\left| \sum_{i=1}^{N} e^{iwki} \right|^2 = N$, we condonce

$$\frac{d\omega}{dAd\omega} = c |E_0(\omega) \sum_{i=1}^{N} e^{i\omega ki}|^2 = > \frac{d\omega}{dAd\omega} = c E_0(\omega) N$$

which derives into

d) By contrast, show that if all the particles are in a region much smaller than a unveloight and they emit their pulses simultaneously, then the measured spectrum will be N2 times that spectrum of an individual pulse
For this case we argue that, $2>> d$ $60 \text{ that means } t-ti \approx t \text{ which also means } 6i=0,$
Hoefere, $E(\omega) = E(\omega) \sum_{i=1}^{N} e^{i\omega(i)}(N) = E(\omega)N$ 50 for Hi,
$\frac{dw}{dAdw} = c \hat{E}(w) ^2 = \frac{dw}{dAdw} = c \hat{E}(w)N ^2$
Which derives into
Which derives into $ \frac{dw}{dA dw} = N^2 \frac{dW}{dA dw} $ Lesingle culse