

# **Majorana Neutrinos & TeV Phenomenology**

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# Outline

- Introduction
- Models with Majorana neutrinos
- TeV Phenomenology: (i)  $\nu$  mass generations,  
(ii)  $0\nu\beta\beta$  decays, and (iii) other physics
- Summary

## ● Introduction

### Experiments on solar neutrinos

$$\Delta m_{21}^2 = 7.59^{+0.20}_{-0.18} \times 10^{-5} \text{ eV}^2$$

$$\Delta m_{sun}^2 = |m_2|^2 - |m_1|^2 > 0$$

### Neutrinos born in Cosmic ray collisions and on earth

$$|\Delta m_{31}^2| = \begin{cases} 2.45 \pm 0.09 \times 10^{-3} \text{ eV}^2 \\ 2.34^{+0.10}_{-0.09} \times 10^{-3} \text{ eV}^2 \end{cases}$$

normal hierarchy,  
inverted hierarchy,

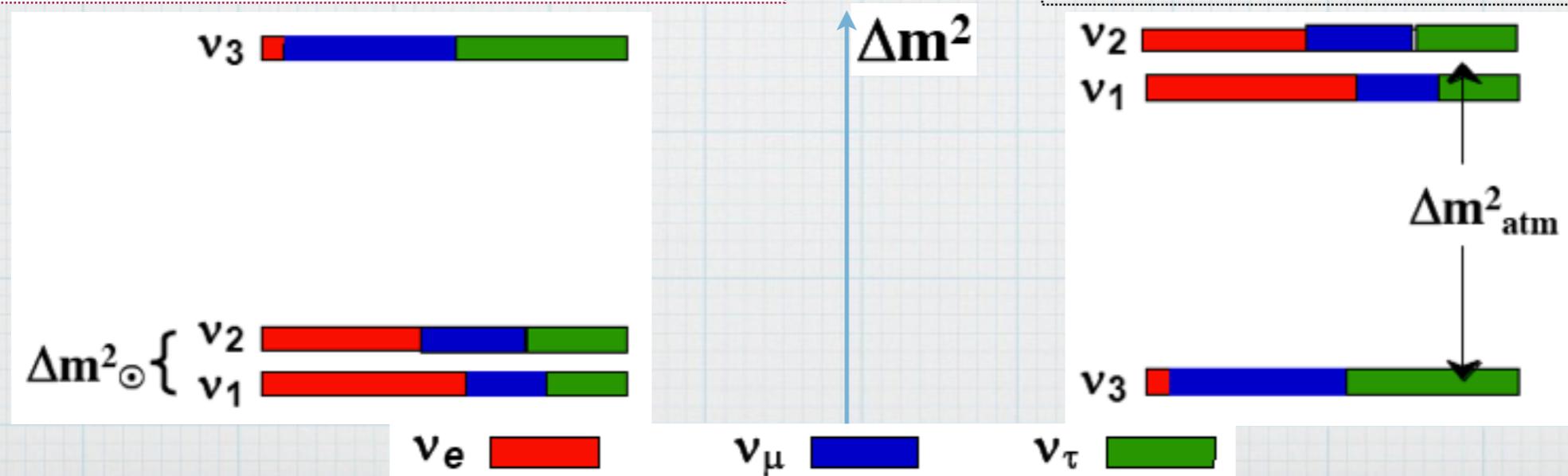
$$\Delta m_{atm}^2 = |\Delta m_{31}^2|$$

$$\Delta m_{31}^2 = |m_3|^2 - |m_1|^2$$

#### Normal hierarchy

$$|m_{\nu_1}| < |m_{\nu_2}| \ll |m_{\nu_3}|$$

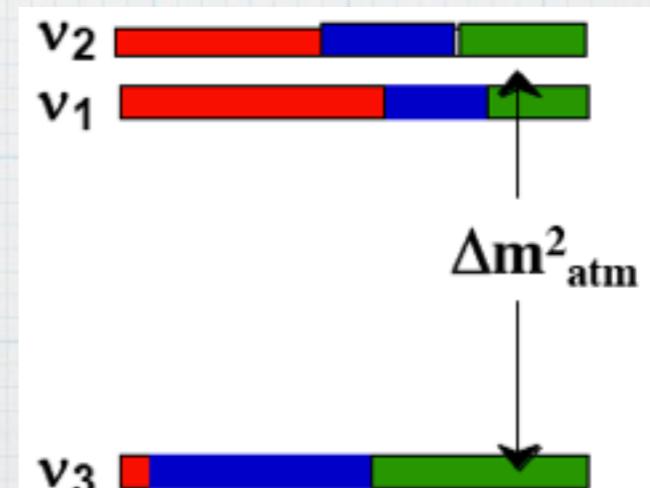
$$m_1 \simeq 0, m_2^2 \simeq \Delta m_{\odot}^2, \text{ and } m_3^2 \simeq \Delta m_{atm}^2.$$



#### Inverse hierarchy

$$|m_{\nu_1}| \simeq |m_{\nu_2}| \gg |m_{\nu_3}|$$

$$m_1^2 \simeq m_2^2 \simeq \Delta m_{atm}^2 \gg m_3^2$$



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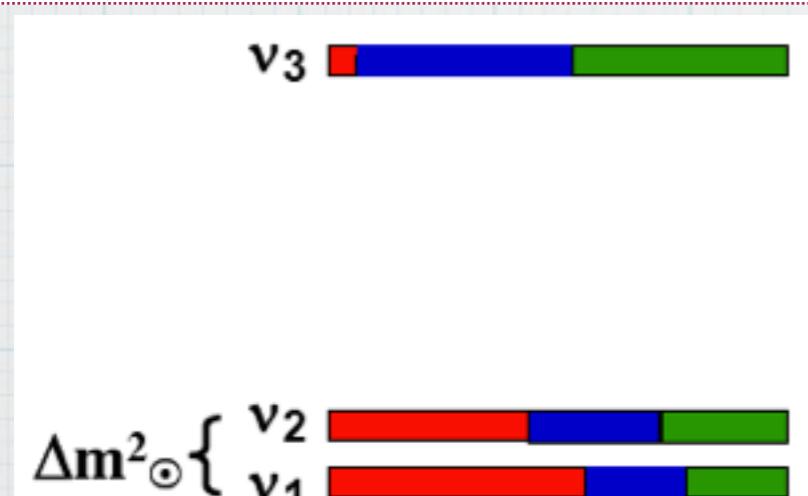
$m_{\nu_e} < 2.3$  eV (95% C.L.)

KATRIN **0.2 eV**

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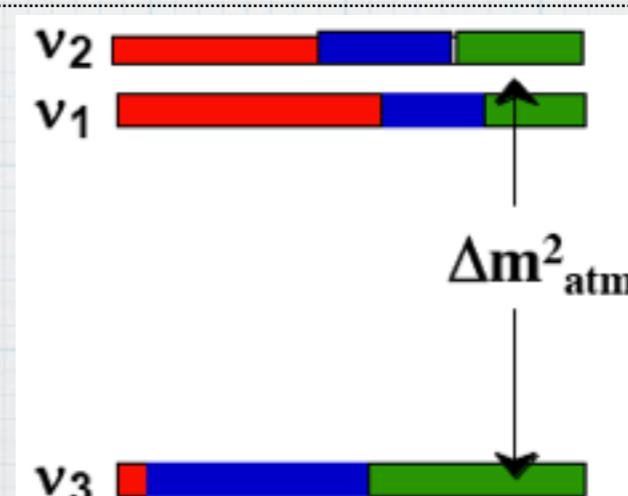
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$m_{\nu_e} < 2.3 \text{ eV}$  (95% C.L.)

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The best bound to their absolute values of the masses comes from Cosmology



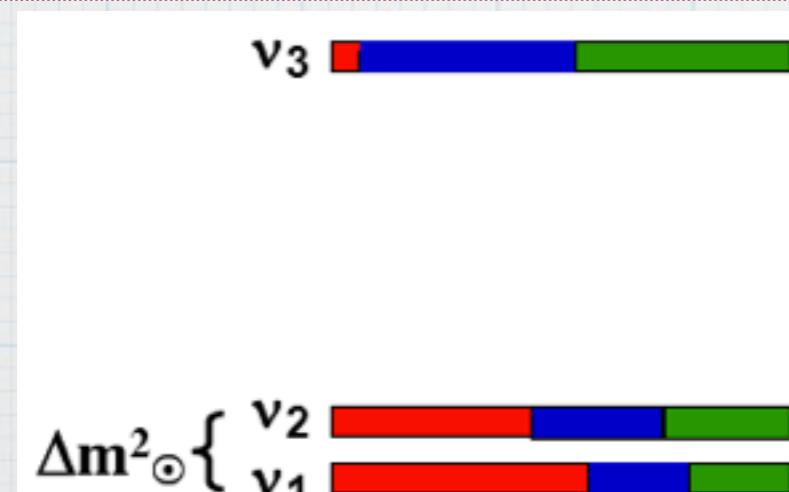
$$\sum_i m_{\nu_i} <$$

0.25 eV 95% CL (Planck+other data)

**Normal hierarchy**

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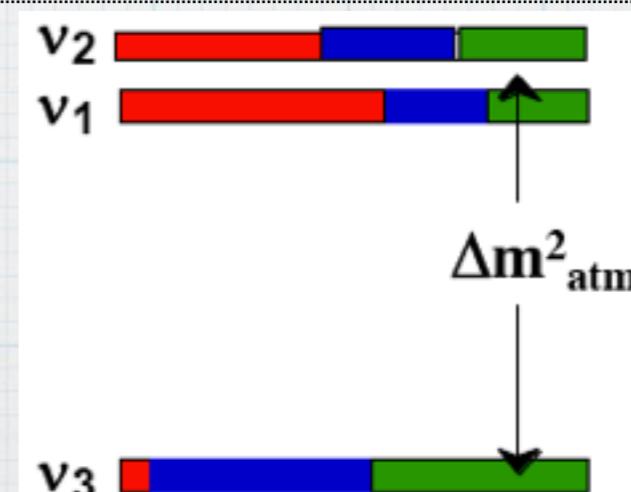
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$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \underbrace{\begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix}}_{\text{BF}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \rightarrow V_{PMNS} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix}$$

A global fit yields

$$\sin^2 \theta_{12} = 3.07_{-0.16}^{+0.18} \times 10^{-1}, (16\%)$$

$$\sin^2 \theta_{13} = 2.41 \pm 0.25 \times 10^{-2}, (10\%)$$

$$\sin^2 \theta_{23} = 3.86_{-0.21}^{+0.24} \times 10^{-1}, (21\%)$$

$$\delta/\pi = 1.08_{-0.31}^{+0.28} \text{ rad},$$

BF

$$\begin{aligned} \theta_{12} &\approx 33.7^\circ \\ \theta_{23} &\approx 38.4^\circ \\ \theta_{13} &\approx 8.93^\circ \\ \delta &\approx \pi \end{aligned}$$

## Bimaximal Matrix

$$\cancel{\begin{pmatrix} \frac{\sqrt{2}}{2} & \frac{\sqrt{2}}{2} & 0 \\ -\frac{1}{2} & \frac{1}{2} & \frac{\sqrt{2}}{2} \\ \frac{1}{2} & -\frac{1}{2} & \frac{\sqrt{2}}{2} \end{pmatrix}}$$

$$\theta_{12}=45^\circ, \theta_{23}=45^\circ, \theta_{13}=0$$

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Daya-Bay

$$\theta_{12}=35.3^\circ, \theta_{23}=45^\circ, \theta_{13}=0$$

$$|U_{PMNS}| \sim \begin{pmatrix} 0.8 & 0.5 & 0.2 \\ 0.4 & 0.6 & 0.7 \\ 0.4 & 0.6 & 0.7 \end{pmatrix}$$

$$|V_{CKM}| \sim \begin{pmatrix} 1 & 0.2 & 0.004 \\ 0.2 & 1 & 0.04 \\ 0.008 & 0.04 & 1 \end{pmatrix}$$

*Neutrino oscillations measure  $m^2$  but they do not provide information about the absolute neutrino spectrum and cannot distinguish pure Dirac and Majorana neutrinos.*

## Origin of the neutrino masses: Dirac or Majorana?



Dirac neutrino mass:

$$\mathcal{L}_D = -m_D \overline{\nu_L} \nu_R + \text{h.c.}$$



Majorana neutrino mass:

$$\mathcal{L}_M = -m_M \overline{\nu^c} \nu + \text{h.c.}$$

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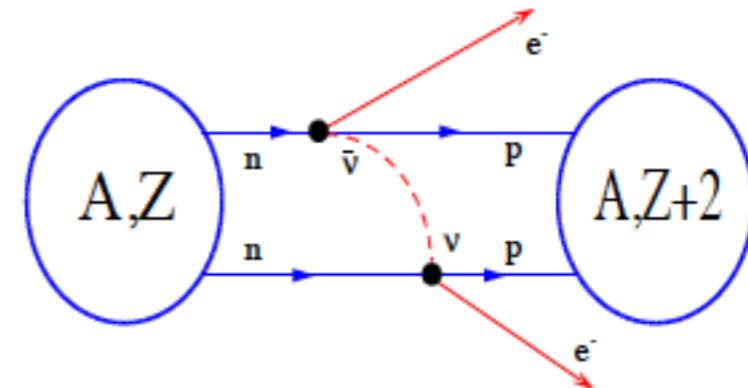
$$\nu \leftrightarrow \bar{\nu}$$

😊 the lepton number L is conserved

- the lepton number L is violated

Nuclear  $0\nu\beta\beta$ -decay

**FORBIDDEN  
IN THE SM.**



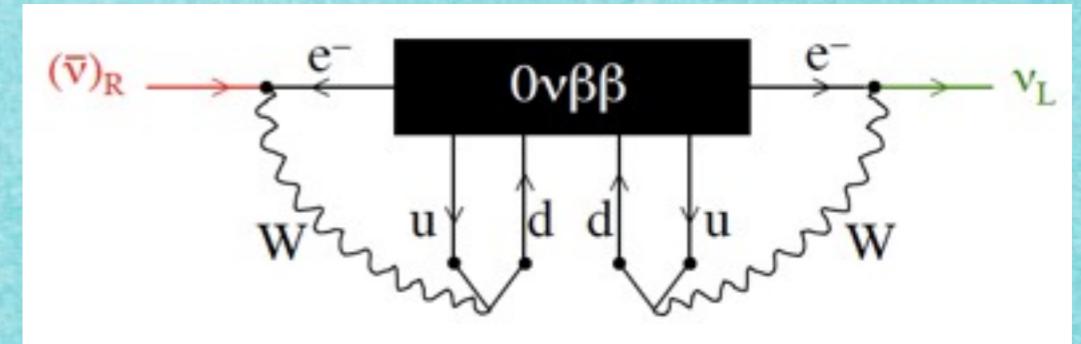
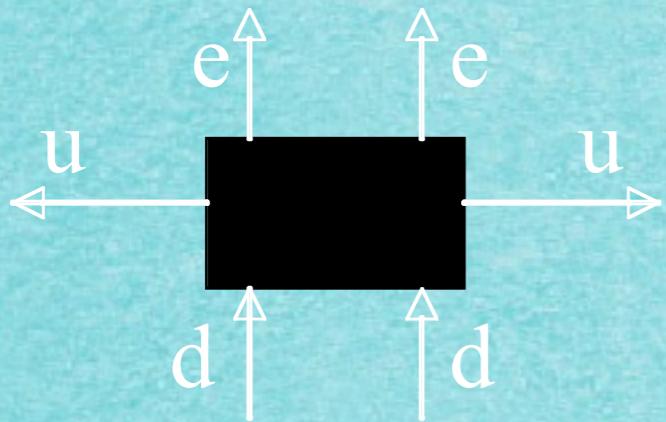
The present limit is given by  
[H.V.Klapdor-Kleingrothaus]

$$|\langle m_\nu \rangle| \equiv \left| \sum U_{ei}^2 m_i \right| < 0.2 \text{ eV}$$

# “Black Box” theorem

J. Schechter and J.W.F. Valle, Phys.Rev. D 25, 2951 (1982)

``Any mechanism inducing the  $0\nu\beta\beta$  decay produces an effective Majorana neutrino mass term, which must therefore contribute to this decay.”



**$0\nu\beta\beta$  decay**

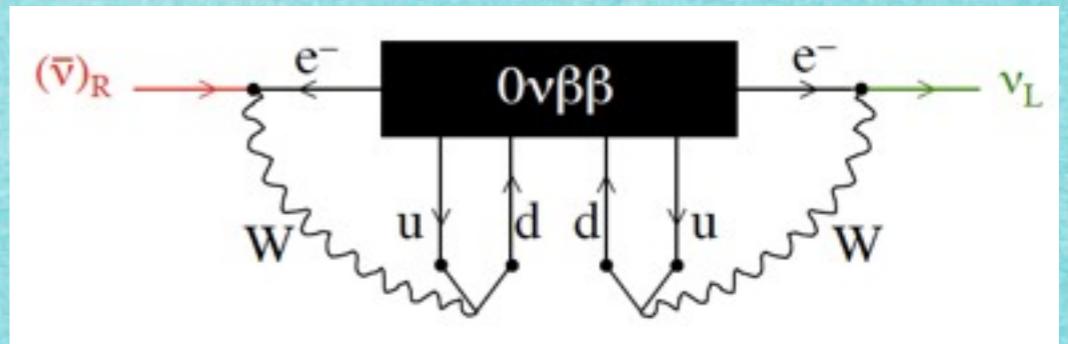
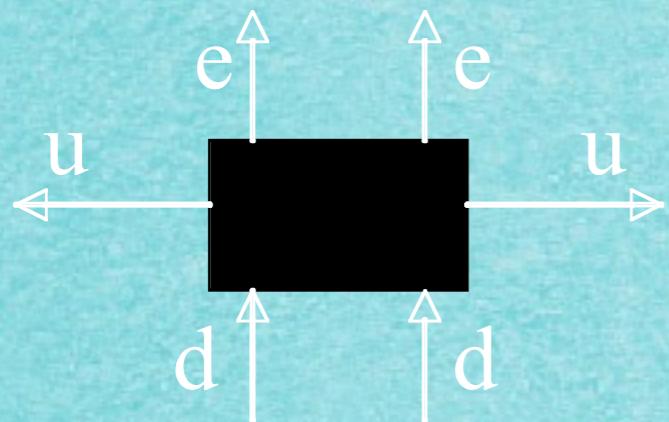


**Majorana neutrino mass**

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Majorana neutrino mass



The theorem does not state if the mechanism for  $0\nu\beta\beta$  from  $m_\nu$  is the dominant one.

In some models, the dominant contributions to  $0\nu\beta\beta$  are generated without directly involving  $\nu_M$ .

# In the SM:

	$SU(3) \otimes SU(2) \otimes U(1)$
$L_a = (\nu_a, l_a)^T$	(1, 2, -1)
$e_a^c$	(1, 1, 2)
$Q_a = (u_a, d_a)^T$	(3, 2, 1/3)
$u_a^c$	( $\bar{3}$ , 1, -4/3)
$d_a^c$	( $\bar{3}$ , 1, 2/3)
$\Phi$	(1, 2, 1)

Table 1: Matter and scalar multiplets of the Standard Model (SM)

- A  $SU(2)$  doublet fundamental scalar Higgs field is employed to give masses to **BOTH** the  $SU(2) \times U(1)$  gauge bosons and fermions.

$$H = \begin{pmatrix} 0 \\ \frac{v+h^0}{\sqrt{2}} \end{pmatrix}, \quad v = 247 \text{ GeV}$$

- Higgs fermion interaction

$$y_e (\overline{\nu_{eL}} \quad \overline{e_L}) \begin{pmatrix} 0 \\ \frac{v+h^0}{\sqrt{2}} \end{pmatrix} e_R + h.c. \rightarrow \frac{y_e v}{\sqrt{2}} \bar{e} e + \frac{y_e v}{\sqrt{2}} \bar{e} e h^0$$

- Fermion mass  $m_f = \frac{y_f v}{\sqrt{2}}$  and  $\bar{f} f H$  coupling is proportional to fermion mass



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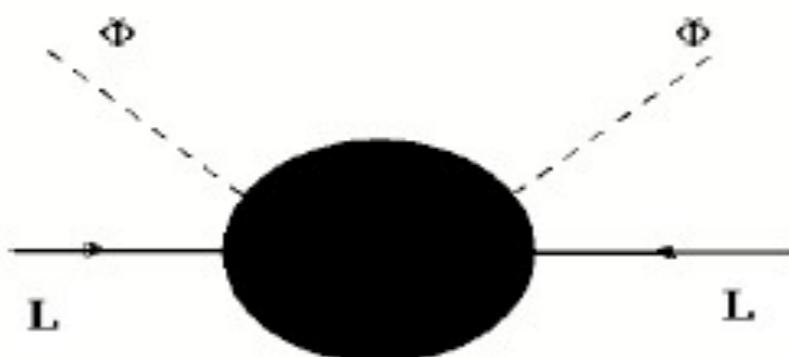
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Dimension five operator responsible for neutrino mass

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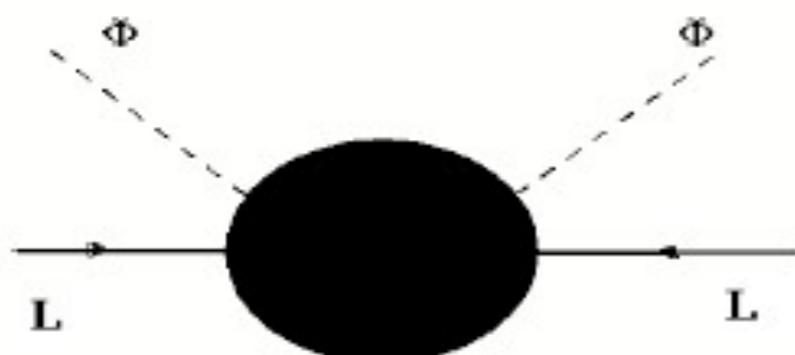
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↓  
SSB

$$m_\nu = \lambda_0 \frac{\langle \Phi \rangle^2}{M_X}, \quad (\text{Majorana})$$

For  $\lambda_0 \sim 1$ ,  $\langle \Phi \rangle \sim 100$  GeV,  $M_X \sim M_P \rightarrow m_\nu \sim 10^{-6}$  eV (too small)

**BSM:** (a) If the right handed neutrinos  $\nu_R$  exist:  $\nu_R = (1,1,0)$

$$\mathcal{L}_Y = Y_\nu \bar{L} \Phi \nu_R + h.c. \Rightarrow m_\nu^D = Y_\nu \langle \Phi \rangle$$

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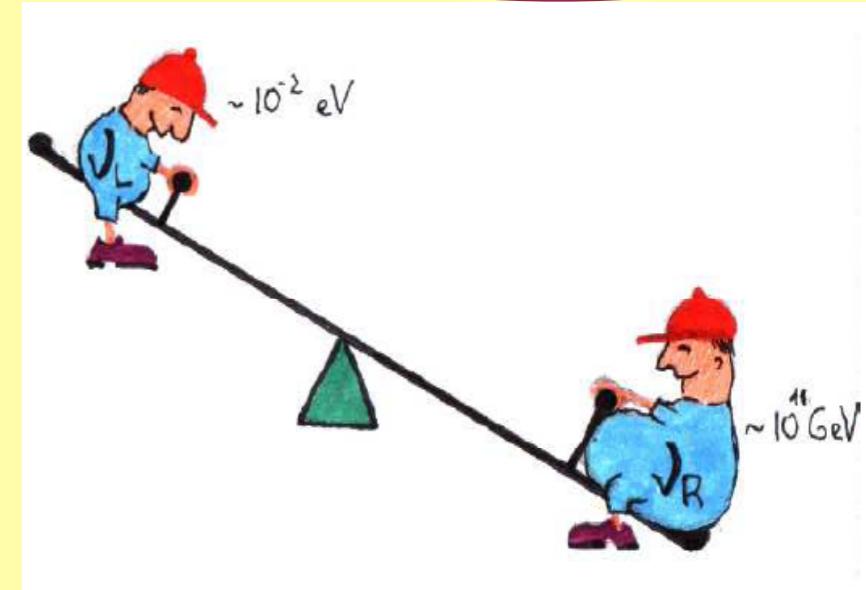
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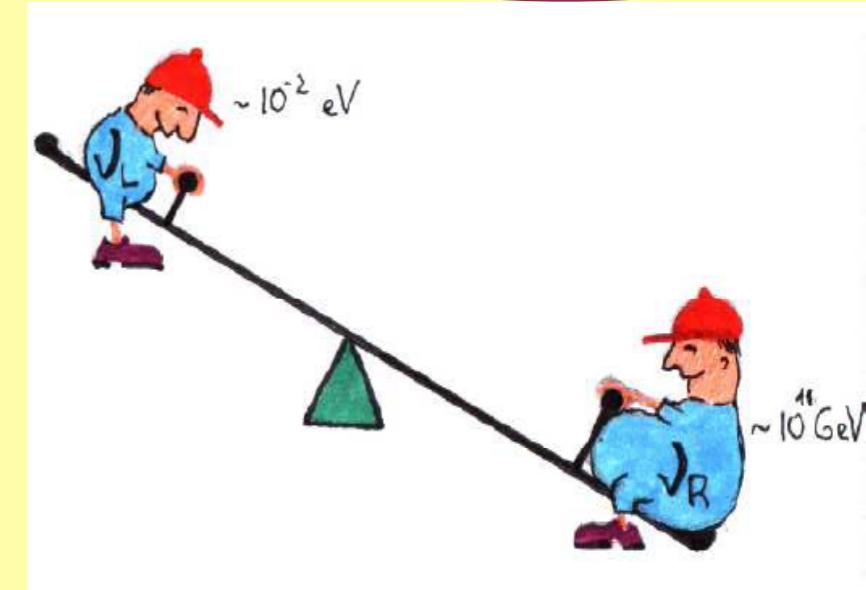
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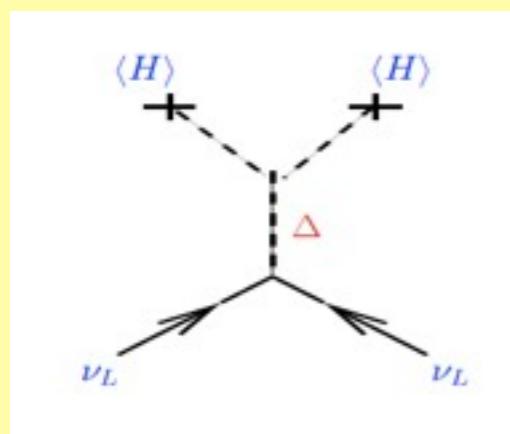
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(c) Without  $\nu_R$ :

Majorana : tree level



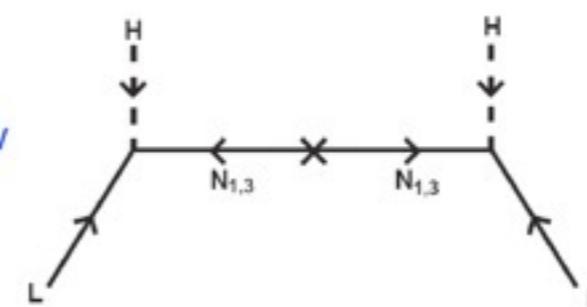
Type II seesaw

Schechter & Valle, 1980, 1982  
Cheng & Li, 1980  
Mohapatra, Senjanovic, 1981  
...

Minkowski 1977; ...

Foot, Lew, He, Joshi 1989

Type (I,III) seesaw



$N_1 : (1, 1, 0)$   
 $N_3 : (1, 3, 0)$

$$\Delta \equiv \begin{pmatrix} H^- & -\sqrt{2} H^0 \\ \sqrt{2} H^{--} & -H^- \end{pmatrix} = (1, 3, 2) \quad \text{scalar triplet}$$

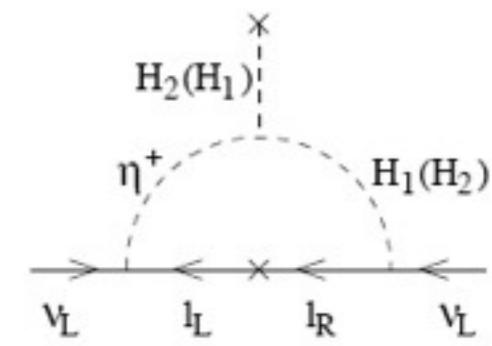
$$M_\nu = \sqrt{2} Y_\Delta \langle \Delta \rangle = Y_\Delta \frac{\mu_\Delta v^2}{M_\Delta^2}$$

# • Majorana : loop level

1980

- Zee model (with charged scalar singlet and additional scalar doublets).

$$l^T \hat{f}_i \sigma_2 l \eta^+ + \sum_{i=1,2} \bar{l} \hat{f}_i e H_i,$$



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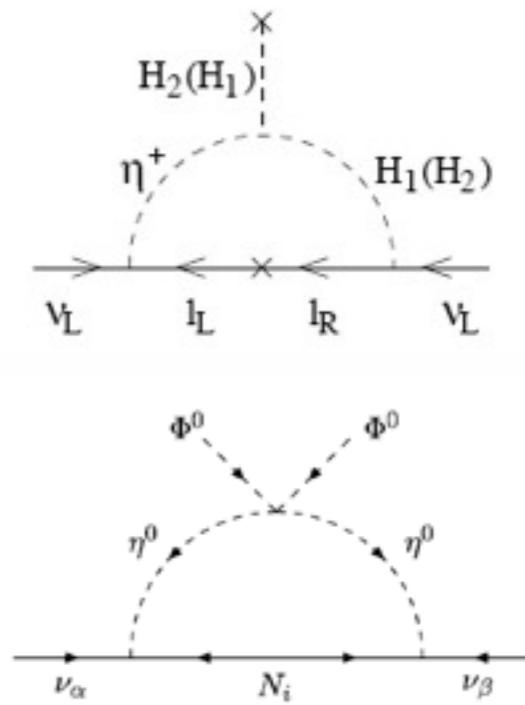
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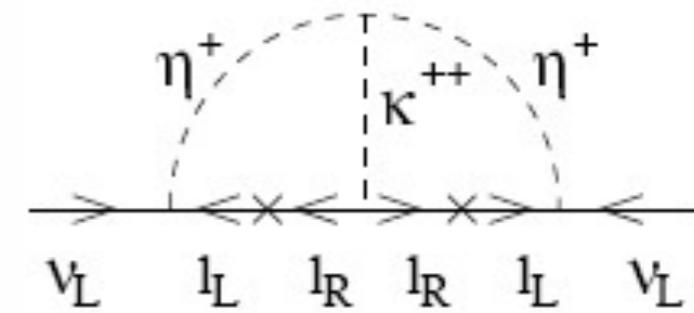
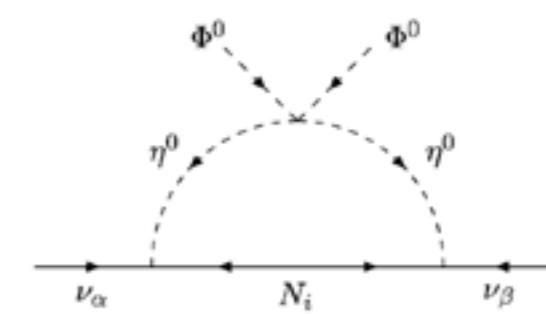
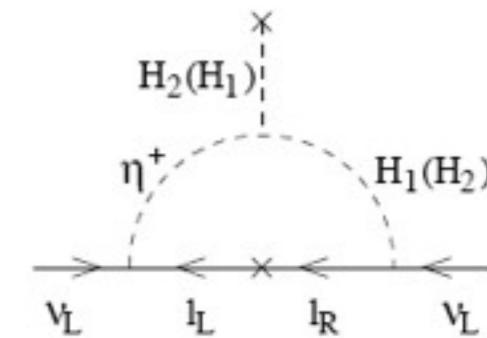
- Ma model (with fermion singlet and additional scalar doublet).

[1986](#)      [1988](#)

- Zee-Babu model (with doubly and singly charged scalars).

$$l^T \hat{f} l \eta^+ + l_R^T \hat{h} l_R k^{++}$$

[2006](#)



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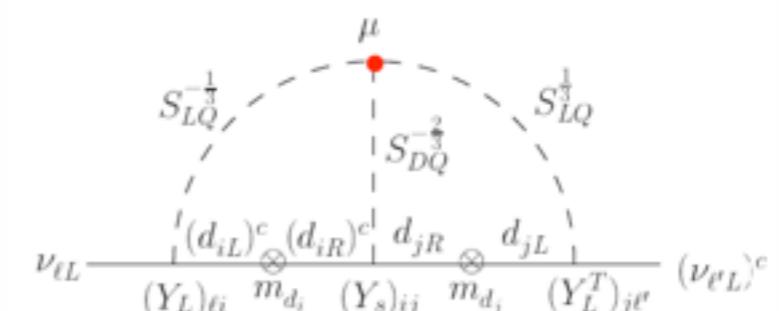
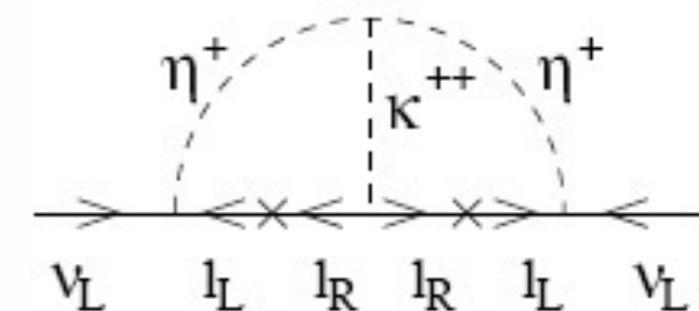
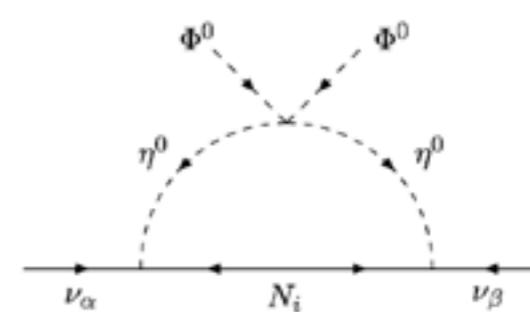
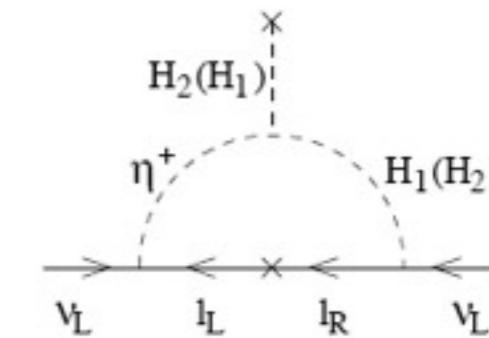
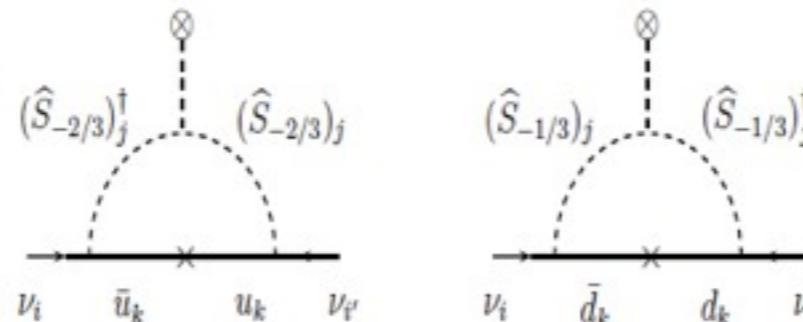
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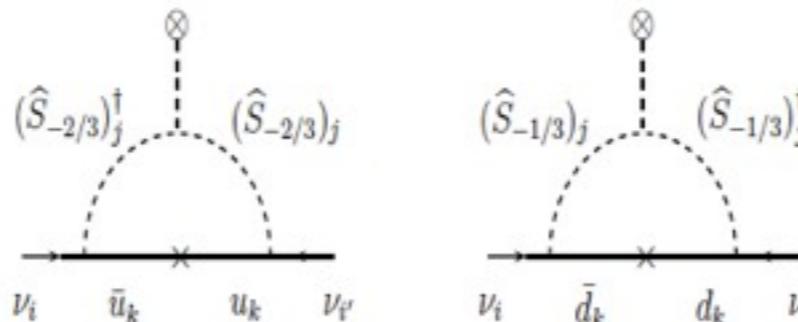
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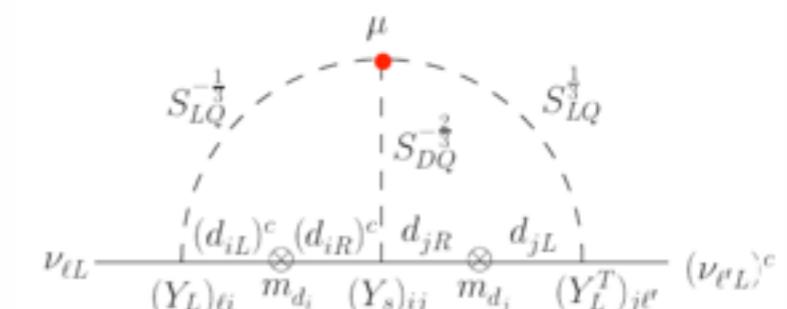
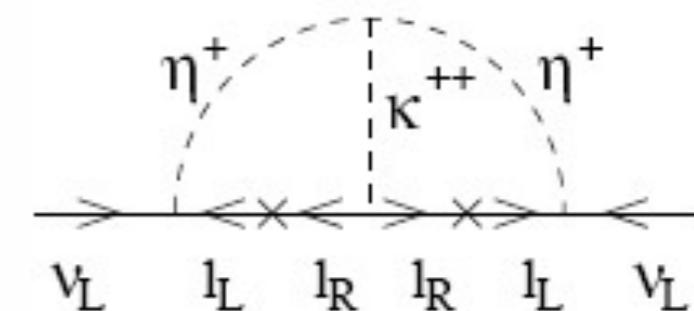
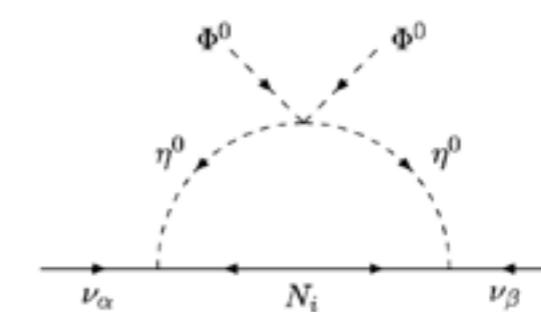
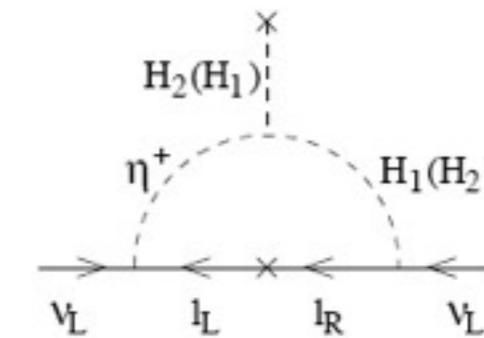
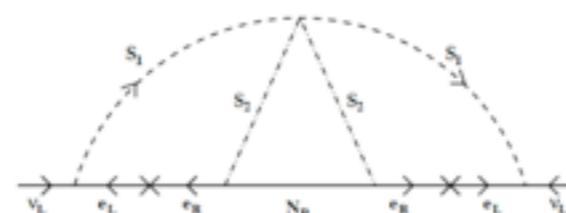
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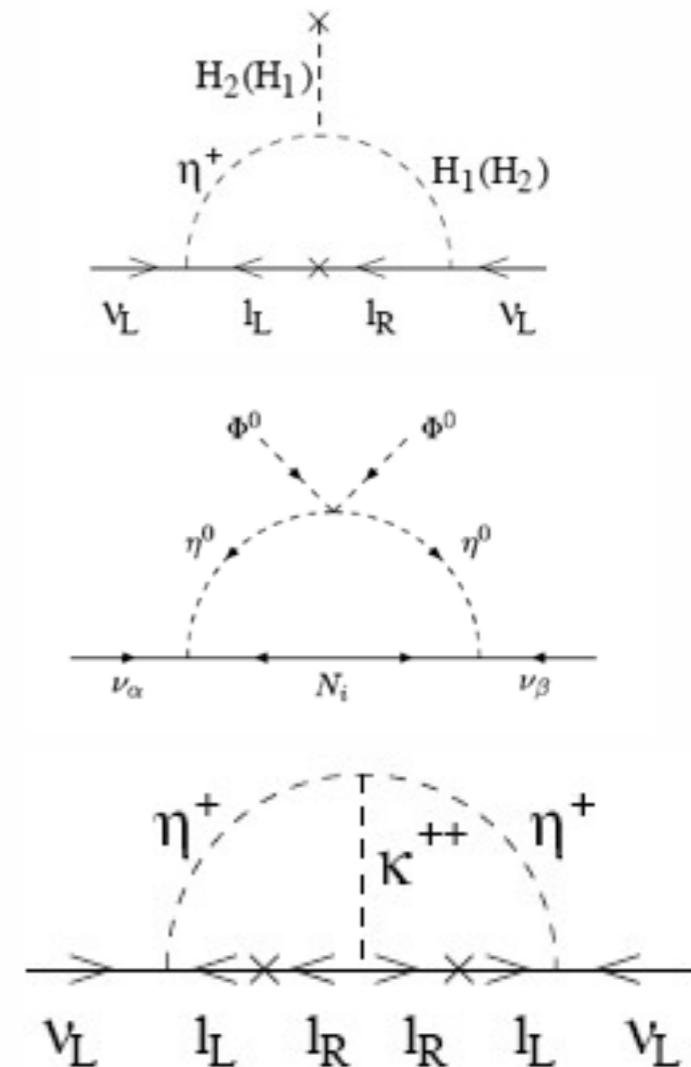
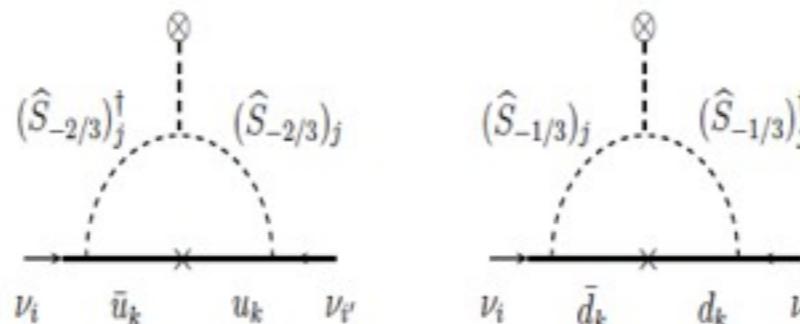
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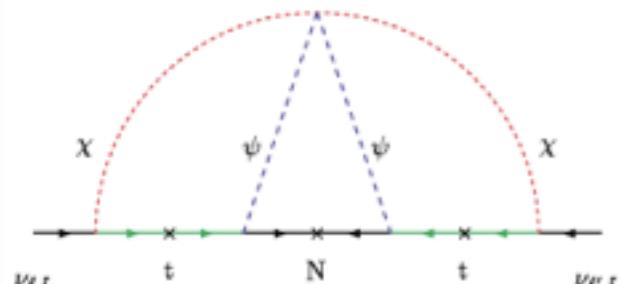
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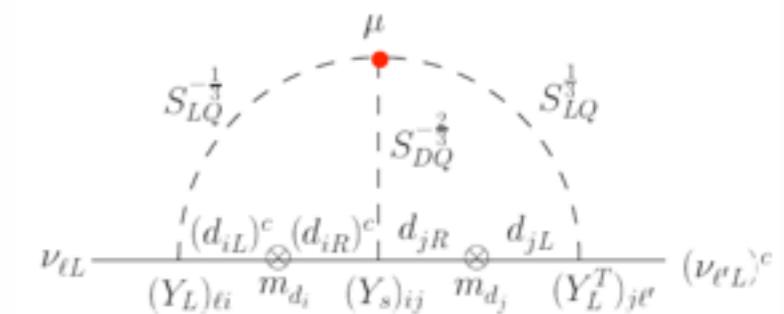
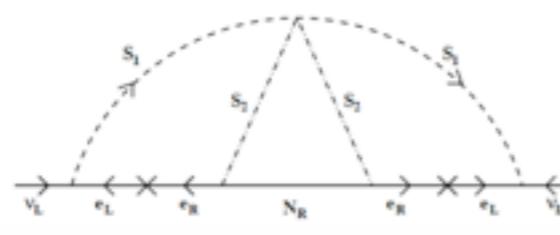


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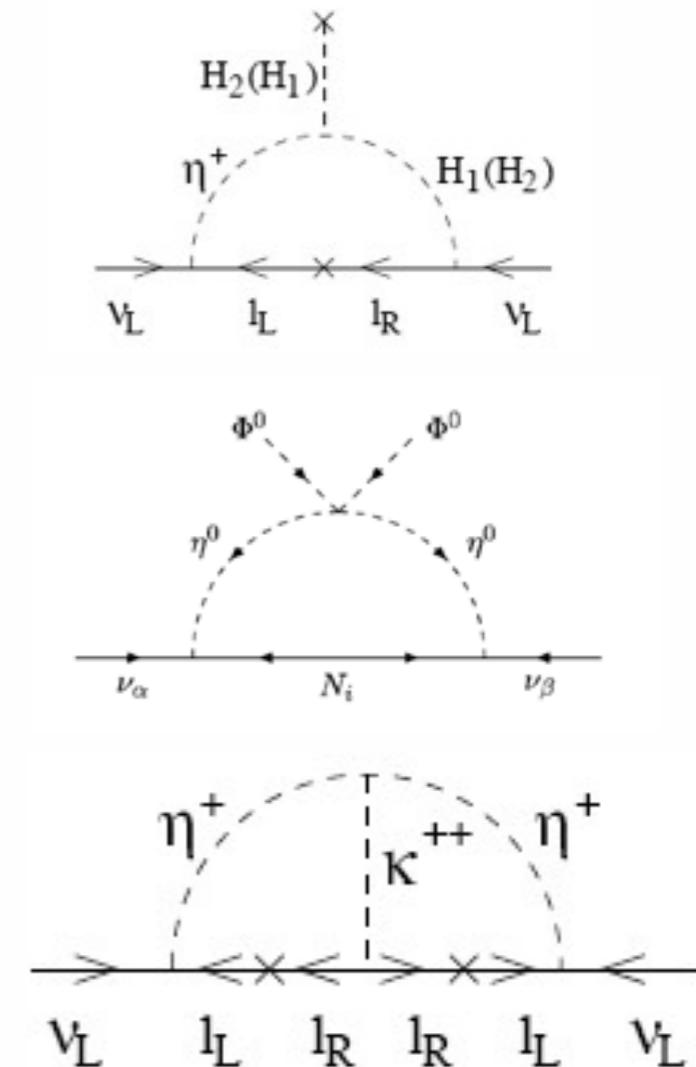
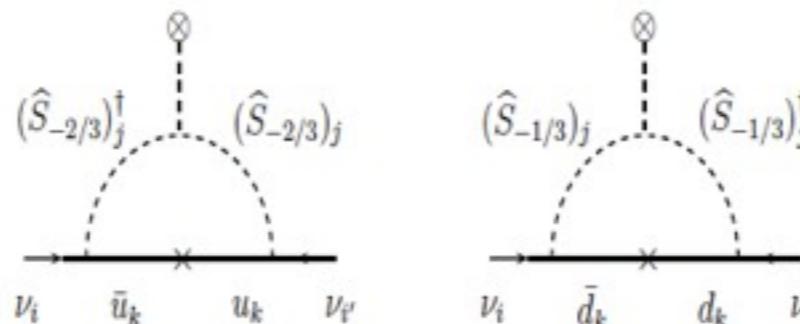
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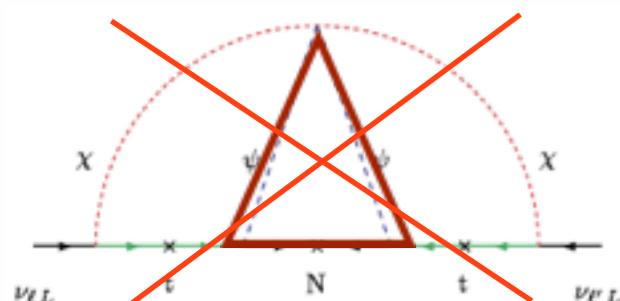
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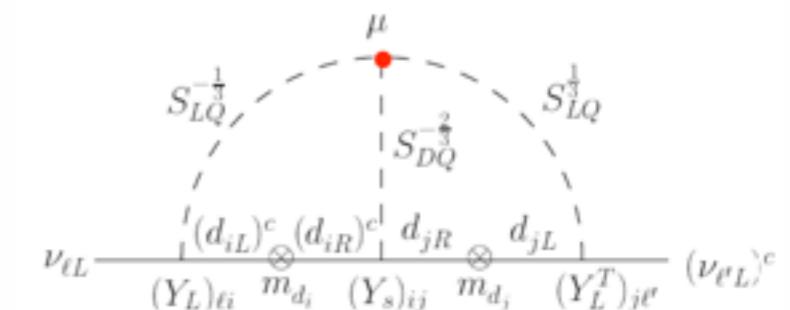
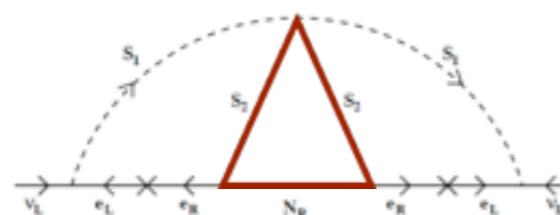


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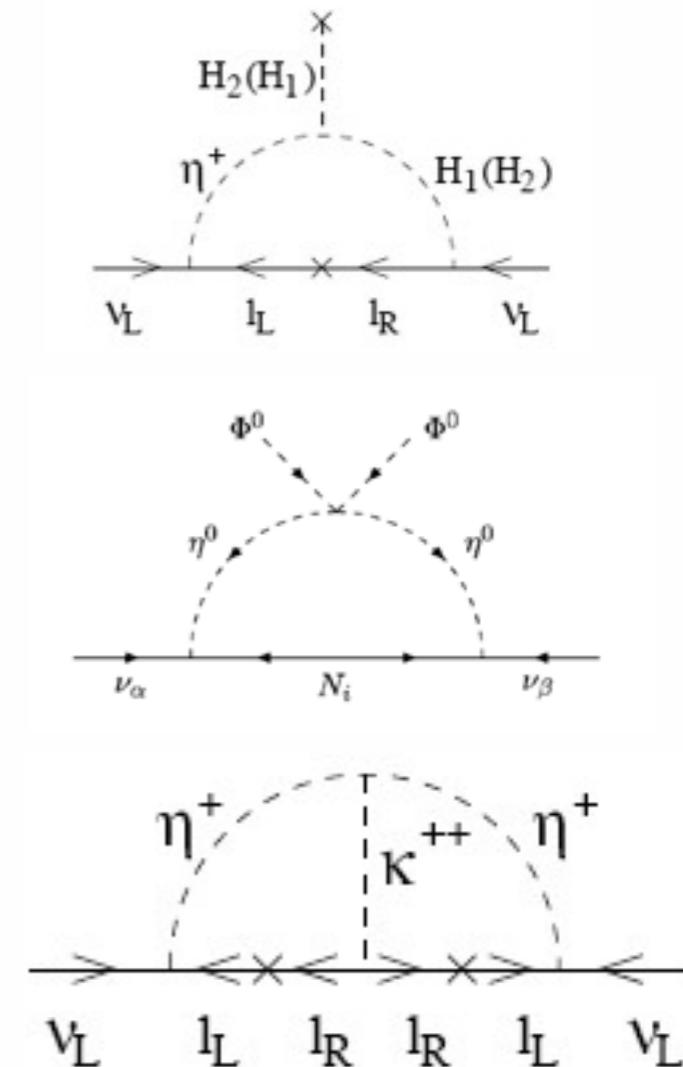
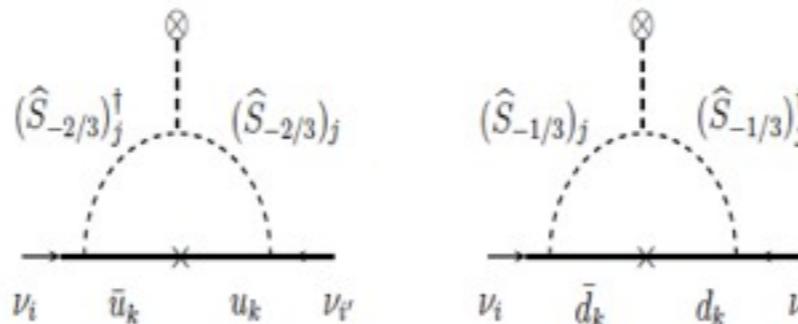
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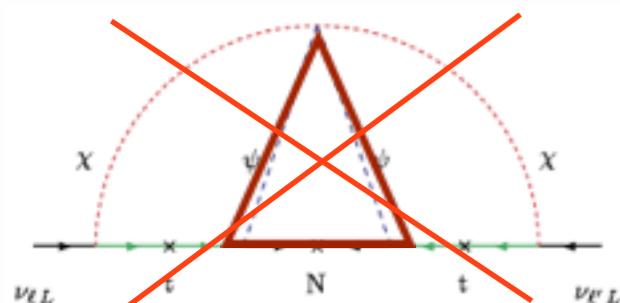
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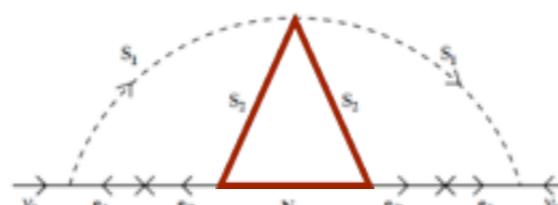


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**Suppressed  $0\nu\beta\beta$  in these models!**

$$\nu_{\ell L} - \frac{(d_{iL})^c (d_{iR})^c}{(Y_L)_{\ell i}} \frac{d_{jR}}{m_{d_i}} \frac{\otimes d_{jL}}{(Y_s)_{ij}} \frac{1}{m_{d_j}} \frac{(Y_L^T)_{j\ell}}{(Y_L^T)_{j\ell}} (\nu_{\ell' L})^c$$

- Models with Majorana Neutrinos:

C.S.Chen+CQG+J.N.Ng,  
PRD75,053004(07)

	$SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$
$L_a = (\nu_a, l_a)^T_L$	(1, 2, -1)
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$\Phi$	(1, 2, 1)

Table 1: Matter and scalar multiplets of the Standard Model (SM)

**No  $\nu_R$  added**

New scalars: a triplet  $T$  (1,3,2) + a singlet  $\Psi$  (1,1,4)

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New Yukawa term:

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No Yukawa coupling for the triplet:

~~LLT~~

Forbidden by some symmetry\*

**\*Symmetry:** two Higgs doublets ( $\Phi_1$  and  $\Phi_2$ )  
with  $Z_2$ -symmetry or T-parity

CS.Chen,CQG,PRD82,105004(2010)

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**Without Symmetry:**

$$\xi(1,N,2) + \Psi(1,1,4)$$

~~LL~~ $\xi$  if  $N > 3$

We will consider higher dimensional multiplets so that  
**NO LL-like term** is allowed in the Yukawa interactions.

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e.g. for  $N=5$  

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• The scalar potential reads

$$V(\Phi, \xi, \Psi) = -\mu_\Phi^2 |\Phi|^2 + \lambda_\Phi |\Phi|^4 + \mu_\xi^2 |\xi|^2 + \lambda_\xi^\alpha |\xi|_\alpha^4 + \mu_\Psi^2 |\Psi|^2 + \lambda_\Psi |\Psi|^4 + \lambda_{\Phi\xi}^\beta (|\Phi|^2 |\xi|^2)_\beta + \lambda_{\Phi\Psi} |\Phi|^2 |\Psi|^2 + \lambda_{\xi\Psi} |\xi|^2 |\Psi|^2 + [\mu \xi \xi \Psi + h.c.]$$

No  $N=4, 6, 8, 10, \dots$ , even dimensions  
due to their antisymmetric products

$N=5, 7, \dots$ , odd dimensions

## Constraints on the models:

VEVs:  $\langle \phi^0 \rangle \equiv \frac{v}{\sqrt{2}}$  and  $\langle T^0 \rangle \equiv \frac{v_T}{\sqrt{2}}$ .

$$M_W^2 = \frac{g^2}{4} (v^2 + 2v_T^2), \quad M_Z^2 = \frac{g^2}{4 \cos^2 \theta_W} (v^2 + 4v_T^2),$$

$$\rho = 1.0002^{+0.0007}_{-0.0004} \quad \rightarrow \quad v_T < 4.41 \text{ GeV}$$

Two doubly charged scalars:

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Mass eigenstates:

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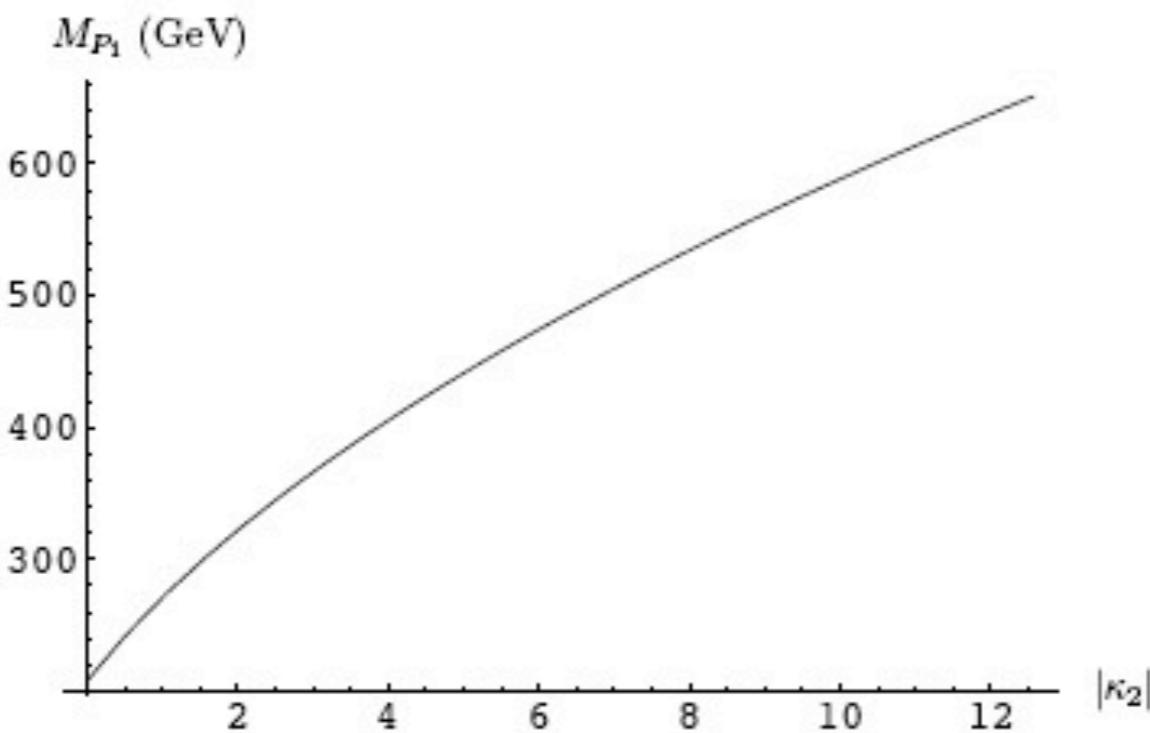
$$\begin{pmatrix} P_1^{\pm\pm} \\ P_2^{\pm\pm} \end{pmatrix} = \begin{pmatrix} \cos \delta & \sin \delta \\ -\sin \delta & \cos \delta \end{pmatrix} \begin{pmatrix} T^{\pm\pm} \\ \Psi^{\pm\pm} \end{pmatrix}$$

$$\sin 2\delta = \left[ 1 + \left( \frac{2m^2 + (2\lambda'_T + \rho)v_T^2}{2\lambda v^2} + \frac{\kappa_2 + \kappa_\Psi}{2\lambda} - \frac{\omega}{\lambda} \right)^2 \right]^{-\frac{1}{2}}$$

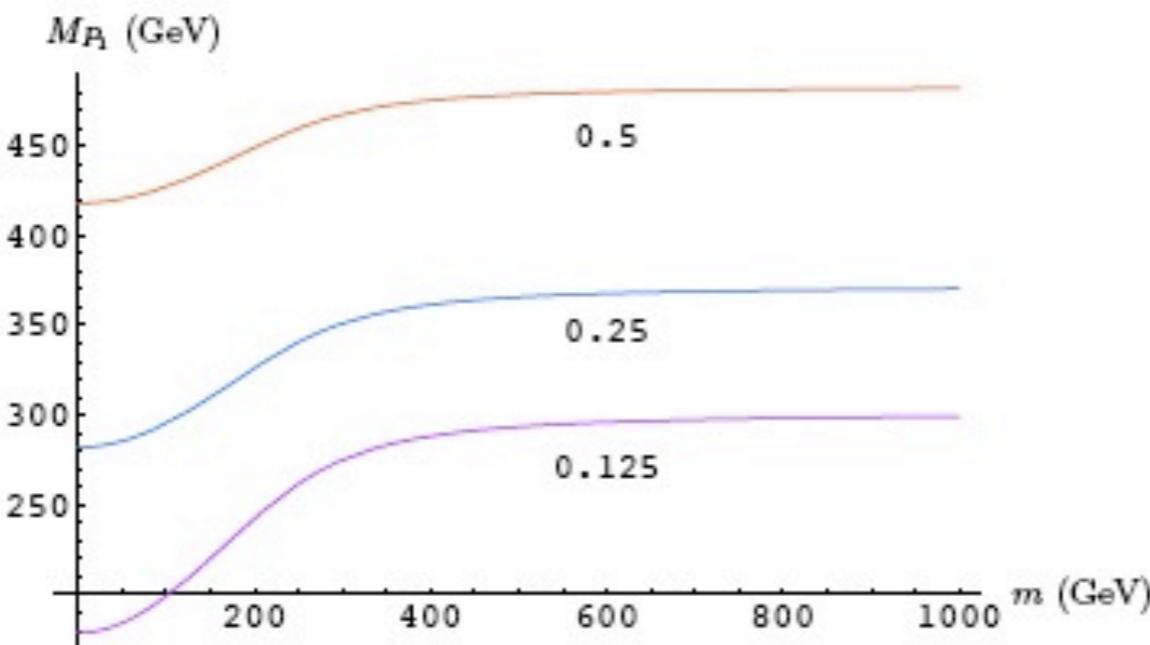
$$M_{P_{1,2}}^2 = \frac{1}{2} \left[ a + c \mp \sqrt{4b^2 + (c-a)^2} \right]$$

$$\omega \equiv \frac{M}{\sqrt{2}v_T}$$

$$a = \frac{1}{2}(2\omega - \kappa_2)v^2 - \lambda'_T v_T^2, \quad b = \frac{1}{2}\lambda v^2, \quad c = m^2 + \frac{1}{2}(\kappa_\Psi v^2 + \rho v_T^2).$$



**Figure 1:** Maximum value of  $M_{P_1}$  for  $v_T = M = 4 \text{ GeV}$ , and  $|\lambda'_T|$  set to  $4\pi$ .



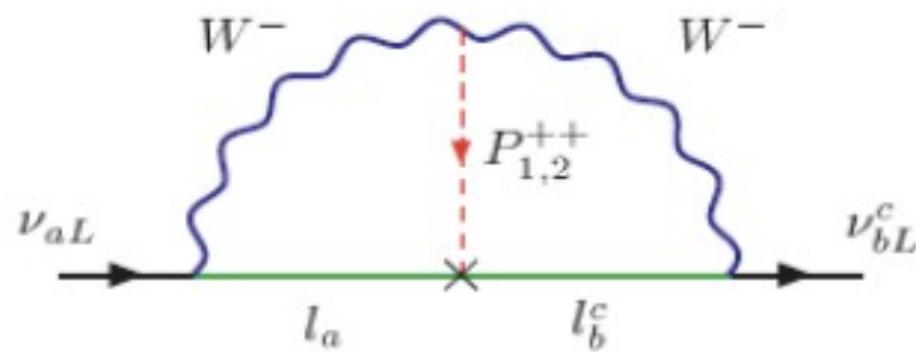
**Figure 2:**  $M_{P_1}$  as a function of  $m$  for  $|\kappa_2| = 0.5, 0.25, 0.125$  in units of  $4\pi$ , with  $v_T = M = 4 \text{ GeV}$  and  $\lambda = -\lambda'_T = 1$ .

The  $P_1$  state is well within the reach of the LHC;  
 $P_2$  will be too heavy to be of interest to the LHC.

Current LHC limit:  
 200~400 GeV

- TeV Phenomenology:

### (i) $\nu$ mass generations:

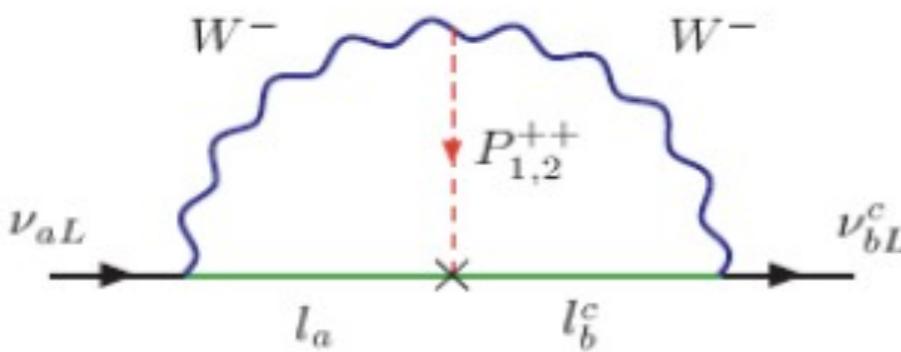


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$$(m_\nu)_{ab} = \frac{1}{\sqrt{2}} g^4 m_a m_b v_T Y_{ab} \sin(2\delta) [I(M_W^2, M_{P_1}^2, m_a, m_b) - I(M_W^2, M_{P_2}^2, m_a, m_b)]$$

$$I(M_W^2, M_{P_i}^2, m_a^2, m_b^2) = \int \frac{d^4 q}{(2\pi)^4} \int \frac{d^4 k}{(2\pi)^4} \frac{1}{k^2 - m_a^2} \frac{1}{k^2 - M_W^2} \frac{1}{q^2 - M_W^2} \frac{1}{q^2 - m_b^2} \frac{1}{(k-q)^2 - M_{P_i}^2}.$$

$$M_{P_{1,2}} > M_W$$

$$I(M_W^2, M_{P_i}^2, 0, 0) \sim \frac{1}{(4\pi)^4} \frac{1}{M_{P_i}^2} \log^2 \left( \frac{M_W^2}{M_{P_i}^2} \right)$$

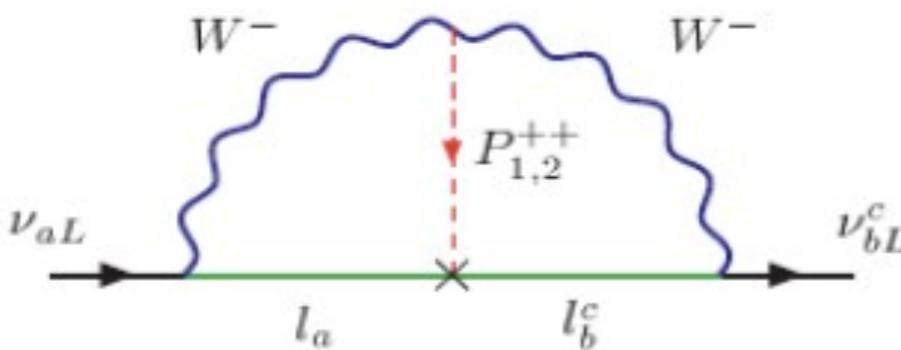
$$\begin{aligned} m_\nu &= \tilde{f}(M_{P_1}, M_{P_2}) \times \begin{pmatrix} m_e^2 Y_{ee} & m_e m_\mu Y_{e\mu} & m_e m_\tau Y_{e\tau} \\ m_e m_\mu Y_{e\mu} & m_\mu^2 Y_{\mu\mu} & m_\tau m_\mu Y_{\mu\tau} \\ m_e m_\tau Y_{e\tau} & m_\tau m_\mu Y_{\mu\tau} & m_\tau^2 Y_{\tau\tau} \end{pmatrix} \\ &= f(M_{P_1}, M_{P_2}) \times \begin{pmatrix} 2.6 \times 10^{-7} Y_{ee} & 5.4 \times 10^{-5} Y_{e\mu} & 9.1 \times 10^{-4} Y_{e\tau} \\ 5.4 \times 10^{-5} Y_{e\mu} & 1.1 \times 10^{-2} Y_{\mu\mu} & 0.19 Y_{\mu\tau} \\ 9.1 \times 10^{-4} Y_{e\tau} & 0.19 Y_{\mu\tau} & 3.17 Y_{\tau\tau} \end{pmatrix} \end{aligned}$$

$$\tilde{f}(M_{P_1}, M_{P_2}) = \frac{\sqrt{2} g^4 v_T \sin(2\delta)}{128\pi^4} \left[ \frac{1}{M_{P_1}^2} \log^2 \left( \frac{M_W}{M_{P_1}} \right) - \frac{1}{M_{P_2}^2} \log^2 \left( \frac{M_W}{M_{P_2}} \right) \right]$$

$$f = \tilde{f} \times (1 \text{ GeV}^2)$$

# • TeV Phenomenology:

## (i) $\nu$ mass generations:



The neutrino masses are generated radiatively at two-loop level

$$a, b = e, \mu, \tau.$$

$$(m_\nu)_{ab} = \frac{1}{\sqrt{2}} g^4 m_a m_b v_T Y_{ab} \sin(2\delta) [I(M_W^2, M_{P_1}^2, m_a, m_b) - I(M_W^2, M_{P_2}^2, m_a, m_b)]$$

$$I(M_W^2, M_{P_i}^2, m_a^2, m_b^2) = \int \frac{d^4 q}{(2\pi)^4} \int \frac{d^4 k}{(2\pi)^4} \frac{1}{k^2 - m_a^2} \frac{1}{k^2 - M_W^2} \frac{1}{q^2 - M_W^2} \frac{1}{q^2 - m_b^2} \frac{1}{(k-q)^2 - M_{P_i}^2}.$$

$$M_{P_{1,2}} > M_W$$

$$I(M_W^2, M_{P_i}^2, 0, 0) \sim \frac{1}{(4\pi)^4} \frac{1}{M_{P_i}^2} \log^2 \left( \frac{M_W^2}{M_{P_i}^2} \right)$$

$$\begin{aligned} m_\nu &= \tilde{f}(M_{P_1}, M_{P_2}) \times \begin{pmatrix} m_e^2 Y_{ee} & m_e m_\mu Y_{e\mu} & m_e m_\tau Y_{e\tau} \\ m_e m_\mu Y_{e\mu} & m_\mu^2 Y_{\mu\mu} & m_\tau m_\mu Y_{\mu\tau} \\ m_e m_\tau Y_{e\tau} & m_\tau m_\mu Y_{\mu\tau} & m_\tau^2 Y_{\tau\tau} \end{pmatrix} \\ &= f(M_{P_1}, M_{P_2}) \times \begin{pmatrix} 2.6 \times 10^{-7} Y_{ee} & 5.4 \times 10^{-5} Y_{e\mu} & 9.1 \times 10^{-4} Y_{e\tau} \\ 5.4 \times 10^{-5} Y_{e\mu} & 1.1 \times 10^{-2} Y_{\mu\mu} & 0.19 Y_{\mu\tau} \\ 9.1 \times 10^{-4} Y_{e\tau} & 0.19 Y_{\mu\tau} & 3.17 Y_{\tau\tau} \end{pmatrix} \end{aligned}$$

**normal hierarchy:**

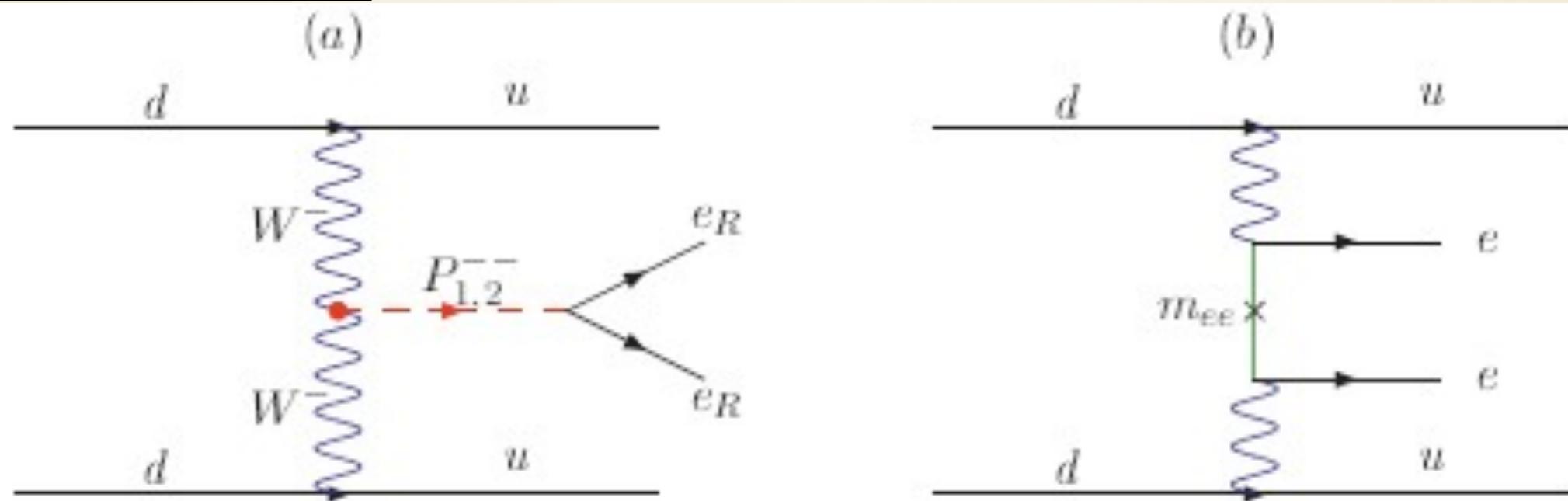
$$\begin{pmatrix} \varepsilon' & \varepsilon & \varepsilon \\ \varepsilon & 1+\eta & 1+\eta \\ \varepsilon & 1+\eta & 1+\eta \end{pmatrix}$$

$$\tilde{f}(M_{P_1}, M_{P_2}) = \frac{\sqrt{2} g^4 v_T \sin(2\delta)}{128\pi^4} \left[ \frac{1}{M_{P_1}^2} \log^2 \left( \frac{M_W}{M_{P_1}} \right) - \frac{1}{M_{P_2}^2} \log^2 \left( \frac{M_W}{M_{P_2}} \right) \right]$$

$$f = \tilde{f} \times (1 \text{ GeV}^2)$$

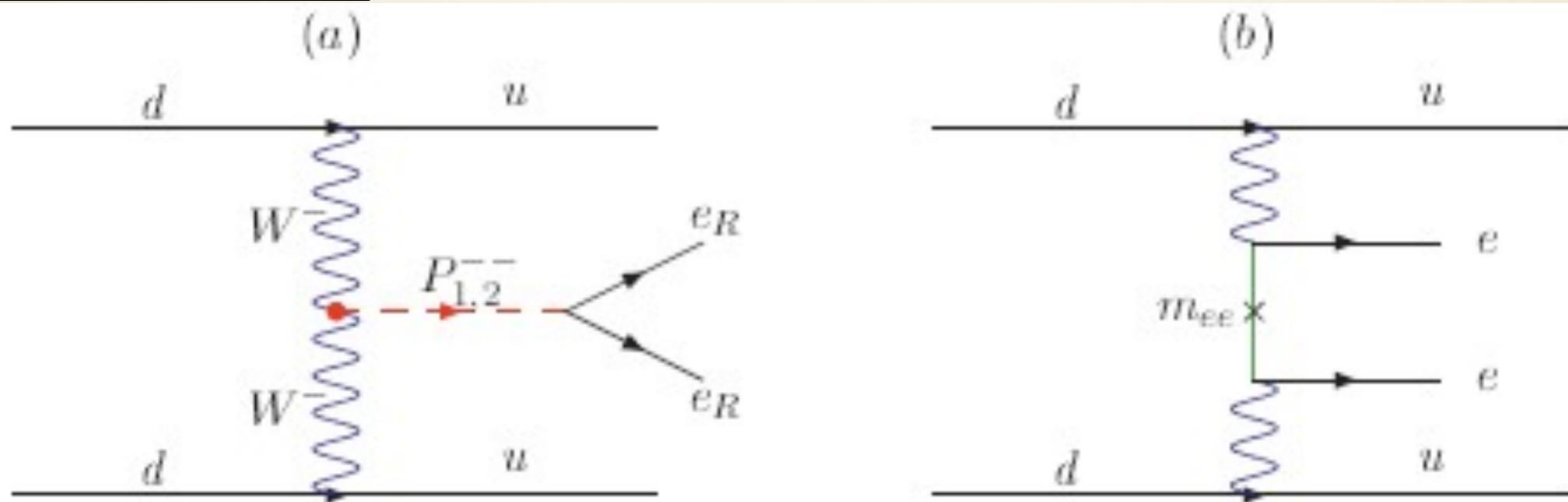
$$\begin{aligned} Y_{ee} < 0.17, & \quad Y_{e\mu} < 0.2, & \quad Y_{e\tau} < 0.2 \\ Y_{\mu\mu} < 3.5, & \quad Y_{\mu\tau} < 0.2, & \quad Y_{\tau\tau} < 0.02 \end{aligned}$$

## (ii) $0\nu\beta\beta$ decays:



**Figure 9:**  $0\nu\beta\beta$  decays via exchange of: (a) doubly charged Higgs and (b) light Majorana neutrinos.

## (ii) $0\nu\beta\beta$ decays:

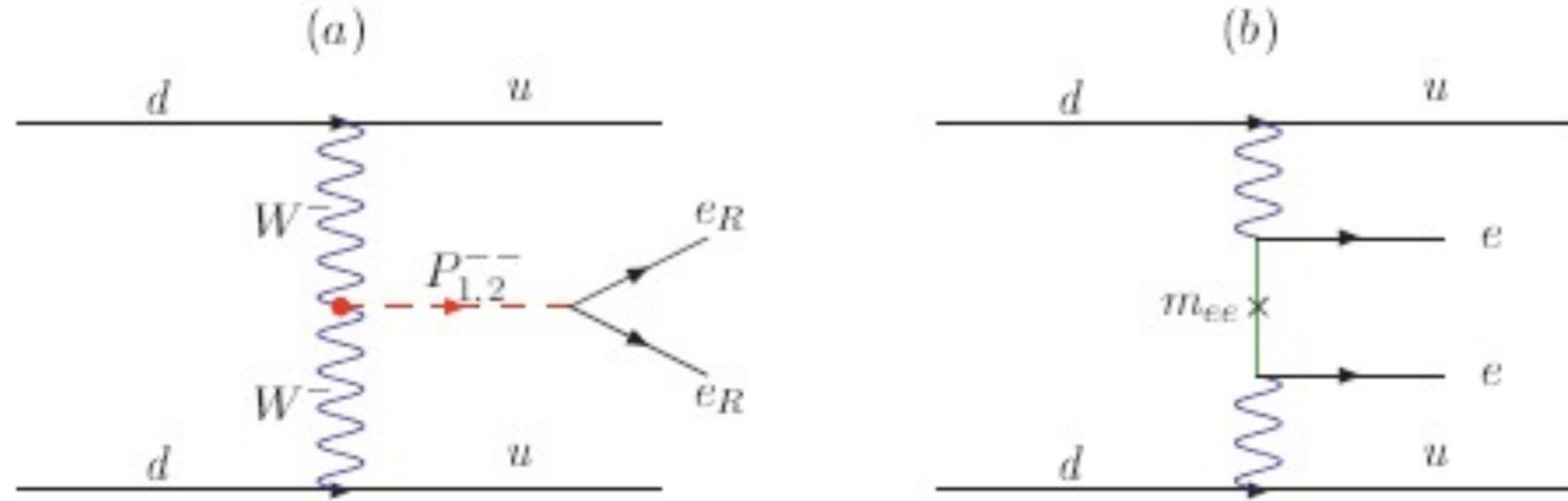


**Figure 9:**  $0\nu\beta\beta$  decays via exchange of: (a) doubly charged Higgs and (b) light Majorana neutrinos.

$$A_{P_{1,2}^{-+}} \sim \frac{g^4 Y_{ee} v_T \sin 2\delta}{16\sqrt{2} M_W^4} \left( \frac{1}{M_{P_1}^2} - \frac{1}{M_{P_2}^2} \right)$$

$$A_\nu \sim \frac{g^4}{M_W^4} \frac{m_{ee}}{\langle p \rangle^2} \quad \langle p \rangle \sim 0.1 \text{ GeV}$$

## (ii) $0\nu\beta\beta$ decays:



**Figure 9:**  $0\nu\beta\beta$  decays via exchange of: (a) doubly charged Higgs and (b) light Majorana neutrinos.

$$A_{P_{1,2}^{--}} \sim \frac{g^4 Y_{ee} v_T \sin 2\delta}{16\sqrt{2} M_W^4} \left( \frac{1}{M_{P_1}^2} - \frac{1}{M_{P_2}^2} \right) \quad \gg \quad A_\nu \sim \frac{g^4}{M_W^4} \frac{m_{ee}}{\langle p \rangle^2} \quad \langle p \rangle \sim 0.1 \text{ GeV}$$

$$A_\nu / A_{P_{1,2}^{--}} \lesssim 10^{-7}$$

The smallness of this ratio is due to the fact that in our model,  $m_{ee}$  is suppressed not only by a two-loop factor, it is also suppressed by the electron mass factor  $(m_e/M_W)^2$  coming from the doubly charged scalar coupling.



**Black box theorem** is irrelevant as  $0\nu\beta\beta$  dominantly arises from the SD contribution

$$\frac{C_{ab}^{(9)}}{\tilde{\Lambda}} \overline{\ell^c}_{R_a} \ell_{R_b} W_\mu^+ W^{+\mu}$$



$$\mathcal{O}^9 \equiv C_{ab}^{(9)} \overline{\ell^c}_{R_a} \ell_{R_b} \left[ (D_\mu H)^T i\sigma_2 H \right]^2$$

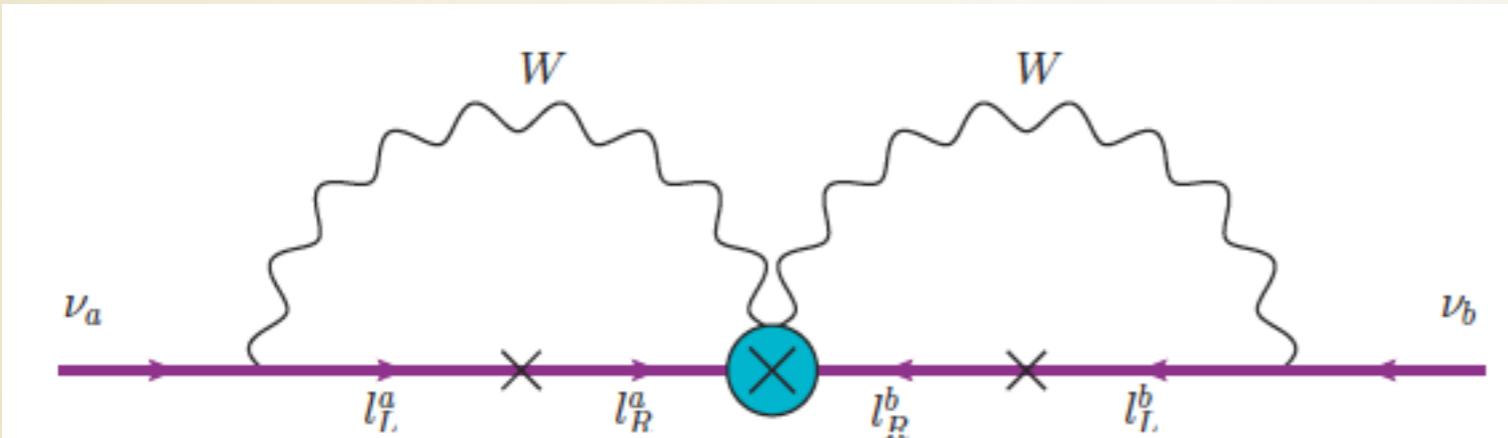
**dimension-9  
L violating O**

$$\frac{C_{ab}^{(9)}}{\tilde{\Lambda}} \overline{\ell^c}_{R_a} \ell_{R_b} W_\mu^+ W^{+\mu}$$



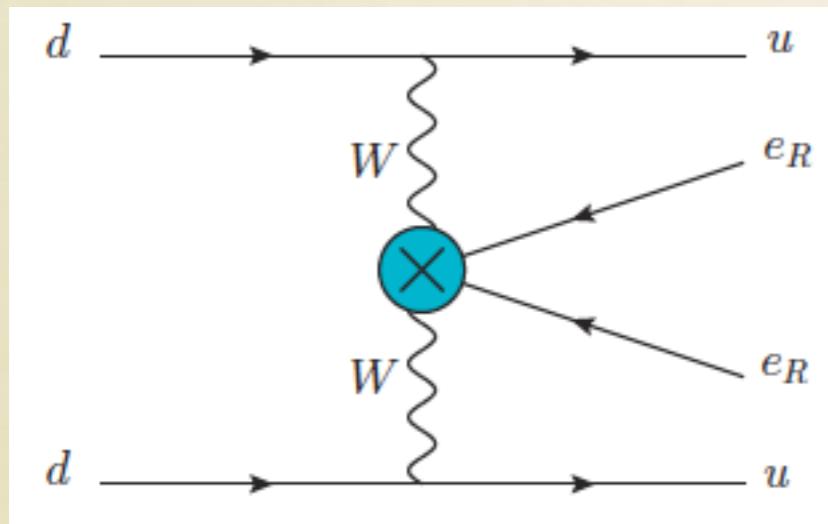
$$\mathcal{O}^9 \equiv C_{ab}^{(9)} \overline{\ell^c}_{R_a} \ell_{R_b} \left[ (D_\mu H)^T i\sigma_2 H \right]^2$$

**dimension-9  
L violating O**



$$m_{ab}^\nu \sim \left( \frac{1}{16 \pi^2} \right)^n C_{ab}^{(9)} \frac{m_{l_a} m_{l_b}}{\Lambda}$$

**n = # of loops**



$$\mathcal{L}_{0\nu\beta\beta} = \frac{G_F^2}{2m_p} \epsilon_3 J^\mu J_\mu \bar{e}(1 - \gamma_5)e^c$$

$$J^\mu = \bar{u}\gamma^\mu(1 - \gamma_5)d$$

$$\epsilon_3 = -2m_p \mathcal{A}_{0\nu\beta\beta}^{\text{SD}}$$

### (iii) Other physics:

## Multi Charged Scalars

### a. Lepton flavor physics:

1. Muonium anti-muonium conversion  $\mu^+ e^- - \mu^- e^+$   $H_{M\bar{M}} = \frac{Y_{ee} Y_{\mu\mu}}{2 M_{--}^2} \bar{\mu} \gamma^\mu e_R \bar{\mu} \gamma_\mu e_R + h.c.,$
2. Effective  $e^+ e^- \rightarrow l^+ l^-$ ,  $l = e, \mu, \tau$ , contact interactions  $\frac{Y_{ee}^2}{M_{--}^2} \bar{e}_R \gamma^\mu e_R \bar{e}_R \gamma_\mu e_R$
3. Rare  $\mu \rightarrow 3e$  decays and its  $\tau$  counterparts
4. Radiative flavor violating charged leptonic decays  $Br(\mu \rightarrow e\gamma) = \frac{\alpha}{3\pi G_F^2} \sum_{l=e,\mu,\tau} \left( \frac{Y_{l\mu} Y_{le}}{M_{--}^2} \right)^2$

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### b. Doubly charged scalars at the LHC:

#### 1 Production of the doubly charged Higgs

WW fusion processes similar to  $0\nu\beta\beta$  decays

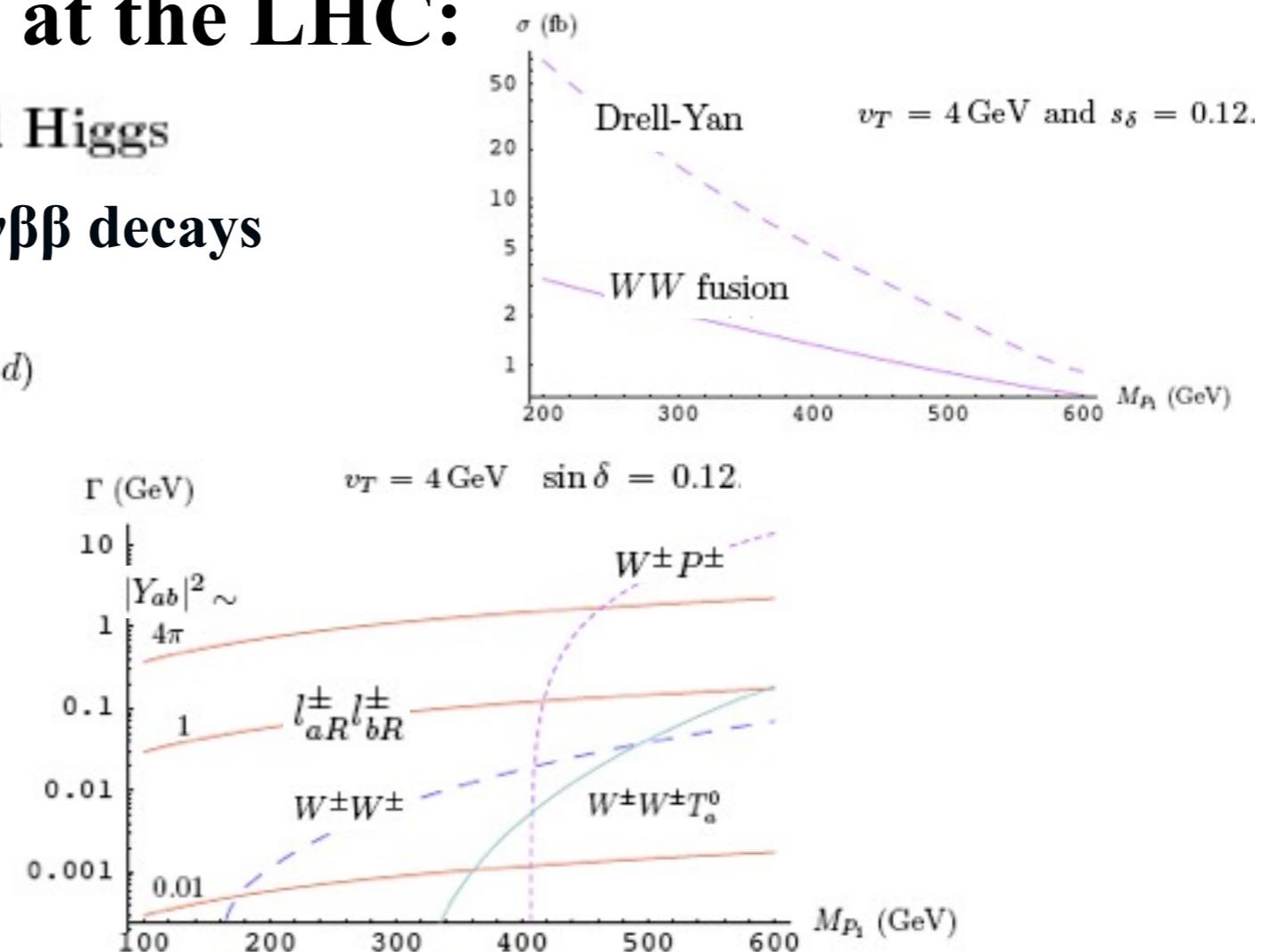
+Drell-Yan Annihilation processes

$$q\bar{q} \rightarrow \gamma^*, Z^* \rightarrow P_1^{++} P_1^{--} \quad (q = u, d)$$

#### 2 The decay of $P_1^{\pm\pm}$

- (1)  $P_1^{\pm\pm} \rightarrow l_{aR}^\pm l_{bR}^\pm$  ( $a, b = e, \mu, \tau$ ),
- (2)  $P_1^{\pm\pm} \rightarrow W^\pm W^\pm$ ,
- (3)  $P_1^{\pm\pm} \rightarrow P^\pm W^\pm$ ,
- (4)  $P_1^{\pm\pm} \rightarrow P^\pm P^\pm$ ,
- (5)  $P_1^{\pm\pm} \rightarrow W^\pm W^\pm X^0$ ,  $X^0 = T_a^0, h^0, P^0$
- (6)  $P_1^{\pm\pm} \rightarrow P^\pm P^\pm X^0$ .

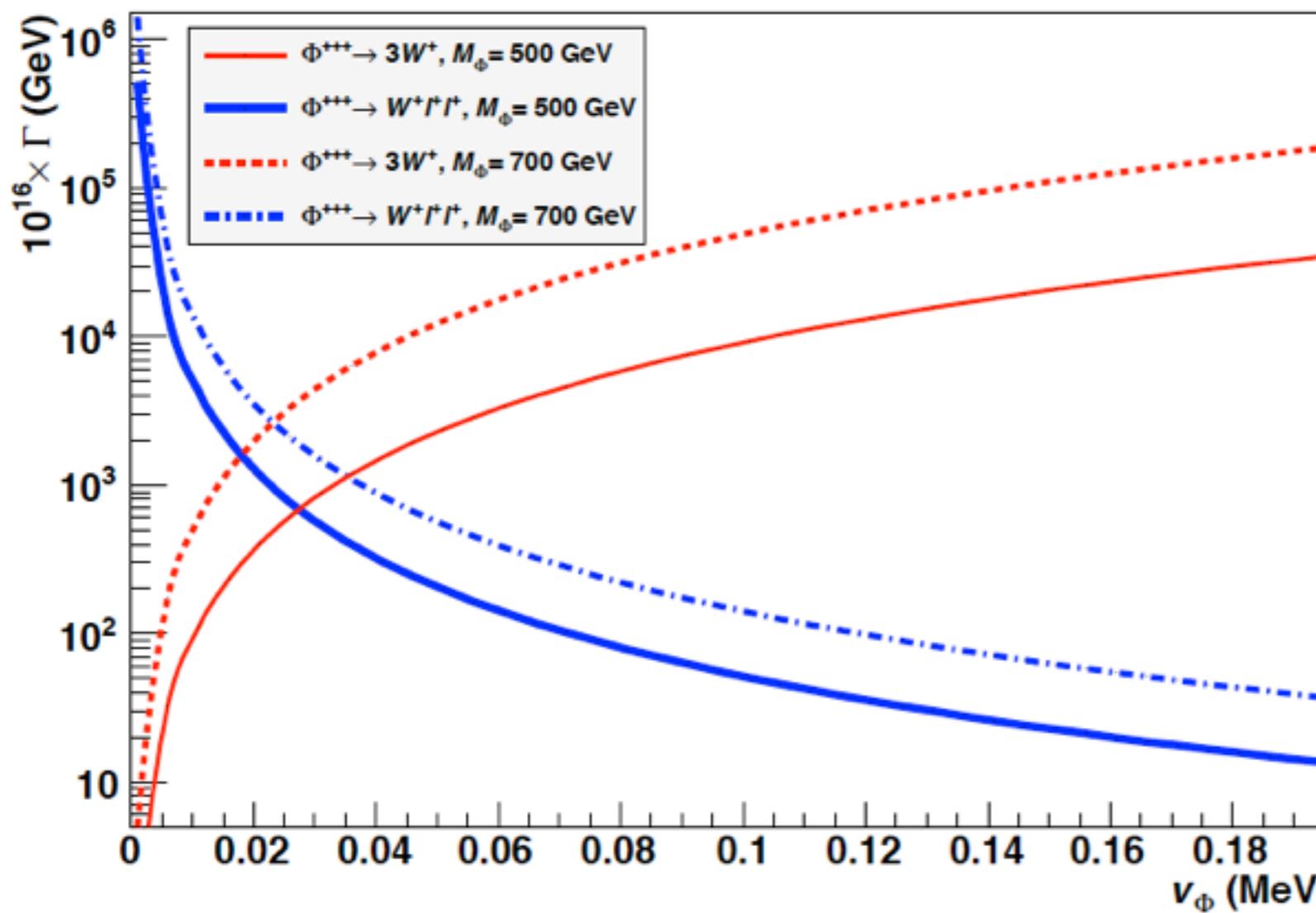
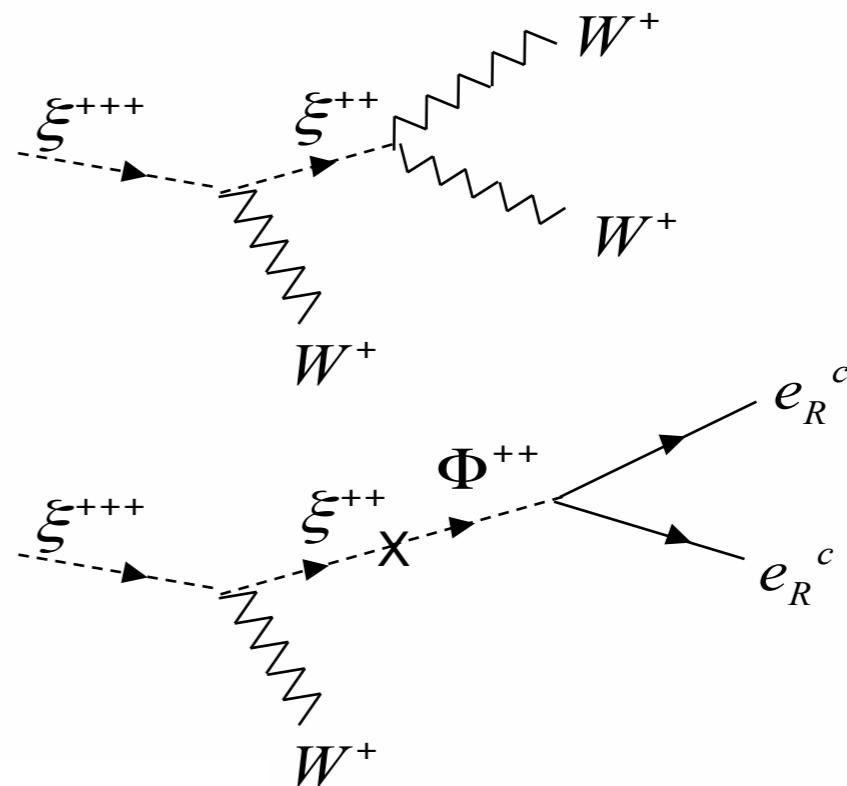
(4) and (6) are not allowed in our model



### c. Triply charged scalar decays:

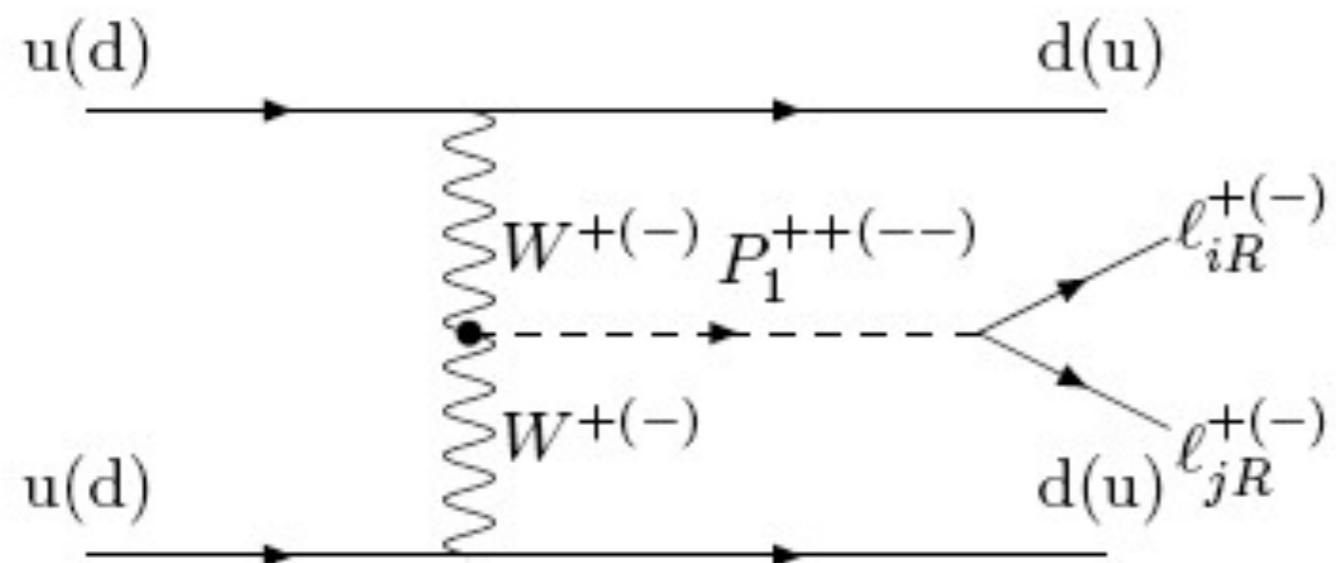
$$\Gamma(\Phi^{+++} \rightarrow 3W) = \frac{3g^6}{2048\pi^3} \frac{v_\Phi^2 M_\Phi^5}{m_W^6}$$

$$\Gamma(\Phi^{+++} \rightarrow W^+ \ell^+ \ell^+) = \frac{g^2}{6144\pi^3} \frac{M_\Phi \sum_i m_i^2}{v_\Phi^2}$$



## d. Same-sign single dilepton signatures:

$$pp \rightarrow \ell_i^\pm \ell_j^\pm X \xrightarrow{JJ}$$



*Chen, CQG, Zhuridov,  
Eur.Phys.J.C60,119(2009)*

$$\frac{d\sigma_{\pm}^{pp}}{d \cos \theta} = A (\lambda_1^{ij})^2 H_{\pm}^{pp}$$

$$A = \frac{G_F^4 M_W^6}{2^7 \pi^5} = 50 \text{ ab}, \quad \lambda_1^{ij} = \sqrt{2 - \delta_{ij}} |Y_{ij}| c_\delta s_\delta,$$

$$H_{\pm}^{pp} = \left( \frac{v_T}{M_W} \right)^2 \int_{z_0}^1 \frac{dz}{z} \int_z^1 \frac{dy}{y} \int_y^1 \frac{dx}{x} p_{\pm}(x, xs) p_{\pm}\left(\frac{y}{x}, \frac{y}{x}s\right) l\left(\frac{z}{y}\right) h\left(\frac{s}{M_{P_1}^2} z\right)$$

### Remarks:

(a) In our model, the final state charged leptons are right-handed. Hence, in principle, helicity measurements can be used to distinguish between our model and those whose doubly charged Higgs coupling only to left-handed leptons.

(b)  $P_1^{\pm\pm}$  will directly produce spectacular lepton # violating signals from like-sign dileptons such as  $e\mu$ ,  $e\tau$  and  $\mu\tau$ .

## e. Multi charged scalar contributions to $H \rightarrow \gamma\gamma$ and $H \rightarrow Z\gamma$ :

$$\Gamma(H \rightarrow \gamma\gamma) = \frac{G_F \alpha^2 m_H^3}{128\sqrt{2}\pi^3} \left| \sum_f N_f^c Q_f^2 A_{1/2}(\tau_f) + A_1(\tau_W) + \sum_{I_3} (I_3 + 1)^2 \frac{v}{2} \frac{\mu_s}{m_s^2} A_0(\tau_s) \right|^2,$$

$$\Gamma(H \rightarrow Z\gamma) = \frac{G_F \alpha \alpha_Z m_H^3}{64\sqrt{2}\pi^3} \left| \sum_f N_f^c Q_f^Z Q_f A_{1/2}^Z(\tau_f, \lambda_f) + Q_W^Z A_1^Z(\tau_W, \lambda_W) + \sum_{I_3} Q_s^Z Q_s \frac{v}{2} \frac{\mu_s}{m_s^2} A_0^Z(\tau_s, \lambda_s) \right|^2, \quad -\frac{(n-1)}{2} \leq I_3 \leq \frac{n-1}{2}$$

$\xi = (1, N, 2)$  with  $N=3, 5, \dots$

e.g.  $\xi = (\xi^{+++}, \xi^{++}, \xi^+, \xi^0, \xi^-)^T$  for  $N=5$

$I_3 = (-N+3)/2$  to  $(N+1)/2$

Chen, CQG, Huang, Tsai,  
 $PRD87,077702$  (2013);  
 $PRD87,075019$  (2013)

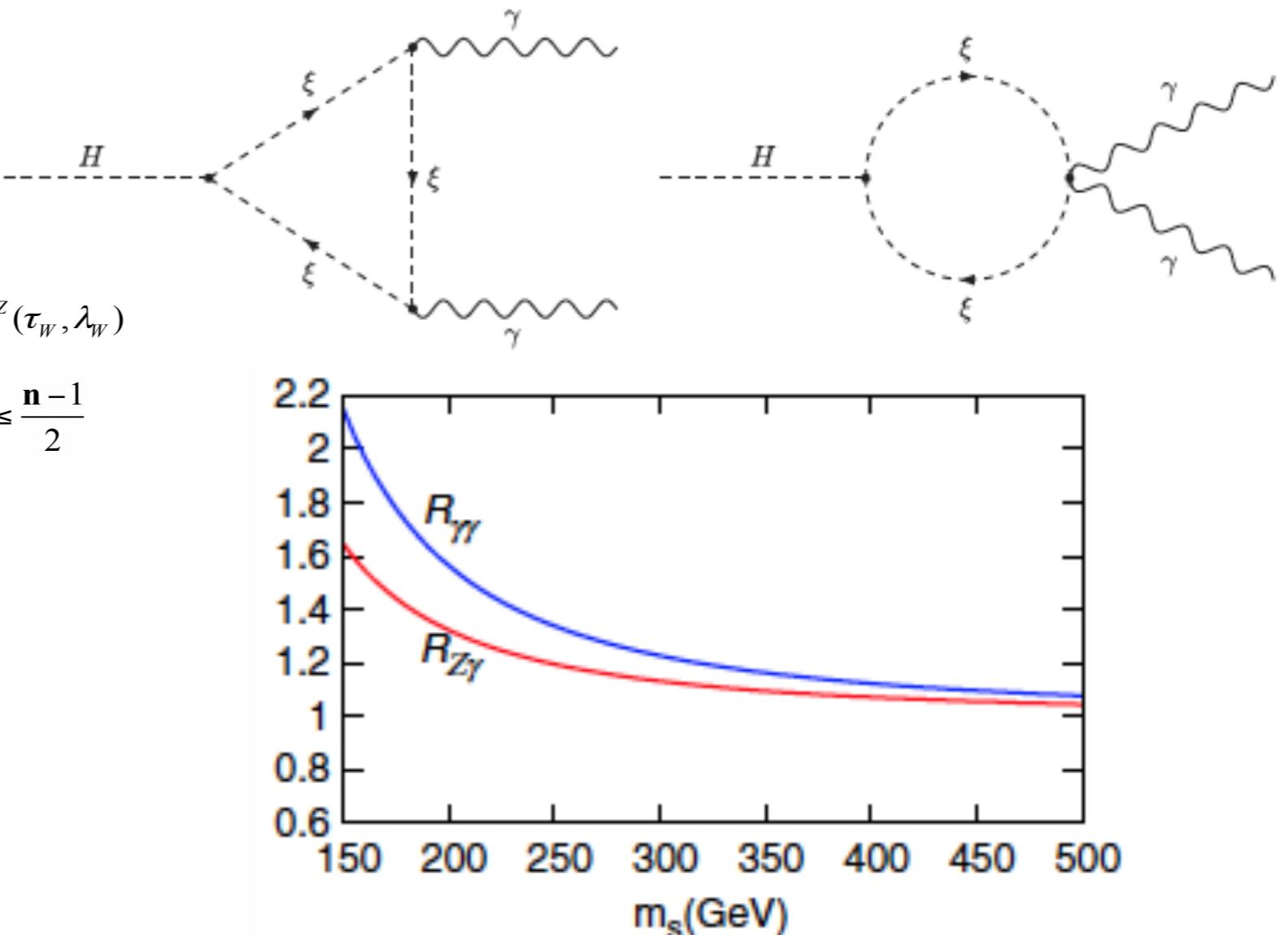


FIG. 4 (color online).  $R_{\gamma\gamma} \equiv \Gamma(H \rightarrow \gamma\gamma)/\Gamma(H \rightarrow \gamma\gamma)_{SM}$  and  $R_{Z\gamma} \equiv \Gamma(H \rightarrow Z\gamma)/\Gamma(H \rightarrow Z\gamma)_{SM}$  as functions of the degenerate mass factor  $m_s$  of the multicharged scalar states with  $n = 5$  and the universal trilinear coupling to Higgs,  $\mu_s = -100$  GeV.

The correlation among  $H \rightarrow \gamma\gamma$  and  $H \rightarrow Z\gamma$  strongly depends on the Gauge representation of charged scalars

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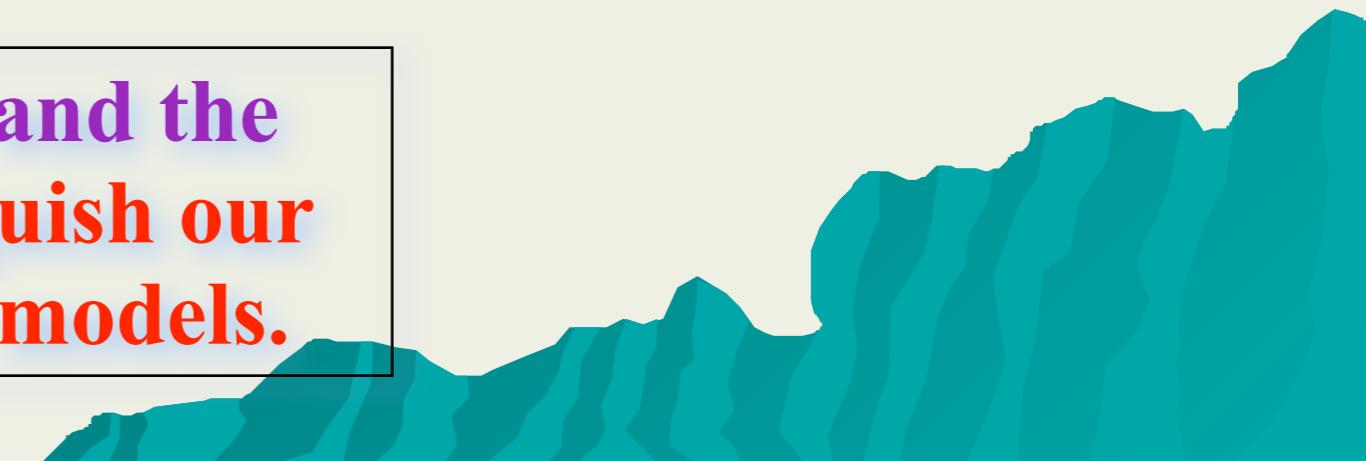
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Future data on  $0\nu\beta\beta$  decays and the LHC searches would distinguish our models from other neutrino models.



# 2nd International Workshop on *Particle Physics and Cosmology after Higgs and Planck*

## 後希格斯與普朗克粒子物理與宇宙學國際研討會

October 8-11, 2014 - National Center for Theoretical Sciences, NTHU, Hsinchu, Taiwan

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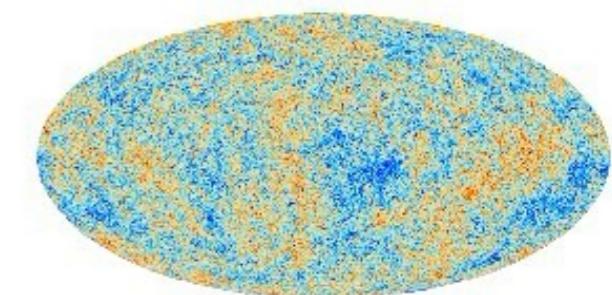
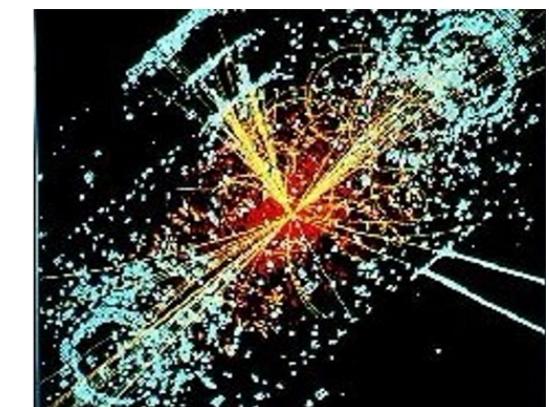
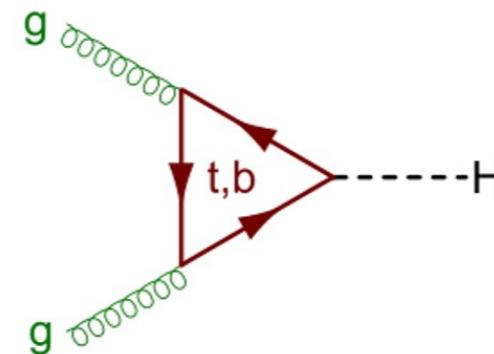
2013

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### Topics

- Higgs and Collider Physics
- Flavor and Neutrino Physics
- CP Violation
- Dark Matter and Dark Energy
- Gravity
- Inflation



Thank you!

謝謝！