

Dark Matter, Dark Energy & Neutrino Mass

暗物质，暗能量和中微子质量

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Lecture 1: Introduction to Particle Physics and Cosmology

Lecture 2: Some Basic Backgrounds of the Standard Model of Particle Physics and Cosmology

Lecture 3: Neutrino Mass Generation

Lecture 4: Theoretical Understanding of Dark Matter Detections

Lecture 5: Dark Energy and Gravitational Waves

Lecture 3: Neutrino Mass Generation

Outline

- Introduction
- A brief overview of neutrino mass generation
- **A special class of models to generate M_ν**
 - Neutrino mass generation
 - $0\nu\beta\beta$ decays
 - Other physics

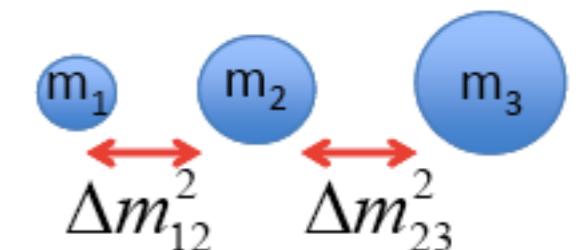
● Introduction

Weak eigenstate ($\alpha = e, \mu, \tau$) ————— $|\nu_\alpha\rangle = \sum_i U_{\alpha i} |\nu_i\rangle$ ————— Mass eigenstate ($i = 1, 2, 3$)

• PMNS (Pontecorvo-Maki-Nakagawa-Sakata) matrix

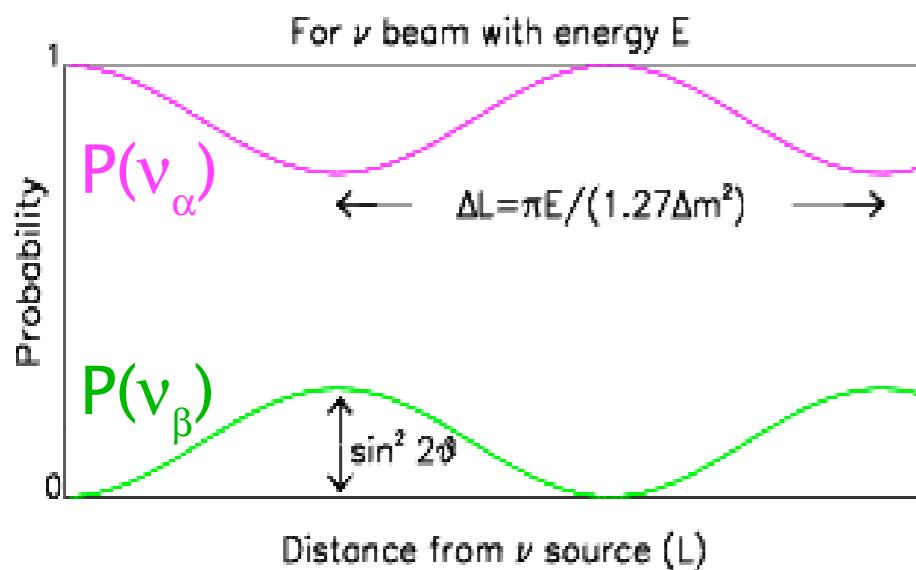
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U_{\text{PMNS}} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}}_{\text{Atmospheric}} \underbrace{\begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta} \sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta} \sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix}}_{\text{Reactor}} \underbrace{\begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\text{Solar}}$$



Two-neutrino case:

$$P(\nu_\alpha \rightarrow \nu_\beta) = |\langle \nu_\beta | \nu(t) \rangle|^2 = \sin^2 2\theta \sin^2 \left(\frac{1.27 \Delta m^2 [\text{eV}^2] L [\text{km}]}{E [\text{GeV}]} \right)$$



$$P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - P(\nu_\alpha \rightarrow \nu_\beta)$$

disappearance of ν_α

$$P(\nu_\alpha \rightarrow \nu_\beta) = \sin^2(2\theta) \sin^2 \left(\frac{\Delta m_{12}^2 L}{4E} \right)$$

appearance of ν_β

θ	: mixing angle
Δm^2	: mass squared difference
L [km]	: the distance traveled
E (GeV)	: the energy of neutrino

中微子振盪

如果中微子有質量，則不同類中微子間會產生振盪現象。

舉 ν_μ 及 ν_τ 為例：

$$\begin{pmatrix} \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

ν_μ, ν_τ ：弱作用本徵態

ν_1, ν_2 ：質量本徵態

在 $t=0$ 時， ν_μ 在大氣中產生，則

$$|\nu_\mu(0)\rangle \equiv |\nu_\mu\rangle = \cos \theta |\nu_1(0)\rangle + \sin \theta |\nu_2(0)\rangle$$

$$|\nu_\tau\rangle = -\sin \theta |\nu_1(0)\rangle + \cos \theta |\nu_2(0)\rangle$$

到了時間 t ，上述狀態演變為

$$|\nu_\mu(t)\rangle = \exp(-iE_1 t / \hbar) \cos \theta |\nu_1(0)\rangle + \exp(-iE_2 t / \hbar) \sin \theta |\nu_2(0)\rangle,$$

Using $t \approx L$ and $E - P_i = E - \sqrt{E^2 - m_i^2} \approx m_i^2/2E$,

如果 $m_1 \neq m_2$ ，則 $E_1 \neq E_2$ ，因此 $|\nu_\mu(t)\rangle$ 不再垂直於 $|\nu_\tau\rangle$ ！
 $\Delta m_{21}^2 = m_2^2 - m_1^2$

$$P(\nu_\alpha \Rightarrow \nu_\beta) = \sin^2(2\theta) \sin^2\left(\frac{\Delta m_{12}^2 L}{4E}\right)$$

$m_1 \neq m_2$



Neutrino Oscillations



Neutrinos have masses and mix with each other.

$$P(\nu_l \rightarrow \nu_{l'}) = \sum_j |U_{lj}|^2 |U_{l'j}|^2 + 2 \sum_{j>k} |U_{l'j} U_{lj}^* U_{lk} U_{l'k}^*|^2 \cos\left(\frac{\Delta m_{jk}^2}{2p} L - \varphi_{l'l;jk}\right)$$

$$\Delta m_{jk}^2 = m_j^2 - m_k^2$$

$$\varphi_{l'l;jk} = \arg(U_{l'j} U_{lj}^* U_{lk} U_{l'k}^*)$$

ν_{atm} SK UP-DOWN ASYMMETRY

θZ -, L/E- dependences of μ -like events

Dominant $\nu_{\mu} \rightarrow \nu_{\tau}$ **K2K, MINOS, T2K; CNGS (OPERA)** $\longrightarrow |\Delta m_{32}^2|, \sin^2 \theta_{23}$

ν_{\odot} **Homestake, Kamiokande, SAGE, GALLEX/GNO**
Super-Kamiokande, SNO, BOREXINO; KamLAND

$\longrightarrow \Delta m_{21}^2, \sin^2 \theta_{12}$

$\bar{\nu}_e$ (from reactors): **Daya Bay, RENO, Double Chooz**

Dominant $\nu_e \rightarrow \nu_{\mu,\tau}$ $\longrightarrow \sin^2 \theta_{13}$

T2K, MINOS (ν_{μ} from accelerators): $\nu_{\mu} \rightarrow \nu_e$

Experiments on solar neutrinos

$$\Delta m_{21}^2 = \left(7.54^{+0.26}_{-0.22} \right) \times 10^{-5} \text{ eV}^2$$

$$\Delta m_{sun}^2 = |m_2|^2 - |m_1|^2 > 0$$

Neutrinos born in Cosmic ray collisions and on earth

$$|\Delta m_{31}^2| = \begin{cases} (2.43 \pm 0.06) \times 10^{-3} \text{ eV}^2 & \text{normal hierarchy,} \\ (2.38 \pm 0.06) \times 10^{-3} \text{ eV}^2 & \text{inverted hierarchy,} \end{cases}$$

$$\Delta m_{atm}^2 = |\Delta m_{31}^2|$$

Normal hierarchy

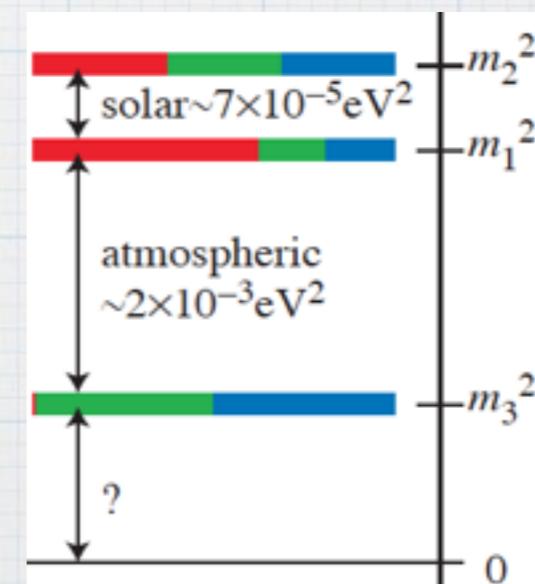
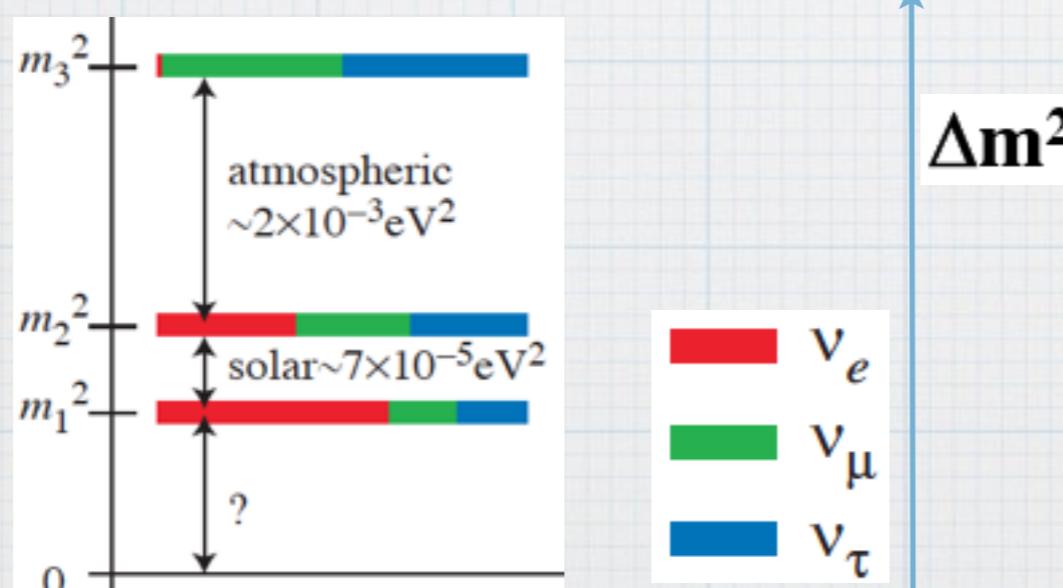
$$|m_{\nu_1}| < |m_{\nu_2}| \ll |m_{\nu_3}|$$

Inverted hierarchy

$$|m_{\nu_1}| \simeq |m_{\nu_2}| \gg |m_{\nu_3}|$$

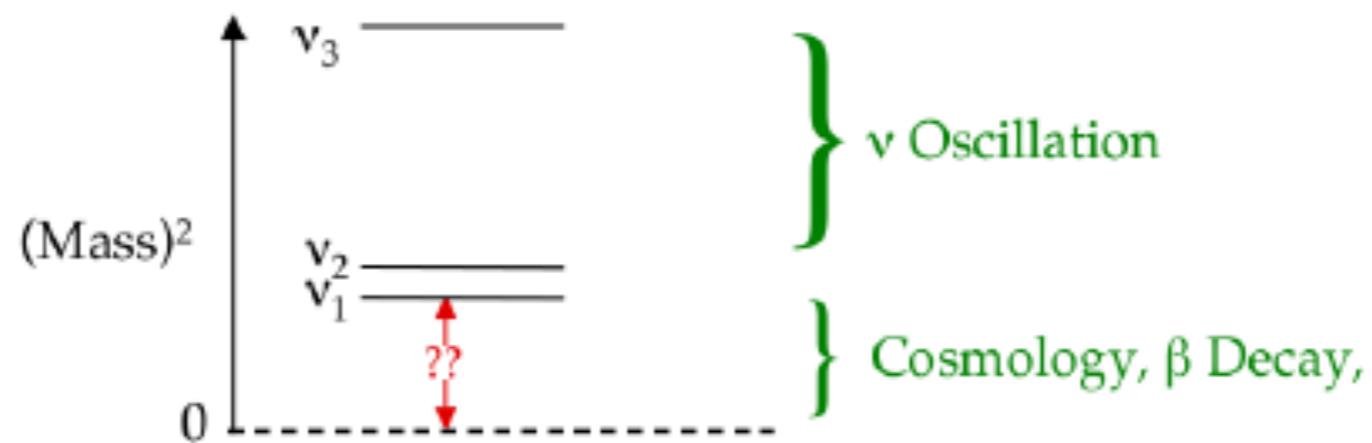
$$m_1 \simeq 0, m_2^2 \simeq \Delta m_{\odot}^2, \text{ and } m_3^2 \simeq \Delta m_{atm}^2.$$

$$m_1^2 \simeq m_2^2 \simeq \Delta m_{atm}^2 \gg m_3^2$$



Absolute Neutrino Mass Scale

The Absolute Scale of Neutrino Mass



How far above zero
is the whole pattern?

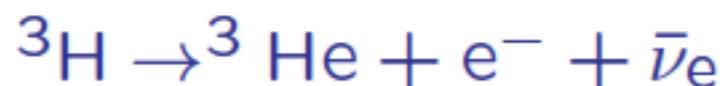
Oscillation Data $\Rightarrow \sqrt{\Delta m_{\text{atm}}^2} < \text{Mass[Heaviest } \nu_i]$

The Troitzk and Mainz ${}^3\text{H}$ β -decay experiments

$m_{\bar{\nu}_e} < 2.05 \text{ eV}$ (95% C.L.)

KATRIN

0.2 eV



Improved β energy resolution requires a **BIG** β spectrometer.

KATRIN

2006年11月25日
德國



Improved β energy resolution requires a **BIG** β spectrometer.

KATRIN



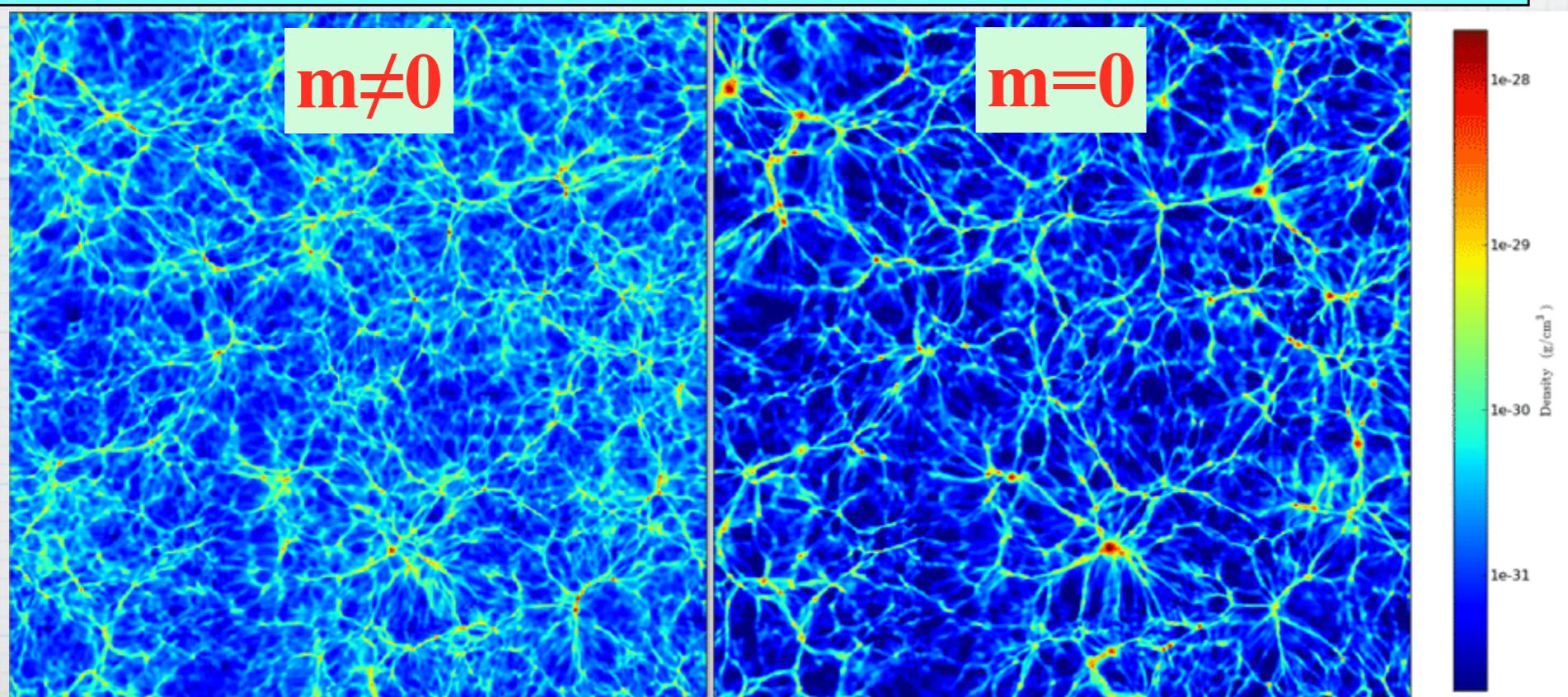
The cosmological bound on m_ν

- Number density

$$n_\nu = \int \frac{d^3 p}{(2\pi)^3} f_\nu(p, T_\nu) = \frac{3}{11} n_\gamma = \frac{6\zeta(3)}{11\pi^2} T_{CMB}^3$$

- Energy density

$$\rho_{\nu_i} = \int \sqrt{p^2 + m_{\nu_i}^2} \frac{d^3 p}{(2\pi)^3} f_\nu(p, T_\nu) \rightarrow \begin{cases} \frac{7\pi^2}{120} \left(\frac{4}{11}\right)^{4/3} T_{CMB}^4 & \text{Massless} \\ m_{\nu_i} n_\nu & \text{Massive } m_\nu \gg T \end{cases}$$



The cosmological bound on m_ν

- Number density

$$n_\nu = \int \frac{d^3 p}{(2\pi)^3} f_\nu(p, T_\nu) = \frac{3}{11} n_\gamma = \frac{6\zeta(3)}{11\pi^2} T_{CMB}^3$$

Contribution to the energy density of the Universe

$$\Omega_\nu h^2 = 1.7 \times 10^{-5}$$

$$\Omega_\nu h^2 = \frac{\sum m_i}{94.1 \text{ eV}}$$

Massless

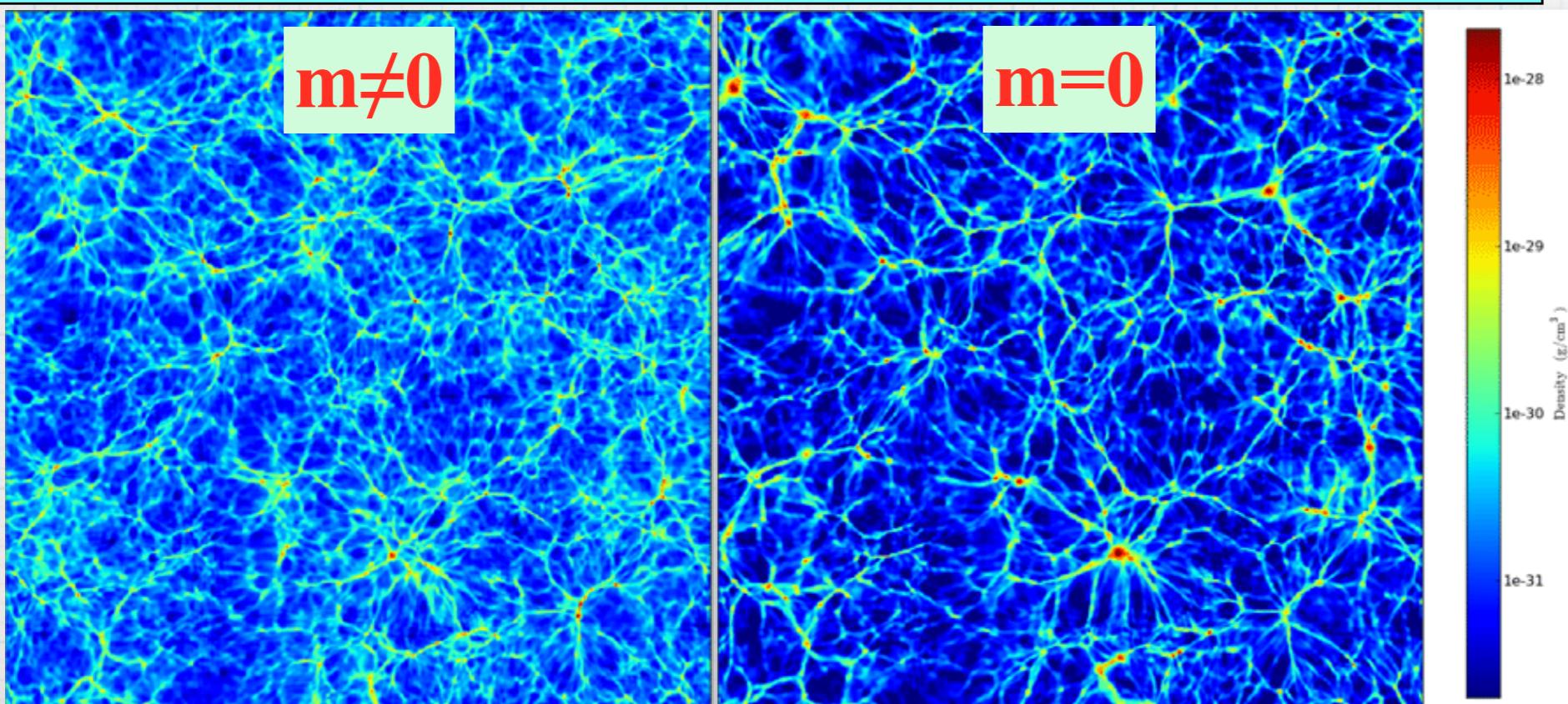
Massive $m_\nu \gg T$

The best bound to absolute values of neutrino masses from Cosmology



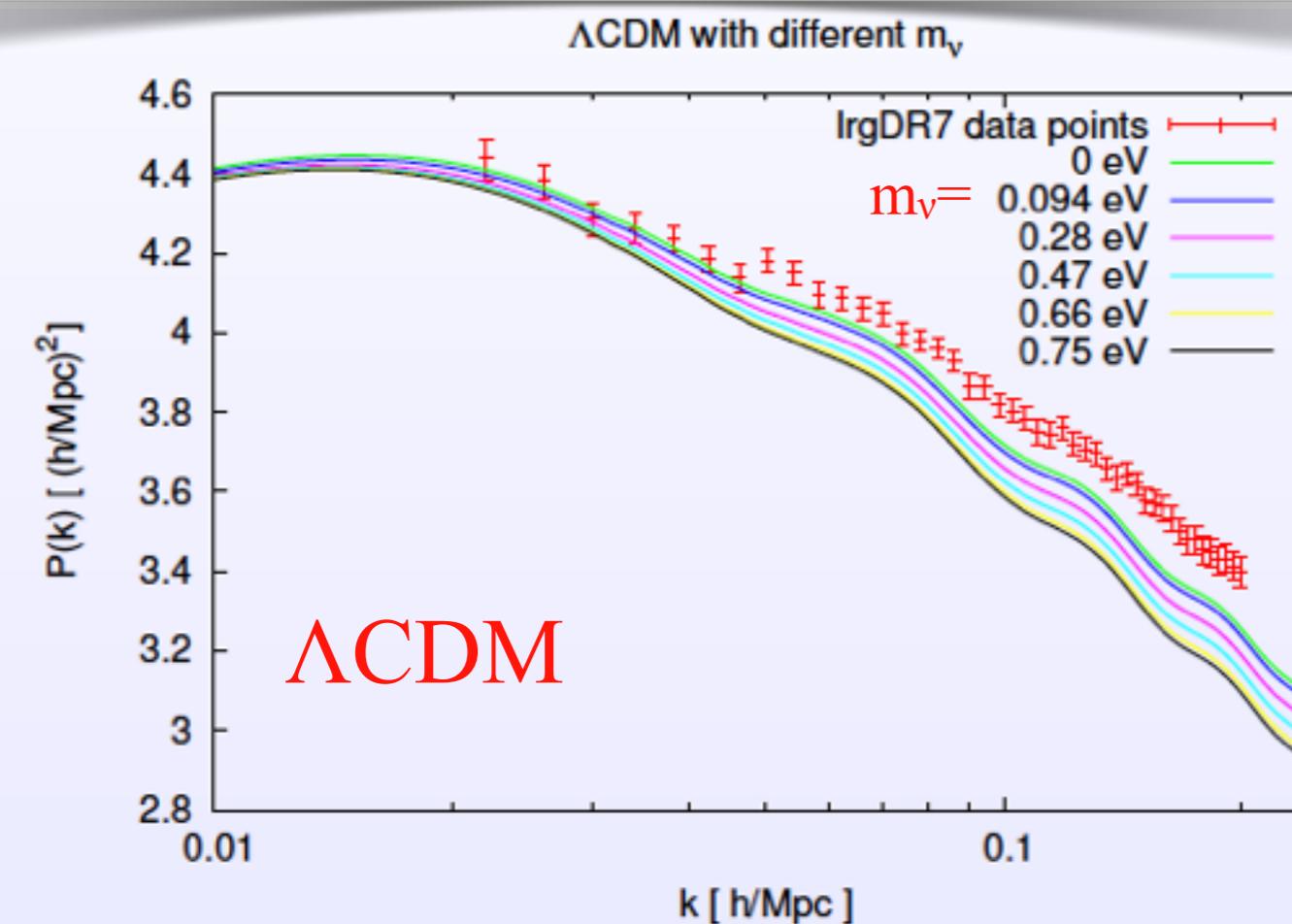
$$\sum_i m_{\nu_i} < 0.23 \text{ eV}$$

95%CL (Planck+other data)

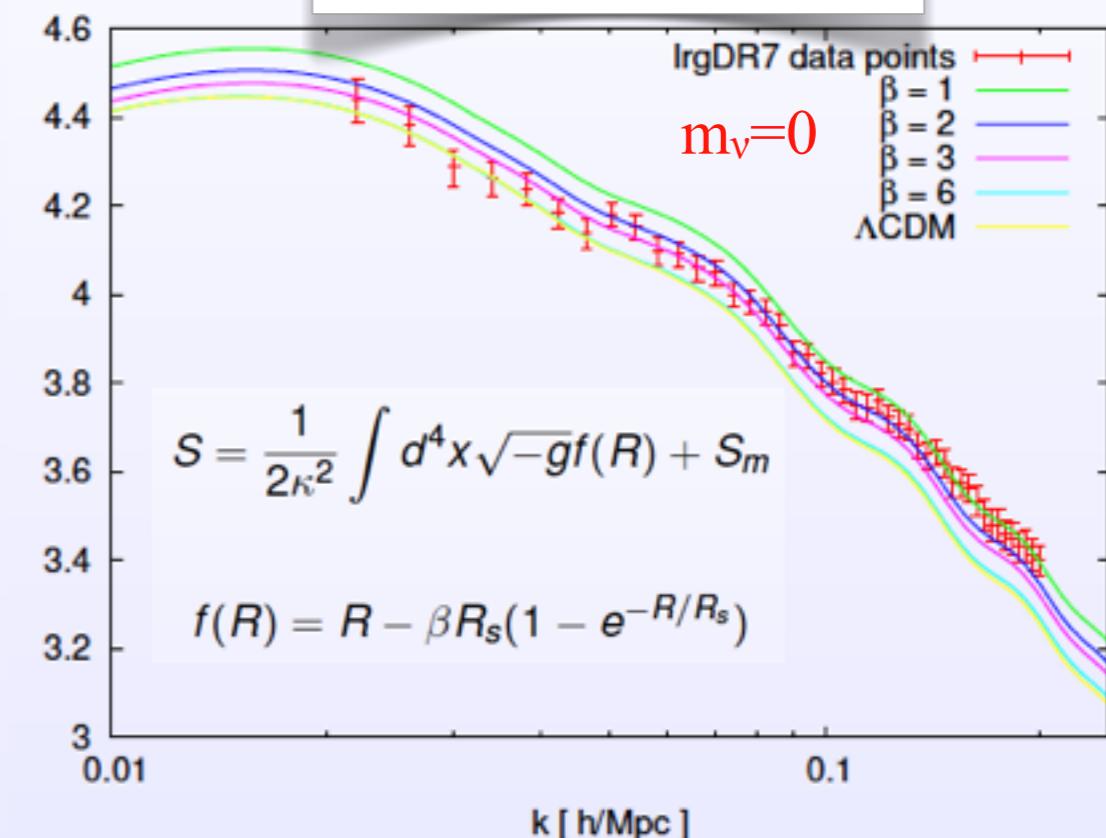


ΛCDM

Matter power spectrum in Λ CDM and $f(R)$

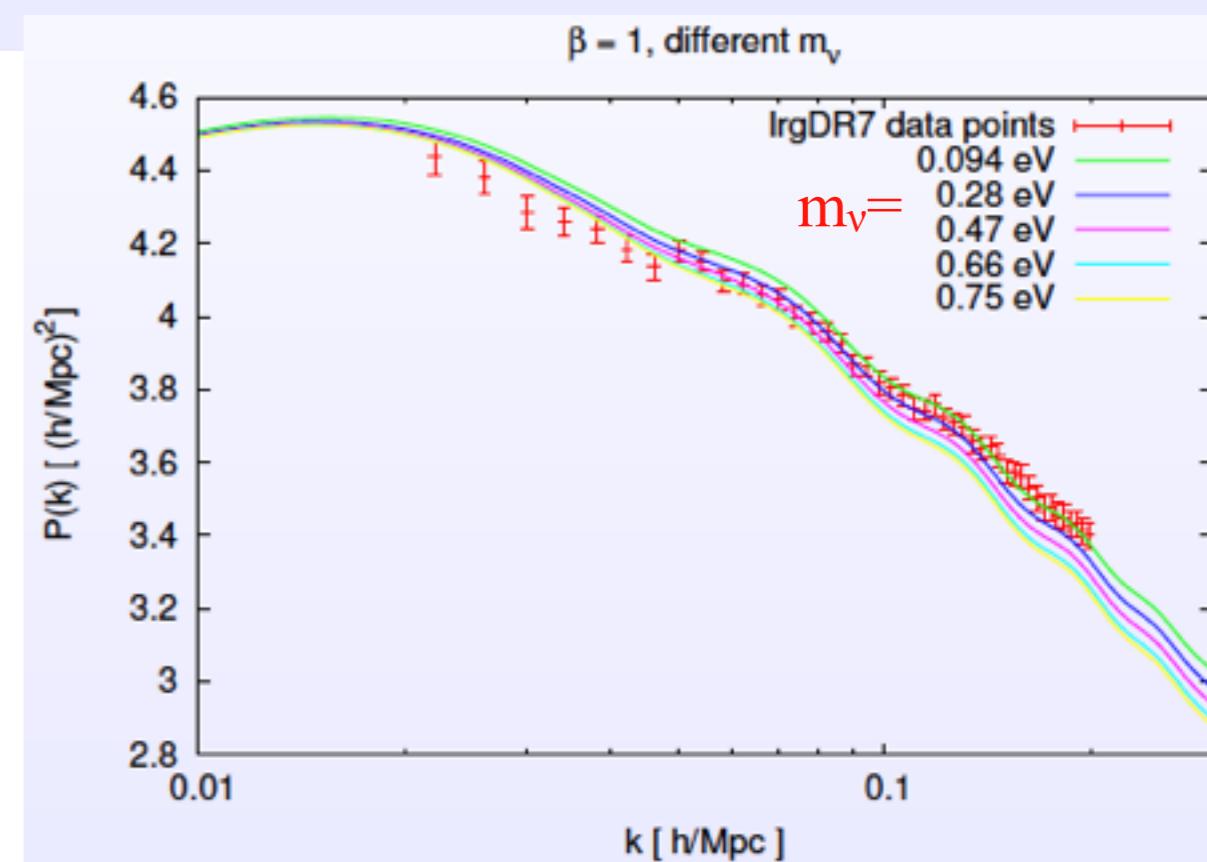


Exponential $f(R)$ model



$f(R)$ model	Σm_ν
Λ CDM	< 0.200 eV
Starobinsky	$0.248^{+0.203}_{-0.232}$ eV
Exponential	< 0.214 eV

CQG+CC.Lee, J.L.Shen,
PLB740,285(2015)



In terms of the PMNS mixing matrix

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \underbrace{\begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix}}_{\rightarrow} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$V_{PMNS} = \begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{bmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\alpha_{21}/2} & 0 \\ 0 & 0 & e^{i\alpha_{31}/2} \end{pmatrix}$$

PDG2016

$$\sin^2 \theta_{12} = 0.308 \pm 0.017, \sin^2 \theta_{23} = 0.437^{+0.033}_{-0.023}, \sin^2 \theta_{13} = 0.0234^{+0.020}_{-0.019}, \delta = (1.39^{+0.38}_{-0.27})\pi$$

NH

$$\sin^2 \theta_{12} = 0.308 \pm 0.017, \sin^2 \theta_{23} = 0.455^{+0.039}_{-0.031}, \sin^2 \theta_{13} = 0.0240^{+0.019}_{-0.020}, \delta = (1.31^{+0.29}_{-0.33})\pi$$

IH

Bimaximal Matrix

$$\begin{pmatrix} \frac{\sqrt{2}}{2} & \frac{\sqrt{2}}{2} & 0 \\ -\frac{1}{2} & \frac{1}{2} & \frac{\sqrt{2}}{2} \\ \frac{1}{2} & -\frac{1}{2} & \frac{\sqrt{2}}{2} \end{pmatrix}$$

$$\theta_{12}=45^\circ, \theta_{23}=45^\circ, \theta_{13}=0$$

Tribimaximal Matrix

$$\begin{pmatrix} \frac{-\sqrt{6}}{3} & \frac{\sqrt{3}}{3} & 0 \\ -\frac{\sqrt{6}}{6} & \frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \\ \frac{\sqrt{6}}{6} & -\frac{\sqrt{3}}{3} & \frac{\sqrt{2}}{2} \end{pmatrix}$$

Daya-Bay

center values

$$\begin{aligned} \theta_{12} &\approx 34^\circ \\ \theta_{23} &\approx 42^\circ \\ \theta_{13} &\approx 9^\circ \\ \delta &\approx 1.4\pi \end{aligned}$$

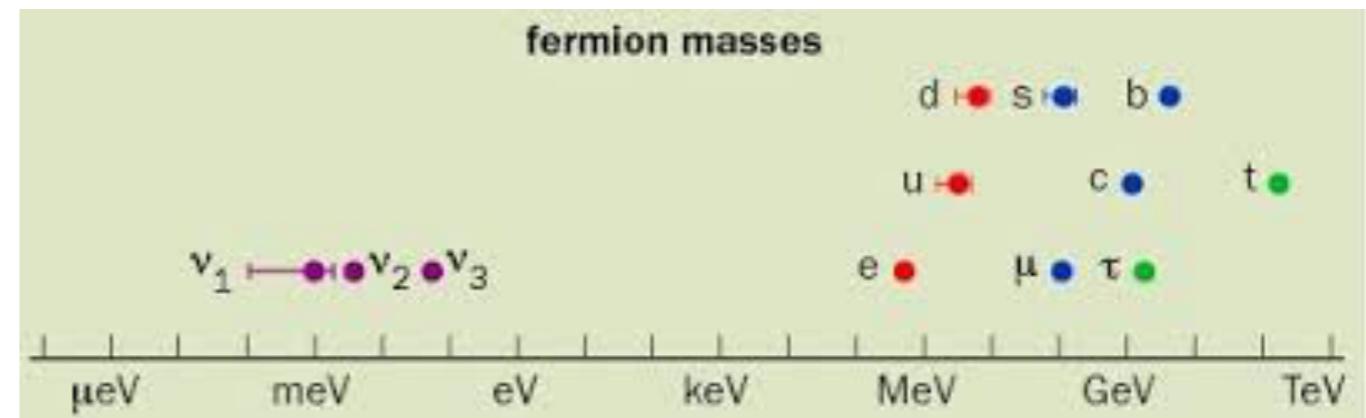
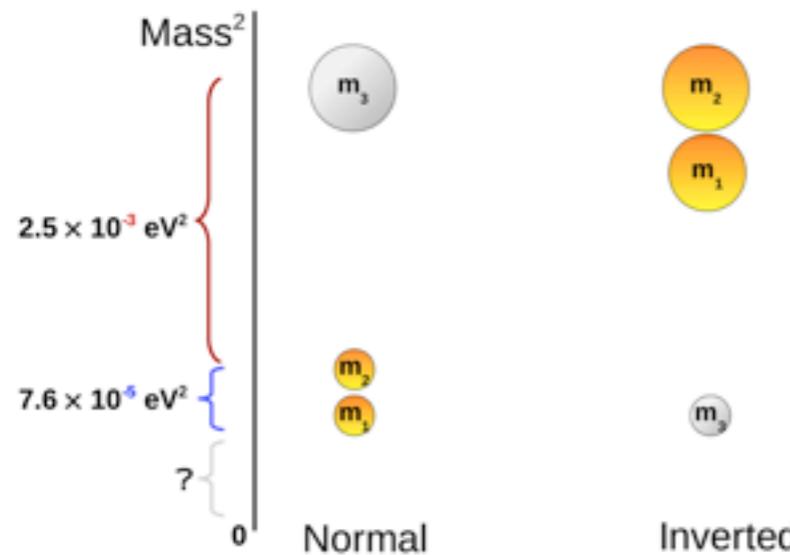
$$\theta_{12}=35.3^\circ, \theta_{23}=45^\circ, \theta_{13}=0$$

$$|U_{PMNS}| \sim \left(\begin{array}{ccc} \text{yellow} & \text{yellow} & \cdot \\ \text{light green} & \text{light green} & \cdot \\ \text{light green} & \text{light green} & \cdot \\ \cdot & \cdot & \cdot \end{array} \right)$$

$$|V_{CKM}| \sim \left(\begin{array}{ccc} \text{red} & \cdot & \cdot \\ \cdot & \text{orange} & \cdot \\ \cdot & \cdot & \text{orange} \\ \cdot & \cdot & \cdot \end{array} \right)$$

• A brief overview of neutrino mass generation

Fermion Mass Problem



$$m_{\nu_j} \lll m_{e, \mu, \tau}, m_q, q = u, c, t, d, s, b$$

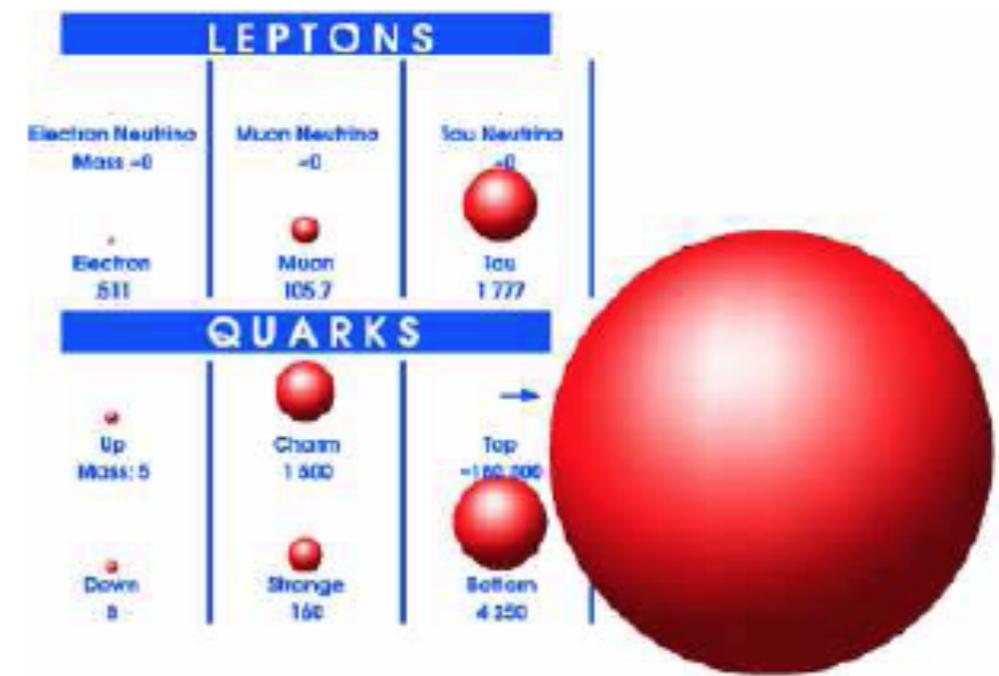
$$\text{For } m_{\nu_j} \lesssim 1 \text{ eV: } m_{\nu_j}/m_{l,q} \lesssim 10^{-6}$$

Questions:

Where do the

- quark mass hierarchy
 - small neutrino masses
 - small quark mixings and
 - large lepton mixings
- originate from?

如果中微子有質量，它們的質量為什麼遠小於相對的帶電輕子及夸克的質量？



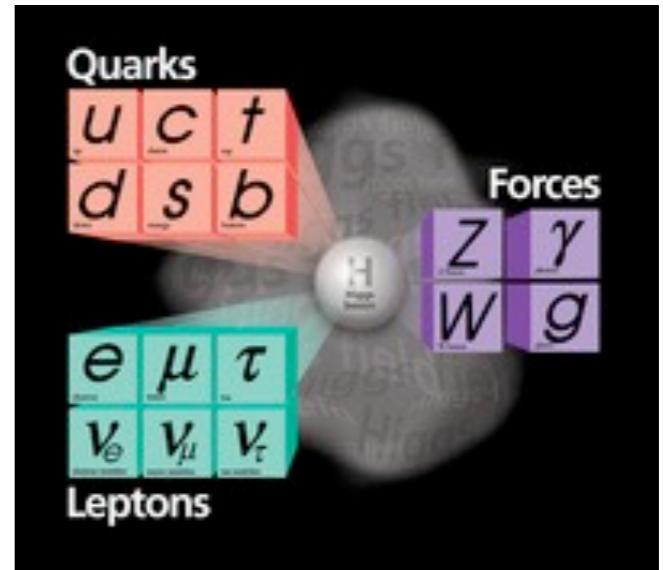
- The standard model: $SU(3)_C \times SU(2)_L \times U(1)_Y$

$$Q_L : \begin{pmatrix} u \\ d \end{pmatrix}_L \quad \begin{pmatrix} c \\ s \end{pmatrix}_L \quad \begin{pmatrix} t \\ b \end{pmatrix}_L \quad L_L : \begin{pmatrix} \nu_e \\ e \end{pmatrix}_L \quad \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}_L \quad \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}_L$$

$$U_R : \quad u_R \quad \quad c_R \quad \quad t_R$$

$$D_R : \quad d_R \quad \quad s_R \quad \quad b_R \quad \quad E_R : \quad e_R \quad \quad \mu_R \quad \quad \tau_R$$

Higgs: H^0 Gauge Bosons: W^\pm, Z, γ, g



Yukawa interactions: $Y = \sum_{i,j} h_{ij}^d \bar{Q}_L \phi D_R + h_{ij}^u \bar{Q}_L \tilde{\phi} U_R + h_{ij}^e \bar{L}_L \phi E_R + h.c.$

$$\Phi = \Phi_0 = (-\mu^2/2\lambda)^{1/2}$$

SSB

$$V_L^{d+} M_d V_R^d = M_d^{diag.}, \quad D_{L(R)j} = (V_{L(R)}^d)_{ji} D'_{L(R)i}$$

$$V_L^{u+} M_u V_R^u = M_u^{diag.}, \quad U_{L(R)j} = (V_{L(R)}^u)_{ji} U'_{L(R)i}$$

■ What about neutrinos?

■ Do neutrinos get their masses like charged fermions?

Neutrino masses: Dirac or Majorana

Dirac neutrino mass:

$$\mathcal{L}_D = -m_D \bar{\nu}_L \nu_R + \text{h.c.}$$



Introduce ν_R
(not in the SM)

Majorana neutrino mass:

$$\mathcal{L}_M = -m_M \bar{\nu}^c \nu + \text{h.c.} \quad \nu \leftrightarrow \bar{\nu}$$



FORBIDDEN
IN THE SM.

😊 the lepton number L is conserved

• the lepton number L is violated

Neutrino oscillations measure Δm^2 , but they do not provide information about the absolute neutrino spectrum and cannot distinguish pure Dirac and Majorana neutrinos.

In the SM:

- No Dirac mass term
(no right-handed neutrino).
- No Majorana mass term either
(ν_L is an SU(2) doublet).

	$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$
$L_a = (\nu_a, l_a)^T_L$	(1, 2, -1)
$e_a^c L$	(1, 1, 2)
$Q_a = (u_a, d_a)^T_L$	(3, 2, 1/3)
$u_a^c L$	($\bar{3}$, 1, -4/3)
$d_a^c L$	($\bar{3}$, 1, 2/3)
Φ	(1, 2, 1)

Table 1: Matter and scalar multiplets of the Standard Model (SM)

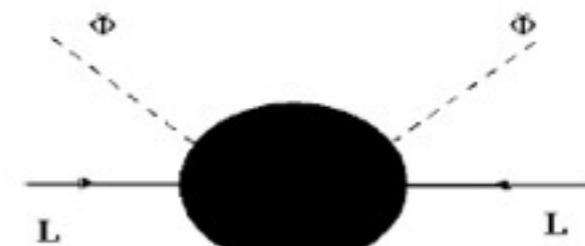
■ Effective Dim-5 operator:

$$O = (\lambda_0/M_X) L \Phi L \Phi$$

↓ SSB

$$m_\nu = \lambda_0 \frac{\langle \Phi \rangle^2}{M_X},$$

(Majorana)



Dimension five operator responsible for neutrino mass

For $\lambda_0 \sim 1$, $\langle \Phi \rangle \sim 100$ GeV, $M_X \sim M_P \rightarrow m_\nu \sim 10^{-6}$ eV (too small)

$$\Delta m_{21}^2 \sim 7 \times 10^{-5} \text{ eV}^2$$

$$|\Delta m_{31}^2| \sim 2 \times 10^{-3} \text{ eV}^2$$

Neutrino masses beyond the SM:

■ If there are right handed neutrinos ν_R : $\nu_R = (1, 1, 0)$

$$\mathcal{L}_Y = Y_\nu \bar{L} \Phi \nu_R + h.c. \Rightarrow m_\nu^D = Y_\nu \langle \Phi \rangle$$

The observed neutrino masses would require $Y_\nu \leq 10^{-13} - 10^{-12}$

(unnatural)?

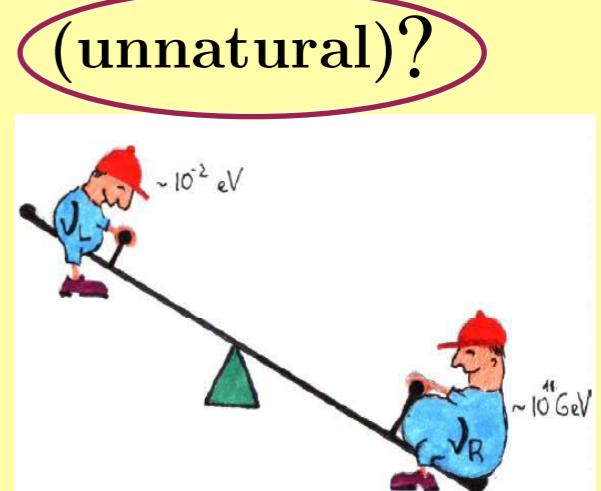
■ Majorana mass for ν_R :

$$M_R \nu_R^T C^{-1} \nu_R + h.c.$$

Type-I see-saw mechanism:

$$\mathcal{M}_\nu = -m_D^T M_R^{-1} m_D.$$

(naturally small?+Majorana)



The Seesaw Mechanism

The Seesaw mechanism refers to the neutrino mass matrix of the form:

$$L_m = -\frac{1}{2}(\nu_L^c, \nu_R) \begin{pmatrix} 0 & m_D \\ m_D^T & M_R \end{pmatrix} \begin{pmatrix} \nu_L \\ \nu_R^c \end{pmatrix}$$

For one generation, if $M_R \gg m_D$, the diagonal masses are

$$m_\nu \approx -m_D M_R^{-1} m_D^T, \quad m_N \approx M_R$$

$$\begin{aligned} m_{\nu_1} &= m_e^2/M_R \\ m_{\nu_2} &= m_\mu^2/M_R \\ m_{\nu_3} &= m_\tau^2/M_R \end{aligned}$$

How large M_R needs to be?

For $m_{\nu_1} = 0.1\text{eV}, M_R = 2.5\text{TeV}$

For $m_{\nu_2} = 0.1\text{eV}, M_R = 10^8\text{GeV}$

For $m_{\nu_3} = 0.1\text{eV}, M_R = 3 \times 10^{10}\text{GeV}$.

A very nice way to explain why light neutrino masses are so much lighter than their charged lepton partners

(Minkowski (1977); Gell-Mann, Ramond, and Slansky (1979); Yanagida (1979); Glashow (1980); Mohapatra and Senjanovic(1980))

Inverse Seesaw

Basis (ν, ν^c, S) :

$$M_\nu = \begin{pmatrix} 0 & \textcolor{blue}{m}_D & 0 \\ \textcolor{blue}{m}_D^T & 0 & M \\ 0 & M^T & \mu \end{pmatrix},$$

Mohapatra & Valle,
1986

After EWSB, the effective light neutrino mass matrix is given by

$$M_\nu = \textcolor{blue}{m}_D M^{T^{-1}} \mu M^{-1} \textcolor{blue}{m}_D^T.$$

“Inverse” seesaw, because:

$$\textcolor{red}{M}_\nu \Rightarrow 0 \quad \text{IF} \quad \mu \Rightarrow 0$$

Effective Operators

d = 5:

$$\mathcal{O}_W \propto \frac{c_{ij}}{\Lambda} (L_i H)(L_j H)$$

Weinberg, 1980

d = 7:

$$\mathcal{O}_2 \propto LLL e^c H$$

Babu & Leung, 2001

$$\mathcal{O}_3 \propto LLQ d^c H$$

de Gouvea & Jenkins, 2007

$$\mathcal{O}_4 \propto LL\bar{Q}\bar{u}^c H$$

$$\mathcal{O}_8 \propto L\bar{e}^c \bar{u}^c d^c H$$

$$\mathcal{O} \propto (LH)(LH)(H_u H_d)$$

d = 9:

$$\mathcal{O}_5 \propto LLQ d^c H H H^\dagger$$

$$\mathcal{O}_6 \propto LL\bar{Q}\bar{u}^c H H^\dagger H$$

$$\mathcal{O}_7 \propto LQ\bar{e}^c \bar{Q} H H H^\dagger$$

$$\mathcal{O}_9 \propto LLL e^c L e^c \quad .$$

$$\mathcal{O}_{10} \propto LLL e^c Q d^c$$

$$\mathcal{O}_{11} \propto LLQ d^c Q d^c$$

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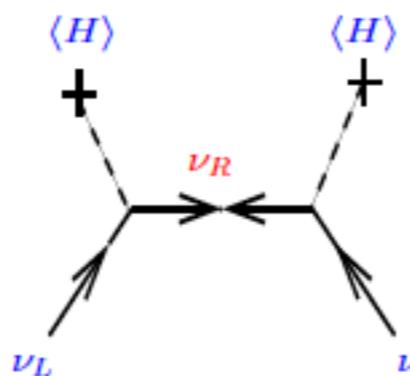
Example realization:

$d = 5$:

$$\mathcal{O}_W \propto \frac{c_{ij}}{\Lambda} (L_i H)(L_j H)$$

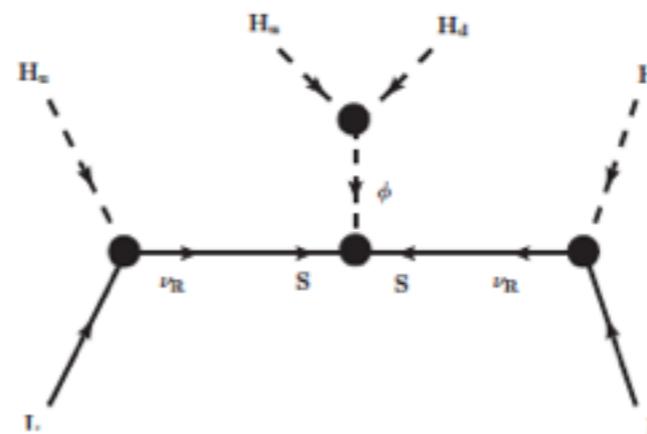
$$\Lambda \simeq M_{\nu R_k}$$

$$c_{ij} \propto Y_{ik}^\nu Y_{jk}^\nu$$

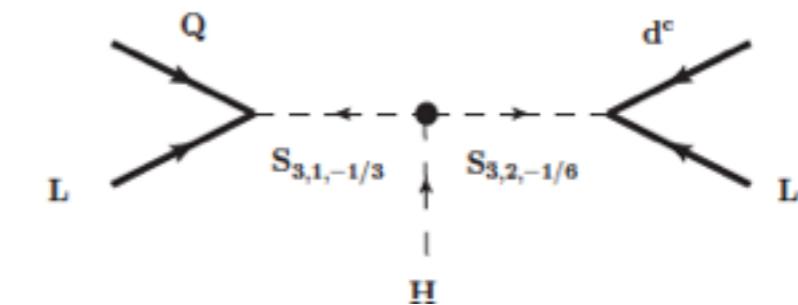
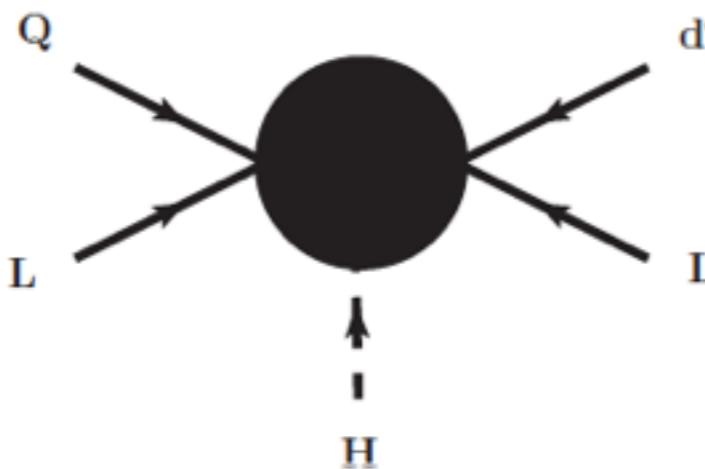


$d = 7$:

$$\mathcal{O} \propto (LH)(LH)(H_u H_d)$$

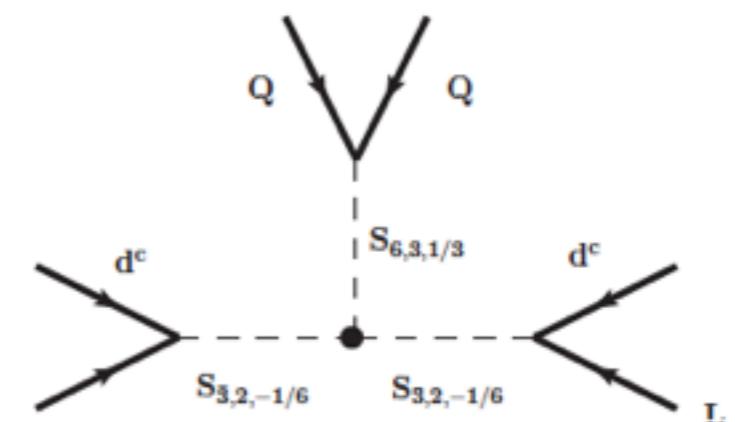
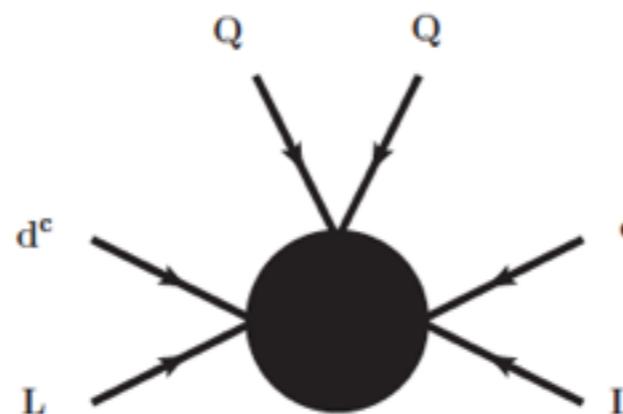


$$\mathcal{O}_3 \propto LLQd^cH$$

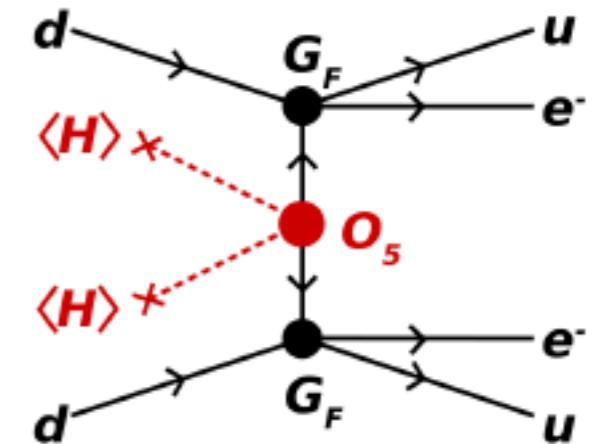


$S_{3,1,-1/3}$ - singlet leptoquark
 $S_{3,2,1/6}$ - doublet leptoquark

$$d = 9: \quad \mathcal{O}_{11} \propto LLQd^cQd^c$$



$0\nu\beta\beta$ decay:



● 產生 Majorana 中微子質量簡介

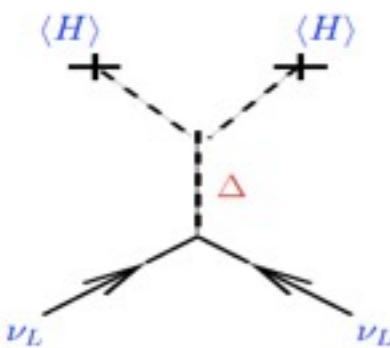
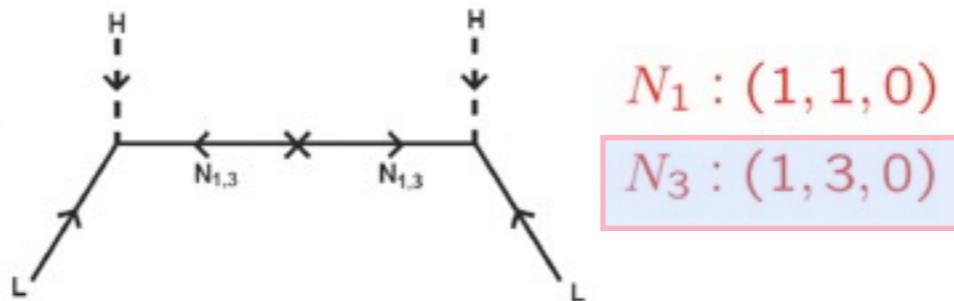
Without v_R

Tree level

Minkowski 1977; ...

Xiao-Gang He et al 1989

Type (I, III) seesaw



$$\Delta \equiv \begin{pmatrix} H^- & -\sqrt{2} H^0 \\ \sqrt{2} H^{--} & -H^- \end{pmatrix} = (1, 3, 2) \quad \text{scalar triplet}$$

$$\mathcal{L}_{\text{TypeII}} = \text{Tr}(D_\mu \Delta)^\dagger (D^\mu \Delta) - (Y_\nu l_L^T C i\sigma_2 \Delta l_L + \text{h.c.}) - V(H, \Delta),$$

Type II seesaw

Schechter & Valle, 1980, 1982

Cheng & Li, 1980

Mohapatra, Senjanovic, 1981

$$V(H, \Delta) = M_\Delta^2 \text{Tr} \Delta^\dagger \Delta + (\mu H^T i\sigma_2 \Delta^\dagger H + \text{h.c.}) + \lambda_1 (H^\dagger H) \text{Tr} \Delta^\dagger \Delta \\ + \lambda_2 (\text{Tr} \Delta^\dagger \Delta)^2 + \lambda_3 \text{Tr} (\Delta^\dagger \Delta)^2 + \lambda_4 H^\dagger \Delta \Delta^\dagger H.$$

$$M_\nu = \sqrt{2} Y_\nu v_\Delta$$

$$v_\Delta = \frac{\mu v^2}{\sqrt{2} M_\Delta^2}$$

$$M_\Delta \sim 250 \text{ GeV}, \mu \sim 0.1 \text{ eV}, Y_\nu \sim 1$$

—————> $M_\nu \sim 0.1 \text{ eV}$

Type III seesaw

Foot, Lew, X.G.He and Joshi, 1989

$$\Sigma = \begin{pmatrix} \Sigma_L^0/\sqrt{2} & \Sigma_L^+ \\ \Sigma_L^- & -\Sigma_L^0/\sqrt{2} \end{pmatrix} \quad \text{the triplet } \Sigma = (1, 3, 0)$$

$$\mathcal{L} = Tr[\bar{\Sigma} i \not{D} \Sigma] - \frac{1}{2} Tr[\bar{\Sigma} M_\Sigma \Sigma^c + \bar{\Sigma^c} M_\Sigma^* \Sigma] - \tilde{H}^\dagger \bar{\Sigma^c} \sqrt{2} Y_\Sigma L_L - \bar{L_L} \sqrt{2} Y_\Sigma^\dagger \Sigma^c \tilde{H}$$

The mass terms are given by:

$$\mathcal{L}_{\Psi} = -(\bar{l}_R \bar{\Psi}_R) \begin{pmatrix} m_l & 0 \\ Y_\Sigma v & M_\Sigma \end{pmatrix} \begin{pmatrix} l_L \\ \Psi_L \end{pmatrix} - (\bar{\nu}_L^c \bar{\Sigma}_L^{0c}) \begin{pmatrix} 0 & Y_\Sigma^T v / 2\sqrt{2} \\ Y_\Sigma v / 2\sqrt{2} & M_\Sigma / 2 \end{pmatrix} \begin{pmatrix} \nu_L \\ \Sigma_L^0 \end{pmatrix}$$

$$v \equiv \sqrt{2}\langle\phi^0\rangle = 246 \text{ GeV. } \Psi \equiv \Sigma_L^{+c} + \Sigma_L^-$$

$$M_\Sigma \sim 100 \text{ GeV} \Rightarrow Y_\Sigma \sim 10^{-7}$$

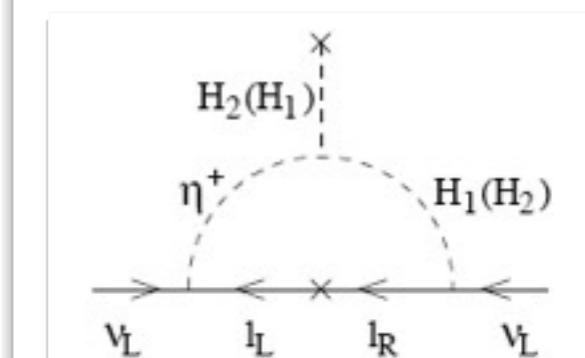
Loop level

1-loop:

1980

a. Zee model (with charged scalar singlet and additional scalar doublets).

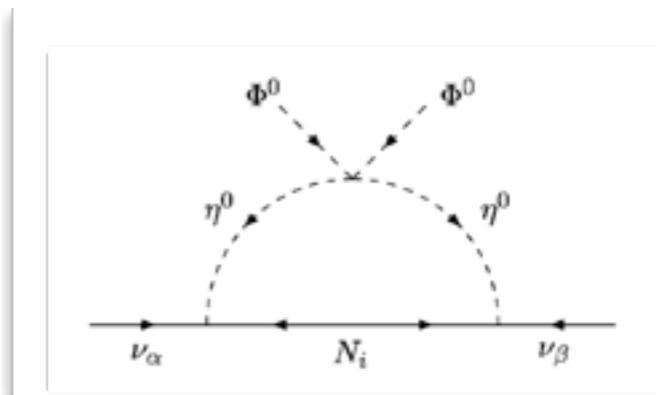
$$l^T \hat{f}_i \sigma_2 l \eta^+ + \sum_{i=1,2} \bar{l} \hat{f}_i e H_i;$$



2006

b. Ma model (with fermion singlet N_i and additional scalar doublet η).

$$h_{\alpha i} (\nu_\alpha \eta^0 - l_\alpha \eta^+) N_i$$

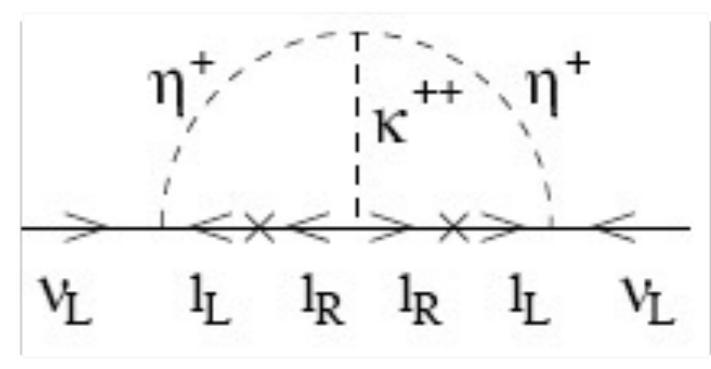


2-loop:

1986 1988

Zee-Babu model (with doubly and singly charged scalars).

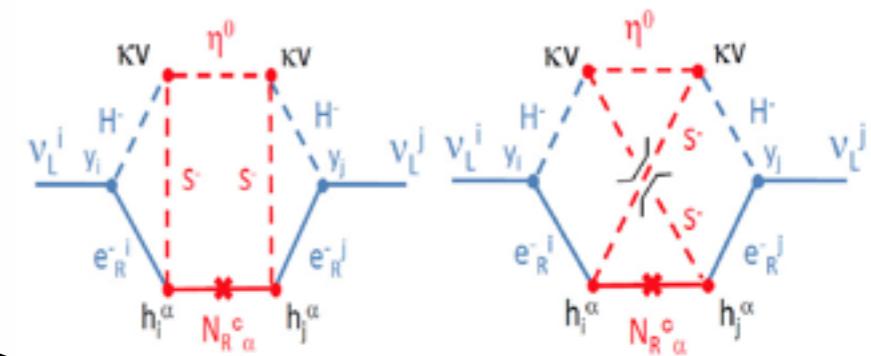
$$l^T \hat{f} l \eta^+ + l_R^T \hat{h} l_R k^{++}$$



3-loop:

	Q^i	u_R^i	d_R^i	L^i	e_R^i	Φ_1	Φ_2	S^\pm	η	N_R^α
Z_2 (exact)	+	+	+	+	+	+	+	-	-	-
\tilde{Z}_2 (softly broken)	+	-	-	+	+	+	-	+	-	+

M.Aoki,S.Kanemura,O.Seto, PRL102,051805(2009)

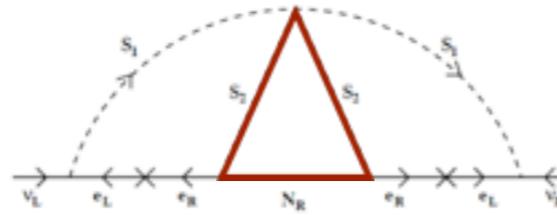
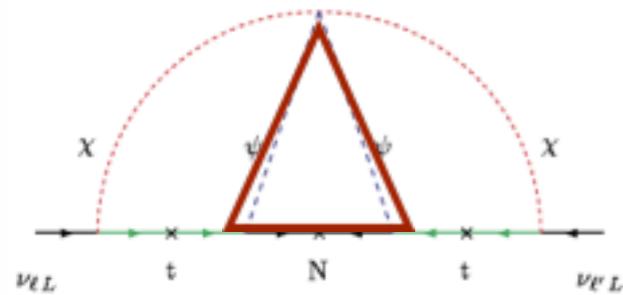


Dark Matter

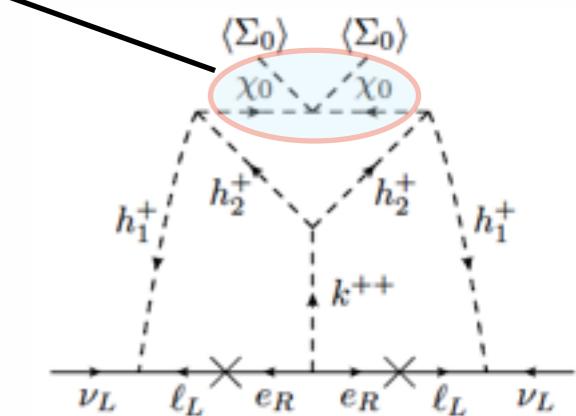
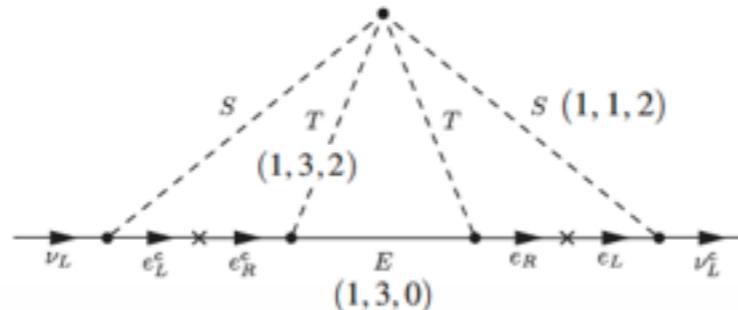
Top quark as a dark portal

L.Krauss,S.Nasri, and M.Trodden, 2002

John N. Ng, Alejandro de la Puente 2013

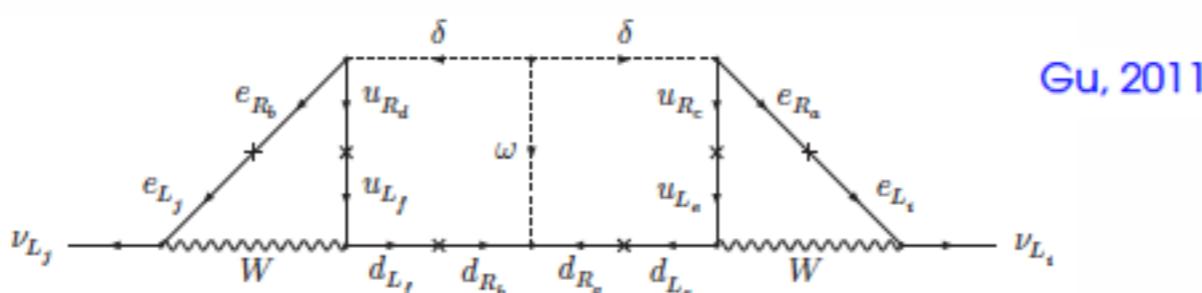


C.S.Chen,K.L.McDonald,S.Nasri, 2014



Hatanaka,Nishiwaki,Okada,Orikasa,
arXiv:1412.8664

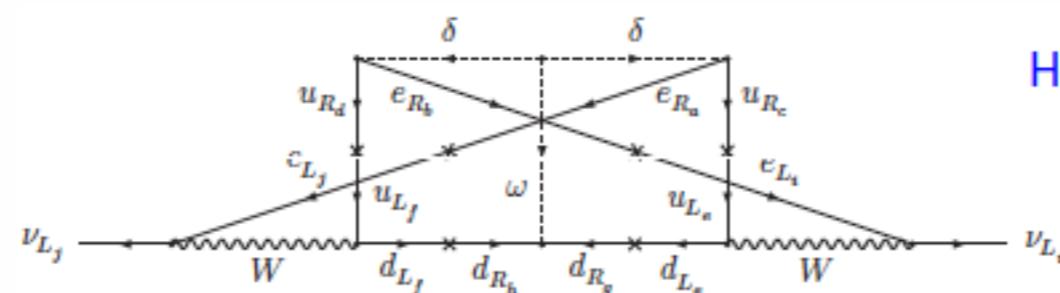
4-loop:



Gu, 2011

$$m_\nu \simeq 10^{-8} \text{ eV}$$

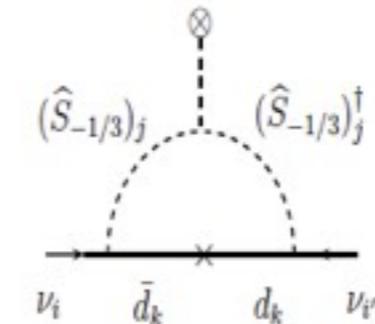
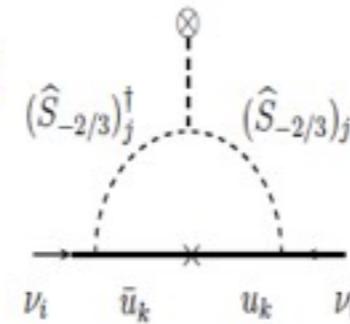
... because $d = 9$ 4-loop
Needs (Quasi)-Dirac ν 's
to explain oscillation data



HeLo et al., 2015

Other models with loops:

Hirsch et al. 1996, Aristizabal et al. 2008
Leptoquarks



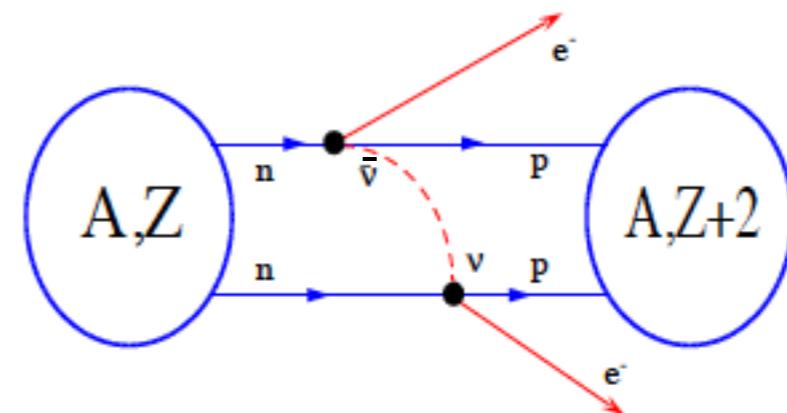
$$\nu_{\ell L} \frac{\mu}{(Y_L)_{\ell i} \frac{(d_{iL})^c}{m_{d_i}} \frac{(d_{iR})^{c\dagger}}{(Y_s)_{ij}} \frac{d_{jR}}{m_{d_j}} \frac{d_{jL}}{(Y_L^T)_{j\ell'}} (\nu_{\ell'L})^c}{S_{LQ}^{-\frac{1}{2}} S_{DQ}^{-\frac{3}{2}}} S_{LQ}^{\frac{1}{2}}$$

Suppressed $0\nu\beta\beta$ in all these loop models!

Double beta decay: $(A, Z) \rightarrow (A, Z + 2) + e^- + e^- + \bar{\nu}_e + \bar{\nu}_e$

Neutrinoless double beta decay:

0νββ $(A, Z) \rightarrow (A, Z + 2) + e^- + e^-$

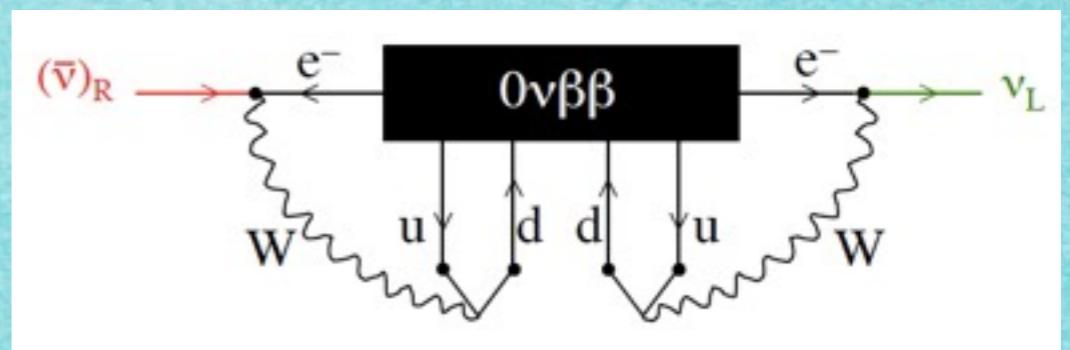
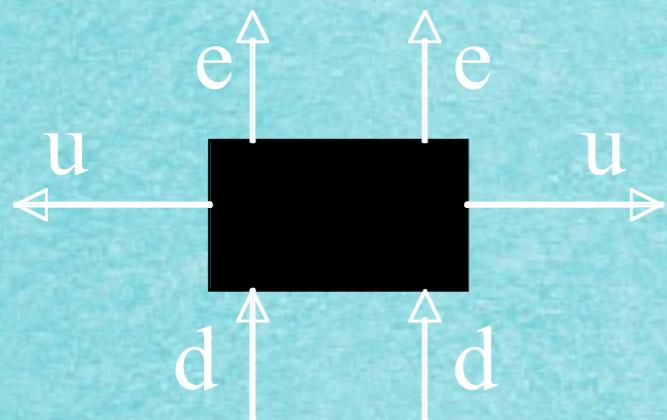


Majorana nature of ν $\nu \leftrightarrow \bar{\nu}$

“Black Box” theorem

J. Schechter and J.W.F. Valle, Phys.Rev. D 25, 2951 (1982)

``Any mechanism inducing the $0\nu\beta\beta$ decay produces an effective Majorana neutrino mass term, which must therefore contribute to this decay.”



$0\nu\beta\beta$ decay



Majorana neutrino mass



The multi-loop with $0\nu\beta\beta$ to m_ν is too small,
 $\sim \mathcal{O}(10^{-25})$ eV.

M.Duerr, M.Lindner, A.Merle, JHEP1106, 091 (2011)

The theorem does not state if the mechanism for $0\nu\beta\beta$ from M_ν is the dominant one.

In some models, the dominant contributions to $0\nu\beta\beta$ are generated without directly involving ν_M or M_ν .

● A special class of models to generate M_v

C.S.Chen+CQG+J.N.Ng,
PRD75,053004(2007)

	$SU(3)_C \otimes SU(2)_L \otimes U(1)_Y$
$L_a = (\nu_a, l_a)^T_L$	(1, 2, -1)
e_{aL}^c	(1, 1, 2)
$Q_a = (u_a, d_a)^T_L$	(3, 2, 1/3)
u_{aL}^c	($\bar{3}$, 1, -4/3)
d_{aL}^c	($\bar{3}$, 1, 2/3)
Φ	(1, 2, 1)

No ν_R added

Table 1: Matter and scalar multiplets of the Standard Model (SM)

New scalars: a triplet T (1,3,2) + a singlet Ψ (1,1,4)

$$\begin{aligned}
 V(\phi, T, \psi) = & -\mu^2 \phi^\dagger \phi + \lambda_\phi (\phi^\dagger \phi)^2 - \mu_T^2 \text{Tr}(T^\dagger T) + \lambda_T [\text{Tr}(T^\dagger T)]^2 + \lambda'_T \text{Tr}(T^\dagger T T^\dagger T) + m^2 \Psi^\dagger \Psi + \lambda_\Psi (\Psi^\dagger \Psi)^2 \\
 & + \kappa_1 \text{Tr}(\phi^\dagger \phi T^\dagger T) + \kappa_2 \phi^\dagger T T^\dagger \phi + \kappa_\Psi \phi^\dagger \phi \Psi^\dagger \Psi + \rho \text{Tr}(T^\dagger T \Psi^\dagger \Psi) \\
 & + [\lambda(\phi^T T \phi \Psi^\dagger) - M(\phi^T T^\dagger \phi) + \text{H.c.}]
 \end{aligned}$$

New Yukawa term:

$$Y_{ab} \bar{l}_{aR}^c l_{bR} \Psi$$

lepton # for Ψ is 2

No Yukawa coupling for the triplet:

~~LLT~~

Forbidden by a symmetry

Φ_1 and Φ_2

Z_2 or T -parity: $\Phi_1 \rightarrow \Phi_1$; $\Phi_2 \rightarrow -\Phi_2$; $T \rightarrow -T$; $L \rightarrow L$

C.S.Chen+CQG,
PRD82,105004(2010)

***Symmetry:** two Higgs doublets (Φ_1 and Φ_2)
with Z_2 -symmetry or T-parity

CS.Chen,CQG,PRD82,105004(2010)

T-parity: $\Phi_1 \rightarrow \Phi_1 ; \Phi_2 \rightarrow -\Phi_2 ; T \rightarrow -T ; L \rightarrow L$

~~LLT~~

$$\begin{aligned} V = & -\mu_1^2 \phi_1^\dagger \phi_1 + \lambda_1 (\phi_1^\dagger \phi_1)^2 - \mu_2^2 \phi_2^\dagger \phi_2 + \lambda_2 (\phi_2^\dagger \phi_2)^2 \\ & - \mu_T^2 \text{Tr}(T^\dagger T) + \lambda_T [\text{Tr}(T^\dagger T)]^2 + \lambda'_T \text{Tr}(T^\dagger T T^\dagger T) \\ & + m^2 \Psi^\dagger \Psi + \lambda_\Psi (\Psi^\dagger \Psi)^2 + \kappa_{\phi_1} \text{Tr}(\phi_1^\dagger \phi_1 T^\dagger T) \\ & + \kappa'_{\phi_1} \phi_1^\dagger T T^\dagger \phi_1 + \kappa_{\Psi_1} \phi_1^\dagger \phi_1 \Psi^\dagger \Psi \\ & + \kappa_{\phi_2} \text{Tr}(\phi_2^\dagger \phi_2 T^\dagger T) + \kappa'_{\phi_2} \phi_2^\dagger T T^\dagger \phi_2 \\ & + \kappa_{\Psi_2} \phi_2^\dagger \phi_2 \Psi^\dagger \Psi + \lambda_3 \phi_1^\dagger \phi_1 \phi_2^\dagger \phi_2 \\ & + \lambda_4 \phi_1^\dagger \phi_2 \phi_2^\dagger \phi_1 + \rho \text{Tr}(T^\dagger T \Psi^\dagger \Psi) + (M \phi_1^T T^\dagger \phi_2 \\ & + \lambda_5 \phi_1^\dagger \phi_2 \phi_1^\dagger \phi_2 + \lambda \tilde{\phi}_1^\dagger T \tilde{\phi}_2^* \Psi + \text{H.c.}), \end{aligned}$$

No effect for other couplings

Chen,CQG,Huang,Tsai, PRD87,077702 (2013)

Without Symmetry:

$\xi(1,N,2) + \Psi(1,1,4)$



~~LL~~ if $N > 3$

There are only three possible new renormalizable Yukawa interactions involving SM fermions with non-Higgs scalars:

$$\cancel{f_{ab} \bar{L}_a^c L_b S},$$

Zee model

$$y_{ab} \bar{\ell}_{R_a}^c \ell_{R_b} \Psi,$$

Zee-Babu model

$$\cancel{g_{ab} \bar{L}_a^c L_b T}$$

Type-II seesaw

(Colorless scalars)

We will consider higher dimensional multiplets so that
NO LL-like term is allowed in the Yukawa interactions.

We replace $s=(1,1,2)$ and $T=(1,3,2)$ by $\xi=(1,N,2)$

Chen, CQG, Huang, Tsai,
PRD 87, 077702 (2013)

$N > 3$ ($= 4, 5, 6, 7, \dots$) is the quantum # under $SU(2)_L$
and $Y=2$ is the hypercharge with $Q_{em} = I_3 + Y/2$

→ Multi High Charged Scalars e.g. for $N=5$ $\xi = (\xi^{+++}, \xi^{++}, \xi^+, \xi^0, \xi^-)^T$



The scalar potential reads

$$V(\Phi, \xi, \Psi) = -\mu_\Phi^2 |\Phi|^2 + \lambda_\Phi |\Phi|^4 + \mu_\xi^2 |\xi|^2 + \lambda_\xi^\alpha |\xi|_\alpha^4 + \mu_\Psi^2 |\Psi|^2 + \lambda_\Psi |\Psi|^4 \\ + \lambda_{\Phi\xi}^\beta (|\Phi|^2 |\xi|^2)_\beta + \lambda_{\Phi\Psi} |\Phi|^2 |\Psi|^2 + \lambda_{\xi\Psi} |\xi|^2 |\Psi|^2 \\ + [\mu \xi \xi \Psi + h.c.]$$

No $N=4, 6, 8, 10, \dots$, even dimensions
due to their antisymmetric products



$N=5, 7, \dots$, odd dimensions

Some detail calculations: doublet $\Phi(1,2,-1)$ + triplet $T(1,3,2)$ + singlet $\Psi(1,1,4)$

$$\phi = \begin{pmatrix} \phi^0 \\ \phi^- \end{pmatrix}_{-1} \quad T = \begin{pmatrix} T^0 & \frac{T^-}{\sqrt{2}} \\ \frac{T^-}{\sqrt{2}} & T^{--} \end{pmatrix}_{-2} \quad \psi_4^{++}$$

The most general potential is

$$\begin{aligned}
 V(\phi, T, \psi) = & -\mu^2 \phi^\dagger \phi + \lambda_\phi (\phi^\dagger \phi)^2 - \mu_T^2 Tr(T^\dagger T) + \lambda_T [Tr(T^\dagger T)]^2 + \lambda'_T Tr(T^\dagger T T^\dagger T) \\
 & + m^2 \psi^\dagger \psi + \lambda_\psi (\psi^\dagger \psi)^2 + \kappa_{T1} Tr(\phi^\dagger \phi T^\dagger T) + \kappa_{T2} \phi^\dagger T T^\dagger \phi + \kappa_\psi \phi^\dagger \phi \psi^\dagger \psi \\
 & + \rho Tr(T^\dagger T \psi^\dagger \psi) + \lambda (\phi^\dagger T \phi \psi + h.c.) \\
 = & -\mu^2 (\phi^+ \phi^- + |\phi^0|^2) + \lambda_\phi (|\phi^0|^4 + 2|\phi^0|^2 \phi^+ \phi^- + (\phi^+ \phi^-)^2) \\
 & - \mu_T^2 (|T^0|^2 + T^{++} T^{--} + T^+ T^-) \\
 & + (\lambda_T + \lambda'_T) [|T^0|^4 + 2|T^0|^2 T^+ T^- + (T^{++} T^{--})^2 + 2T^{++} T^{--} T^+ T^-] \\
 & + (\lambda_T + \frac{\lambda'_T}{2}) (T^+ T^-)^2 + \lambda'_T (T^0 T^{++} T^- T^- + T^{0*} T^{--} T^+ T^+) \\
 & + m^2 (\psi^{++} \psi^{--}) + \lambda_\psi (\psi^{++} \psi^{--})^2 \\
 & + (\kappa_{T1} + \kappa_{T2}) (|\phi^0|^2 |T^0|^2 + \phi^+ \phi^- T^{++} T^{--}) \\
 & + (\kappa_{T1} + \frac{\kappa_{T2}}{2}) (|\phi^0|^2 T^+ T^- + \phi^+ \phi^- T^+ T^-) + \kappa_{T1} (|T^0|^2 \phi^+ \phi^- + |\phi^0|^2 T^{++} T^{--}) \\
 & - \frac{\kappa_{T2}}{\sqrt{2}} (\phi^0 T^0 \phi^+ T^- + \phi^0 T^+ \phi^+ T^{--} + h.c.) + \kappa_\psi (|\phi^0|^2 + \phi^+ \phi^-) \psi^{++} \psi^{--} \\
 & + \rho (|T^0|^2 + T^{++} T^{--} + T^+ T^-) \psi^{++} \psi^{--} \\
 & + \lambda (T^{0*} \phi^- \phi^- \psi^{++} + \sqrt{2} \phi^{0*} T^- \phi^- \psi^{++} + \phi^{0*} \phi^{0*} T^{--} \psi^{++} + h.c.) \tag{4}
 \end{aligned}$$

The vacuum energy

$$-\frac{\mu^2 v^2}{2} + \frac{\lambda_\phi v^4}{4} - \frac{\mu_T^2 v_T^2}{2} + (\lambda_T + \lambda'_T) \frac{v_T^4}{4}$$

Tadpole terms

$$+[-\mu^2 v + \lambda_T v^3 + (\kappa_{T1} + \kappa_{T2}) \frac{vv_T^2}{2}] \phi_1 + [-\mu_T^2 v_T + (\lambda_T + \lambda'_T) v_T^3 + (\kappa_{T1} + \kappa_{T2}) \frac{v^2 v_T}{2}] T_1$$

The mass terms

$$\begin{aligned} & -\mu^2 \left(\frac{\phi_1^2 + \phi_2^2}{2} + \phi^+ \phi^- \right) + \frac{\lambda_T}{2} (3v^2 \phi_1^2 + v^2 \phi_2^2 + 2v^2 \phi^+ \phi^-) - \mu_T^2 \left(\frac{T_1^2 + T_2^2}{2} + T^+ T^- + T^{++} T^{--} \right) \\ & + \frac{\lambda_T + \lambda'_T}{4} (6v_T^2 T_1^2 + 2v_T^2 T_2^2 + 4v_T^2 T^+ T^-) + m^2 (\psi^{++} \psi^{--}) \\ & + \frac{\kappa_{T1} + \kappa_{T2}}{4} [v^2 (T_1^2 + T_2^2) + v_T^2 (\phi_1^2 + \phi_2^2) + 4vv_T \phi_1 T_1] + (\kappa_{T1} + \frac{\kappa_{T2}}{2}) \frac{v^2}{2} T^+ T^- \\ & + \kappa_{T1} \left(\frac{v_T^2}{2} \phi^+ \phi^- + \frac{v^2}{2} T^{++} T^{--} \right) - \frac{\kappa_{T2}}{\sqrt{2}} \frac{vv_T}{2} (\phi^+ T^- + \phi^- T^+) + \frac{\lambda v^2}{2} (T^{--} \psi^{++}) \\ & + \left(\frac{\kappa_\psi v^2}{2} + \frac{\rho v_T^2}{2} \right) \psi^{++} \psi^{--} \end{aligned} \tag{7}$$

Trilinear couplings

$$\begin{aligned} & \frac{\lambda_\phi}{4} (4v\phi_1^3 + 4v\phi_1\phi_2^2 + 8v\phi_1\phi^+\phi^-) + \frac{\lambda_{T1} + \lambda'_T}{4} (4v_T T_1^3 + 4v_T T_1 T_2^2 + 8v_T T_1 T^+ T^-) \\ & + \frac{\lambda'_T}{\sqrt{2}} (v_T T^{++} T^- T^- + v_T T^{--} T^+ T^+) + \frac{\kappa_{T1} + \kappa_{T2}}{2} [v\phi_1(T_1^2 + T_2^2) + v_T T_1(\phi_1^2 + \phi_2^2)] \\ & + (\kappa_{T1} + \frac{\kappa_{T2}}{2}) v\phi_1 T^+ T^- + \kappa_{T1} (v_T T_1 \phi^+ \phi^- + v\phi_1 T^{++} T^{--}) \\ & - \frac{\kappa_{T2}}{\sqrt{2}} [\frac{v}{2} (T_1 + iT_2) \phi^+ T^- + \frac{v_T}{2} (\phi_1 + i\phi_2) \phi^+ T^- + \frac{v}{\sqrt{2}} T^+ \phi^+ T^{--} + h.c.] \\ & + (\kappa_\psi v\phi_1 + \rho v_T T_1) \psi^{++} \psi^{--} \\ & + \lambda [\frac{v_T}{\sqrt{2}} \phi^- \phi^- \psi^{++} + v T^- \phi^- \psi^{++} + v(\phi_1 - i\phi_2) T^{--} \psi^{++} + h.c.] \end{aligned}$$

The quartic couplings

$$\begin{aligned}
& \frac{\lambda_\phi}{4} [\phi_1^4 + \phi_2^4 + \phi_1^2 \phi_2^2 + 4\phi_1^2 \phi^+ \phi^- + 4\phi_2^2 \phi^+ \phi^- + 4(\phi^+ \phi^-)^2] \\
& + (\lambda_T + \lambda'_T) [\frac{1}{4}(T_1^4 + T_2^4 + T_1^2 T_2^2) + (T_1^2 + T_2^2) T^+ T^- + (T^{++} T^{--})^2 + 2T^{++} T^{--} T^+ T^-] \\
& + (\lambda_T + \frac{\lambda'_T}{2})(T^+ T^-)^2 + \frac{\lambda'_T}{\sqrt{2}} [(T_1 + iT_2) T^{++} T^- T^- + (T_1 - iT_2) T T^{--} T^+ T^+] \\
& + (\kappa_{T1} + \kappa_{T2}) [\frac{1}{4}(\phi_1^2 + \phi_2^2)(T_1^2 + T_2^2) + \phi^+ \phi^- T^{++} T^{--}] \\
& + (\kappa_{T1} + \frac{\kappa_{T2}}{2})(\frac{\phi_1^2 + \phi_2^2}{2} T^+ T^- + \phi^+ \phi^- T^+ T^-) + \frac{\kappa_{T1}}{2} [(T_1^2 + T_2^2) \phi^+ \phi^- + (\phi_1^2 + \phi_2^2) T^{++} T^{--}] \\
& - \frac{\kappa_{T2}}{\sqrt{2}} [(\frac{\phi_1 T_1 - \phi_2 T_2}{2} + i \frac{\phi_2 T_1 + \phi_1 T_2}{2}) \phi^+ T^- + \frac{\phi_1 + i \phi_2}{\sqrt{2}} T^+ \phi^+ T^{--} + h.c.] \\
& + [\kappa_\psi \phi^+ \phi^- + \frac{\kappa_\psi}{2} (\phi_1^2 + \phi_2^2) + \frac{\rho}{2} (T_1^2 + T_2^2) + \rho T^{++} T^{--} + \rho T^+ T^-] \psi^{++} \psi^{--} \\
& + \lambda [\frac{T_1 - iT_2}{\sqrt{2}} \phi^- \phi^- \psi^{++} + (\phi_1 - i \phi_2) T^- \phi^- \psi^{++} + \frac{\phi_1^2 - \phi_2^2 - 2i \phi_1 \phi_2}{2} T^{--} \psi^{++} + h.c.] \quad (9)
\end{aligned}$$

Let

$$\phi^0 = \frac{v + \phi_1 + i\phi_2}{\sqrt{2}} \quad \text{and} \quad T^0 = \frac{v_T + T_1 + iT_2}{\sqrt{2}}$$

Mass mixing matrices

$$\begin{aligned}
& \left(\begin{array}{cc} \phi_1 & T_1 \end{array} \right) \left(\begin{array}{cc} -\frac{\mu^2}{2} + \frac{3}{2}\lambda_\phi v^2 + \frac{(\kappa_{T1}+\kappa_{T2})}{4} v_T^2 & \frac{(\kappa_{T1}+\kappa_{T2})}{2} v v_T \\ \frac{(\kappa_{T1}+\kappa_{T2})}{2} v v_T & -\mu_T^2 + \frac{3}{2}(\lambda_T + \lambda'_T) v_T^2 + \frac{(\kappa_{T1}+\kappa_{T2})}{4} v^2 \end{array} \right) \left(\begin{array}{c} \phi_1 \\ T_1 \end{array} \right) \\
& \left(\begin{array}{cc} \phi_2 & T_2 \end{array} \right) \left(\begin{array}{cc} -\frac{\mu^2}{2} + \frac{\lambda_\phi}{2} v^2 + \frac{(\kappa_{T1}+\kappa_{T2})}{4} v_T^2 & 0 \\ 0 & -\frac{\mu_T^2}{2} + \frac{(\lambda_T + \lambda'_T)}{2} v_T^2 + \frac{(\kappa_{T1}+\kappa_{T2})}{4} v^2 \end{array} \right) \left(\begin{array}{c} \phi_2 \\ T_2 \end{array} \right) \\
& \left(\begin{array}{cc} \phi^+ & T^+ \end{array} \right) \left(\begin{array}{cc} -\mu^2 + \lambda_\phi v^2 + \frac{\kappa_{T1}}{2} v_T^2 & -\frac{\kappa_{T2}}{2\sqrt{2}} v v_T \\ -\frac{\kappa_{T2}}{2\sqrt{2}} v v_T & -\mu_T^2 + (\lambda_T + \lambda'_T) v_T^2 + (\frac{\kappa_{T1}}{2} + \frac{\kappa_{T2}}{4}) v^2 \end{array} \right) \left(\begin{array}{c} \phi^- \\ T^- \end{array} \right) \\
& \left(\begin{array}{cc} T^{++} & \psi^{++} \end{array} \right) \left(\begin{array}{cc} -\mu_T^2 + \lambda_T v_T^2 v_T^2 + \frac{\kappa_{T1}}{2} v^2 & \frac{\lambda}{2} v^2 \\ \frac{\lambda}{2} v^2 & m^2 + \frac{\kappa_\psi}{2} v^2 + \frac{\rho}{2} v_T^2 \end{array} \right) \left(\begin{array}{c} T^{--} \\ \psi^{--} \end{array} \right)
\end{aligned}$$

The covariant derivative

$$D_\mu T = \partial_\mu T - i\frac{g}{2}([\vec{\sigma} \cdot \vec{W}_\mu T] + [\vec{\sigma} \cdot \vec{W}_\mu]^t) - i\frac{g'}{2}YB_\mu T$$

$$\begin{aligned} Tr[(D_\mu T)^\dagger (D^\mu T)] &= (\partial_\mu T^{0*})(\partial^\mu T^0) + (\partial_\mu T^-)(\partial^\mu T^+) + (\partial_\mu T^{--})(\partial^\mu T^{++}) \\ &+ (\partial^\mu T^0)[ig(T^{0*}W_\mu^3 + T^+W_\mu^-) - ig'B_\mu T^{0*}] + h.c. \\ &+ (\partial^\mu T^+)[-ig(T^{--}W_\mu^+ + T^0W_\mu^-) + ig'B_\mu T^-] + h.c. \\ &+ (\partial^\mu T^{++})[-ig(W_\mu^-T^- - T^{--}W_\mu^3) + ig'B_\mu T^{--}] + h.c. \\ &+ [g(T^{0*}W_\mu^3 + T^+W_\mu^-) - g'B_\mu T^{0*}]^2 \\ &+ [g(W_\mu^+T^{--} + T^0W_\mu^-) - ig'B_\mu T^-]^2 \\ &+ [g(W_\mu^-T^- - T^{--}W_\mu^3) - g'B_\mu T^{--}]^2 \end{aligned}$$

$$\begin{aligned} (D_\mu \phi)^\dagger (D^\mu \phi) &= (\partial_\mu \phi^0)(\partial^\mu \phi^{0*}) + (\partial_\mu \phi^+)(\partial^\mu \phi^-) \\ &+ \frac{1}{2}(\partial_\mu \phi^+)[ig(W_\mu^3 \phi^- - \sqrt{2}W_\mu^- \phi^0) + ig'B_\mu \phi^-] + h.c. \\ &+ \frac{1}{2}(\partial_\mu \phi^0)[ig(\sqrt{2}W^{-\mu} \phi^+ + W^{3\mu} \phi^{0*}) - ig'B^\mu \phi^{0*}] + h.c. \\ &+ \frac{1}{4}[g(\sqrt{2}W_\mu^- \phi^0 - W_\mu^3 \phi^-) - g'B^\mu \phi^-]^2 \\ &+ \frac{1}{4}[g(\sqrt{2}W_\mu^- \phi^+ - W_\mu^3 \phi^{0*}) - g'B_\mu \phi^{0*}]^2 \end{aligned}$$

$$\begin{aligned} (D_\mu \psi)^\dagger (D^\mu \psi) &= (\partial_\mu \psi^{--})(\partial^\mu \psi^{++}) + 2ig'B_\mu \psi^{--}(\partial^\mu \psi^{++}) + h.c. \\ &+ 4g'^2 B_\mu B^\mu \psi^{--} \psi^{++} \end{aligned}$$

The masses of gauge bosons are

$$M_W^2 W_\mu^+ W^{-\mu} = \frac{g^2}{4}(v^2 + 2v_T^2) W_\mu^+ W^{-\mu}$$

$$M_W^2 = \frac{g^2}{4}(v^2 + 2v_T^2)$$

The mixing matrix

$$\frac{1}{2} \begin{pmatrix} W_\mu^3 & B_\mu \end{pmatrix} \begin{pmatrix} \frac{g^2}{8}(v^2 + 4v_T^2) & -\frac{gg'}{8}(v^2 - 4v_T^2) \\ -\frac{gg'}{8}(v^2 - 4v_T^2) & \frac{g'^2}{8}(v^2 + 4v_T^2) \end{pmatrix} \begin{pmatrix} W^{3\mu} \\ B^\mu \end{pmatrix}$$

$$\begin{pmatrix} Z_\mu \\ A_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta_W & -\sin \theta_W \\ \sin \theta_W & \cos \theta_W \end{pmatrix} \begin{pmatrix} W_\mu^3 \\ B_\mu \end{pmatrix}$$

$$\begin{pmatrix} Z_\mu \\ A_\mu \end{pmatrix} = \begin{pmatrix} \cos \theta_W & -\sin \theta_W \\ \sin \theta_W & \cos \theta_W \end{pmatrix} \begin{pmatrix} W_\mu^3 \\ B_\mu \end{pmatrix}$$

$$\tan 2\theta_W = \frac{2\frac{g'}{g}\frac{v^2 - 4v_T^2}{v^2 + 4v_T^2}}{1 - \frac{g'^2}{g^2}}$$

$$M_Z^2 = \frac{g^2}{4 \cos^2 \theta_W} (v^2 + 4v_T^2)$$

$$\rho = \frac{M_W^2}{M_Z^2 \cos^2 \theta_W} = \frac{v^2 + 2v_T^2}{v^2 + 4v_T^2}$$

$$\begin{pmatrix} G^- \\ P^- \end{pmatrix} = \frac{1}{\sqrt{v^2 + 2v_T^2}} \begin{pmatrix} v & \sqrt{2}v_T \\ -\sqrt{2}v_T & v \end{pmatrix} \begin{pmatrix} \phi^- \\ T^- \end{pmatrix}$$

Gauge-scalar trilinear interactions :

$$W_\mu^+ - T^+ - T^0 : \quad \frac{ig}{\sqrt{2}} W^{-\mu} [(\partial_\mu T_1) T^+ - (\partial_\mu T^+) T_1] + \frac{g}{\sqrt{2}} W^{-\mu} [(\partial_\mu T^+) T_2 - (\partial_\mu T_2) T^+] + h.c.$$

$$W_\mu^+ - T^+ - T^{--} : \quad ig W^{+\mu} [(\partial_\mu T^{--}) T^+ - (\partial_\mu T^+) T^{--}] + h.c.$$

$$Z_\mu(A_\mu) - T^+ - T^-$$
$$ig' \sin \theta_W (\partial_\mu T^-) T^+ Z^\mu - ig \sin \theta_W (\partial_\mu T^-) T^+ A^\mu + h.c$$

$$A_\mu - T^{++} - T^{--}$$
$$-i \frac{g}{\cos \theta_W} (\partial_\mu T^{--}) T^{++} A^\mu$$
$$Z_\mu - T^0 - T^0$$
$$\frac{g}{\cos \theta_W} Z^\mu [(\partial_\mu T_1) T_2 - (\partial_\mu T_2) T_1]$$

$$A_\mu - \phi^+ - \phi^-$$
$$-\frac{ig}{2 \cos \theta} (\partial_\mu \phi^-) \phi^+ A^\mu + h.c.$$

$$W_\mu^+ - \phi^- - \phi^0$$
$$\frac{ig}{2} W_\mu^+ [(\partial^\mu \phi^-) \phi_1 - (\partial^\mu \phi_1) \phi^-] + \frac{g}{2} W_\mu^+ [(\partial^\mu \phi^-) \phi_2 - (\partial^\mu \phi_2) \phi^-] + h.c.$$

Gauge-scalar trilinear interactions :

$W_\mu^+ - T^+ - T^0 :$	$\frac{ig}{\sqrt{2}} W^{-\mu} [(\partial_\mu T_1) T^+ - (\partial_\mu T^+) T_1] + \frac{g}{\sqrt{2}} W^{-\mu} [(\partial_\mu T^+) T_2 - (\partial_\mu T_2) T^+] + h.c.$
$W_\mu^+ - T^+ - T^{--} :$	$ig W^{+\mu} [(\partial_\mu T^{--}) T^+ - (\partial_\mu T^+) T^{--}] + h.c.$
$Z_\mu(A_\mu) - T^+ - T^-$	$ig' \sin \theta_W (\partial_\mu T^-) T^+ Z^\mu - ig \sin \theta_W (\partial_\mu T^-) T^+ A^\mu + h.c$
$A_\mu - T^{++} - T^{--}$	$-i \frac{g}{\cos \theta_W} (\partial_\mu T^{--}) T^{++} A^\mu$
$A_\mu - \phi^+ - \phi^-$	$-\frac{ig}{2 \cos \theta} (\partial_\mu \phi^-) \phi^+ A^\mu + h.c.$
$W_\mu^+ - \phi^- - \phi^0$	$\frac{ig}{2} W_\mu^+ [(\partial^\mu \phi^-) \phi_1 - (\partial^\mu \phi_1) \phi^-] + \frac{g}{2} W_\mu^+ [(\partial^\mu \phi^-) \phi_2 - (\partial^\mu \phi_2) \phi^-] + h.c.$
$Z_\mu - \phi^0 - \phi^0$	$\frac{g}{2 \cos \theta_W} Z^\mu [(\partial_\mu \phi_1) \phi_2 - (\partial_\mu \phi_2) \phi_1]$
$T^+ - W_\mu^- - Z_\mu$	$\frac{g^2 + g^2 \sin^2 \theta_W}{\cos \theta_W} \left(\frac{v_T}{\sqrt{2}} \right) Z^\mu W_\mu^- T^+ + h.c.$
$T^0 - W^{+\mu} - W_\mu^-$	$g^2 v_T T_1 W^{+\mu} W_\mu^-$
$\phi^0 - W^{+\mu} - W_\mu^-$	$\frac{g^2 v}{2} W^{+\mu} W_\mu^- \phi_1$
$\phi^0 - Z^\mu - Z_\mu$	$\frac{g^2 v}{4 \cos^2 \theta_W} Z^\mu Z_\mu \phi_1$
$T^{--} - W^{+\mu} - W_\mu^+$	$\frac{g^2}{\sqrt{2}} v_T T^{--} W^{+\mu} W_\mu^+ + h.c.$
$W^{-\mu} - A_\mu - \phi^+$	$-\frac{g^2 v}{4 \cos \theta_W} W^{-\mu} A_\mu \phi^+ + h.c$
$W^{-\mu} - Z_\mu - \phi^+$	$\frac{g^2 v}{4 \cos \theta_W} W^{-\mu} Z_\mu \phi^+ + h.c.$

$$T^+ - T^- - Z_\mu - Z_\mu$$

$$-g'^2 \sin^2 \theta_W T^+ T^- Z^\mu Z_\mu$$

$$T^+ - T^- - A^\mu - A_\mu$$

$$g'^2 \cos^2 \theta_W T^+ T^- A^\mu A_\mu$$

$$T^+ - T^- - Z^\mu - A_\mu$$

$$-2g'^2 \sin \theta_W \cos \theta_W T^+ T^- Z^\mu A_\mu$$

$$T^{--} - T^+ - W^{+\mu} - Z_\mu$$

$$gg' \sin \theta_W W^{+\mu} Z_\mu T^+ T^{--} + h.c.$$

$$T^{--} - T^+ - W^{+\mu} - A_\mu$$

$$-(gg' \cos \theta_W + \frac{g^2}{\cos \theta_W}) W^{+\mu} A_\mu T^+ T^{--} + h.c.$$

$$T^0 - T^0 - W^{+\mu} - W_\mu^-$$

$$\frac{g^2}{2}(T_1^2 + T_2^2) W^{+\mu} W_\mu^-$$

$$T^0 - T^+ - W^{-\mu} - Z_\mu$$

$$\frac{g^2}{\sqrt{2}}(\frac{1 + \sin^2 \theta_W}{\cos \theta_W}) T^+ Z_\mu W^{-\mu} (T_1 + iT_2) + h.c.$$

$$T^0 - T^+ - W^{-\mu} - A_\mu$$

$$-\frac{g^2 \sin \theta_W}{\sqrt{2}} T^+ A_\mu W^{-\mu} (T_1 + iT_2) + h.c$$

$$T^0 - T^{--} - W^{+\mu} - W_\mu^+$$

$$\frac{g^2}{\sqrt{2}}(T_1 - iT_1) T^{--} W^{+\mu} W_\mu^+ + h.c.$$

$$T^0 - T^0 - Z_\mu - Z^\mu$$

$$\frac{g^2}{2 \cos^2 \theta_W} (T_1^2 + T_2^2) Z^\mu Z_\mu$$

$$T^+ - T^- - W^{+\mu} - W_\mu^-$$

$$2 g^2 W^{+\mu} W_\mu^- T^+ T^-$$

$$T^{++} - T^{--} - W^{+\mu} - W_\mu^-$$

$$g^2 W^{+\mu} W_\mu^- T^{++} T^{--}$$

$$T^{++} - T^{--} - A^\mu - A_\mu$$

$$\frac{g^2}{\cos^2 \theta_W} A^\mu A_\mu T^{++} T^{--}$$

$$\phi^+ - \phi^- - A^\mu - A_\mu$$

$$\frac{g^2}{4 \cos^2 \theta_W} A^\mu A_\mu \phi^+ \phi^-$$

$$\phi^0 - \phi^0 - W^{+\mu} - W_\mu^-$$

$$\frac{g^2}{4} W^{+\mu} W_\mu^- (\phi_1^2 + \phi_2^2)$$

$$\phi^0 - \phi^+ - W^{-\mu} - A_\mu$$

$$-\frac{g^2}{4 \cos \theta_W} W^{-\mu} A_\mu \phi^+ (\phi_1 + i \phi_2) + h.c.$$

$$\phi^0 - \phi^0 - Z^\mu - Z_\mu$$

$$\frac{g^2}{8 \cos^2 \theta_W} Z^\mu Z_\mu (\phi_1^2 + \phi_2^2)$$

$$\phi^+ - \phi^- - W^{-\mu} - W_\mu^+$$

$$\frac{g^2}{2} W^{+\mu} W_\mu^- \phi^+ \phi^-$$

$$\phi^0 - \phi^+ - W^{-\mu} - Z_\mu$$

$$\frac{g^2}{4 \cos \theta_W} W^{-\mu} Z_\mu \phi^+ (\phi_1 + i \phi_2) + h.c.$$

$$Z^\mu (A^\mu) - Z_\mu (A_\mu) - \psi^{++} - \psi^{--}$$

$$4g'^2 \sin^2 \theta_W Z^\mu Z_\mu \psi^{++} \psi^{--} + 4g'^2 \cos^2 \theta_W A^\mu A_\mu \psi^{++} \psi^{--} - 8g'^2 \sin \theta_W \cos \theta_W A^\mu Z_\mu$$

Constraints on the models:

VEVs: $\langle \phi^0 \rangle \equiv \frac{v}{\sqrt{2}}$ and $\langle T^0 \rangle \equiv \frac{v_T}{\sqrt{2}}$.

$$M_W^2 = \frac{g^2}{4}(v^2 + 2v_T^2), \quad M_Z^2 = \frac{g^2}{4 \cos^2 \theta_W}(v^2 + 4v_T^2),$$

$$\rho = 1.0002^{+0.0007}_{-0.0004} \quad \rightarrow \quad v_T < 4.41 \text{ GeV}$$

Two doubly charged scalars:

$$T = \begin{pmatrix} T^0 & \frac{T^-}{\sqrt{2}} \\ \frac{T^-}{\sqrt{2}} & T^{--} \end{pmatrix} \quad \text{and} \quad \Psi^{++}$$

Mass eigenstates:

or for N=5 $\xi = (\xi^{+++}, \xi^{++}, \xi^+, \xi^0, \xi^-)^T$

$$\begin{pmatrix} P_1^{\pm\pm} \\ P_2^{\pm\pm} \end{pmatrix} = \begin{pmatrix} \cos \delta & \sin \delta \\ -\sin \delta & \cos \delta \end{pmatrix} \begin{pmatrix} T^{\pm\pm} \\ \Psi^{\pm\pm} \end{pmatrix}$$

$$\sin 2\delta = \left[1 + \left(\frac{2m^2 + (2\lambda'_T + \rho)v_T^2}{2\lambda v^2} + \frac{\kappa_2 + \kappa_\Psi}{2\lambda} - \frac{\omega}{\lambda} \right)^2 \right]^{-\frac{1}{2}}$$

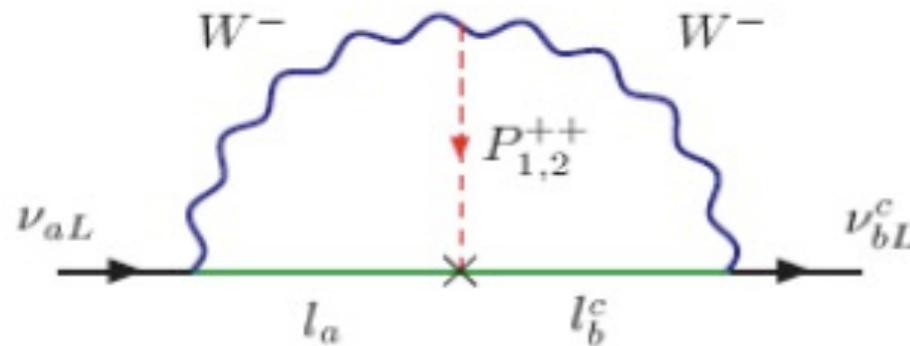
$$M_{P_{1,2}}^2 = \frac{1}{2} \left[a + c \mp \sqrt{4b^2 + (c-a)^2} \right]$$

$$\omega \equiv \frac{M}{\sqrt{2}v_T}$$

$$a = \frac{1}{2}(2\omega - \kappa_2)v^2 - \lambda'_T v_T^2, \quad b = \frac{1}{2}\lambda v^2, \quad c = m^2 + \frac{1}{2}(\kappa_\Psi v^2 + \rho v_T^2).$$

● Neutrino mass generation:

The neutrino masses are generated radiatively at two-loop level



$$a, b = e, \mu, \tau.$$

$$(m_\nu)_{ab} = \frac{1}{\sqrt{2}} g^4 m_a m_b v_T Y_{ab} \sin(2\delta) [I(M_W^2, M_{P_1}^2, m_a, m_b) - I(M_W^2, M_{P_2}^2, m_a, m_b)]$$

$$I(M_W^2, M_{P_i}^2, m_a^2, m_b^2) = \int \frac{d^4 q}{(2\pi)^4} \int \frac{d^4 k}{(2\pi)^4} \frac{1}{k^2 - m_a^2} \frac{1}{k^2 - M_W^2} \frac{1}{q^2 - M_W^2} \frac{1}{q^2 - m_b^2} \frac{1}{(k-q)^2 - M_{P_i}^2}.$$

$$M_{P_{1,2}} > M_W$$

$$I(M_W^2, M_{P_i}^2, 0, 0) \sim \frac{1}{(4\pi)^4} \frac{1}{M_{P_i}^2} \log^2 \left(\frac{M_W^2}{M_{P_i}^2} \right)$$

$$\begin{aligned} m_\nu &= \tilde{f}(M_{P_1}, M_{P_2}) \times \begin{pmatrix} m_e^2 Y_{ee} & m_e m_\mu Y_{e\mu} & m_e m_\tau Y_{e\tau} \\ m_e m_\mu Y_{e\mu} & m_\mu^2 Y_{\mu\mu} & m_\tau m_\mu Y_{\mu\tau} \\ m_e m_\tau Y_{e\tau} & m_\tau m_\mu Y_{\mu\tau} & m_\tau^2 Y_{\tau\tau} \end{pmatrix} \\ &= f(M_{P_1}, M_{P_2}) \times \begin{pmatrix} 2.6 \times 10^{-7} Y_{ee} & 5.4 \times 10^{-5} Y_{e\mu} & 9.1 \times 10^{-4} Y_{e\tau} \\ 5.4 \times 10^{-5} Y_{e\mu} & 1.1 \times 10^{-2} Y_{\mu\mu} & 0.19 Y_{\mu\tau} \\ 9.1 \times 10^{-4} Y_{e\tau} & 0.19 Y_{\mu\tau} & 3.17 Y_{\tau\tau} \end{pmatrix} \end{aligned}$$

normal hierarchy:

$$\begin{pmatrix} \varepsilon' & \varepsilon & \varepsilon \\ \varepsilon & 1+\eta & 1+\eta \\ \varepsilon & 1+\eta & 1+\eta \end{pmatrix}$$

$$\tilde{f}(M_{P_1}, M_{P_2}) = \frac{\sqrt{2} g^4 v_T \sin(2\delta)}{128\pi^4} \left[\frac{1}{M_{P_1}^2} \log^2 \left(\frac{M_W}{M_{P_1}} \right) - \frac{1}{M_{P_2}^2} \log^2 \left(\frac{M_W}{M_{P_2}} \right) \right]$$

$$f = \tilde{f} \times (1 \text{ GeV}^2)$$

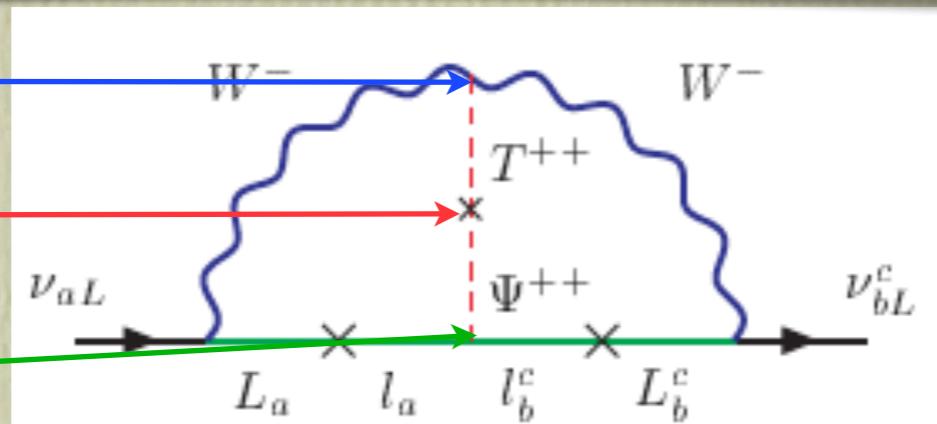
$$\begin{aligned} Y_{ee} < 0.17, & \quad Y_{e\mu} < 0.2, & \quad Y_{e\tau} < 0.2 \\ Y_{\mu\mu} < 3.5, & \quad Y_{\mu\tau} < 0.2, & \quad Y_{\tau\tau} < 0.02 \end{aligned}$$

The neutrino masses are generated radiatively at two-loop level

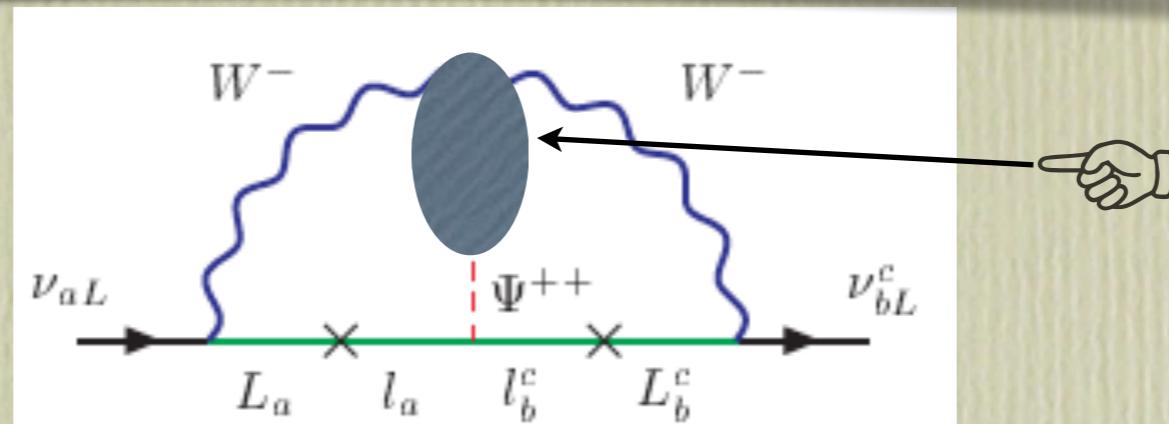
$$Tr[(D_\mu T)^\dagger(D^\mu T)]$$

$$\lambda \Phi^T T \Phi \Psi$$

$$Y_{ab} \overline{l_{aR}^c} l_{bR} \Psi$$

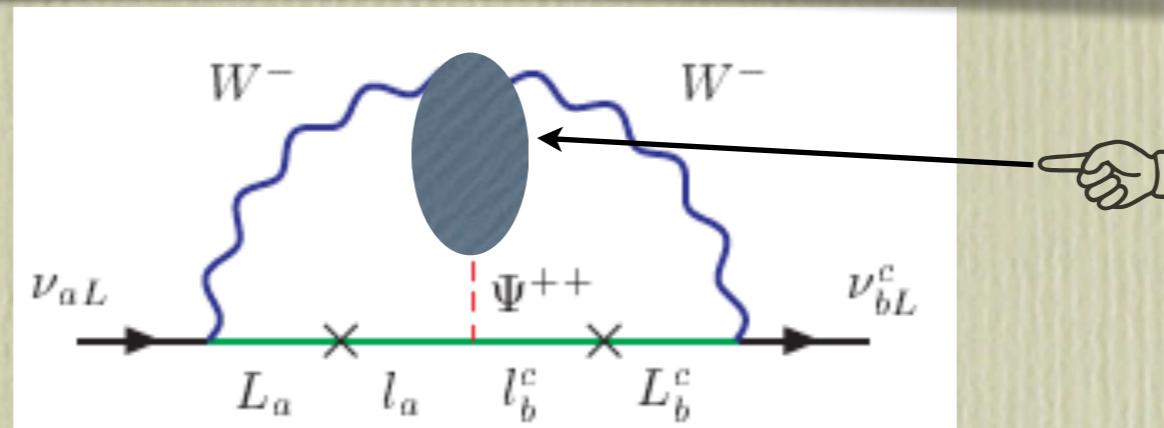


The neutrino masses are generated radiatively at two-loop level



$W^\pm W^\pm \Psi^\mp \mp$

The neutrino masses are generated radiatively at two-loop level



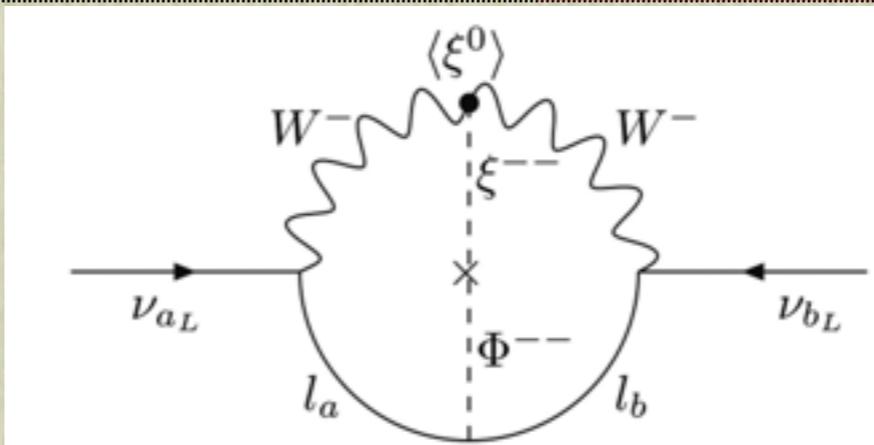
$W^\pm W^\pm \Psi^\mp \mp$

New model: a multiplet ξ (1,5,2) + a singlet Φ (1,1,4)

Chen,CQG,Huang,Tsai,
PRD87, 077702 (2014)

Without Symmetry:

~~$I\!\!L\xi$~~



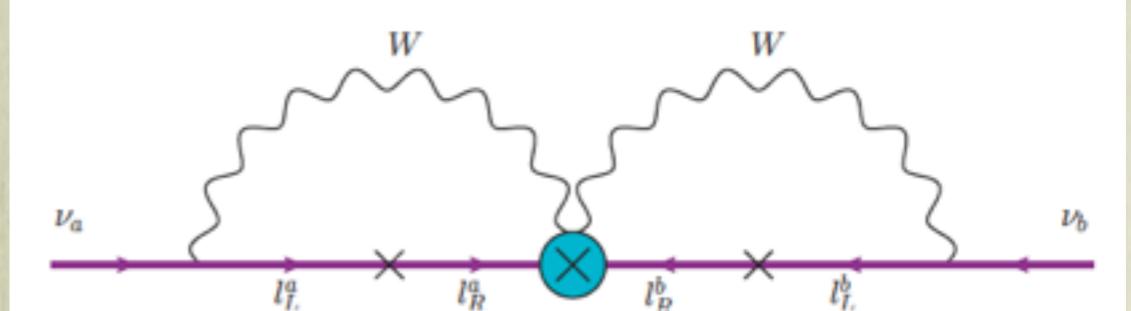
M.Gustafsson,J.M.No,M.A.Rivera,
PRD90, 013012 (2014)

dimension-9
L violating O

$$\mathcal{O}^9 \equiv C_{ab}^{(9)} \bar{\ell}^c R_a \ell_{R_b} \left[(D_\mu H)^T i\sigma_2 H \right]^2$$



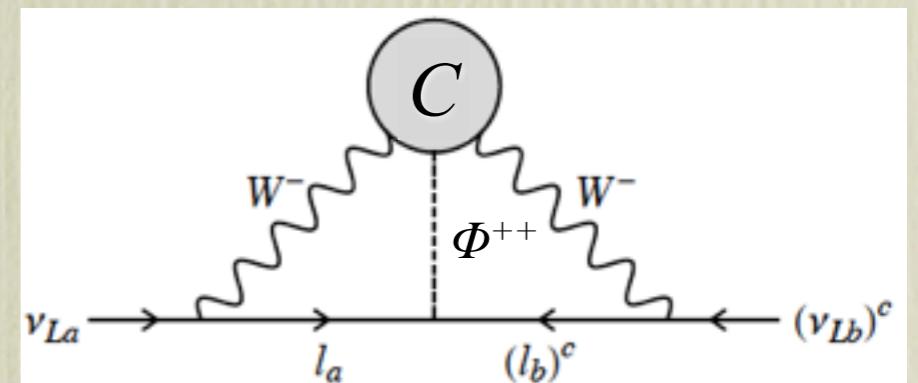
$$\frac{C_{ab}^{(9)}}{\tilde{\Lambda}} \bar{\ell}^c R_a \ell_{R_b} W_\mu^+ W^{+\mu}$$



S.F.King,A.Merle,L.Panizzi,
JHEP1411, 124 (2014)

dimension-7 O

$$\begin{aligned} \mathcal{O}_7^{(a)} &= \Phi (H \otimes H)_{\underline{3}} [(D_\mu H) \otimes (D^\mu H)]_{\underline{3}} \\ \mathcal{O}_7^{(b)} &= \Phi [(D_\mu H) \otimes H]_{\underline{1}} [(D^\mu H) \otimes H]_{\underline{1}} \\ \mathcal{O}_7^{(c)} &= \Phi [(D_\mu H) \otimes H]_{\underline{3}} [(D^\mu H) \otimes H]_{\underline{3}} \end{aligned}$$



$$(M_\nu)_{ab} \propto m_a m_b Y_{ab}$$

For $Y_{ab} \sim O(1)$ $(M_\nu)_{ee} \ll (M_\nu)_{e\mu} \ll (M_\nu)_{e\tau} \ll (M_\nu)_{\mu\mu} \ll (M_\nu)_{\mu\tau} \ll (M_\nu)_{\tau\tau}$

→ **Normal hierarchy**

Z.-z.Xing,PLB530,159(2002);PLB539,85(2002);
Frampton,Glashow,Marfatia,PLB536,79(2002);
W.L.Guo,Z.-z.Xing,PRD67,053002(2003);
.....

With $(M_\nu)_{ee} \simeq (M_\nu)_{e\mu} \simeq 0$ and the center values of PDG2014 :

$$\sin^2 \theta_{12} = 0.308 \pm 0.017, \sin^2 \theta_{23} = 0.437^{+0.033}_{-0.023}, \sin^2 \theta_{13} = 0.0234^{+0.0020}_{-0.0019},$$

$$\Delta m_{21}^2 = (7.54^{+0.26}_{-0.22}) \times 10^{-5} \text{ eV}, \Delta m_{32}^2 = (2.43 \pm 0.06) \times 10^{-3} \text{ eV},$$

CQG+L.H.Tsai, Annals Phys. 365,210 (2016)
CQG, MPLA30,1530018(2015)

$$M_\nu \simeq \begin{pmatrix} 0 & 0 & 1.0 e^{-i\eta} \\ 0 & 2.4 e^{i(\frac{\pi}{2}+\eta)} & 2.3 e^{i\frac{\pi}{2}} \\ 1.0 e^{-i\eta} & 2.3 e^{i\frac{\pi}{2}} & 2.8 e^{i(\frac{\pi}{2}+\frac{2\eta}{3})} \end{pmatrix} \times 10^{-2} \text{ eV}$$

Dirac and Majorana phases:

$$\delta = \frac{3}{2}\pi - \frac{3}{2}\eta, \alpha_{21} = \pi + \frac{3}{2}\eta, \alpha_{31} = \frac{3}{2}\pi - \frac{1}{2}\eta, (\eta \simeq 0.07\pi)$$

Agree well with
Global Best-Fit

$$\delta/\pi = 1.39^{+0.38}_{-0.27}$$

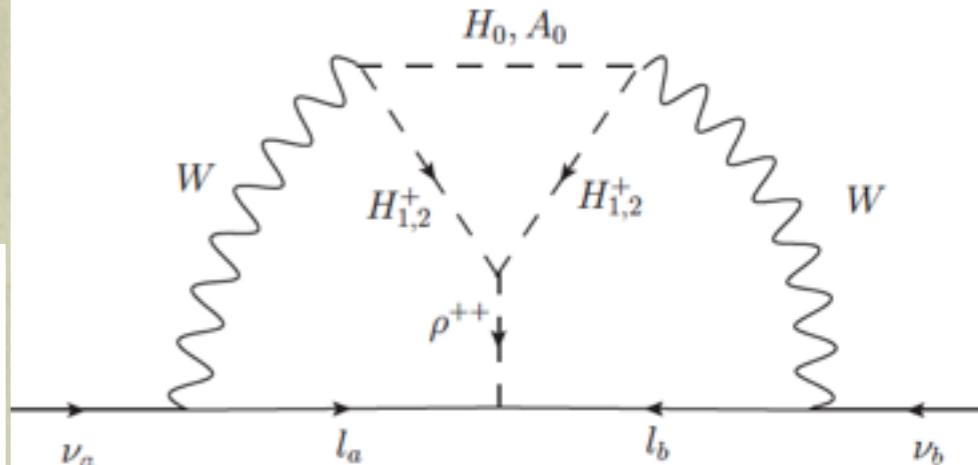
At three-loop level

THE COCKTAIL MODEL FOR NEUTRINO MASSES

*M.Gustafsson,J.M.No,M.A.Rivera,
PRL110, 211802 (2013); Erratum,
PRL112, 259902 (2014).*

	$SU(2)_L$	$U(1)_Y$	Z_2
Φ_2	2	1	-
S^+	1	2	-
ρ^{++}	1	4	+

$$-\mathcal{L}_{\text{dark}} = \frac{\lambda_5}{2}(\Phi_1^\dagger \Phi_2)^2 + \kappa_1 \Phi_2^T i\sigma_2 \Phi_1 S^- + \kappa_2 \rho^{++} S^- S^- \\ + \xi \Phi_2^T i\sigma_2 \Phi_1 S^+ \rho^{--} + C_{ab} \bar{\ell}^c{}_{aR} \ell_{bR} \rho^{++} + \text{h.c.},$$



$$(m_\nu)_{ab} = (x_a C_{ab} x_b) \frac{s_{2\beta}}{(16\pi^2)^3} (\mathcal{A}_1 \mathcal{I}_1 + \mathcal{A}_2 \mathcal{I}_2),$$

$$\mathcal{A}_1 = \frac{[\kappa_2 s_{2\beta} + (\xi v) c_{2\beta}]}{m_\rho^2} \frac{(\Delta m_+^2)^2 \Delta m_0^2}{m_\rho^2 v^2},$$

$$\mathcal{A}_2 = \frac{\xi v}{m_\rho^2} \frac{\Delta m_+^2 \Delta m_0^2}{v^2},$$

$$\Delta m_+^2 = m_{H_2}^2 - m_{H_1}^2, \quad \Delta m_0^2 = m_{A^0}^2 - m_{H^0}^2$$

$$x_a = m_a/v, \quad I_1 \sim I_2 \sim O(1)$$

For $C_{ab} \sim O(1)$

$$m_\nu = \begin{pmatrix} \approx 0 & \approx 0 & 10.1 \\ \approx 0 & -5.01 & 0.0980 \\ 10.1 & 0.0980 & -4.77 \end{pmatrix} \times 10^{-3} + i \begin{pmatrix} \approx 0 & \approx 0 & 0.23 \\ \approx 0 & -2.37 & -2.33 \\ 0.23 & -2.33 & -2.74 \end{pmatrix} \times 10^{-2} \text{ eV}$$

$$C_{ab} = \begin{pmatrix} \leq \mathcal{O}(10^{-2}) & \leq \mathcal{O}(10^{-2}) & e^{0.224i} \\ \leq \mathcal{O}(10^{-2}) & 1.90 \times 10^{-1} e^{-1.78i} & 1.08 \times 10^{-2} e^{-1.56i} \\ e^{0.224i} & 1.08 \times 10^{-2} e^{-1.56i} & 7.73 \times 10^{-4} e^{-1.74i} \end{pmatrix} \times |C_{e\tau}|$$

• $0\nu\beta\beta$ decays:

Y_{ee}

C.S.Chen+CQG+J.N.Ng,
PRD75,053004(2007)

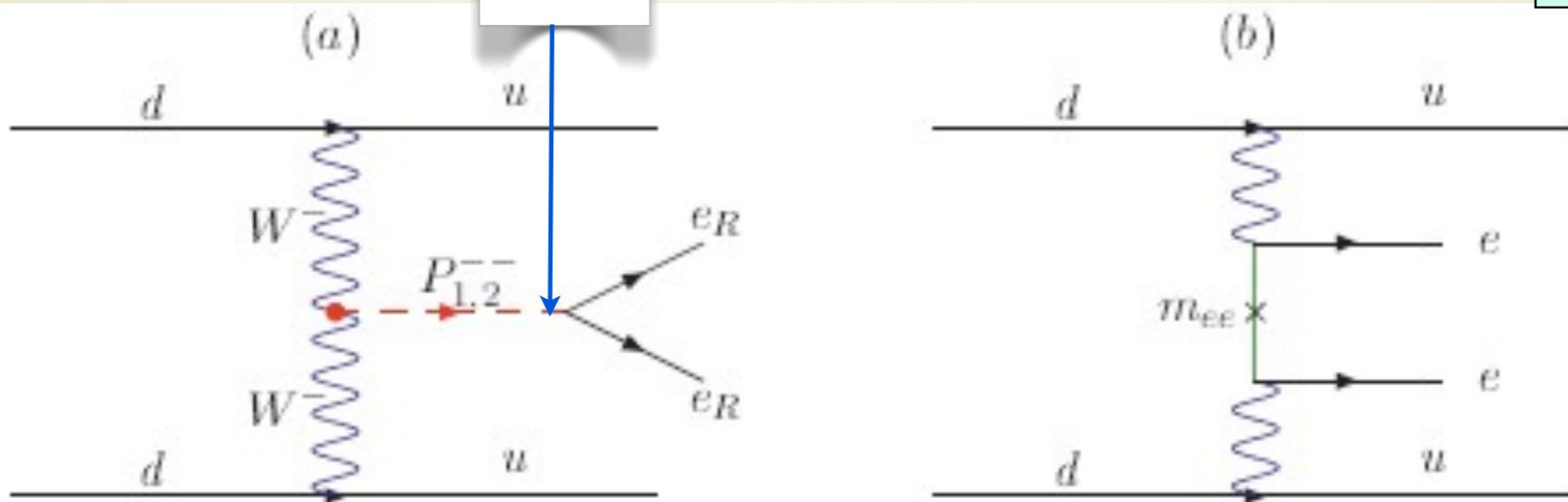


Figure 9: $0\nu\beta\beta$ decays via exchange of: (a) doubly charged Higgs and (b) light Majorana neutrinos.

$$A_{P_{1,2}^{--}} \sim \frac{g^4 Y_{ee} v_T \sin 2\delta}{16\sqrt{2} M_W^4} \left(\frac{1}{M_{P_1}^2} - \frac{1}{M_{P_2}^2} \right)$$

>>

$$A_\nu \sim \frac{g^4}{M_W^4} \frac{m_{ee}}{\langle p \rangle^2}$$

$\langle p \rangle \sim 0.1 \text{ GeV}$

$$A_\nu / A_{P_{1,2}^{--}} \lesssim 10^{-7}$$

握手 Black box theorem is irrelevant as $0\nu\beta\beta$ dominantly arises from the SD contribution

No other strong constraint on Y_{ee} except the rate of $0\nu\beta\beta$ itself. So the rate of $0\nu\beta\beta$ can be very large, which would correlate with the LHC searches.

• $0\nu\beta\beta$ decays:

Y_{ee}

C.S.Chen+CQG+J.N.Ng,
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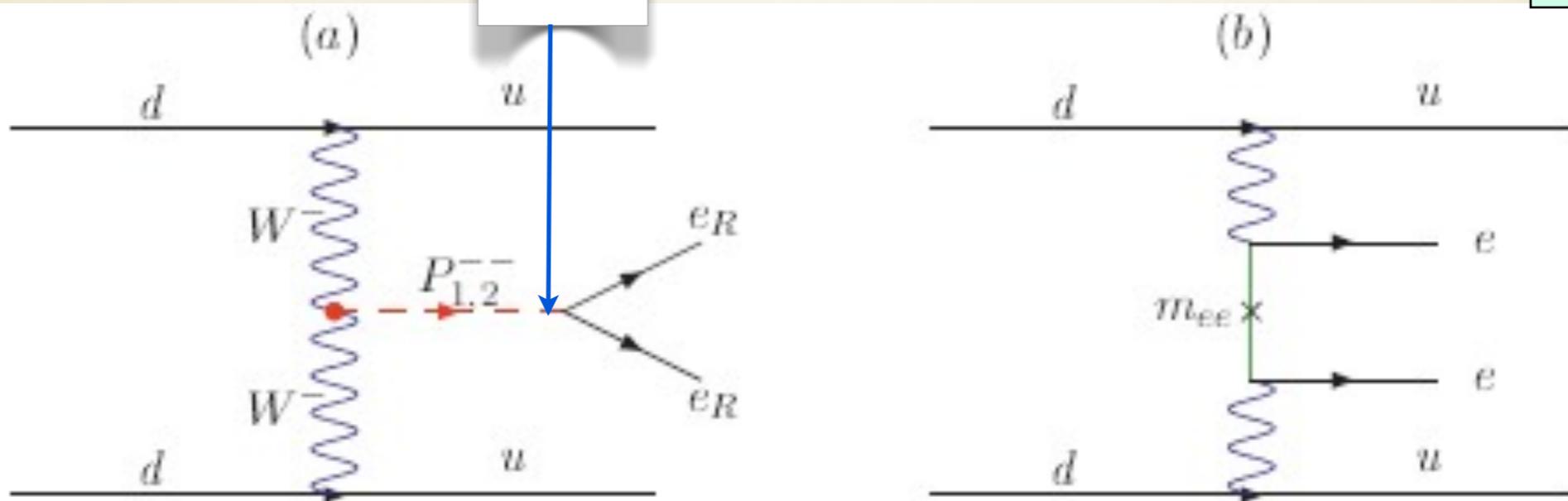


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$$\langle p \rangle \sim 0.1 \text{ GeV}$$

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The smallness of this ratio is due to the fact that in our model, m_{ee} is suppressed not only by a two-loop factor, it is also suppressed by the electron mass factor $(m_e/M_W)^2$ coming from the doubly charged scalar coupling.

握手 Black box theorem is irrelevant as $0\nu\beta\beta$ dominantly arises from the SD contribution

No other strong constraint on Y_{ee} except the rate of $0\nu\beta\beta$ itself. So the rate of $0\nu\beta\beta$ can be very large, which would correlate with the LHC searches.

Currently, $Y_{ee} < 0(10^{-2})$ for $M_P \sim 0(1) \text{ TeV}$

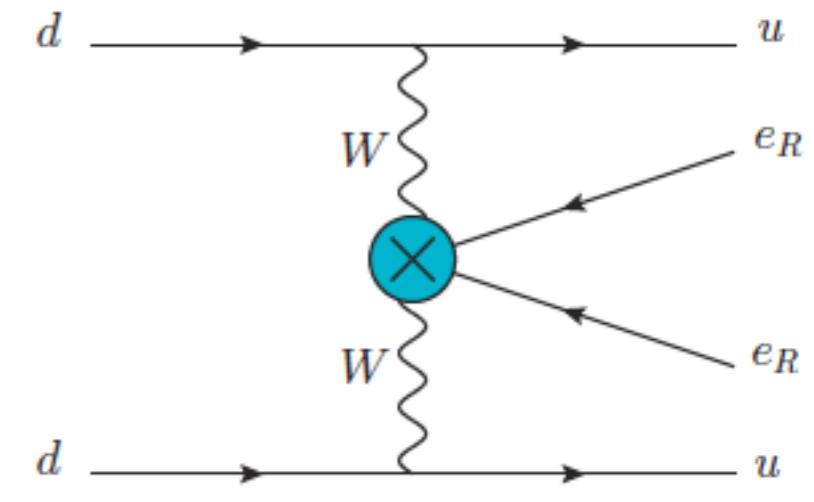
Dimension-9 L violating O

$$\frac{C_{ab}^{(9)}}{\tilde{\Lambda}} \bar{\ell}^c R_a \ell_{R_b} W_\mu^+ W^{+\mu}$$

M.Gustafsson,J.M.No,M.A.Rivera,
PRD90, 013012 (2014)

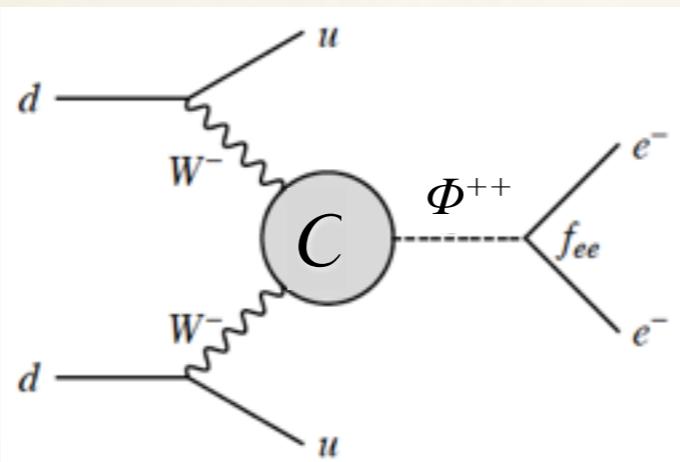
$$\mathcal{L}_{0\nu\beta\beta} = \frac{G_F^2}{2m_p} \epsilon_3 J^\mu J_\mu \bar{e}(1 - \gamma_5)e^c$$

$$J^\mu = \bar{u}\gamma^\mu(1 - \gamma_5)d \quad \epsilon_3 = -2m_p \mathcal{A}_{0\nu\beta\beta}^{\text{SD}}$$



Dimension-7 O

S.F.King,A.Merle,L.Panizzi,
JHEP1411, 124 (2014)



$$\mathcal{L}_{0\nu\beta\beta}^{\text{eff}} = \frac{C f_{ee}}{4M_\phi^2 \Lambda^3} J_{L\mu} J_L^\mu \bar{e}(1 - \gamma_5)e^c$$

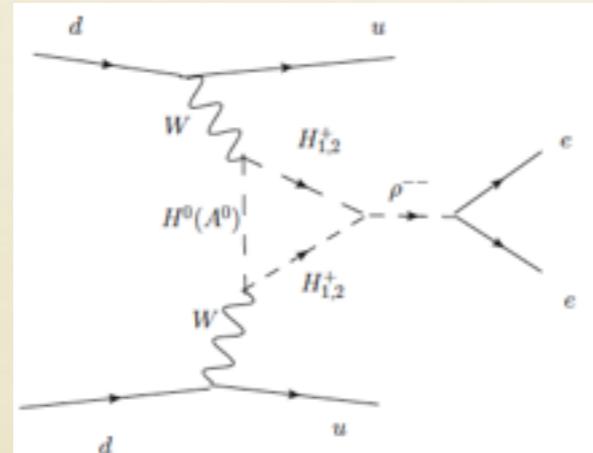


$$\frac{C f_{ee}}{M_\phi^2 \Lambda^3} < 4.0 \times 10^{-3} \text{TeV}^{-5}$$

Cocktail model

Gustafsson,No,Rivera,
PRD90, 013012 (2014)

CQG,D.Huang,L.H.Tsai,
PRD90, 113005 (2014)



	$> T_{\text{exp}}(10^{25} \text{yr})$	$ C_{ee} _{\text{max}}$
GERDA-1(⁷⁶ Ge) [22]	2.1	0.0015
KamLAND-Zen(¹³⁶ Xe) [23]	1.9	0.0011
NEMO-3(¹⁵⁰ Nd) [24]	0.0018	0.0060
CUORICINO(¹³⁰ Te) [25]	0.3	0.0016
NEMO-3(⁸² Se) [26, 27]	0.036	0.0059
NEMO-3(¹⁰⁰ Mo) [27]	0.11	0.0021

$$\mathcal{A}_{0\nu\beta\beta}^{\text{loop}} = \frac{\Delta m_+^2 s_{2\theta^+}}{8\pi^2 m_\rho^2} C_{ee} \{ [\Delta m_+^2 s_{2\theta^+} - \xi v (c_{\theta^+}^2 m_{H_2^+}^2 + s_{\theta^+}^2 m_{H_1^+}^2)] [F_{H_1^+, H_2^+, H_0} - F_{H_1^+, H_2^+, A_0}] - \xi v [m_{H_0}^2 F_{H_1^+, H_2^+, H_0} - m_{A_0}^2 F_{H_1^+, H_2^+, A_0}] \}$$

• Other physics:

Multi Charged Scalars

a. Lepton flavor physics:

1. Muonium anti-muonium conversion $\mu^+ e^- - \mu^- e^+$ $H_{M\bar{M}} = \frac{Y_{ee} Y_{\mu\mu}}{2 M_{--}^2} \bar{\mu} \gamma^\mu e_R \bar{\mu} \gamma_\mu e_R + h.c.,$
2. Effective $e^+ e^- \rightarrow l^+ l^-$, $l = e, \mu, \tau$, contact interactions $\frac{Y_{ee}^2}{M_{--}^2} \bar{e}_R \gamma^\mu e_R \bar{e}_R \gamma_\mu e_R$
3. Rare $\mu \rightarrow 3e$ decays and its τ counterparts
4. Radiative flavor violating charged leptonic decays $Br(\mu \rightarrow e\gamma) = \frac{\alpha}{3\pi G_F^2} \sum_{l=e,\mu,\tau} \left(\frac{Y_{l\mu} Y_{le}}{M_{--}^2} \right)^2$

b. Doubly charged scalars at the LHC:

1 Production of the doubly charged Higgs

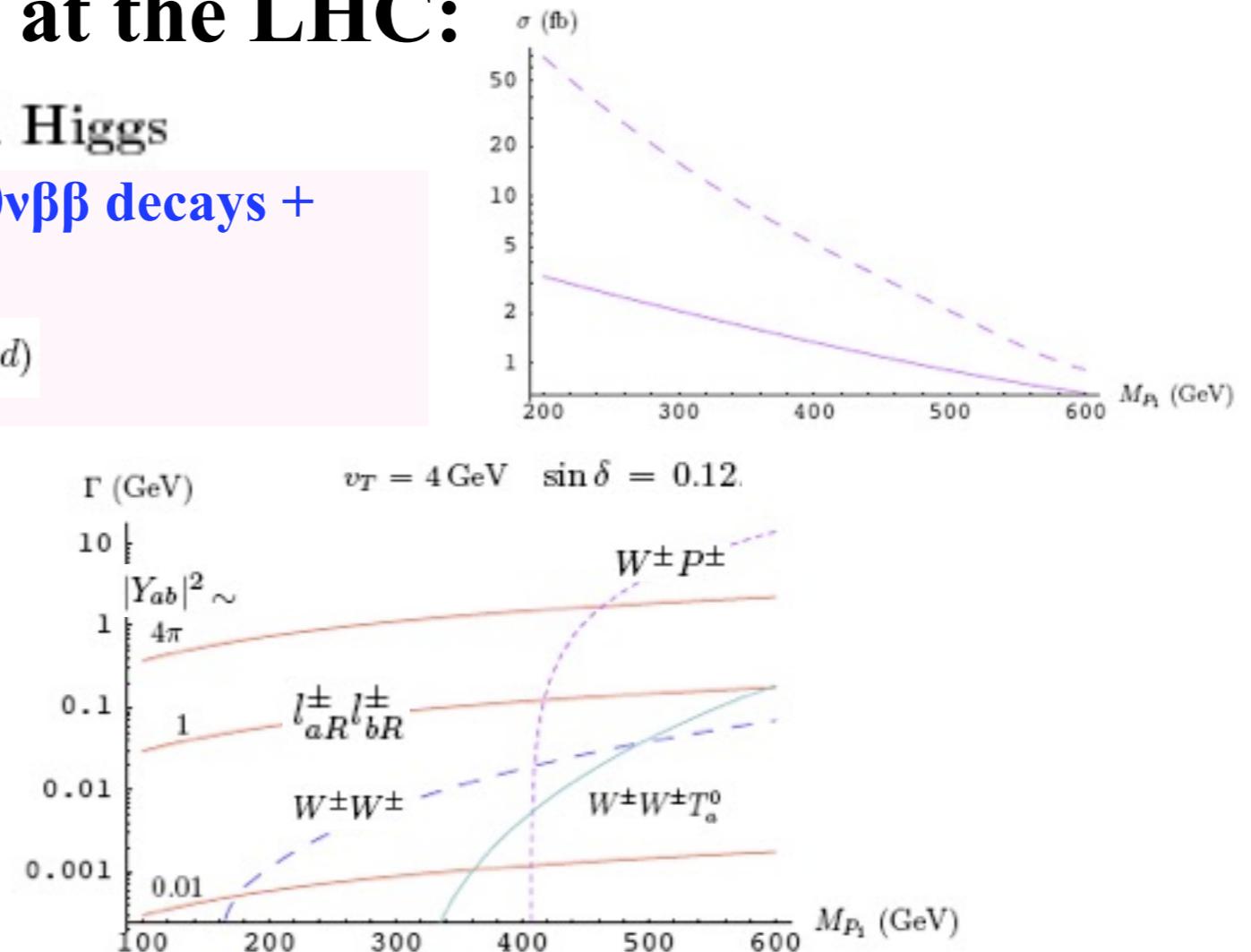
The WW fusion processes similar to $0\nu\beta\beta$ decays + the Drell-Yan annihilation processes:

$$q\bar{q} \rightarrow \gamma^*, Z^* \rightarrow P_1^{++} P_1^{--} \quad (q = u, d)$$

2 The decay of $P_1^{\pm\pm}$

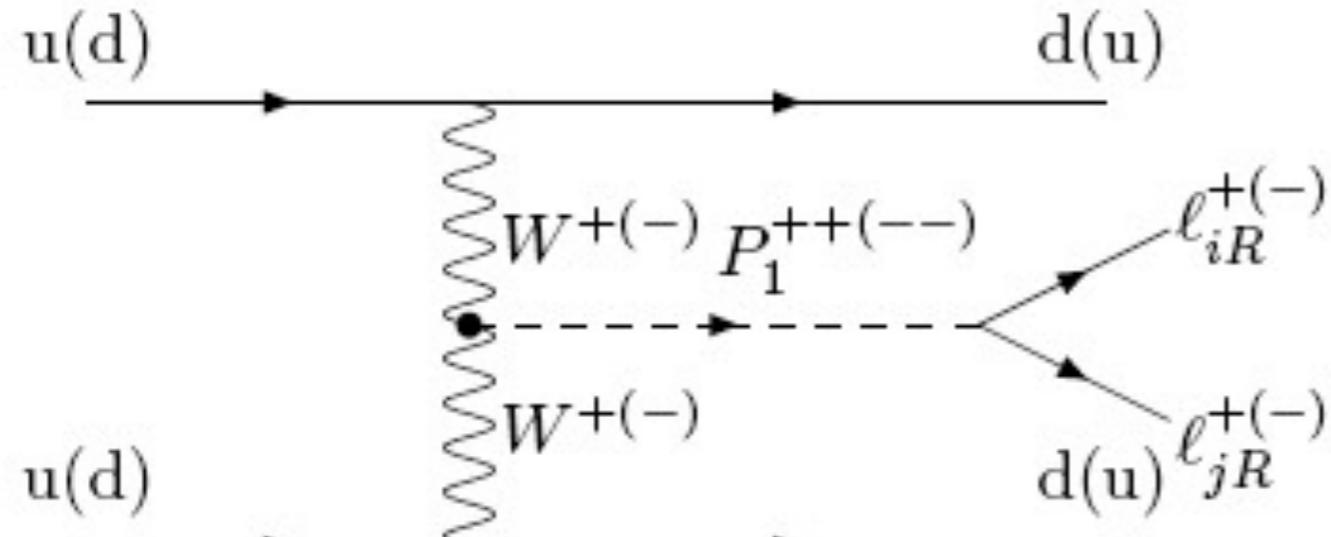
- (1) $P_1^{\pm\pm} \rightarrow l_{aR}^\pm l_{bR}^\pm$ ($a, b = e, \mu, \tau$),
- (2) $P_1^{\pm\pm} \rightarrow W^\pm W^\pm$,
- (3) $P_1^{\pm\pm} \rightarrow P^\pm W^\pm$,
- (4) $P_1^{\pm\pm} \rightarrow P^\pm P^\pm$,
- (5) $P_1^{\pm\pm} \rightarrow W^\pm W^\pm X^0$, $X^0 = T_a^0, h^0, P^0$
- (6) $P_1^{\pm\pm} \rightarrow P^\pm P^\pm X^0$.

(4) and (6) are not allowed in our model



c. Same-sign single dilepton signatures:

$$pp \rightarrow \ell_i^\pm \ell_j^\pm X \xrightarrow{JJ}$$



*Chen, CQG, Zhuridov,
Eur.Phys.J.C60,119(2009)*

$$\frac{d\sigma_\pm^{pp}}{d \cos \theta} = A (\lambda_1^{ij})^2 H_\pm^{pp}$$

$$A = \frac{G_F^4 M_W^6}{2^7 \pi^5} = 50 \text{ ab}, \quad \lambda_1^{ij} = \sqrt{2 - \delta_{ij}} |Y_{ij}| c_\delta s_\delta,$$

$$H_\pm^{pp} = \left(\frac{v_T}{M_W} \right)^2 \int_{z_0}^1 \frac{dz}{z} \int_z^1 \frac{dy}{y} \int_y^1 \frac{dx}{x} p_\pm(x, xs) p_\pm \left(\frac{y}{x}, \frac{y}{x}s \right) l \left(\frac{z}{y} \right) h \left(\frac{s}{M_{P_1}^2} z \right)$$

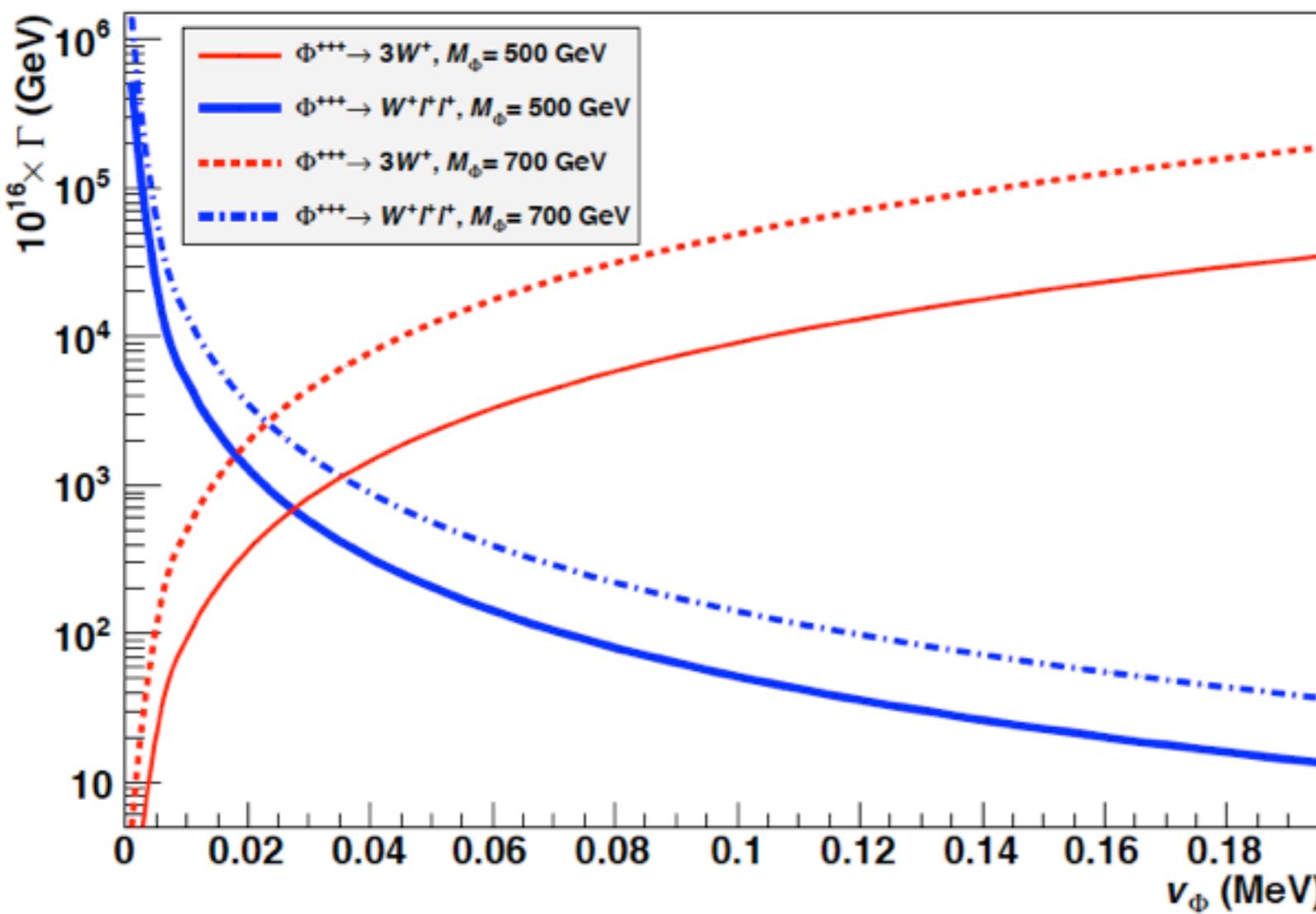
Remarks:

- (i) In our model, the final state charged leptons are right-handed. Hence, in principle, helicity measurements can be used to distinguish between our model and those whose doubly charged Higgs coupling only to left-handed leptons (LLT).
- (ii) $P_1^{\pm\pm}$ will directly produce spectacular lepton # violating signals from like-sign dileptons such as $e\mu$, $e\tau$ and $\mu\tau$.

d. Triply charged scalar decays:

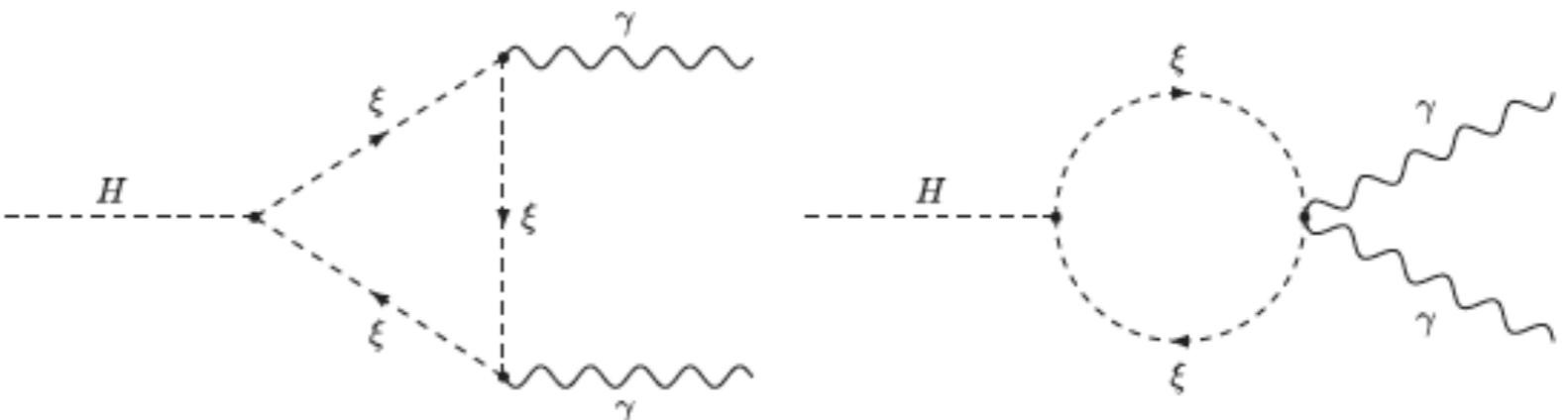
$$\Gamma(\Phi^{+++} \rightarrow 3W) = \frac{3g^6}{2048\pi^3} \frac{v_\Phi^2 M_\Phi^5}{m_W^6}$$

$$\Gamma(\Phi^{+++} \rightarrow W^+ \ell^+ \ell^+) = \frac{g^2}{6144\pi^3} \frac{M_\Phi \sum_i m_i^2}{v_\Phi^2}$$



e. Multi charged scalar contributions to $H \rightarrow \gamma\gamma$ and $H \rightarrow Z\gamma$:

$$\Gamma(H \rightarrow \gamma\gamma) = \frac{G_F \alpha^2 m_H^3}{128\sqrt{2}\pi^3} \left| \sum_f N_f^c Q_f^2 A_{\frac{1}{2}}(\tau_f) + A_1(\tau_W) + \sum_{I_3} (I_3 + 1)^2 \frac{v}{2} \frac{\mu_s}{m_s^2} A_0(\tau_s) \right|^2,$$



$\xi=(1, N, 2)$ with $N=3, 5, \dots$

e.g. $\xi = (\xi^{+++}, \xi^{++}, \xi^+, \xi^0, \xi^-)^T$ for $N=5$

$I_3=(-N+3)/2$ to $(N+1)/2$

Chen, CQG, Huang, Tsai,
PRD87, 077702 (2013)

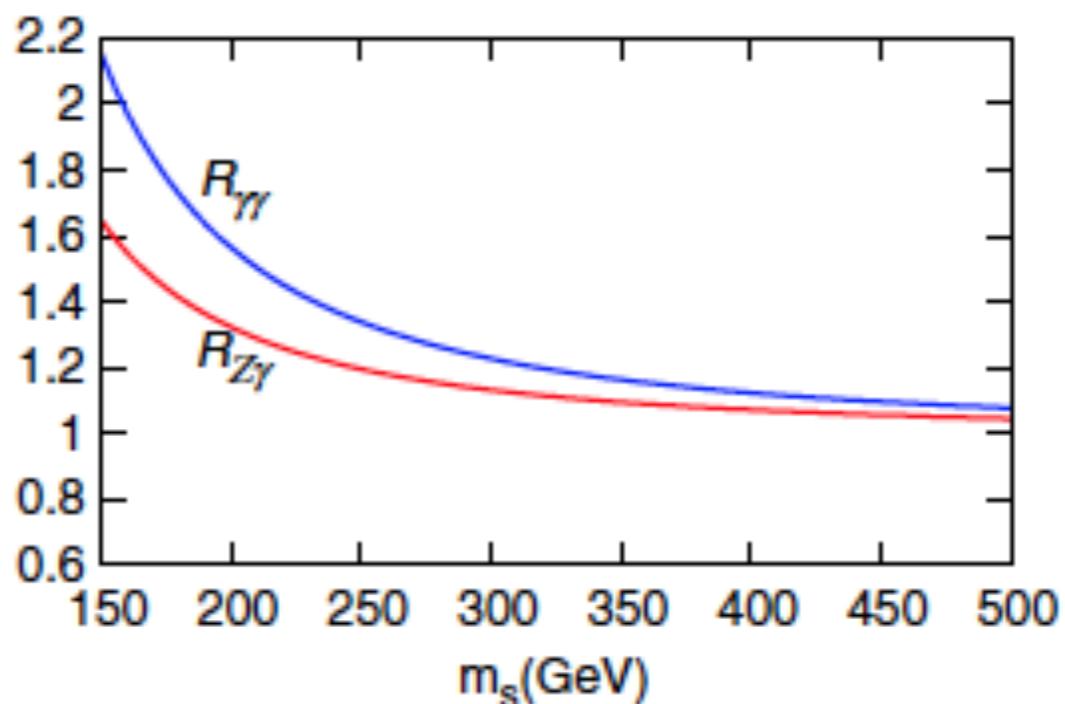


FIG. 4 (color online). $R_{\gamma\gamma} \equiv \Gamma(H \rightarrow \gamma\gamma)/\Gamma(H \rightarrow \gamma\gamma)_{\text{SM}}$ and $R_{Z\gamma} \equiv \Gamma(H \rightarrow Z\gamma)/\Gamma(H \rightarrow Z\gamma)_{\text{SM}}$ as functions of the degenerate mass factor m_s of the multicharged scalar states with $n = 5$ and the universal trilinear coupling to Higgs, $\mu_s = -100$ GeV.

😊 Open questions in neutrino physics:

- 1. What are the masses of the neutrino mass eigenstates (ν_i)?**
- 2. Are the neutrino mass eigenstates Dirac or Majorana particles?**
- 3. If $0\nu\beta\beta$ is observed, is ν a Majorana particle?**
- 4. Can we understand the mixing angles in the neutrino sector? Is there a symmetry behind them?**
- 5. Is there CP violation in the neutrino sector?**
- 6. Others: sterile neutrino, dark radiation?**

New field: Astro-Neutrino Physics

A real window for new physics

2015 Nobel Prize for
neutrino oscillation
Icecube see high energy
cosmic neutrinos

Daya Bay; Reno see θ_{13}

K2K confirms
atmospheric
oscillations

KamLAND confirms
solar oscillations

Nobel Prize for neutrino
astroparticle physics!

SNO shows solar
oscillation to active flavor

Super K confirms solar
deficit and "images" sun

Super K sees evidence of atmos-
pheric neutrino oscillations

Nobel Prize for $\bar{\nu}$ discovery!

LSND sees possible indication
of oscillation signal

Nobel prize for discovery
of distinct flavors!

Kamioka II and IMB see
supernova neutrinos

Kamioka II and IMB see
atmospheric neutrino anomaly

SAGE and Gallex see the solar deficit
LEP shows 3 active flavors

Kamioka II confirms solar deficit

2 distinct flavors identified
Davis discovers
the solar deficit

Reines & Cowan
discover
(anti)neutrinos

Pauli Predicts the Neutrino
Fermi's theory of weak interactions

1930

1955

1980

2005

2015

The Growing Excitement of Neutrino Physics



謝謝！