

## ISSIAtomicData/phase2\_20161006/02\_test

The file you want:

```
test_intensities_fe_13.fits.h5
```

This file contains 1000 ‘fake’ intensities with known density and path lengths. It was saved with

```
nrl_save_hdf, logn=logn, logt_max=logt_max, emissivity=emissivity, wavelength=wavelength, $
intensities=intensities, intensities_error=intensities_error, $
n_intensities=n_intensities, $
time_stamp=time_stamp, logn_obs=logn_obs, ds_obs=ds_obs, $
file=intensity_file
```

The routines:

- TEST\_INTENSITIES\_FE\_13: This routine generates fake intensities using an assumed electron density and path length. Some nuances
  - We select random densities uniformly on the range 8.5 to 9.5.
  - We estimate the path length from a theoretical approximation derived from a steady, uniform heating model (Martens et al. 2000, equation 24).

$$ds = \frac{2.56 \times 10^8}{P_0} \quad (1)$$

where  $P_0 = 2k_b n_e T_e$  is in dyne  $\text{cm}^{-2}$  and  $ds$  is in cm. The intensity is

$$I_\lambda = \epsilon_\lambda(n_e, T_e) n_e^2 ds \quad (2)$$

where we use the default CHIANTI emissivity to compute the intensity.

- Statistical uncertainties are added assuming the EIS pre-flight effective areas, a 60 s exposure time, and the 2 arcsec slit.
- FIT\_TEST\_INTENSITIES\_FE\_13: This routine finds the best-fit density and path length for a given set of fake intensities. For example,

```
model log_n = 9.76 +- 0.013 [9.76]
model log_ds = 7.95 +- 0.027 [7.95]
chi2 = 2.8
```

Line	Imodel	Iobs	SigmaI	dI/Sigma	dI/I
196.525	463.45	462.16	7.41	0.17	0.3
200.021	500.55	508.83	10.08	0.82	1.6
201.121	555.42	550.57	12.81	0.38	0.9
202.044	796.81	787.23	18.37	0.52	1.2
203.165	216.98	230.17	11.83	1.12	5.7
203.826	2511.51	2504.69	44.59	0.15	0.3
209.916	137.61	149.46	18.76	0.63	7.9

Looping over all of the realizations of CHIANTI yields a distribution that looks like this.

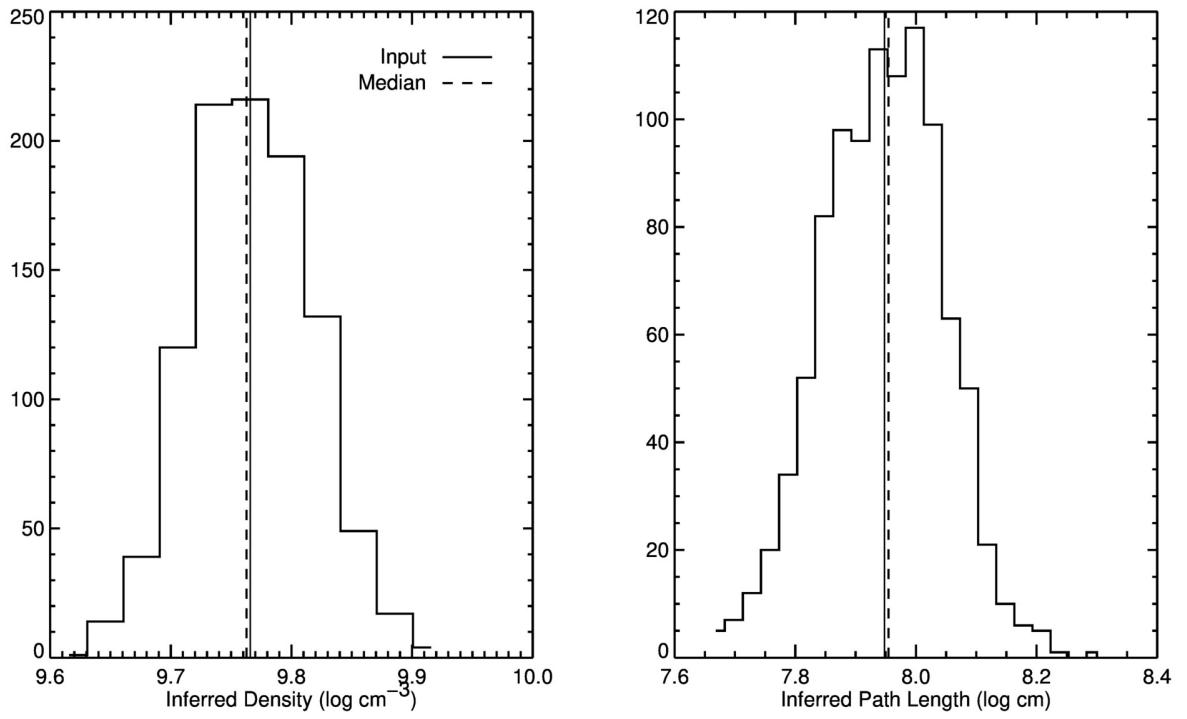


Figure 1: The distribution of inferred density and path lengths for a single set of input intensities. Note that normally distributed errors have been added to the intensities.