ISSIAtomicData/phase2_20161006/04_observed

The file that you want:

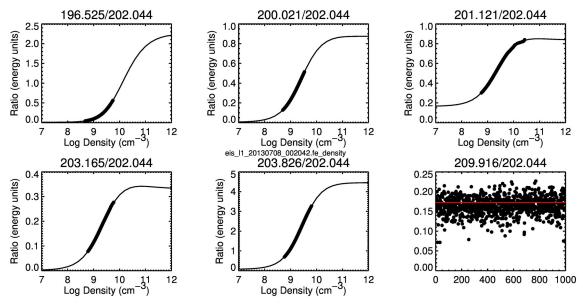
```
eis_11_20130708_002042.fe_density.h5
```

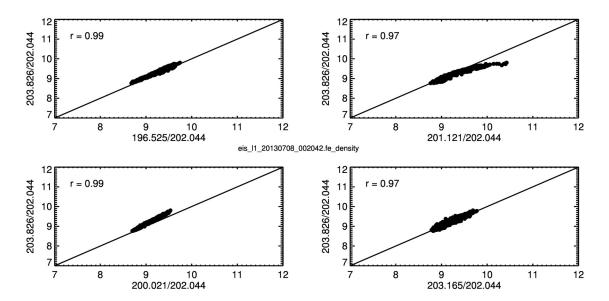
This file contains sets of observed intensities from an EIS observation of an active region (eis_l1_20130708_002042). Intensities for the following Fe XIII lines are saved: 196.525, 200.021, 201.121, 202.044, 203.165, 203.826, 209.916.

The data are saved with the following call. Note that the emssivities are also saved in this file for completeness.

```
nrl_save_hdf, intensities=intensities_out, intensities_error=intensities_error_out, $
index=index_out, eis_files=files, eis_nx=fit.nx, eis_ny=fit.ny, $
n_intensities = n_intensities, n_lines=nfiles, ints_204_min=ints_204_min, $
logn=logn, emissivity=emissivity, wavelength=wavelength, $
time_stamp=time_stamp, $
file=opf+'.h5'
```

- fe_13_density.pro: This routine reads and organizes the intensities. Some nuances:
 - 1000 sets of intensities and the corresponding statistical uncertainties are saved.
 - The intensities are stored in ascending order of wavelength.
 - The index variable saves the position of the intensities in the original raster. Additional information is contained in the file fe_13_density.txt
 - The default atomic data is used to compute densities from each of the density sensitive line ratios. The 209.916/202.044 is not sensitive to density so the observed and theoretical ratio for each observation is plotted.





• fe_13_fit_intensities.pro: This routine takes a single set of intensities and find the best-fit density and path length. For example,

Here is the result of using each of the 1000 realizations of CHIANTI to fit the single set of observations.

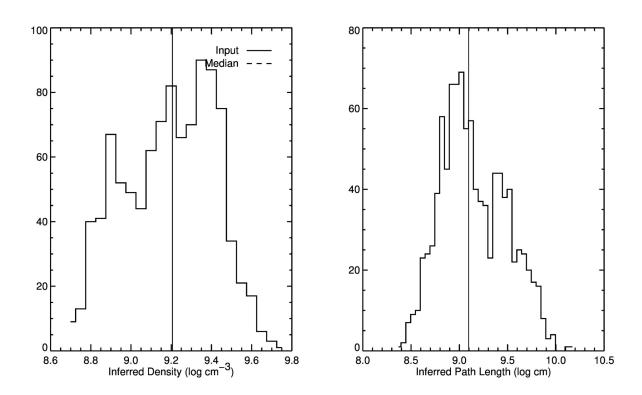


Figure 1: Distributions of density and path length.