

# SPECIAL THEORY OF RELATIVITY

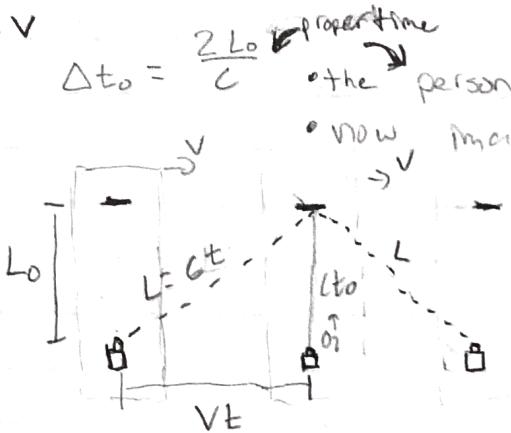
- physics are the same in all inertial RF (reference frame)
- nothing can go faster than light,  $c$  is the same in all RFs

## Time Dilation

$$ct = L$$

$$\Delta t = \frac{L}{c}$$

(the long distance)



imagine this light clock is on a rocket

$\Delta t_0 = \frac{2L_0}{c}$  proper time

the person on the rocket measures this time

now imagine you are watching this rocket move

Lets take a look at that time

$ct$  = distance light travels in t sec

$vt$  = distance rocket travels in t sec

$c\Delta t_0$  = distance light travels

relative to the person on the rocket

\*the reason you can say  $L_0 = c\Delta t_0$  is because  $c = \frac{L_0}{\Delta t_0}$  no matter what with  $\Delta t_0$  being relative to rocket now

- so if  $c$  can't change and  $L_0$  can't change  $\Delta t_0$  has to change

- but both observers can agree on  $L_0$

$$(ct)^2 = (vt)^2 + (c\Delta t_0)^2 \rightarrow \Delta t = \sqrt{\frac{\Delta t_0}{1 - \frac{v^2}{c^2}}}$$

me watching I can get these two

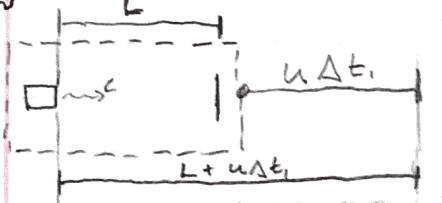
you get same  $L_0$  can't change and  $L_0 = c\Delta t_0$  recent change also

$\Delta t_0$  time person on rocket measures

$c$  moves less distance in inertial RF compared to immobile so only thing that can change is time

## Length Contraction

\* $\Delta t_1$  is time to hit mirror  $\Delta t_2$  is time to reflect & return

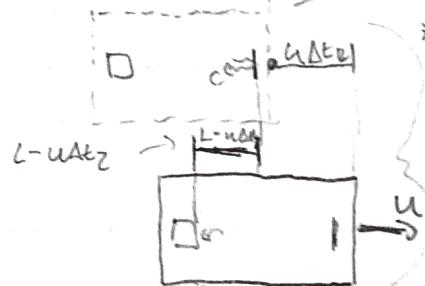


light had to travel  $L + u\Delta t_1$  (where this is how far the rocket traveled in  $\Delta t_1$ )

& the ship is running away from it

total distance light traveled =  $c\Delta t_1 = L + u\Delta t_1$

$$\Delta t_1 = \frac{L}{c-u}$$



the ship is running into the light, so in the time  $\Delta t_2$  for it to return to the emitter, the ship travels

$u\Delta t_2$ . So the total distance light traveled is  $L - u\Delta t_2$

$$c\Delta t_2 = L - u\Delta t_2 \rightarrow \Delta t_2 = \frac{L}{c+u}$$

$$\text{the total return time} = \Delta t = \Delta t_1 + \Delta t_2 = \frac{L}{c-u} + \frac{L}{c+u}$$

Now we have  $\Delta t$  & because

and  $\Delta t = \frac{L}{c}$  set them equal, solve for  $L$ ,

$$L = L_0 \sqrt{1 - \frac{v^2}{c^2}}$$

length contraction when moving

Time dilation and length contraction.  
length contraction in one RF is a time dilation in another

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \frac{v^2}{c^2}}} \quad L = L_0 \sqrt{1 - \frac{v^2}{c^2}}$$

moon hitting the earth

REMEMBER  $c$  has to be one

in all RF  $c = 30^9$

$$C = \frac{L_0}{\Delta t_0}$$

$$C = \frac{L}{\Delta t}$$

$$C = \frac{L_0}{\Delta t_0} \text{ and } C = \frac{L}{\Delta t}$$

don't make sense

**Time dilation:** person is at rest watching the moon

- the moon measures its lifetime (proper life) to be  $(2.2 \cdot 10^{-6} s)$
- the person measures the height of earth (proper length) to be  $(100 \text{ km})$



Find the minimum speed  $v$  has to travel to reach Earth.

person is stationary ref frame

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \frac{v^2}{c^2}}} \quad (2.2 \cdot 10^{-6} s)$$

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \frac{v^2}{c^2}}} \text{ solving for } v \rightarrow 333 \cdot 10^{-6} s = \frac{2.2 \cdot 10^{-6}}{\sqrt{1 - \frac{v^2}{c^2}}} \Rightarrow v = .999978 c$$

Recall:  $\Delta t = \frac{L}{c}$  long length or  
length person sees  $v$  go, in this case  $L = L_0$

$$\Delta t = \frac{L_0}{c} = \frac{100 \text{ km}}{3 \cdot 10^8} = 333 \cdot 10^{-6} s \text{ (use } c \text{ cause it must live for at least this amount of time)}$$

**Length contraction:**

→ person moving with moon (or Earth flying toward us)

→ moon stationary ref frame

• Earth speed is  $v = .999978 c$

• for observer on Earth we travel  $100 \text{ km}$  ( $L_0$ )

height of from  
atmosphere out  
of pos

$$L = L_0 \sqrt{1 - \frac{v^2}{c^2}} = 66 \text{ km or } 660 \text{ m}$$

\* moving means moving relative to stationary observer

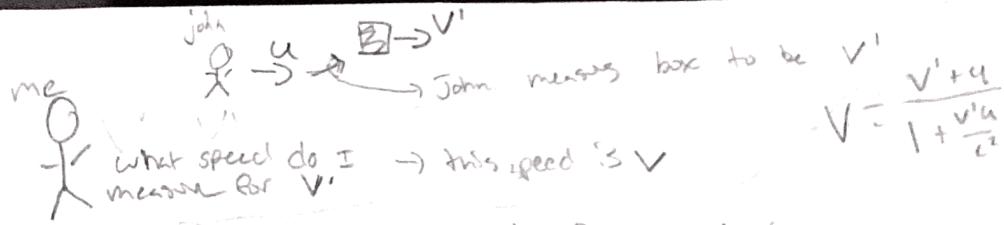
**time dilation:** figure out what the moving thing's  $\Delta t_0$  is (things time measure when moving)

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - \frac{v^2}{c^2}}} \quad \text{Velocity of moving thing}$$

**length contraction:**  $L_0$  length a person watching at rest measures.

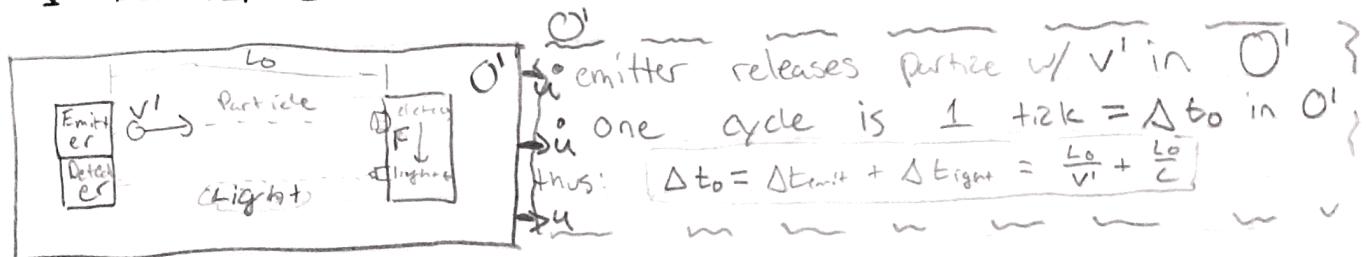
length things  
moving less

$$L = L_0 \sqrt{1 - \frac{v^2}{c^2}} \quad \text{Velocity of moving thing}$$



## Consequences of Special Relativity

### VELOCITY ADDITION



for  $O$  (and  $O'$  is moving at  $u$ ) it sees at speed  $v$  (not  $v'$ )

now just for particle to reach  $F$  in  $\Delta t_1$ ,

$$\rightarrow ① v \Delta t_1 = L + u \Delta t_1 \rightarrow D_{tot} = D_{contracted} + Du$$

[Contracted length back]  
[we see it moving]

$$② \text{ now for light side get time in } \Delta t_2 \text{ seconds}$$

$$c \Delta t_2 = L - u \Delta t_2 \rightarrow D_{tot} = D_{light} - Du$$

{doesn't have to go the whole shuttle path since emission}

$$\Delta t_1 = \frac{L}{v-u} \rightarrow \Delta t_2 = \frac{L}{c-u} \rightarrow \Delta t = \frac{L}{v-u} + \frac{L}{c-u} \text{ (for } O\text{)}$$

$$1. \text{ Use time dilation formula to plug } \Delta t \text{ and use } \Delta t_0 = \frac{L_0}{v'} + \frac{L_0}{c}$$

$$2. \text{ then use Length contraction formula to relate } L \text{ & } L_0$$

↓ you'll get  $\frac{v'u}{c^2}$

$$V = \frac{v' + u}{1 + \frac{v'u}{c^2}}$$

$V$  = speed  $O$  sees of particle

$v'$  = speed  $O'$  sees of particle

$u$  = speed of  $O'$  in  $O$  RF

notice if  
 $\frac{v'u}{c^2}$  is small  
you get normal  
velocity addition

(only for velocity components in direction of  $u$ )

\* If  $O'$  measures the particle  $v' = c$  then both observe  $c$ , required by Einstein

(Example)

- $O$   $v'$   $\rightarrow O$  measures rocket moving away at  $.8c$   
 $\rightarrow u$  fires missile measures its speed  $.6c$  relative to it  
 $\rightarrow$  what speed does  $O$  measure  $v'$ ?

$$V = \frac{v' + u}{1 + \frac{v'u}{c^2}} = .95c, O \text{ only measures } .95c \text{ and not } 1.4c!! \text{ yay!!}$$

# The Relativistic Doppler Effect

Classical Doppler Effect:

$$f' = f \cdot \frac{(v + v_o)}{v + v_s}$$

(emitted by source)      (traveling toward observer)  
 (speed of waves in medium)      (speed of observer)  
 \* relative to medium  
 (velocity of source)

Since velocities don't add like this and light doesn't need a medium

→ consider a source of waves at rest in O. (same velocity)

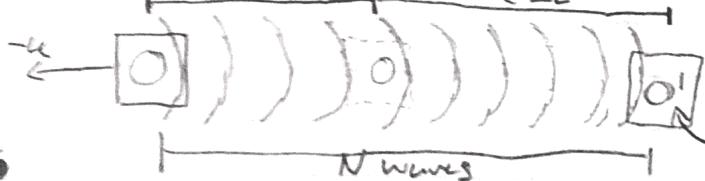
→ O' moves at u relative to source

→ In this situation were in RF of O'

→ Observes the source to emit N waves at freq f. According to O it takes time interval  $\Delta t_0 = \frac{N}{f}$ , this is proper time in RF O

→ In same time interval for O' O moves

→ the total wavelength according to distance O moves  $u\Delta t'$   $\Delta t'$



$O'$  is  $\frac{\text{total length interval}}{\text{number waves}}$

Since  $u = \text{speed of } O \text{ then}$

$O'$  sees waves go away at speed c

assumes over distribution

were in this RF

(FOR O')!

$$[1] \lambda' = \frac{\Delta t' + u\Delta t'}{N}$$

since O sees  $f = \frac{N}{\Delta t_0} = \frac{\text{total waves then}}{\text{total length of wave}}$

$$\lambda' = \frac{\Delta t' + u\Delta t'}{Nf} \leftarrow [O's \text{ freq}]$$

$$[2] \text{ According to } O \quad f' = \frac{c}{\lambda'} \quad \text{algebra}$$

[best way for me to think about it is]  
 $c = \lambda \cdot f$  cause how often they hit times dist between is speed

$$f' = f \frac{\Delta t_0}{\Delta t'} \left( \frac{1}{1 + \frac{u}{c}} \right)$$

time O measures is proper cause it's the time measured by the moving object cause O' is stationary in this RF

[O' is stationary in this RF so the time it measures the time dilate?]

Using time dilation eq for  $\frac{\Delta t_0}{\Delta t'}$ , and algebra

$$f' = f \sqrt{1 - \frac{u^2}{c^2}}$$

NOTE  
!!!!!!

assumes source and observer are separated if approaching then  $f' = f \sqrt{1 + \frac{u^2}{c^2}}$

Example

A galaxy is moving away from the Earth such that the blue hydrogen line at 434nm on Earth is recorded at 600nm. What is speed of the galaxy?

Knowing  $f = \frac{c}{\lambda}$  and  $f' = \frac{c}{\lambda'}$  then

$$\frac{c}{\lambda'} = \frac{c}{\lambda} \sqrt{\frac{1 - \frac{u^2}{c^2}}{1 + \frac{u^2}{c^2}}} \Rightarrow \text{plugging in } u = 0.3c$$

length when source is moving	length when source isn't moving
------------------------------	---------------------------------

~ similar to length contraction

## Lorentz transformation derivation

- An event is something that happens at a definite time and place.

→ lets assign event with coordinates  $(x, t)$  (a fire cracker goes off)  
 → you moving to the right at speed  $u$  give it  $(x', t')$   
 (assumed  $x'=0 + x=0$  are at  $t=t'=0$ )

Lets look at Galilean:

Synchronized  $\rightarrow t=t'=0 \rightarrow$  thus  $t=t'$

you say the event is at  $x' = x - ut$

\* because you've been moving to right at  $u$  the event will be less by  $ut$

Since velocity of light is said to be the same for each observer  
 we have to modify.

Thus if I predict you will say event is at  $x' = x - ut$   
 you will say lengths need to be scaled by  $\gamma$

$$x' = \gamma(x - ut) \quad \text{like wise} \quad x = \gamma(x' + ut)$$

this leaves open times being different

Suppose when our origins coincided we sent a light pulse to set off fire cracker. The light took  $t$  seconds to travel  $x$  meters according to me.  
 According to you it took  $t'$ ,  $x'$  to go but we both agree on  $c$ .

$$\text{thus: } x = ct \quad x' = ct'$$

now → multiply (1) and (2) then use these cause we are transforming between two events.

$$c^2 tt' = \gamma^2 (cc'tt' + cctt' - cct't' - ct't') \quad \text{take out & cancel } ct$$

$$\gamma^2 = \frac{1}{1 - \frac{u^2}{c^2}} \rightarrow \gamma = \sqrt{1 - \frac{u^2}{c^2}} \quad \text{now plug in} \rightarrow x' = \frac{x - ut}{\sqrt{1 - \frac{u^2}{c^2}}}$$

if you isolate  $t'$  remember

$$t' = \frac{t - \frac{ux}{c^2}}{\left(1 - \frac{u^2}{c^2}\right)} \sqrt{1 - \frac{u^2}{c^2}}$$

Using these equations to get Length contraction + time dilation

$$x' = \frac{x - ux}{\sqrt{1 - \frac{u^2}{c^2}}} \quad y' = y \quad z' = z, \quad t' = \frac{t - \left(\frac{u}{c^2}\right)x}{\sqrt{1 - \frac{u^2}{c^2}}}$$

### Length contraction

$$O' \xrightarrow{\text{relative to } O} u \quad O' \text{ measures } L_0 = x'_2 - x'_1$$

$$O \text{ measures } L = x_2 - x_1$$

- definition  
of  
 $L_0$

For  $O$  to make the measurement  $O'$  must measure at same time.

lets say  $O$  measures a flash by  $O'$  at one end then the other  
first flash:  $O'$  measures  $x'_2, t'_2$  and  $O$  measures  $x_1, t_1$   
second flash:  $O'$  measures  $x'_2, t'_2$  and  $O$  measures  $x_2, t_2$

$$x'_1 = \frac{x_1 - ut_1}{\sqrt{1 - \frac{u^2}{c^2}}}$$

$$x'_2 = \frac{x_2 - ut_2}{\sqrt{1 - \frac{u^2}{c^2}}}$$

time you measure it  
now  $O'$  must arrange  
flashes so they appear  
simultaneous

$$\text{Subtracting... } x'_2 - x'_1 = \frac{x_2 - x_1}{\sqrt{1 - \frac{u^2}{c^2}}} - u \left( \frac{t_2 - t_1}{\sqrt{1 - \frac{u^2}{c^2}}} \right)$$

$$L = L_0 \sqrt{1 - \frac{u^2}{c^2}}$$

The length that the observer who doesn't need to measure at the same time measures

### Time dilation

$$\beta = \sqrt{1 - \frac{u^2}{c^2}}$$

$$t' = \frac{t - ux}{\beta}$$

$O' \rightarrow$  imagine  $O'$  has a clock ticks  
each tick occurs at same place

$$x'_2 - x'_1 = 0 \in \text{definition of } \Delta t_0!$$

$O'$

$$x = \frac{xt}{\beta} \rightarrow x = x'\beta + ut$$

$$\text{thus } \Delta x = \Delta x' \beta + \Delta tu$$

$$\Delta t_0 = \Delta t' = \frac{\Delta t - u \frac{\Delta x}{c^2}}{\beta}$$

proper

$$\beta \Delta t_0 = \Delta t - \frac{u \Delta x}{c^2} \rightarrow \beta \Delta t_0 = \Delta t \left( 1 - \frac{u^2}{c^2} \right)$$

thus

$$\boxed{\Delta t = \frac{\Delta t_0}{1 - \frac{u^2}{c^2}}}$$

proper time is the time for observer, with speed  $u$ , relative to another observer  $\text{measures}$ , and  $\Delta x$  between ticks is zero

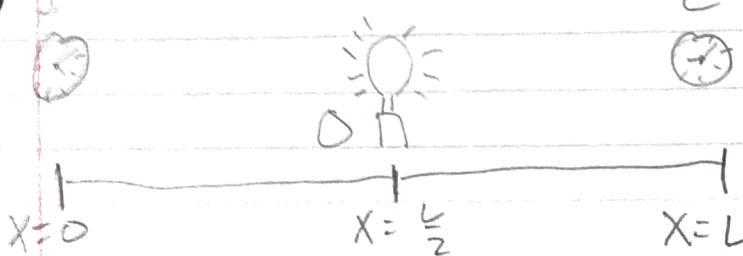
proper length is the length the observer at rest relative to another observer with speed  $u$  measures, and  $\Delta t$  between measurements is zero

# Simultaneity & clock synchronization

important notes!

location of  $O'$  doesn't matter  
but velocity does

clock of velocity is important  
if he moves other way  
the order is reversed



- flash of light starts two clocks simultaneously in  $O'$
- $O'$  sees  $2$  ahead of  $1$   
hence why

light travels  $\frac{L}{2c}$  to reach the clock

→ viewing from moving observer  $O'$

in RF  $O$  it measures:

clock [1] at  $x_1=0$ ,  $t_1 = \frac{L}{2c}$   
clock [2] at  $x_2=L$ ,  $t_2 = \frac{L}{c}$

using Lorentz transformation, we find  $O'$  observes clock 1 to run

$$t'_1 = \frac{t_1 - (\frac{v}{c^2})x_1}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{L/2c}{\sqrt{1 - \frac{v^2}{c^2}}}$$

$$t'_2 = \frac{t_2 - (\frac{v}{c^2})x_2}{\sqrt{1 - \frac{v^2}{c^2}}} = \frac{L/2c - (\frac{v}{c^2})L}{\sqrt{1 - \frac{v^2}{c^2}}}$$

thus  $t'_2$  is smaller than  $t'_1$  so clock 2 appears to run slow earlier

$$\Delta t' = t'_2 - t'_1 = \frac{UL/c^2}{\sqrt{1 - \frac{v^2}{c^2}}}$$

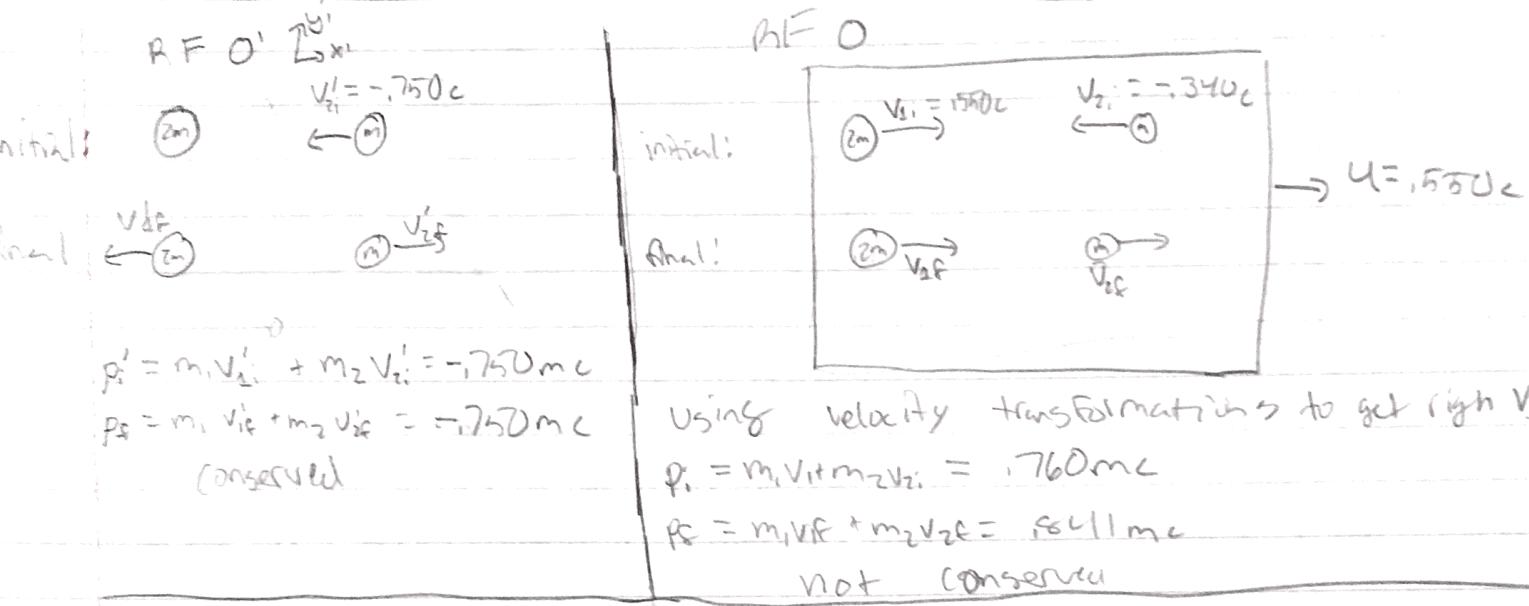
- $O'$  observes both clocks to run slow due to the dilation
- $O'$  also observes clock 2 ahead of clock 1

so 2 events simultaneous in one RF isn't in another

# NO DERIVATION

time

## Relativity with momentum and energy



$$p_i' = m_1 v_{1i}' + m_2 v_{2i}' = -1.500mc$$

$$p_f = m_1 v_{1f} + m_2 v_{2f} = -1.500mc$$

(conserved)

Using velocity transformations to get right  $v_i'$

$$p_i = m_1 v_{1i} + m_2 v_{2i} = 1.760mc$$

$$p_f = m_1 v_{1f} + m_2 v_{2f} = 1.500mc$$

not conserved

Have to find a new formula to satisfy Einstein.

$$\vec{p} = \frac{m\vec{v}}{\sqrt{1-\frac{v^2}{c^2}}} \rightarrow p_x = \frac{mv_x}{\sqrt{1-\frac{v^2}{c^2}}} \quad p_y = \frac{mv_y}{\sqrt{1-\frac{v^2}{c^2}}}$$

$v$  is velocity of the particle as measured in inertial RF,

$$p_i = \frac{m_1 v_{1i}}{\sqrt{1-\frac{v_{1i}^2}{c^2}}} + \frac{m_2 v_{2i}}{\sqrt{1-\frac{v_{2i}^2}{c^2}}} = \frac{m_1(0)}{\sqrt{1-0}} + \frac{m_2(-0.750c)}{\sqrt{1-(0.750)^2}} = -1.134mc$$

conserved

$$p_f = \frac{m_1 v_{1f}}{\sqrt{1-\frac{v_{1f}^2}{c^2}}} + \frac{m_2 v_{2f}}{\sqrt{1-\frac{v_{2f}^2}{c^2}}} = -1.134mc$$

Same for momentum in O

**Energy** → If you blindly apply old formulas to this example, it won't work

$$K = \frac{mc^2}{\sqrt{1-\frac{v^2}{c^2}}} - mc^2 \quad K \rightarrow \infty \quad V \rightarrow c$$

or

$$K = E - E_0 \quad \text{and} \quad E = \frac{mc^2}{\sqrt{1-\frac{v^2}{c^2}}} \quad \text{and} \quad E_0 = mc^2$$

manipulating momentum & energy

$$E = \sqrt{(pc)^2 + (mc^2)^2}$$

Diagram illustrating the relationship between energy and momentum. A right-angled triangle represents the energy-momentum relation. The vertical leg is labeled  $mc^2$ , the horizontal leg is labeled  $pc$ , and the hypotenuse is labeled  $E$ . The angle between the vertical leg and the hypotenuse is labeled  $\theta$ .

total energy measured  
in RF in which  
particle is at rest

$$P = \frac{mv}{\sqrt{1-\frac{v^2}{c^2}}}, E = \frac{mc^2}{\sqrt{1-\frac{v^2}{c^2}}}, K = E - mc^2$$

$$mass = \frac{mc^2}{c^2} \quad E = \text{MeV}$$

Very useful b/c  
velocity  $v$  is constant  
 $c + p + E$  are!

$$E^2 = (pc)^2 + (mc^2)^2, \text{min rel}^2 = \left(\frac{E}{c}\right)^2 - \left(\frac{pc}{c^2}\right)^2$$

$$E_{\text{tot}}^2 = m^2 c^2 (1 + \frac{p^2}{c^2})$$

$$p = \frac{mc}{c}$$

## Relativistic energy and momentum

→ examples, examples,  
makes calc easy

$$\text{Momentum of a proton w/ } v_p = 860c \rightarrow P = \frac{mv_p}{\sqrt{1-\frac{v_p^2}{c^2}}} \rightarrow pc = \frac{mc^2}{\sqrt{1-\frac{v_p^2}{c^2}}}$$

$$E_{\text{opt}} = m_p c^2 = 938 \text{ MeV}$$

$$pc = 1580 \text{ MeV} \rightarrow p = 1580 \frac{\text{MeV}}{c}$$

$$\text{Relativistic Energy} \rightarrow E = \sqrt{(pc)^2 + (mc^2)^2} = \sqrt{(1580)^2 + (938)^2} \quad c \text{ cancel out!}$$

$$K = E - E_{\text{opt}} = 899 \text{ MeV}$$



$$I_{\text{sun}} = 1.4 \cdot 10^3 \frac{\text{W}}{\text{m}^2} \text{ at earth}$$

at what rate the mass of the sun is decreasing.

imagine the sphere with "r" so the total sun output is  $(4\pi r^2)(I_{\text{sun}})$   
 $= 4 \cdot 10^{26} \text{ J/s}$

Sun loses energy by a loss in rest energy,

$$\frac{\Delta E_0}{c^2} = \frac{\Delta m}{c^2} = 4.4 \cdot 10^9 \text{ kg/s} \leftarrow \text{billion kg every second!}$$

## conservation laws in decays and collisions

→ a loss in kinetic energy is accompanied by a gain in rest energy

→ but the total relativistic energy / Kinetic + rest energy is constant.

$$m_K \xrightarrow[\text{decay}]{} K \rightarrow m_\pi 0 \rightarrow m_\pi \quad K_K = 77 \text{ MeV} \quad m_\pi = 139.6 \frac{\text{MeV}}{c^2}$$

$$m_K = 497.7 \frac{\text{MeV}}{c^2} \quad p_\pi = 381.6 \frac{\text{MeV}}{c}$$

a) find momentum and total relativistic energy of unknown particle ( $m_x$ )

$$E_K = K_K + m_K c^2 = 574.7 \text{ MeV} \quad \left. \begin{array}{l} \text{for } \pi \\ \text{meson} \end{array} \right\} E_\pi = \sqrt{(p_\pi)^2 + (m_\pi c^2)^2} = 406.3 \text{ MeV}$$

$$p_K c = \sqrt{E_K^2 - m_K^2 c^2} = 287.4 \frac{\text{MeV}}{c} \quad \left. \begin{array}{l} \text{now use momentum to solve problem} \\ \text{and } p_K \end{array} \right\}$$

(initial) (final)

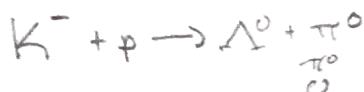
$$p_K = p_\pi + p_x \rightarrow p_x = p_K - p_\pi \rightarrow p_x = -941.2 \frac{\text{MeV}}{c}$$

$$E_{\text{initial}} = E_{\text{final}} \rightarrow E_K = E_\pi + E_x \rightarrow E_x = E_K - E_\pi = 574.7 - 406.3$$

$$E_x = 168.4 \text{ MeV}$$

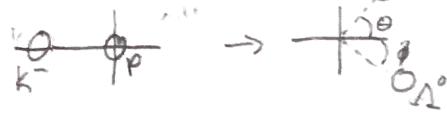
$$\text{you can find mass by using } m_x c^2 = \sqrt{E_x^2 - (cp_x)^2} = 139.6 \frac{\text{MeV}}{c^2}$$

$m_x = p_x \text{ meson}$



$$m_p = 938.3 \frac{MeV}{c^2} \quad m_K = 938.3 \frac{MeV}{c^2}$$

$$m_\Lambda = 1115.7 \frac{MeV}{c^2} \quad m_\pi = 135.0 \frac{MeV}{c^2}$$



$$K_K = 152.4 \text{ MeV} \quad K_\Lambda = ?$$

$$K_\pi = 254.8 \text{ MeV} \quad \theta = ?$$

$$\phi = ?$$

ex 2

$$E_{\text{initial}} = E_K + E_p = (K_K + m_K c^2) + m_p c^2$$

$$E_{\text{final}} = E_\Lambda + E_\pi = (K_\Lambda + m_\Lambda c^2) + (K_\pi + m_\pi c^2)$$

Einitial = Efinal everything known except  $K_\Lambda$

$$K_\Lambda = 78.9 \text{ MeV}$$

to get direction information use conservation of mom

$$p_{\text{initial}} = \text{just } m_K \Rightarrow (p_K = \sqrt{E_K^2 - (m_K c^2)^2}) \text{ and } E_K = K_K c + m_K c^2$$

$$p_{\text{initial}} = p_K = \sqrt{116.8 \frac{MeV}{c}}$$

$$(p_K = \sqrt{E_K^2 - (m_K c^2)^2} = 426.9 \text{ and } (p_\pi = \sqrt{E_\pi^2 - (m_\pi c^2)^2})$$

$$E_K = K_K + m_K c^2$$

$$E_\pi = K_\pi + m_\pi c^2$$

momentum or two particles is...

$$\left[ \begin{array}{l} p_{x\text{f}} = p_K \cos\theta + p_\pi \cos\phi \\ p_{y\text{f}} = p_K \sin\theta - p_\pi \sin\phi \\ p_{x\text{i}} = p_K \end{array} \right] \quad \left[ \begin{array}{l} p_{y\text{f}} = p_K \sin\theta - p_\pi \sin\phi \\ p_{y\text{i}} = 0 \end{array} \right] \quad \begin{array}{l} \text{set } p_{x\text{f}} = p_{x\text{i}} \\ \text{set } p_{y\text{i}} = p_{y\text{f}} \end{array}$$

$$\phi = \cos^{-1} \left( \frac{p_x^2 + p_\pi^2 - p_K^2}{2 p_K p_\pi} \right) = 65.7^\circ \quad \theta = \sin^{-1} \left( \frac{p_\pi \sin\phi}{p_K} \right) = 51.3^\circ$$

$$\textcircled{2} \quad p + p \rightarrow p + p + p + \bar{p}$$

initial threshold kinetic energy for  $\cancel{p} \rightarrow \cancel{p}$

$E_p + M_p c^2$  each particle has same  $p$  +  $E'_p$

more as unit

total initial energy of 2  $p$  is

$$E_p + M_p c^2$$

highest energy but unchanged in momentum

total final energy of 4  $p$ 's

$$E_p + E'_p + E_p + E'_p = 4 E'_p$$

$$\text{so: } E_p + M_p c^2 = 4 E'_p$$

conservation of momentum  $\rightarrow$

$$p_{p+0} = 4 p'_p \rightarrow \sqrt{E_p^2 - (M_p c^2)^2} = 4 \sqrt{E'_p^2 - (M_p c^2)^2}$$

do algebraic

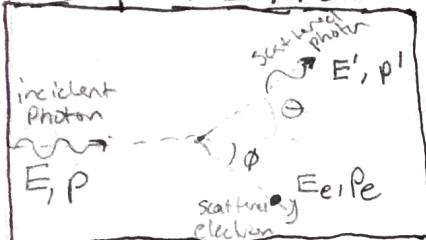
$$E_p = 7 M_p c^2$$

$$k_p = \sqrt{E_p^2 - (M_p c^2)^2} = 6 M_p c^2$$

$$= 5.626 \text{ GeV}$$

# Interaction of Photons with electrons, and atoms, and other combinations!

## Compton Effect



Photon:  $E = hf = \frac{hc}{\lambda}$  and  $p = \frac{E}{c}$  [initial]  
 $E' = hc/\lambda'$   $p' = E'/c$  and moves with angle  $\Theta$  [final]

electron:  
 total Energy =  $E_e$ , and tot momentum  $p_e$  with angle  $\phi$   
 rest energy  $\rightarrow m_e c^2$   $\star 3$  eq + 4 unknowns

$$\text{Initial} = \text{Final} \rightarrow E + m_e c^2 = E' + E_e$$

$$p_{x\text{ initial}} = p_{x\text{ final}} \rightarrow p = p_e \cos \phi + p' \cos \Theta \rightarrow p_e \cos \phi = p - \cos \Theta$$

$$p_{y\text{ initial}} = p_{y\text{ final}} \rightarrow 0 = p_e \sin \phi + p' \sin \Theta \rightarrow p_e \sin \phi = -p' \sin \Theta$$

$$p_e^2 \cos^2 \phi = p^2 + (p')^2 \cos^2 \Theta - 2pp' \cos \Theta \rightarrow \sin^2 \phi + \cos^2 \phi = 1$$

$$\rightarrow \text{square both momentum eq} \rightarrow p_e^2 \sin^2 \phi = p_e^2 \sin^2 \Theta$$

$$\rightarrow \text{Combine: } p_e^2 = p^2 + (p')^2 - 2pp' \cos \Theta$$

Now use Energy  $E + m_e c^2 = E' + E_e \rightarrow E + m_e c^2 - E' = E_e = \sqrt{(p_e c)^2 + (m_e c^2)^2}$

then:  $(E + m_e c^2 - E')^2 = (p_e c)^2 + (m_e c^2)^2$  Plug in and ALOE BRAH!

$$\rightarrow \frac{1}{E'} - \frac{1}{E} = \frac{1}{m_e c^2} (1 - \cos \Theta) \quad \text{or} \quad \lambda' - \lambda = \frac{\hbar}{m_e c} (1 - \cos \Theta)$$

Also

Kinetic energy of electron is what is left over between both photons

$$K_e = E - E'$$

algebra

$$\tan \phi = \frac{\sin \phi}{\cos \phi} = \frac{p_e \sin \phi}{p_e \cos \phi} = \frac{p_e \sin \phi}{p - p' \cos \Theta}$$

recall  $x: p = p_e \cos \phi + p' \cos \Theta$

$$y: p_e \sin \phi = p_e \sin \Theta$$

$$\frac{p_e \sin \phi}{p - p' \cos \Theta} = \frac{p_e \sin \Theta}{E - p_e \cos \Theta} = \frac{E_2 \sin \Theta}{E_1 - E_2 \cos \Theta}$$

Ex. X-ray w/  $\lambda = 2400 \text{ nm}$  and  $\lambda' = \lambda + \frac{\hbar}{m_e c^2} (1 - \cos \Theta)$  d)  $\phi = \tan^{-1} \left( \frac{E' \sin \Theta}{E - E' \cos \Theta} \right)$

Compton scattered at  $\Theta = 60^\circ$  b)  $E' = \frac{h\nu}{\lambda'}$  -

Find: a) scattered wavelength  $\lambda'$

b) energy of scattered X-ray photons

c)  $K_e = E - E'$

d)  $K_e + \text{kinetic energy of electron}$

## Photon + atom

$$E_i = E_f + K + E$$

(initial energy of atom)      (final energy of atom)      (recoil energy of atom)      (Photon energy emitted)  
very small

has to have recoil energy because of conservation of momentum.

atom absorbs photon w/ Energy  $E$ .

$$E_i + E = E_f + K$$

## Bremsstrahlung and X-ray Production

when an electron is accelerated it radiates electromagnetic Energy. Or emits photons

- beam of electrons accelerated by a  $\Delta V$ , thus experience loss in pot. eng.  $-e\Delta V$ , gain  $K = e\Delta V$
- when they hit a target they make many collisions with the atoms
- momentum transferred to atom, electron slow, Photon emitted  
(recoil energy of atom so small can be neglected)

$k_e$  of electron before collision,  $k'_e$  of electron after.

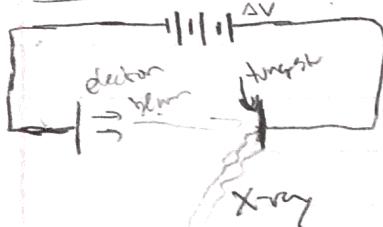
$$\text{photon energy } E = hf = \frac{hc}{\lambda}$$

$$K_e - K'_e = \text{leftover energy} = \frac{hc}{\lambda}$$

electron will make many collisions entree  $\Delta V$  from target  $\Delta$  (small energy) top to  $hf_{max} = K$   
meaning electron loses all its energy  $K$  in a single encounter.

the smallest wavelength

is determined by max Energy loss



"inverse photoelectric effect"

$$K_e - K'_e = \frac{hc}{\lambda} \rightarrow K_e = \frac{hc}{\lambda_{min}}$$

(no energy after it loses all its  $K_e$ )

$$\lambda_{min} = \frac{hc}{K_e}$$

$$K_e = e\Delta V$$

$$\lambda_{min} = \frac{hc}{e\Delta V}$$

Pair Production and Annihilation.  $\text{Photon} \rightarrow \text{electron} + \text{positron}$  or  $\text{electron} + \text{positron} \rightarrow \text{photon}$

$$(m_e c^2 + K_+) + (m_e c^2 + K_-) = E_{ph+} + E_{ph-}$$

$$\text{for positron} + \text{electron at rest} \rightarrow E_{ph+} = E_{ph-} = m_e c^2$$

Why 2 photons? momentum conservation requires at least two photons

Q's

- Why are photons emitted
- Can you create a machine that "cancels out" light waves
- Why don't light all around mess with each other