

Electromagnetic Waves

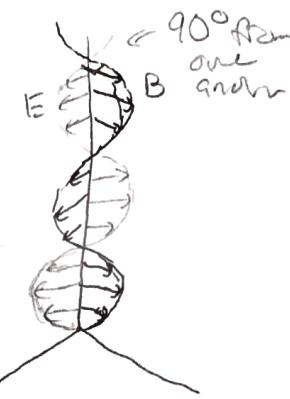
(if charges are accelerated or current carries) over time

$$\vec{E} = E_0 \sin(kz - \omega t)$$

produces EM wave

$$\vec{B} = B_0 \sin(kz - \omega t)$$

$$k = \frac{2\pi}{\lambda}, \omega = 2\pi f, C = \lambda f, \frac{\omega}{k} = C! \quad B_0 = \frac{E_0}{C}$$



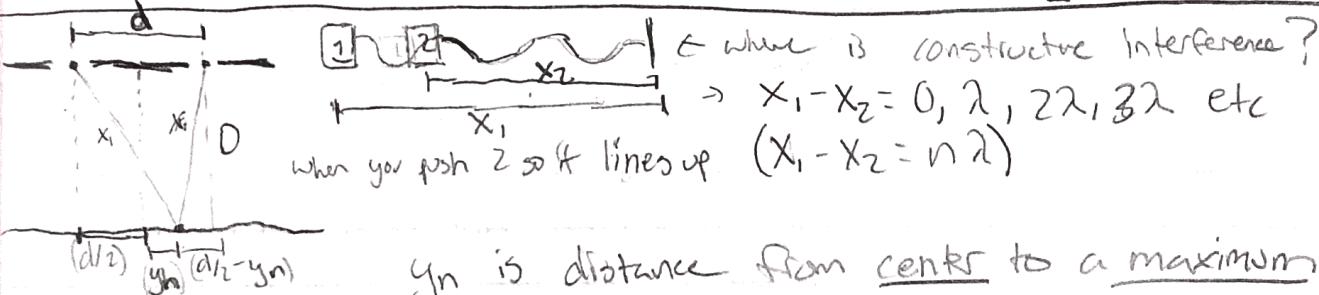
Power Area

flux: $\vec{S} = \frac{1}{\mu_0} \vec{E} \times \vec{B}$, for a plane wave: $\vec{S} = \frac{1}{\mu_0} E_0 B_0 \sin^2(kz - \omega t) \hat{k} \left(\frac{W^2}{m^2} \right)$

power: $P = S \cdot A \Rightarrow P = \frac{1}{\mu_0 C} E_0^2 A \sin^2(kz - \omega t) \sim$ using $B_0 = \frac{E_0}{C} \sim$

the 'intensity' ($S = \frac{P}{A}$) fluctuates very quickly (vis light 10^{15} cycles/s) we observe avg cycles

T is observation time: $P_{av} = \frac{1}{T} \int_0^T P dt \Rightarrow I = \frac{P_{av}}{A} = \frac{1}{2\mu_0 C} \frac{E_0^2}{C} \left[\begin{array}{l} \text{increase in intensity} \\ \text{means increase in} \\ \text{Electric field} \end{array} \right]$

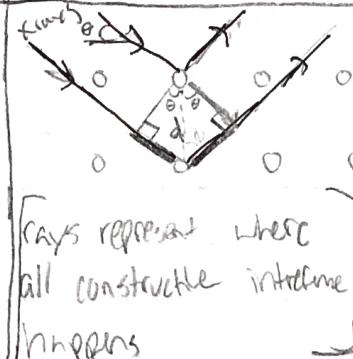
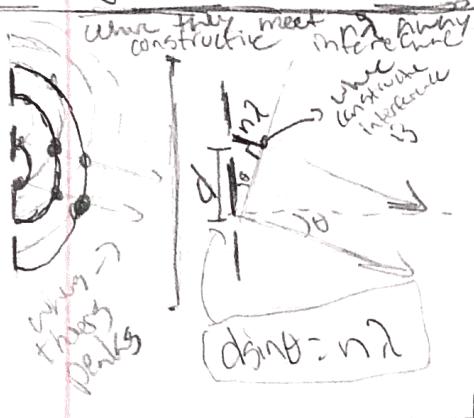


y_n is distance from center to a maximum

$$X_1^2 = D^2 + \left(\frac{d}{2} + y_n\right)^2, \quad X_2^2 = D^2 + \left(\frac{d}{2} - y_n\right)^2 \xrightarrow{\text{approx}} y_n = \frac{X_1^2 - X_2^2}{2d} = \frac{(X_1 + X_2)(X_1 - X_2)}{2d}$$

$X_1 \approx D, X_2 \approx D$ if usually D is 1 m and d, $y_n \approx 1 \text{ mm}$

$$y_n = (X_1 - X_2) \frac{D}{d} \rightarrow \text{constructive interference} \rightarrow y_n = n \frac{\lambda}{2}$$



R_2 travels $2ds \sin \theta$ more

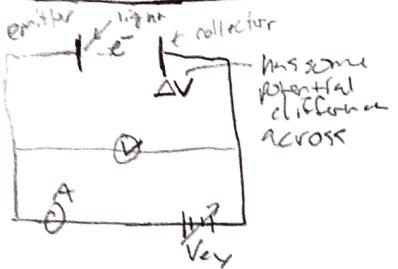
$$R_1 + 2ds \sin \theta = R_2$$

$$2ds \sin \theta = \lambda \rightarrow \text{lets say } R_1 = 0.2, R_2 = 2 \rightarrow 2ds \sin \theta = 2$$

if it was one more d down (another layer of atoms) that length would be $2(2ds \sin \theta)$ but it'd be the second wavelength so it'd cancel & idk why general form is $2ds \sin \theta = n\lambda$ tho?

these lines can be thought of where peaks are, think of those as your X_1 and X_2 in D from above

Photo electric effect



electrons going from emitter to collector

* electrons with $K_e < \Delta U$ get pushed back to emitter

$$\Delta U = -e\Delta V \quad (\text{for growth})$$

Close that amount of K_e

* the voltage point that repels the most energetic electrons is stopping potential

$$K_{max} = eV_s$$

→ an increase in intensity, is an increase in electric field, thus increase in Force
 $\rightarrow F = qE$, increase in force increases K_{max} .

ex: laser w/ $I = 120 \text{ W/m}^2$ $P_{av} = IA$ $A = \pi r^2$
 shot at Na, takes 2.3 eV of energy to release e^- .
 assume e^- confined to a radius of 10nm

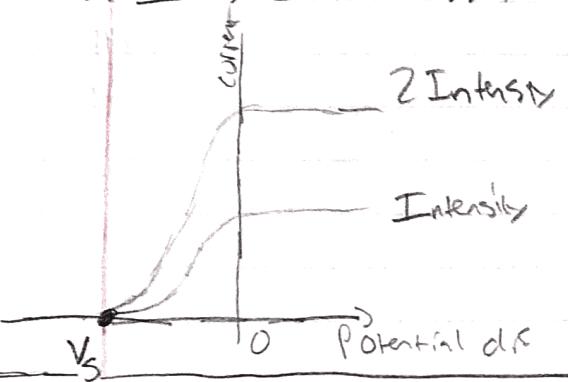
$$P_{av} = \frac{\Delta E_{\text{Energy}}}{\Delta t}$$

$$\Delta t = \frac{\Delta E_{\text{Energy}}}{P_{av}} = \frac{\phi}{IA} = .10 \text{ secs}$$

work function, how much it takes to pull an electron off,

how long will take to release an electron

WHERE THIS CLASSICAL THEORY FAILS:



from $K_{max} = eV_s$, doubling intensity increases K_{max} , but if you increase K_{max} V_s should change. But it doesn't.

classical says a brighter light (more intense) will knock the e^- off harder generating more of a current, meaning you'll need higher V_s . If you think about it as particle, brighter means more photons so only one electron will come off not greater energy.

Quantum Theory of Photoelectric effect.

Light wave not continuously distributed but instead in small little bundles (quanta)

Energy of a photon associated with some λ , or $E = hf$ or $E = \frac{hc}{\lambda}$ $h = 6.67 \cdot 10^{-34} \text{ J}\cdot\text{s}$

if photons energy from relativity, $\beta E = pc$ (rest mass = 0) thus $p = \frac{h}{\lambda}$

After strike, a single photon delivered to electron, if photon energy $> \phi$ then electrons produced

doubling intensity means twice as many photons strike, thus twice as many electrons. But they will all have the same energy.

$$K_{max} = hf - \phi$$

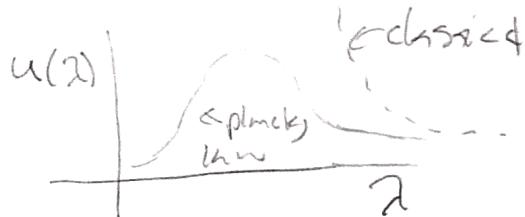
↑
total energy
↓
energy need to release

so left overs of the two is how much energy electron has

if a photon supplies exactly ϕ , the corresponding f_c (cut off f_c), there is no excess K_{max} so

$0 = hf_c - \phi \Rightarrow hf_c = \phi \Rightarrow \lambda_c = \frac{h}{\phi}$

cut off λ_c represents the largest wavelength for which the photoelectric effect can be observed for with a work function ϕ .



Thermal Radiation

• black body absorbs all radiation incident on it, reflects none, $\frac{1}{4}u(\lambda)$

$I(\lambda)$ = intensity as a function of λ

$$I = \int_0^{\infty} I(\lambda) d\lambda \rightarrow \text{graph of } I(\lambda) \text{ vs } \lambda \quad \begin{array}{l} \text{max} \\ \downarrow \\ d\lambda \end{array}$$

stephan boltzmann constant

$\frac{dI(\lambda)}{d\lambda} = 0$ [Stephans Law] $I(T) = \int_0^{\infty} I(\lambda) d\lambda = \sigma T^4 \leftarrow \text{for blackbody}$

Wien's Displacement law: λ_{max} at which $I(\lambda)$ releases λ_{max}

$\lambda_{\text{max}} T = 2,898 \cdot 10^{-3} \text{ m}\cdot\text{K}$

if not a perfect black body it behaves as a proportion of a black body

$$I = \epsilon \sigma T^4, \epsilon = 1 \text{ blackbody}, \epsilon < 1 \text{ non ideal blackbody}$$

blackbody (the hole), walls are in thermal eqq

$$u(\lambda) = \text{energy density} \rightarrow I(\lambda) \propto c u(\lambda)$$

$$I(\lambda) = \text{radiation intensity}$$

* if only $1/4$ EM radiation exits blackbody then $\boxed{I(\lambda) = \frac{1}{4} c u(\lambda)}$

$n = 4$ (goes $1 \rightarrow \infty$) if has 4 peaks over distant L

$$\lambda = \frac{2L}{n}$$

for a small interval of $d\lambda$, n can be treated as cont. func.

$$N(\lambda)d\lambda = 2 \left| \frac{dn}{d\lambda} \right| d\lambda = 4L \frac{d\lambda}{\lambda^2} \left[\text{for standing wave} \right]$$

due to both oscillations of light (polarization)

number of standing waves between $\lambda + d\lambda$

$$3D: n = \frac{4\pi L^3}{3} \left(\frac{1}{\lambda} \right)^3, N(\lambda)d\lambda = 2 \left| \frac{dn}{d\lambda} \right| = \frac{8\pi V}{\lambda^4} d\lambda$$

$$u = \frac{(\text{number of modes})(E_{\text{av}})}{\text{volume}}$$

$$E_{\text{av}} = \frac{1}{N} \int_0^{\infty} E e^{-\frac{E}{kT}} dE = kT$$

small number var

for one degree of freedom, its $\frac{1}{2}kT$
 (energy per mode: kT
 1 degree of freedom per mode)

$$u(\lambda) = \frac{8\pi kT}{\lambda^4}$$

$$\frac{du}{d\lambda}?$$

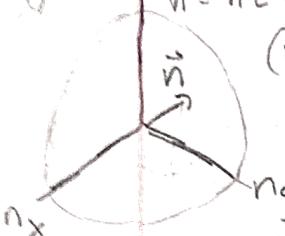
Better derivation of $\alpha(\lambda)$

mode \rightarrow when you have all the components of the wave be sinusoidal at a fixed f

now we need to get N modes that can meet this condition so you need to count all possible combos of \vec{n} . (approx. is treating it like its the volume of a sphere in space)

$$(1) \quad N^2 = n_1^2 + n_2^2 + n_3^2 \quad \vec{n} = \sqrt{n_1^2 + n_2^2 + n_3^2}$$

(Volume of n 's) = $\frac{4\pi}{3} (n_1^2 + n_2^2 + n_3^2)^{3/2}$



Now does the # Modes change per $\Delta\lambda$?

$$\text{volume} = \left(\frac{4}{3}\right) \left(\frac{4\pi}{3}\right) \left(\frac{4L^2}{32^3}\right)^{3/2} = \frac{8\pi L^3}{32^3}$$

this # of modes (N)

$$\rightarrow \frac{dN}{d\lambda} = \left| \frac{d}{d\lambda} \left[\frac{8\pi L^3}{32^3} \right] \right| = \frac{8\pi L^3}{32^3} \frac{1}{\lambda^4}$$

from the wave eqn $\frac{\partial^2 E}{\partial x^2} + \frac{\partial^2 E}{\partial y^2} + \frac{\partial^2 E}{\partial z^2} = \frac{\partial^2 E}{\partial \lambda^2}$
and knowing this eqn must give amplitude = 0 on walls, to form standing wave the boundary cond. are $E = E_0 \sin(n_1 \pi x) \sin(n_2 \pi y) \sin(n_3 \pi z)$
Substituting this into wave eqn $\sin(\frac{2\pi n_i \lambda}{L}) = 0 \Rightarrow n_i^2 = \frac{4L^2}{\lambda^2}$

$$N_1^2 + N_2^2 + N_3^2 = \frac{4L^2}{\lambda^2}$$

$$\frac{\partial^2 E}{\partial x^2} + \frac{\partial^2 E}{\partial y^2} + \frac{\partial^2 E}{\partial z^2} = \frac{\partial^2 E}{\partial \lambda^2}$$

\leftarrow square (black body)

modes is only positive but we used negative parts of sphere.

in 3D we only want (+)

so mult by $1/8$

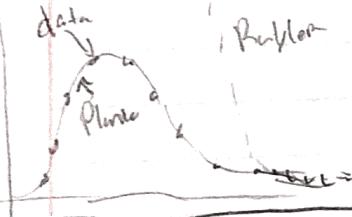
for each polarization, mult by 2,

$$\frac{du}{d\lambda} = \frac{1}{\text{volume}} \left(\frac{dE}{d\lambda} \right) = \frac{1}{(\text{volume})} \frac{K_i dN}{d\lambda}$$

$$= \frac{8\pi K_i}{\lambda^4}$$

E_{av} = Each Number of modes

works only in high f , ultraviolet catastrophe



$$\text{Planck guessed } \alpha(\lambda) = \frac{8\pi}{\lambda^4} \frac{hc/\lambda}{(e^{hc/\lambda kT} - 1)}$$

$$E = hf \rightarrow \text{from Einstein}$$

$$\text{but Energy at a level } B. \quad E_B = m hf$$

these particles have

a distribution given by

Bottom Prob. dist. (P_m)

from the definition of P_m

$\sum m P_m = 1$

$\sum m e^{-hf/mkT} = 1$

Davidson & Germer,
experiments looking to confirm idea through diffraction.

De Broglie's Hypothesis

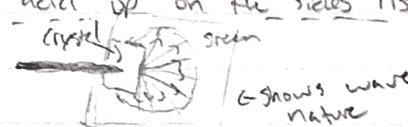
applying $E=hf$, Einstein's idea for energy of light in packets, to particles.

de Broglie said eh lets do it $p = \frac{h}{\lambda}$ or $\lambda = \frac{h}{p}$ for a particle

↳ this wave nature shows itself in diffraction experiments.

$a \sin \theta = n\lambda$ for peaks at certain angles * more intense in middle where each adds up on the sides its only itself -
Doing this on an electron

instead of through a slit, its on a crystal like



shows wave nature

now relative → a beam of electrons is accelerated from rest through a ΔV , acquiring a $K = e\Delta V$
and $p = \sqrt{2mk}$

For Davidson & Germer:

nx comes from the ~~infron~~ const. interfr. at 2π for example!

$\frac{\lambda}{d} \sin \theta = 2\pi$ so if its a wave you'll get the $\frac{U}{E} = \Delta V$, $\Delta U = -\Delta K$ (more on this later)

Peaks happening at that "beaten" and thus will be a maximum.

$(dsin\theta = n\lambda) \rightarrow$ they used wave $w/d = .215 \text{ nm}$, Peak at 50° , ($n=1$ no other peaks)

→ so $\lambda = c \sin \theta = .165 \text{ nm}$, they accelerate e^- through a ΔV of 54V so $K = e\Delta V = 54 \text{ eV}$

and a $p = \sqrt{2mk} = \frac{1}{c} \sqrt{2mc^2 k} = \frac{1}{c} \sqrt{2(51000 \text{ eV})(54 \text{ eV})} = \frac{1}{c} 7430 \text{ eV}$

for de Broglie $\lambda = \frac{hc}{pc} = \frac{1240 \text{ eV} \cdot \text{nm}}{7430 \text{ eV}} = .167 \text{ nm}$

notice how close $.165 \text{ nm}$ & $.167 \text{ nm}$ are, shows wave nature of particles. (if $\sin \theta$ goes over 90° you get max n , or the incident angle)

Double Slit w/ particles

from double slit ego $y_n = n\lambda \frac{D}{d}$ we will see how neutrons do,

$$\Delta y = y_{n+1} - y_n = \lambda \frac{D}{d} \rightarrow \lambda = \frac{\Delta y d}{D}$$

distance from max + center

$-\frac{dy}{d}$

max

$\frac{dy}{d}$

y_n

if you plug

diffraction = 1.89 nm for neutron and de broglie = 1.85 nm

ex through what ΔV do you have to acc electrons to see a 12 nm virus, the diameter is our de Broglie λ

$$p = \frac{h}{\lambda}; E_{tot} = \sqrt{(pc)^2 + (mc^2)^2} \rightarrow \sqrt{E_{tot}^2 - (mc^2)^2} = pc, E_{tot} = K + mc^2$$

$$pc = \sqrt{(K + mc^2)^2 - (mc^2)^2} \text{ if you mult out you get } pc = \sqrt{2Kmc^2 + K^2}, \text{ and } pc = \frac{hc}{\lambda}$$

$$\sqrt{K^2 + 2Kmc^2} = \frac{hc}{\lambda}, \text{ solve for } K \text{ & algebra by quadratic formula you get}$$

only plus root makes sense

$$K = -mc^2 \pm \sqrt{(mc^2)^2 + (\frac{hc}{\lambda})^2} \rightarrow K = -mc^2 + \sqrt{(mc^2)^2 + (\frac{hc}{\lambda})^2} \text{ recall } \Delta U = -\Delta K$$

$$\frac{\Delta U}{E} = \Delta V \text{ so } \Delta V = \frac{\Delta K}{e} = \frac{1}{e} (-mc^2 + \sqrt{(mc^2)^2 + (\frac{hc}{\lambda})^2})$$

Classical Uncertainty

Heisenberg's Uncertainty Principle

idea: $\lambda = \frac{h}{p}$ and it works for particles. If you have a small momentum the more its spread out and the less you know its location. If you have big p the less spreadout and the more you know its location.

Δ size, Δ uncertainty, Δ can be uncertainty?

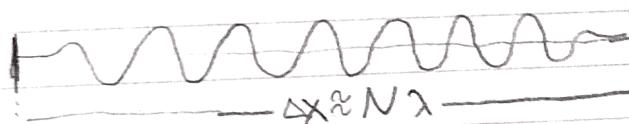
① since our λ is so short, all uncertainty $\Delta\lambda$

this our uncertainty $\Delta\lambda \sim \varepsilon\lambda$ making our uncertainty roughly as a fraction of actual $\Delta x \approx \lambda \rightarrow \Delta x \sim \lambda$ ex: $10 \pm 1\text{m}$, $\lambda = 10\text{m}$, $\Delta\lambda = \pm 1\text{m}$ so $\Delta\lambda \sim \varepsilon\lambda$, and $\varepsilon = 1$ in this case.

② the size of our Δx in this case is: $\Delta x \approx 2$ and $\Delta\lambda = \varepsilon\lambda$ the product is $(\Delta x)(\Delta\lambda) \sim \varepsilon^2 \lambda^2$ $\rightarrow 10\text{m} = 10\text{m}$ checks out (ex $\Delta\lambda = .2$, $\lambda = 7.5$ works out 1.5)

$\downarrow \Delta x$ means $\Delta\lambda$ or our uncertainty when we try to confine the packet.

For a larger wave packet,



$$\Delta\lambda \sim \frac{\varepsilon\lambda}{N}$$

$\Delta x \cdot \Delta \lambda \sim (N\lambda)\left(\frac{\varepsilon\lambda}{N}\right) = \varepsilon\lambda^2$

Same as our smaller case!

Frequency time Uncertainty

Trying to measure period, difficulty knowing where to stop and start cycle.

$\Delta t \sim \varepsilon T$ same logic as above \rightarrow close to $\Delta t = T$ Δt is like our Δx which is like the actual size or period / same formula!

$\Delta t = T$, $\Delta t \cdot \Delta T \sim \varepsilon T^2$ (or $\Delta x \cdot \text{Vave} = \Delta t$, $\Delta T \cdot \text{Vave} = \Delta t$ and $\lambda^2 = \text{Vave} \cdot T^2$) lets get it in f, $f = \frac{1}{T}$ Δf related to ΔT ? not $\Delta f \neq \frac{1}{\Delta T}$, should be directly related

to get relationship $\Delta f = -\frac{\Delta T}{T^2}$, converting to finite intervals, $\Delta f = \frac{\Delta T}{T^2}$ (interval is Δt + 1) plugging into $\Delta t \cdot \Delta T \sim \varepsilon T^2 \rightarrow \Delta f \Delta t \sim \varepsilon$

\rightarrow says longer duration of wave packet more precisely we can measure frequency.

HEISENBERG UNCERTAINTY RELATIONSHIPS

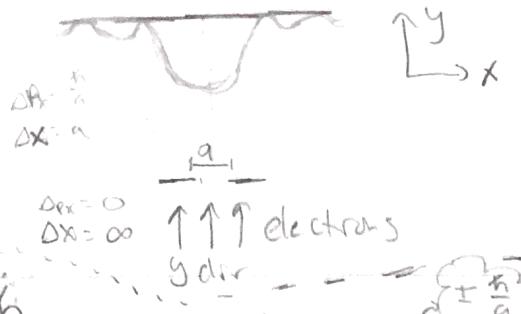
These relationships apply to all waves and thus de Broglie waves.

using $p = \frac{h}{\lambda}$ to find uncertainty $\Delta p = h \frac{\Delta\lambda}{\lambda^2} \rightarrow \Delta p = \frac{h \Delta\lambda}{\lambda^2}$ using $\Delta x \Delta\lambda \sim \varepsilon\lambda^2 \rightarrow \Delta x \Delta p \frac{\lambda^2}{h} \sim \varepsilon\lambda^2 \rightarrow \boxed{\Delta x \Delta p = \varepsilon h}$

the smallest possible value of $\Delta x \Delta p \beta = \frac{h}{4\pi} = \frac{1}{2}\hbar$

so $\Delta x \Delta p_x \geq \frac{\hbar}{2}$ or $\Delta y \Delta p_y \geq \frac{\hbar}{2}$ or $\Delta z \Delta p_z \geq \frac{\hbar}{2}$

Some Examples



electrons before slit
we know $\Delta p_x = 0$ so

$\Delta x = \infty$
electrons after slit

$\Delta x = a$ our uncertainty is confined
to the slit width

→ by $\Delta p \Delta x \sim \hbar$

Δp_x will then become $\frac{\hbar}{\Delta x} = \frac{\hbar}{a}$, so now knows uncertainty in momentum, no longer know if its going straight. So it spreads

→ an angle that specifies where a particle with $p_x = \frac{\hbar}{a}$ lands

$$\theta = \sin \theta \approx \tan \theta = \frac{p_x}{p_y} = \frac{\hbar/a}{p_y}, \text{ using de Broglie } \lambda = \hbar/p_y, \sin \theta = \frac{\lambda}{2\pi a}$$

energy and time uncertainty energy and time uncertainty = magnitude more below

$$\text{from } E = hf \rightarrow \Delta E = h \Delta f \rightarrow \Delta E \Delta t \sim \epsilon h \rightarrow \Delta E \Delta t \geq \frac{\hbar}{2}$$

Idea: particle w/ short lifetime $\Delta t \rightarrow 0$, measurement of its rest energy will be imprecise

A particle that's stable $\Delta t = \infty$ Energy can be measured w/ ∞ precision, $\Delta E = 0$.

Ex) an electron moves

in x_{dir} with $V = 3.6 \cdot 10^6 \text{ m/s}$

can measure its speed with ∞ precision,

$$p_x = m v_x = 3.3 \cdot 10^{-24} \text{ N/s} \cdot \text{kg}$$

$$\text{what } \Delta x = ? \quad \Delta p_x = 0.1 \cdot 3.3 \cdot 10^{-24} \text{ Nm}$$

$$\Delta x \sim \frac{\hbar}{\Delta p_x} = \frac{1.05 \cdot 10^{-34} \text{ J.s}}{3.3 \cdot 10^{-26}} = 3.2 \text{ nm}$$

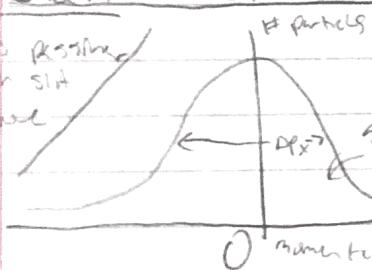
Much more magnet

of p_x

$\Delta p_x = \text{magnitude}$
↓

Statistical Interpretation

partic's passing
through slit
is measured
 p_x



Δp_x is similar to standard deviation

$$\sigma = \sqrt{(A^2)_{\text{avg}} - (A_{\text{avg}})^2}$$

avg is 0, mean is 0, $p_{x,\text{avg}} = 0$

$$\text{By analogy: } \Delta p_x = \sqrt{(p_x)^2_{\text{avg}} - 0} \rightarrow \Delta p_x = \sqrt{(p_x)^2_{\text{avg}}}$$

Ex.

Particle of mass m is
confined to a 1D box

What minimum KE of
this particle?



$$\Delta x = L \rightarrow \Delta x \Delta p \geq \hbar/2 \rightarrow (\Delta p) \geq \frac{\hbar}{2L}$$

$$(\Delta p)^2 = (p^2)_{\text{avg}} - (p_{\text{avg}})^2$$

average x, y components is 0, If squared first, don't cancel

$$E_{\text{Kavg}} = \frac{(p^2)_{\text{avg}}}{2m} = \frac{(\Delta p)^2}{2m} \geq \frac{\hbar^2}{8mL^2}$$

Ex

Nuclear beta decay, electrons are ejected
from the nucleus. Occasionally one escapes.
Charge of Ke should electron have?

$$\Delta x \sim \hbar / \Delta x = 19.7 \text{ MeV/c}$$

(diameter of nucleus)

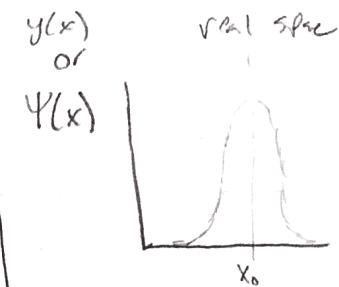
$$k = \sqrt{(pc)^2 + E_{\text{Kavg}}^2} - E_{\text{Kavg}} = 19 \text{ MeV}$$

electrons emitted from nucleus have $k \sim 1 \text{ keV}$ much smaller than the typical spread in energy required by
uncertainty principle for electrons confined to nucleus. Suggests electrons can't be in nucleus.

Wet v. v2 pf 25, estuarine Δx_1 square d₁
Wet v. v2 pf 35

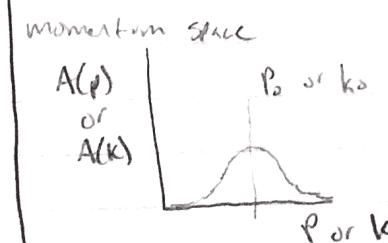
Wave Packets & Quantum

- particle is a wave packet b/c it has finite spatial extent
 - can't be just a simple plane wave with definite P and E must be described as a wave packet of a distribution of momenta.
 - wave packet is the sum of ∞ # of plane waves each weighted by $A(\vec{p})$



$$\Psi(x) = \int_{-\infty}^{\infty} A(k) \cos(kx) dk$$

$A(p)$ is distribution function of p
also correlated to the amplitude when
constructing packets



ex of
Gauss dist

$$\text{using } y(x) = \int_{-\infty}^{\infty} A(k) \cos(kx) dk \text{ w/ } A(k) = C e^{-\frac{a^2}{2}(k-k_0)^2}$$

$$y(x) = \frac{2(\sqrt{\pi})}{\alpha} e^{\frac{-2x^2}{\alpha^2}} \cos(k_0 x)$$

$$|y(x)|^2 = \frac{4C^2\pi}{a^2} e^{-\frac{4x^2}{a^2}} \cos^2(k_0 x)$$

to find probability $\int p(x) dx$

$$\text{probability density} = p(x) = \frac{1}{\sigma\sqrt{2\pi}} e^{-\frac{(x-\mu)^2}{2\sigma^2}}$$

to the center!

by inspection $a = \frac{1}{\partial k} \rightarrow g_k = \frac{1}{a}$ if you square y(k) which squares A(k)

→ probability density too, more later w/ Schrodinger. Just Notice it has sum state as $A(k)$ making it a probability density

~~probabilistic density too, more later w/~~
~~same shape as $A(k)$ making it a probability densit~~

Same shape as A(x) making it a probability
 definition of uncertainty: $\left[\text{comes from real space, } y(x)^2 \right] \left[\text{Amplitude A(x), associated w/ momentum} \right]$

$$\sigma = \text{uncertainty} \quad \text{so} \rightarrow \Delta x = \sigma_x, \quad \Delta k = \sigma_k$$

∴ let's right $p = \frac{b}{\lambda}$ in terms of K , $K = \frac{2\pi}{\lambda}$ so $p = \frac{b}{2\pi} = \frac{\pi K}{2}$

↳ let's right $p = \frac{b}{\lambda}$ in terms of k , $k = \frac{2\pi}{\lambda}$ so $p = \frac{\pi}{\lambda} = \frac{\pi}{\text{NA}} = p$

$$\Delta x \Delta p = \alpha_x \hbar \alpha_p$$

\downarrow \downarrow \downarrow \downarrow

$$\Delta x \Delta p = \left(\frac{a}{2}\right) \times \left(\frac{1}{a}\right)$$

$$\Delta x \Delta p = \left(\frac{\pi}{2}\right) h(\bar{a}) \quad , \quad \Delta x \Delta p = \epsilon$$

For gaussian distribution $\text{Its } \frac{\sigma}{2}$

↑
for any other distribution $> \frac{n}{2}$

Group Velocity & Phase velocity, with Quantum

$$V_p = \frac{\omega}{k} = \frac{2\pi f}{2\pi/\lambda} = \frac{2\pi E/h}{2\pi p/h} = \frac{E}{p}$$

$$V_G = \left(\frac{d\omega}{dk} \right)_{k=0} \rightarrow \left(\frac{dE}{dp} \right)_{p=p_0}$$

V_p depends on p , $E(p)$ or $E(k)$ is called the dispersion relation
 V_G is often evaluated at the center of momentum of $A(p)$

Ex for a low speed particle w/ V_0 , what's phase & group velocity

group

$$E_k = \frac{1}{2} m V_0^2 = \frac{p_0^2}{2m} \rightarrow V_0 = \left. \frac{dE}{dp} \right|_{p=p_0} = \frac{p_0}{m} = V_0$$

*the classical velocity corresponds to the group velocity of the packet, !!!!,

Phase

$$V_p = \frac{E}{p} = \frac{p}{2m} \rightarrow \text{depends on } p, \text{ no classical interpretation makes sense.}$$

Consider a wave packet that represents a confined particle at $t=0$

initial: $\Delta x_0, \Delta p_{x0}$ at time t ,

$$\Delta x_0 = \frac{\Delta p_{x0}}{m}$$

$$X = V_x t \quad V_x = V_{x0} \pm \Delta V_{x0} \rightarrow \Delta x = \sqrt{(\Delta x_0)^2 + (\Delta V_{x0} t)^2}$$

$$\Delta x = \sqrt{(\Delta x_0)^2 + (\Delta V_{x0} t)^2}$$

$$\text{and } \Delta p_{x0} = \frac{\Delta x_0}{\Delta x_0} \rightarrow \Delta x = \sqrt{(\Delta x_0)^2 + \left(\frac{\Delta x_0}{m \Delta x_0} t \right)^2}$$

the narrower we specialize in the wave pulse the quicker it spreads

Construction of Wave Packets

Wave Packets

→ as to as, particle
→ will be found anywhere

- building wave packets, pure sinusoidal waves not useful.
- odd waves of different λ together

$$A = \text{amplitude} \quad k = \text{wave number} = \frac{2\pi}{\lambda}$$

$$\lambda = \text{wavelength}$$

ex add together 2 waves

wave function?

$$y(x) = A_1 \cos(k_1 x) + A_2 \cos(k_2 x) = A_1 \cos\left(\frac{2\pi}{\lambda_1} x\right) + A_2 \cos\left(\frac{2\pi}{\lambda_2} x\right)$$

still \rightarrow gets you wave that has sorta localized spots but not complete, if you add more it gets better.

→ the regions still repeat even if you add a lot so you need to find a acceptable area.

→ if $A_1 = A_2 = A$ then $y(x)$ can be rewritten like this (lots of terms)

$$y(x) = 2A \cos\left(\frac{\pi x}{\lambda_1} - \frac{\pi x}{\lambda_2}\right) \cos\left(\frac{\pi x}{\lambda_1} + \frac{\pi x}{\lambda_2}\right)$$

→ if λ_1 and λ_2 are close together, $\Delta\lambda = \lambda_2 - \lambda_1$ is much smaller than λ_1 or λ_2

$$y(x) = 2A \cos\left(\frac{\Delta\lambda \pi x}{\lambda_{\text{avg}}}^2\right) \cos\left(\frac{2\pi x}{\lambda_{\text{avg}}}\right) \text{ where } \lambda_{\text{avg}} = \frac{\lambda_1 + \lambda_2}{2} \approx \lambda_1 \text{ or } \lambda_2$$

any finite comb of waves with discrete λ will have $\rightarrow \infty$ to ∞ , need to add a part that is large when we want to confine our particle, but goes to 0 as $x \rightarrow \infty$

→ $1/x$ would work $\rightarrow y(x) = \frac{2A}{x} \sin\left(\frac{\Delta\lambda \pi x}{\lambda_{\text{avg}}}^2\right) \cos\left(\frac{2\pi x}{\lambda_{\text{avg}}}\right)$, λ_0 is central wavelength

Otherwise cos blows up at $x=0$

another type is Gaussian func $e^{-\text{some part}}$, while e^x part pulls wave and makes it spread out

$$y(x) = A e^{-2(\Delta\lambda \pi x / \lambda_0)^2} \cdot \cos\left(\frac{2\pi x}{\lambda_0}\right)$$

keeps it localized

$$k = \frac{2\pi}{\lambda} \quad \omega = 2\pi f$$

Generally $y(x) = \sum A_i \cos(k_i x)$

If we have a cont set of wave numbers this sum becomes an integral,

① $y(x) = \int A(k) \cos(kx) dk$ ← integral is carried out over whatever wave numbers is permitted.

(ex if the range is) $\rightarrow k_0 - \Delta k/2$ to $k_0 + \Delta k/2$, cont distro, cent at k_0 , if all have same amplitude A_0

$$y(x) = \frac{2A_0}{\pi} \sin\left(\frac{\Delta k}{2} x\right) \cos(k_0 x) \in \begin{cases} \text{identical to } k_0 = 2\pi k_0 + \Delta k = 2\pi \Delta x / \lambda \\ -(k-k_0)^2 / [2(\Delta k)^2] \end{cases}$$

A better approx. can be made by letting $A(k) = A_0 e^{-\frac{(k-k_0)^2}{2(\Delta k)^2}}$

gives a range of wave numbers wth largest contribution at center, k_0 . And falls off to zero with characteristic width of Δk .

Using ① to the case with k ranging from $-\infty$ to ∞ gives

$$y(x) = A_0 \Delta k \sqrt{2\pi} e^{-\frac{(k-x)^2}{2(\Delta k)^2}} \cos(k_0 x)$$

Motion of wave packet.

* for added $V_{phase} = \frac{v_{wave}}{k_0}$

$y(x, t) = \cos(kx - \omega t)$ lets follow a peak, when $kx - \omega t = 0, 2\pi, 4\pi$

↳ the speed is taking time derivative $\frac{d}{dt}(kx - \omega t) = \frac{d}{dt}(\text{constant})^0 \rightarrow \frac{dx}{dt} = \frac{\omega}{k} = V_{phase}$

When you add waves w/ different λ or k 's

$$y(x, t) = A \cos(k_1 x - \omega_1 t) + A \cos(k_2 x - \omega_2 t)$$

you get

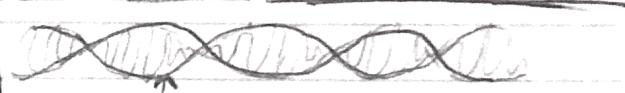
$$\Delta k = k_2 - k_1$$

$$k_m = \frac{k_1 + k_2}{2}$$

$$= 2A \cos\left(\frac{\Delta k}{2} x - \frac{\Delta \omega}{2} t\right) (\cos(k_m x - \omega_m t))$$

tells you about phase

tells you about envelope, velocity



$\cos\left(\frac{\Delta k}{2} x - \frac{\Delta \omega}{2} t\right)$ tells you how its rolling

but how fast?

$$\frac{d}{dt}\left(\frac{\Delta k x}{2} - \frac{\Delta \omega}{2} t\right) = \frac{d}{dt}(\text{const})$$

following peak

at $x = \frac{\Delta k}{2}, \frac{3\Delta k}{2}, \dots$

$\frac{\Delta \omega}{\Delta k} = \frac{dx}{dt} = V_{group}$ use infinitesimals, $\frac{dw}{dk}$,