2D Compressible MHD Code for Tearing Mode

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Abstract

To get more experience on coding fluid codes, I try to write this 2D compressible MHD code for tearing mode (magnetic reconnection).

In fact, I wrote the first version of this code in June 2010. Unfortunately, it doesn't work well. The framework of this code is from [Fu1995], section 7.4 and Appendix 2.

1 2D Compressible MHD Equations

All variables are in (x, z) plane, $\partial/\partial y=0$.

1.1 Original equations

Original equations (no Hall term, η, γ, υ constants):

$$\frac{\partial \rho}{\partial t} = -\mathbf{u} \cdot \nabla \rho - \rho \nabla \cdot \mathbf{u},$$

$$\frac{\partial p}{\partial t} = -\mathbf{u} \cdot \nabla p - \gamma p \nabla \cdot \mathbf{u},$$

$$\rho \frac{\partial \mathbf{u}}{\partial t} = -\rho \mathbf{u} \cdot \nabla \mathbf{u} - \nabla p - \frac{1}{\mu_0} \nabla \frac{B^2}{2} + \frac{1}{\mu_0} \mathbf{B} \cdot \nabla \mathbf{B} + \upsilon \nabla^2 \mathbf{u},$$

$$\frac{\partial \mathbf{B}}{\partial t} = \underbrace{\nabla \times (\mathbf{u} \times \mathbf{B})}_{-\mathbf{u} \cdot \nabla \mathbf{B} - \mathbf{B} \nabla \cdot \mathbf{u} + \mathbf{B} \cdot \nabla \mathbf{u}} - \underbrace{\nabla \times \left[\boldsymbol{\eta} (\nabla \times \mathbf{B}) \right]}_{p \nabla^2 \mathbf{B}}.$$
(0.1)

Where, we have used,

$$\nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t},$$

$$\mathbf{J} = \nabla \times \mathbf{B},$$

$$\mathbf{E} = -\mathbf{u} \times \mathbf{B} + \eta \mathbf{J}.$$
(0.2)

Here, considering that $\nabla \cdot {\bf B} = 0$, we use $A_{_{\scriptscriptstyle Y}}$ instead of $\, {\bf B}$, i.e.,

$$\mathbf{B} = \nabla \times \mathbf{A} = \nabla \times (0, A_{y}, 0) = (-\partial A_{y}/\partial z, 0, \partial A_{y}/\partial x). \tag{0.3}$$

And the last equation in (0.1) changes to,

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$$\nabla \times \left(\frac{\partial \mathbf{A}}{\partial t}\right) = \nabla \times \left\{\mathbf{u} \times (\nabla \times \mathbf{A}) - \frac{\eta}{\eta} \left[\nabla \times (\nabla \times \mathbf{A})\right]\right\},\tag{0.4}$$

Or,

$$\frac{\partial \mathbf{A}}{\partial t} = \mathbf{u} \times (\nabla \times \mathbf{A}) - \frac{\eta}{\eta} [\nabla \times (\nabla \times \mathbf{A})]. \tag{0.5}$$

Rewrite (0.1) and (0.5) to explicit form in 2D (with normalization),

$$\frac{\partial \rho}{\partial t} = -u_{x} \frac{\partial \rho}{\partial x} - u_{z} \frac{\partial \rho}{\partial z} - \rho \left(\frac{\partial u_{x}}{\partial x} + \frac{\partial u_{z}}{\partial z} \right),$$

$$\frac{\partial p}{\partial t} = -u_{x} \frac{\partial p}{\partial x} - u_{z} \frac{\partial p}{\partial z} - \gamma p \left(\frac{\partial u_{x}}{\partial x} + \frac{\partial u_{z}}{\partial z} \right),$$

$$\frac{\partial u_{x}}{\partial t} = -u_{x} \frac{\partial u_{x}}{\partial x} - u_{z} \frac{\partial u_{x}}{\partial z} - \frac{1}{\rho} \frac{\partial}{\partial x} \left(p + \frac{B^{2}}{2} \right) + \frac{1}{\rho} \left(B_{x} \frac{\partial B_{x}}{\partial x} + B_{z} \frac{\partial B_{x}}{\partial z} \right) + \frac{1}{\rho} \upsilon_{m} \left(\frac{\partial^{2} u_{x}}{\partial x^{2}} + \frac{\partial^{2} u_{x}}{\partial z^{2}} \right),$$

$$\frac{\partial u_{z}}{\partial t} = -u_{x} \frac{\partial u_{z}}{\partial x} - u_{z} \frac{\partial u_{z}}{\partial z} - \frac{1}{\rho} \frac{\partial}{\partial z} \left(p + \frac{B^{2}}{2} \right) + \frac{1}{\rho} \left(B_{x} \frac{\partial B_{z}}{\partial x} + B_{z} \frac{\partial B_{z}}{\partial z} \right) + \frac{1}{\rho} \upsilon_{m} \left(\frac{\partial^{2} u_{z}}{\partial x^{2}} + \frac{\partial^{2} u_{z}}{\partial z^{2}} \right),$$

$$\frac{\partial A_{y}}{\partial t} = -u_{x} \frac{\partial A_{y}}{\partial x} - u_{z} \frac{\partial A_{y}}{\partial z} + \eta_{m} \left[\frac{\partial^{2} A_{y}}{\partial x^{2}} + \frac{\partial^{2} A_{y}}{\partial z^{2}} \right].$$
(0.6)

1.2 Normalization

In (0.6), the normalization for density, pressure, length, magnetic field, velocity and time are $\rho_0, p_0, L_0, B_0, u_A = B_0/\sqrt{\mu_0 \rho_0}, \ \tau_A = L_0/u_A, \text{respectively}.$

Parameters are

$$\nu_{m} \to \frac{\nu}{u_{\Lambda}L_{0}\rho_{0}}, \ \eta_{m} \to \frac{\eta}{u_{\Lambda}L_{0}}, \tag{0.7}$$

Lundquist number $S=1/\eta_m$.

1.3 Discrete

In the right hand side of (0.6), for spatial derivative, we use

$$\frac{\partial f}{\partial x} \to \frac{f_{i+1} - f_{i-1}}{2\Delta x},$$

$$\frac{\partial f}{\partial z} \to \frac{f_{j+1} - f_{j-1}}{2\Delta z},$$

$$\frac{\partial^2 f}{\partial x^2} \to \frac{f_{i+1} + f_{i-1} - 2f_i}{\Delta x^2},$$

$$\frac{\partial^2 f}{\partial z^2} \to \frac{f_{j+1} + f_{j-1} - 2f_j}{\Delta z^2}.$$
(0.8)

In the left hand side, for time derivative, we use 4-th order Runge-Kutta scheme.

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2 Code Structure

Variables array x(ni,nj,5), for $~
ho,p,u_{\scriptscriptstyle x},u_{\scriptscriptstyle z},A_{\scriptscriptstyle y}$, respectively.

3 Simulation Results

4 References

[Fu1995] 傅竹风 & 胡友秋, 空间等离子体数值模拟, 安徽科学技术出版社, 1995.

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