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QUANTUM FIELD THEORY

Sep 13, 2020

Before. These notes come from Prof. Paton's QFT notes, Peskin-Schroeder QFT, and "Quantum Field Theory & Condensed Matter" - Shankar.

Enjoy!

Conventions

$$\hbar = c = 1$$

$$[Length] = [time] = [energy] = [mass]$$

$$g_{\mu\nu} = \gamma^{\mu\nu} = \text{diag}(1 - 1 - 1 - 1)$$

$$p^2 = p_\mu p^\mu = E^2 - |\vec{p}|^2 = mc^2 = m$$

$$x_\mu = g_{\mu\nu} x^\nu = (x^0, -\vec{x})$$

$$x^\mu = g^{\mu\nu} x_\nu = (x^0, \vec{x})$$

$$\partial_\mu = \frac{\partial}{\partial x^\mu} = \left(\frac{\partial}{\partial x^0}, \vec{\nabla} \right)$$

$$\epsilon^{0123} = \pm 1 ; \quad \epsilon_{0123} = -1$$

$$\epsilon^{1230} = -1$$

$$E = i \frac{\partial}{\partial x^0} , \quad \vec{p} = -i \vec{\nabla}$$

$$p^\mu = i \partial^\mu$$

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- $\sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}; \quad \sigma^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \quad \sigma^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$

- $\sigma^i \sigma^j = \delta^{ij} + i \epsilon^{ijk} \sigma^k$

- $\sigma^+ = \begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix}; \quad \sigma^- = \begin{pmatrix} 0 & 0 \\ 1 & 0 \end{pmatrix}$

- Heaviside step fn:

$$\theta(x) = \begin{cases} 0 & x \leq 0 \\ 1 & x > 0 \end{cases}$$

- Dirac delta fn: $\delta(x) = \frac{d}{dx} \theta(x)$

- n-dimensional Dirac δ -fn:

$$\int d^n x \delta^n(x) = 1$$

- FT:

$$f(x) = \int \frac{d^4 k}{(2\pi)^4} e^{-ik \cdot x} \tilde{f}(k)$$

$$\tilde{f}(k) = \int d^4 x e^{ik \cdot x} f(x)$$

- $\int d^4 x e^{ik \cdot x} = (2\pi)^4 \delta^{(4)}(k)$

- EM $\Phi = \frac{Q}{4\pi r}$ \leftarrow Coulomb potential

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• Fine structure constant:

$$\alpha = \frac{e^2}{4\pi} = \frac{e^2}{4\pi hc} \approx \frac{1}{137}$$

• Maxwell's eqn:

$$\epsilon^{\mu\nu\rho\sigma} \partial_\nu F_{\rho\sigma} = 0, \quad \partial_\mu F^{\mu\nu} = e j^\nu$$

where

$$A^\mu = (\vec{E}, \vec{A}), \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$

Elements of classical Field Theory

④ Lagrangian Field Theory:

$$S = \int \underline{L} dt = \int \underline{L}(\phi, \partial_\mu \phi) d^4x \quad \left(\underline{L} = L d^4x \right)$$

Principle of least action:

$$0 = \delta S \quad (\text{Lagrangian})$$

$$(\text{E-L eqn}) = \int d^4x \left\{ \frac{\partial \underline{L}}{\partial \phi} \delta \phi + \frac{\partial \underline{L}}{\partial (\partial_\mu \phi)} \delta (\partial_\mu \phi) \right\}$$

$$= \int d^4x \left\{ \frac{\partial \underline{L}}{\partial \phi} \delta \phi - \partial_\mu \left(\frac{\partial \underline{L}}{\partial (\partial_\mu \phi)} \right) \delta \phi + \partial_\mu \left[\frac{\partial \underline{L}}{\partial (\partial_\mu \phi)} \delta \phi \right] \right\}$$

$$\Rightarrow 0 = \partial_\mu \left[\frac{\partial \underline{L}}{\partial (\partial_\mu \phi)} \right] - \frac{\partial \underline{L}}{\partial \phi}$$

FTC \rightarrow term vanishes
at L-mass

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Euler - Lagrange Equations

$$\frac{\partial \mathcal{L}}{\partial \phi} = \partial_m \left\{ \frac{\partial \mathcal{L}}{\partial (\partial_m \phi)} \right\}$$

Ex

$$\mathcal{L} = \dot{\phi}^2 : \quad \frac{\partial \mathcal{L}}{\partial \phi} = 0 ; \quad \frac{\partial \mathcal{L}}{\partial (\partial_m \phi)} = 0 \Rightarrow \ddot{\phi} = 0$$

$$\begin{aligned} \mathcal{L} &= (\partial_m \phi) (\partial^m \phi) \\ \Rightarrow \frac{\partial \mathcal{L}}{\partial \phi} &= 0 ; \quad \frac{\partial \mathcal{L}}{\partial (\partial_m \phi)} = 2 \partial^m \phi \quad \} \Rightarrow \partial^m \phi = 0, \end{aligned}$$

Ex

Klein - Gordon field: (real)

$$\mathcal{L} = \frac{1}{2} (\partial_m \phi) (\partial^m \phi) - \frac{1}{2} m^2 \phi^2$$

$$\therefore \frac{\partial \mathcal{L}}{\partial \phi} = -m^2 \phi$$

$$\therefore \frac{\partial \mathcal{L}}{\partial (\partial_m \phi)} = \partial^m \phi.$$

relativistic particle
of mass m

$$E-L \Rightarrow -m^2 \phi = \partial_m \partial^m \phi = \square \phi$$

$$\rightarrow \boxed{(\square + m^2) \phi = 0}$$

Klein-Gordon Eqn.

$$\text{Ex } \phi = e^{-ipx} \Rightarrow (-p^2 + m^2) = 0 \Rightarrow p^2 = m^2$$

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Noether's Thm

Thm

For every continuous symmetry, there exists a conserved current j^μ which implies a local conservation law

$$\partial_\nu j^\nu = -\vec{\nabla} \cdot \vec{j} \quad \Leftrightarrow \quad \partial_\mu j^\mu = 0$$

Ex Conservation of charge ... (gauss' law)

$$\begin{aligned} Q &= \int j^0 d^3x \\ \Rightarrow \frac{dQ}{dt} &= \int \frac{d j^0}{dt} d^3x = - \int \vec{\nabla} \cdot \vec{j} d^3x \\ &= - \oint \vec{j} \cdot \vec{d}^2 r \end{aligned}$$

Idea Consider continuous transf. \rightarrow infinitesimally
(local)

$$(\star) \quad \phi(x) \rightarrow \phi'(x) = \phi(x) + \alpha \Delta \phi(x)$$

↑
small

(\star) is a symmetry if EOM invariant under (\star) .

$\Rightarrow S$ is invariant.

$\Rightarrow L$ must be invariant, up to $\alpha \partial_\mu J^\mu(x)$,
for some J^μ .

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Let us compare this expectation for ΔL to the result obtained by varying the fields ...

$$\begin{aligned}\alpha \Delta L &= \frac{\partial L}{\partial \phi} (\alpha \Delta \phi) + \left(\frac{\partial L}{\partial (\partial_\mu \phi)} \right) \partial_\mu (\alpha \Delta \phi) \\ &= \alpha \partial_\mu \left(\frac{\partial L}{\partial (\partial_\mu \phi)} \Delta \phi \right) + \alpha \underbrace{\left(\frac{\partial L}{\partial \phi} - \partial_\mu \left\{ \frac{\partial L}{\partial (\partial_\mu \phi)} \right\} \right) \Delta \phi}_{0}\end{aligned}$$

$$\Rightarrow \Delta L = \partial_\mu \left\{ \frac{\partial L}{\partial (\partial_\mu \phi)} \Delta \phi \right\}$$

So $\frac{\partial L}{\partial (\partial_\mu \phi)} \Delta \phi$ is the desired J^μ .

So that $\partial_\mu j^\mu(x) = 0$ where

$$j^\mu = \frac{\partial L}{\partial (\partial_\mu \phi)} \Delta \phi - J^\mu$$

Ex massless KG field

$$L = \frac{1}{2} (\partial_\mu \phi) (\partial^\mu \phi)$$

Corrado transformation:

$$\phi(x) \rightarrow \phi(x) + \alpha, \quad \Delta \phi = 1$$

$$j^\mu = \frac{\partial L}{\partial (\partial_\mu \phi)} \Delta \phi = \partial^\mu \phi.$$

Check $\partial_\mu j^\mu = \partial_\mu \partial^\mu \phi = \square \phi = 0$ since
 $(m^2 + \nabla^2) \phi = 0 \Rightarrow m=0$ \uparrow

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Ex Complex KG field

$$\mathcal{L} = (\partial_m \phi^+) (\partial^m \phi) - m^2 \phi^+ \phi.$$

again, EOM is

$$(m^2 + \Box) \phi = 0.$$

Symmetry: $\phi \rightarrow e^{i\alpha} \phi$.

For infinitesimal drift we have:

$$\begin{aligned}\alpha D\phi &= i\alpha \phi && \text{(Taylor expand)} \\ \alpha D\phi^+ &= -i\alpha \phi^+\end{aligned}$$

One can check that

$$j^\mu = i [(\partial^m \phi^+) \phi - \phi^+ (\partial^m \phi)]$$

is the conserved current.

[Ex]

Space-time translation:

Consider inf. translation:

$$x^m \rightarrow x^m - a^m$$

in field transforms this becomes:

$$\phi(x) \rightarrow \phi(x+a) = \phi(x) + a^m \partial_m \phi(x)$$

Lagrangian is a scalar \Rightarrow must transform the same way:

$$L \rightarrow L + a^m \partial_m L = L + a^v \partial_m (s_v^m L)$$

Compare this to the eqn:

$$L \rightarrow L + a \partial_m J^m,$$

we have

$$J^m = s_v^m L$$

\Rightarrow apply this, we find:

$$J^m = \frac{\partial L}{\partial (\partial_m \phi)} (\partial_m \phi) - s_v^m L$$

value a explicit...

$$T^v_m = \frac{\partial L}{\partial (\partial_v \phi)} \partial_m \phi - s_v^m L$$

\hookrightarrow STRESS-ENERGY TENSOR. (or Energy-Momentum tensor)

Conserved charge \Rightarrow the Hamiltonian

$$H = \int T^{00} d^3x = \int \mathcal{H} d^3x \quad (\text{time-translation})$$

$$\text{momentum} \rightarrow p^i = \int T^{0i} d^3x = - \int \pi \partial_i \phi d^3x \quad (\text{spatial translation})$$

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Ex Klein-Gordon field again

$$\mathcal{L} = \frac{1}{2} (\partial_\mu \phi) (\partial^\mu \phi) - \frac{1}{2} m^2 \phi^2.$$

$$T^{00} = \frac{\partial f}{\partial (\partial_0 \phi)} \partial^0 \phi - g^{00} \mathcal{L} = \dots$$

$$= \frac{1}{2} (\partial_0 \phi) (\partial^0 \phi) - \frac{1}{2} (\partial_i \phi) (\partial^i \phi) + \frac{1}{2} m^2 \phi^2$$

So

$$H = \int T^{00} d^3x = \int \left\{ \frac{1}{2} \left(\frac{\partial \phi}{\partial t} \right)^2 + \frac{1}{2} (\vec{\nabla} \phi)^2 + \frac{1}{2} m^2 \phi^2 \right\} d^3x$$

(note: all terms are positive ... (sum of squares))

→ can't fall into arbitrary negative energy

—————+

THE KLEIN-GORDON FIELD as HARMONIC OSCILLATOR

promote: ϕ, π to operators \Rightarrow impose suitable commutation relations

Reall...

$$[q_i, p_j] = i \delta_{ij}$$

$$[q_i, q_j] = [p_i, p_j] = 0.$$

$$\text{Harmonic oscillator: } H = \frac{1}{2m} p^2 + \frac{1}{2} m \omega^2 q^2$$

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Ladder operators.

- annihilation: $a = \frac{1}{\sqrt{2}} \left(g\sqrt{m\omega} + ip \frac{1}{\sqrt{m\omega}} \right)$

- creation: $a^\dagger = \frac{1}{\sqrt{2}} \left(g\sqrt{m\omega} - ip \frac{1}{\sqrt{m\omega}} \right)$

- $a^\dagger a = \frac{1}{\omega} H - \frac{1}{2} \quad (\Rightarrow H = \omega(a^\dagger a + \frac{1}{2}))$



operator...

- $|0\rangle, a|0\rangle = 0$.

$$|n\rangle = \frac{1}{\sqrt{n!}} \underbrace{a^\dagger \dots a^\dagger}_{n} |0\rangle$$

- $[a^\dagger a, a] = -a$

- $[a^\dagger a, a^\dagger] = a^\dagger$

- $[H, a] = -\omega a ; [H, a^\dagger] = \omega a^\dagger$

a lowers by ω

a^\dagger raises by ω

- $H a^\dagger |n\rangle = (E_n + \omega) a^\dagger |n\rangle$

- $H|0\rangle = \frac{1}{2}\omega|0\rangle \rightarrow E_0 = \frac{1}{2}\omega$

- $E_n = \left(n + \frac{1}{2}\right)\omega$

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For continuous systems ... commutation relations become ..

$$[\phi(x), \pi(y)] = i\delta^{(1)}(x-y)$$

$$[\phi(x), \phi(y)] = [\pi(x), \pi(y)] = 0.$$

Next, the Hamiltonian is now also an operator...
To find $\text{spec}(H)$, Fourier transf $\phi(x)$

$$\rightarrow \phi(x, t) = \int \frac{d^3 p}{(2\pi)^3} e^{i\vec{p} \cdot \vec{x}} \phi(\vec{p}, t)$$

Recall KG eqn: $(m^2 + D)\phi = 0$

$$\Rightarrow \left\{ \frac{\partial^2}{\partial t^2} + (|\vec{p}|^2 + m^2) \right\} \phi(\vec{p}, t) = 0$$

\rightarrow This is the same eqn as SHO with freq

$$\omega_p = \sqrt{|\vec{p}|^2 + m^2}$$

$$\text{Now, } H_{\text{SHO}} = \frac{1}{2}\vec{p}^2 + \frac{1}{2}w^2\phi^2 \quad (m=1)$$

\rightarrow know spectrum! $(n + \frac{1}{2})w$.

$$\phi = \frac{1}{\sqrt{2w}}(at + a) ; \quad \vec{p} = -i\sqrt{\frac{w}{2}}(a - at)$$

$$[a, a^\dagger] = 1.$$

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Since \vec{p} 's were convenient to work in positive space

$$\phi(\vec{x}) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2w_p}} (a_{\vec{p}} e^{i\vec{p} \cdot \vec{x}} + a_{\vec{p}}^\dagger e^{-i\vec{p} \cdot \vec{x}})$$

$$\pi(\vec{x}) = \int \frac{d^3 p}{(2\pi)^3} (-i) \sqrt{\frac{w_p}{2}} (a_{\vec{p}} e^{i\vec{p} \cdot \vec{x}} - a_{\vec{p}}^\dagger e^{-i\vec{p} \cdot \vec{x}})$$

Note

$\left. \begin{array}{l} a_{\vec{p}} \text{ goes with } e^{+i\vec{p} \cdot \vec{x}} \\ a_{\vec{p}}^\dagger \text{ goes with } e^{-i\vec{p} \cdot \vec{x}}. \end{array} \right\}$

9. Try to show that

$$\left\{ \begin{array}{l} [\phi(\vec{x}), \phi(\vec{y})] = 0 = [\pi(\vec{x}), \pi(\vec{y})] \\ [\phi(\vec{x}), \pi(\vec{y})] = i\delta^{(3)}(\vec{x} - \vec{y}). \end{array} \right.$$

* Can re-arrange...

$$\left\{ \begin{array}{l} \phi(\vec{x}) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2w_p}} (a_{\vec{p}} + a_{-\vec{p}}^\dagger) e^{i\vec{p} \cdot \vec{x}} \end{array} \right.$$

$$\left. \begin{array}{l} \pi(\vec{x}) = \int \frac{d^3 p}{(2\pi)^3} (-i) \sqrt{\frac{w_p}{2}} (a_{\vec{p}} - a_{-\vec{p}}^\dagger) e^{+i\vec{p} \cdot \vec{x}}. \end{array} \right.$$

→ set commutation relation between $a_{\vec{p}}$:

$$[a_{\vec{p}}, a_{\vec{p}'}^\dagger] = (2\pi)^3 \delta^{(3)}(\vec{p} - \vec{p}')$$

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Now, can check that

$$\begin{aligned} [\phi(x), \pi(x')] &= \int \frac{d^3 p d^3 p'}{(2\pi)^6} \left(-i\right) \sqrt{\frac{w_p}{w_{p'}}} x e^{i(p-x+p'x')} \\ &\quad \left([a_{-p}^\dagger, a_p] - [a_p, a_{-p}^\dagger] \right) \\ &= i \delta^{(3)}(x-x') \quad \checkmark \end{aligned}$$

• Now, can express Hamilton in terms of ladder ops

recall that \rightarrow KG field, loss time

$$\begin{aligned} H &= \int d^3 x \left\{ \frac{\partial L}{\partial (\partial_0 \phi)} \partial^0 \phi - g^{00} L \right\} \\ &= \int d^3 x \left\{ \frac{1}{2} (\partial_0 \phi)^2 + \frac{1}{2} (\nabla \phi)^2 + \frac{1}{2} m^2 \phi^2 \right\} \end{aligned}$$

To quantize, need to define π ... Turns out that

$$\pi(x) = \frac{\partial f}{\partial (\partial_0 \phi)} = \partial^0 \phi(x) \rightarrow \left(\text{like } p = \frac{\partial f}{\partial \dot{\phi}} \right)$$

$$H = \int d^3 x \left\{ \frac{1}{2} \pi^2 + \frac{1}{2} (\nabla \phi) \cdot (\nabla \phi) + \frac{1}{2} m^2 \phi^2 \right\}$$

with $\pi(x) = \int \frac{d^3 p}{(2\pi)^3} (-i) \sqrt{\frac{w_p}{2}} (a_p - a_{-p}^\dagger) e^{ip \cdot x}$

we get

$$\phi(x) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2w_p}} (a_p + a_{-p}^\dagger) e^{-ip \cdot x}$$

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$$H = \int d^3x \int \frac{d^3p d^3p'}{(2\pi)^6} e^{i(p+p') \cdot x} \left\{ -\frac{\sqrt{w_p w_{p'}}}{4} (a_p - a_{-p}^\dagger) a_{p'}^\dagger a_{-p'} + \frac{-p \cdot p' + m^2}{4\sqrt{w_p w_{p'}}} (a_p + a_{-p}^\dagger) (a_{p'} + a_{-p'}^\dagger) \right\}$$

results in $\delta(p-p')$
⇒ $p=p'$

Some $\delta^{(3)}$
will appear...

$$= \int \frac{d^3p}{(2\pi)^3} w_p (a_p^\dagger a_p + \frac{1}{2} [a_p, a_p^\dagger])$$

Simplifying

$$H = \int \frac{d^3p}{(2\pi)^3} w_p (a_p^\dagger a_p + \frac{1}{2} [a_p, a_p^\dagger])$$

→ can evaluate commutators ...

$$[H, a_p^\dagger] = w_p a_p^\dagger; [H, a_p] = -w_p a_p$$

With H , can find momentum operator ...

KG field → from $p^i = \int d^3x T^{0i} = -\int \pi \partial_i \phi d^3x$, we get

$$\begin{aligned} \vec{P} &= - \int d^3x \pi(x) \nabla \phi(x) \\ &= \int \frac{d^3p}{(2\pi)^3} \vec{p} a_p^\dagger a_p \end{aligned}$$

$E_p \rightarrow 0$

a_p^\dagger creates momentum \vec{p} & energy $w_p = \sqrt{|\vec{p}|^2 + m^2}$.

Excitation: $a_p^\dagger a_q^\dagger \dots |0\rangle$ = "particles".

↳ such excitation at p is a particle.

\Rightarrow get particle statistics -

Consider 2-particle state $|a_p^+ a_q^+ |0\rangle$.

Since $[a_p^+, a_q^+] = 0$, we have

$$[a_p^+ a_q^+ |0\rangle = a_q^+ a_p^+ |0\rangle]$$

\Rightarrow Klein Gordon particles follow Bose-Einstein stats.

* Normalization $\langle 0|0 \rangle = 1$.

$$|p\rangle \propto a_p^+ |0\rangle$$

This $\rightarrow \langle q|p \rangle = (2\pi)^3 \delta^{(3)}(q-p) \rightarrow$ NOT Lorentz inv.

PF Under a Lorentz boost $\rightarrow p'_3 = \gamma(p_3 + \beta E)$

using the identity $E' = \gamma(E + \beta p_3)$

$$\rightarrow \delta(f(x) - f(x_0)) = \frac{1}{|f'(x_0)|} \delta(x - x_0)$$

$$\text{we can write: } \delta^{(3)}(p-q) = \delta^6(p'-q') \cdot \left(\frac{dp'_3}{dp_3} \right)$$

$$\begin{aligned} & \underbrace{\delta(p_1-q_1)\delta(p_2-q_2)\delta(p_3-q_3)}_{\text{same boosted}} \\ &= \delta^{(3)}(p'-q') \cdot \gamma \left(1 + \beta \frac{dE}{dp_3} \right) \\ &= \delta^{(3)}(p'-q') \frac{\gamma}{E} (E + \beta p_3) \\ &= \delta^{(3)}(p'-q') (E'/E) \end{aligned}$$

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$$\text{So } \boxed{\delta^{(3)}(p-q) = \delta^{(3)}(p'-q') \left(\frac{E'}{E}\right)} \rightarrow 1 \Leftrightarrow E = E_p = E'$$

For normalization to work \rightarrow use E_p , not E .

$$\rightarrow \text{define: } |p\rangle = \sqrt{2E_p} a_p^\dagger |0\rangle$$

$$\text{So float } \boxed{\langle p|q\rangle = 2E_p (2\pi)^3 \delta^{(3)}(p-q)}$$

Completeness relation --

$$1 \xrightarrow{\text{particle}} \boxed{1 = \int \frac{dp^3}{(2\pi)^3} |p\rangle \frac{1}{2E_p} \langle p|}$$

RS Interpret $\phi(x)|0\rangle \dots$ we know that a_p^\dagger creates momentum p & energy $E_p = w_p$.

What about operator $\phi(x)$?

$$\phi(x)|0\rangle = \int \frac{dp^3}{(2\pi)^3} \frac{1}{2E_p} e^{-ipx} |p\rangle , \quad a_p^\dagger |0\rangle = \frac{1}{\sqrt{2E_p}} |p\rangle$$

Pf. By defn --

$$\phi(x) = \int \frac{dp^3}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} (a_p e^{ipx} + a_p^\dagger e^{-ipx})$$

$$\rightarrow \phi(x)|0\rangle = \int \frac{dp^3}{(2\pi)^3} \frac{1}{2E_p} e^{-ipx} |p\rangle \quad \square$$

$\rightarrow \phi(x)|0\rangle$ is a lin. superposition of single-particle states

Don't have well-defined momentum.

When nonrelativistic $\rightarrow E_p \approx \text{constant!}$

\Rightarrow $\boxed{\phi(x) \text{ acting on the vacuum, "creates a particle at position } x\text{"}}$

\hookrightarrow Confirm this by computing -

$$\langle 0 | \phi(x) | p \rangle = \langle 0 | \int \frac{d^3 p'}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} (a_p^\dagger e^{ip \cdot x} + a_p^- e^{-ip \cdot x}) \sqrt{2E_{p'}} a_{p'}^\dagger$$

$$\boxed{\langle 0 | \phi(x) | p \rangle = e^{ip \cdot x}}$$

\hookrightarrow Interpretation: position-space representation of the single-particle wfns of the state $|p\rangle$, just like

$$\boxed{\langle n | p \rangle \propto e^{ip \cdot x}} \text{ in QM!}$$

$$\langle 0 | \phi(x) \sim \langle x | \dots \quad (\text{don't take this literally, ofc})$$

Note Hw1, Hw2 we copy, so we'll skip for now.

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THE KLEIN-GORDON FIELD IN SPACETIME

Last time \rightarrow we quantized KG field in the Schrödinger picture.

\rightarrow Now, switch to Heisenberg picture.

Recall... Schrödinger picture:

$U(t) = e^{-iHt}$ is the time evolution.

$| \Psi(t) \rangle = e^{-iHt} | \Psi(0) \rangle$ $\xrightarrow{\text{state evolves in time}}$

\rightarrow In the Heisenberg picture, ... Operators evolve in time

$$\hat{\theta}(t) = U^\dagger(t) \hat{\theta}(0) U(t).$$

So that

$$\langle \Psi_1 | \hat{\theta}(t) | \Psi_2 \rangle = \langle \Psi_1 | \hat{\theta}(t) | \Psi_2 \rangle$$

\downarrow

Heisenberg

\downarrow

Schrödinger.

\rightarrow make the operators ϕ & π time-dependent like this

$$\phi(x) \rightarrow \phi(x, t) = e^{iHt} \phi(x) e^{-iHt}$$

$$\pi(x) \rightarrow \pi(x, t) = e^{iHt} \pi(x) e^{-iHt}$$

Recall Heisenberg eqn of motion $i\frac{\partial}{\partial t} \hat{\theta} = [\hat{\theta}, H]$

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which gives, upon substituting in $\phi(x, t)$, $\pi(x, t)$

$$i \frac{\partial}{\partial t} \phi(x, t) = \left[\phi(x, t), \int d^3x' \left\{ \frac{1}{2}\pi^2 + \frac{1}{2}(\nabla\phi)^2 + \frac{1}{2}m^2\phi^2 \right\} \right]$$

$$(\phi \leftrightarrow \phi) \quad = \int d^3x' \left(i\delta^{(3)}(x-x') \bar{\pi}'(x, t) \right)$$

\rightarrow only nontrivial term is L^2 .

$$= i\bar{\pi}(x, t) \Rightarrow \boxed{\frac{\partial}{\partial t} \phi(x, t) = \bar{\pi}(x, t)}$$

and

$$i \frac{\partial}{\partial t} \pi(x, t) = \left[\pi(x, t), \int d^3x' \left\{ \frac{1}{2}\pi^2 + \frac{1}{2}(\nabla\phi)^2 + \frac{1}{2}m^2\phi^2 \right\} \right]$$

$$= \int d^3x' \left(-i\delta^{(3)}(x-x') (-\nabla^2 + m^2) \phi(x', t) \right)$$

$$(\text{integrate by parts here}) = -i(-\nabla^2 + m^2) \phi(x, t)$$

$$\Rightarrow \boxed{\frac{\partial}{\partial t} \pi(x, t) = (m^2 - \nabla^2) \phi(x, t)}$$

Combining these 2 results we get ...

$$\boxed{\frac{\partial^2}{\partial t^2} \phi(x, t) = (\nabla^2 - m^2) \phi(x, t)}$$

\hookrightarrow rearranging this gives

$$\boxed{(\nabla^2 + m^2) \phi(x, t) = 0} \rightarrow \text{just the KG eqn ...}$$

- Now, can better understand the time dependence of $\phi(x)$, $\pi(x)$ by writing them in terms of creation & annihilation ops.

Recall: $H_{\text{ap}} = a_p (H - E_p) \rightarrow$ from comm. rule.

\rightarrow (proof by induction)

$$H^n a_p = a_p (H - E_p)^n$$

Similarly,

$$H^n a_p^+ = a_p^+ (H + E_p)^n$$

\rightarrow So, we have

$$e^{iEt} a_p e^{-iEt} = a_p e^{-iEpt} \rightarrow \text{from above ...}$$

and

$$e^{iEt} a_p^+ e^{-iEt} = a_p^+ e^{+iEpt}$$

\rightarrow Now we want to write $\phi(x, t)$ in terms of these operators. (since $\phi(x)$ is a sum of a & a^+)

We know that $\phi(x, t) = e^{iEt} \phi(x) e^{-iEt}$.

and from before ...

$$\phi(x) = \phi(x, 0) = \int \frac{dp}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} (a_p e^{ip \cdot x} + a_p^+ e^{-ip \cdot x})$$

Substitute this into $\phi(x, t) = e^{iEt} \phi(x) e^{-iEt}$ we find

(21)

$$\phi(x, t) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left\{ a_p e^{-iE_p t} e^{ip \cdot x} + a_p^+ e^{iE_p t} e^{-ip \cdot x} \right\}$$

now, note that $p^0 = E_p$

$$\Rightarrow p \cdot x = E_p \cdot x^0 - \vec{p} \cdot \vec{x} \Big|_{p^0 = E_p} = E_p t - \vec{p} \cdot \vec{x} \Big|_{p^0 = E_p}$$

So,

$$\boxed{\phi(x, t) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{\sqrt{2E_p}} \left\{ a_p e^{-i\vec{p} \cdot \vec{x}} + a_p^+ e^{+i\vec{p} \cdot \vec{x}} \right\} \Big|_{p^0 = E_p}}$$

Similarly, we can find

$$\boxed{\Pi(x, t) = \int \frac{d^3 p}{(2\pi)^3} (-i) \sqrt{\frac{E_p}{2}} (a_p e^{-i\vec{p} \cdot \vec{x}} - a_p^+ e^{+i\vec{p} \cdot \vec{x}}) \Big|_{p^0 = E_p}}$$

from defn or from $\Pi(x, t) = \frac{\partial}{\partial t} \phi(x, t)$.

Note we can also do everything, but starting from P and not Π . But we won't worry about that.



Causality Note that causality is broken unless without the presence of fields.

→ QFT resolves causality problem.

In our present formalism, still in Heisenberg picture --

-- the amplitude for a particle to propagate from y to x is given by

$$\langle 0 | \phi(x) \phi(y) | 0 \rangle$$

"Two-point correlation function"

Let $\boxed{\mathcal{D}(x-y) = \langle 0 | \phi(x) \phi(y) | 0 \rangle}$

$$\Rightarrow \mathcal{D}(x-y) = \langle 0 | \phi(x) \phi(y) | 0 \rangle$$

$$= \langle 0 | \text{in terms of } a_p, a_p^\dagger \dots | 0 \rangle$$

$$= \langle 0 | a_p^\dagger a_p^\dagger | 0 \rangle =$$

$$= (2\pi)^3 \delta^{(3)}(\mathbf{p} - \mathbf{q})$$

$$= \int \frac{d^3 p}{(2\pi)^3} \frac{1}{2E_p} e^{-ip \cdot (x-y)}$$

More explicitly --

$$\mathcal{D}(x-y) = \langle 0 | \phi(x) \phi(y) | 0 \rangle$$

$$\begin{cases} \mathbf{p}' = \mathbf{p}_p \\ E_p' = E_{p'} \end{cases}$$

$$= \int \frac{d^3 p d^3 p'}{(2\pi)^6} \frac{1}{\sqrt{2E_p}} \frac{1}{\sqrt{2E_{p'}}} \langle 0 | e^{-ip \cdot x} e^{ip' \cdot y} a_p^\dagger a_{p'}^\dagger | 0 \rangle$$

$$= \int \frac{d^3 p d^3 p'}{(2\pi)^6} \left(\frac{1}{\sqrt{2E_p}} \right) \left(\frac{1}{\sqrt{2E_{p'}}} \right) e^{-ip \cdot x - ip' \cdot y} (2\pi)^3 \delta^{(3)}(\mathbf{p} - \mathbf{p}')$$

$$\mathcal{D}(x-y) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{2E_p} e^{-ip \cdot (x-y)} \quad \square$$

Note that this integral is Lorentz-invariant.

→ Now, let us evaluate this integral for some particular values of $x-y$...

(1) Suppose that $x-y = (t, \vec{v}, 0, 0)$, then

$$\mathcal{D}(x-y) = \frac{4\pi}{(2\pi)^3} \int_0^\infty dp \frac{p^2}{2\sqrt{p^2+m^2}} e^{-i\sqrt{p^2+m^2}t}$$

$$(\text{timelike}) = \frac{1}{4\pi^2} \int_m^\infty dE \sqrt{E^2-m^2} e^{-iEt}$$

$$= e^{-imt} \xrightarrow[t \rightarrow \infty]{\text{dominated by region above}} \text{dominated by region above}$$

$$p \approx 0 -$$

(2) Suppose that $x-y = (0, \vec{x}-\vec{y}) = (0, \vec{r})$ then

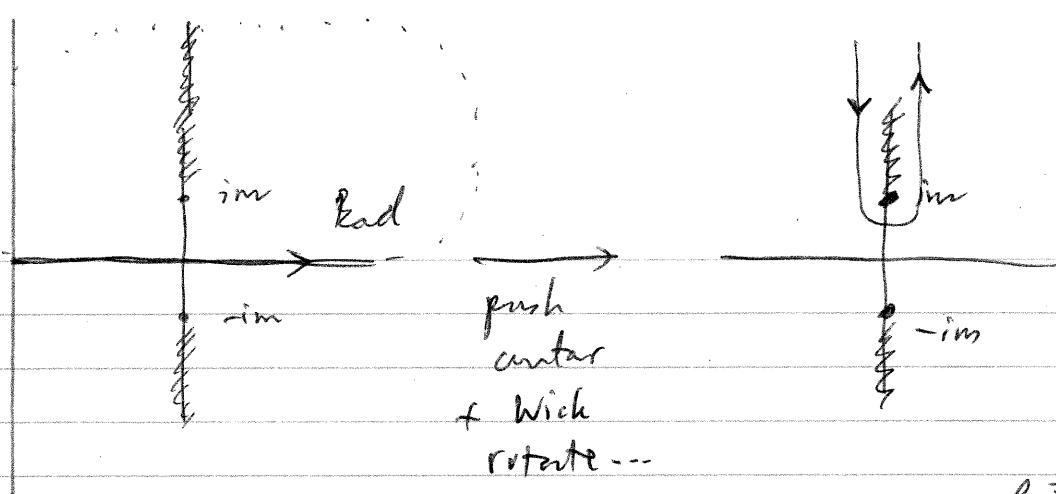
$$\mathcal{D}(x-y) = \int \frac{d^3p}{(2\pi)^3} \frac{1}{2E_p} e^{i\vec{p} \cdot \vec{r}}$$

$$= \frac{2\pi}{(2\pi)^3} \int_0^\infty dp \frac{p^2}{2E_p} \frac{e^{ipr} - e^{-ipr}}{ipr}$$

$$= \frac{-i}{2(2\pi)^2 r} \int_{-\infty}^\infty dp \frac{pe^{ipr}}{\sqrt{p^2+m^2}}$$

→ branch cut @ tim (singularity)

→ must change contour... \rightarrow which route



To get

$$\mathcal{D}(x-y) = \frac{1}{4\pi^2 r} \int_m^\infty d\rho \frac{\rho e^{-i\rho r}}{\sqrt{\rho^2 - m^2}} \quad (\text{Wick Rotate})$$

$$\Rightarrow \boxed{\mathcal{D}(x-y) = \frac{e^{-mr}}{r}} \rightarrow \underline{\text{non zero!}}$$

(more details about integrals like this can be found in Zee's QFT in a nutshell...)

What does it mean for $\mathcal{D}(x-y)$ to be nonzero when $x-y$ is spacelike?

We saw that when $(x-y)^m (x-y)_m = -(\vec{x}-\vec{y})^2 < 0$
is spacelike, cannot have causality between
 $x-y$.

$\mathcal{D}(x-y) \neq 0 \Rightarrow ???$ paradox?

\rightarrow No! To discuss causality, we should ask not whether particles can propagate over spacelike intervals ...

... but whether a "measurement" at one point can affect a measurement at another point where separation from the first is spacelike -

→ No violation of causality b/c no signal can be transmitted → no info is exchanged.

→ Suppose we have a local measurement $\phi(x)$, call this $\phi(x)$. & a local measurement $\phi(y)$, called $\phi(y)$

So long as $[\phi(x), \phi(y)] = 0$, the 2 measurements don't affect one another.

→ measure the field $\phi @ x = @ y$,

If $[\phi(x), \phi(y)] = 0$ when $(x-y)^2 < 0$ then we've proved

$$\rightarrow [\phi(x), \phi(y)] = \int \frac{d^3 p d^3 p'}{(2\pi)^6 2\sqrt{E_p E_{p'}}} \left\{ e^{-ip \cdot x} [a_p, a_{p'}^\dagger] e^{ip' \cdot y} + e^{ip \cdot x} [a_p^\dagger, a_{p'}] e^{-ip' \cdot y} \right\}$$

$$\rightarrow [\phi(x), \phi(y)] = \int \frac{d^3 p d^3 p'}{(2\pi)^6 2\sqrt{E_p E_{p'}}} \times \left\{ [(a_p e^{-ip \cdot x} + a_p^\dagger e^{ip \cdot x}), (a_p^\dagger e^{-ip \cdot y} + a_p^\dagger e^{ip \cdot y})] \right\}$$

$$= \int \frac{d^3 p d^3 p'}{(2\pi)^6 2\sqrt{E_p E_{p'}}} \left\{ e^{-ip \cdot x} [a_p, a_{p'}^\dagger] e^{ip' \cdot y} + e^{ip \cdot x} [a_p^\dagger, a_{p'}] e^{-ip' \cdot y} \right\}$$

$$(2\pi)^3 \delta^3(p-p') - (2\pi)^3 \delta^3(p-p')$$

$$= \int \frac{d^3 p}{(2\pi)^3} \left(\frac{1}{2E_p} \right) \left\{ e^{-ip(x-y)} - e^{-ip(y-x)} \right\} = D(x-y) - D(y-x)$$

$$\text{So } [\phi(x), \phi(y)] = D(x-y) - D(y-x).$$

Since $D(y-x)$ is Lorentz invariant, can take

$$y-x \rightarrow x-y. \text{ (possible b/c } (x-y)^2 < 0)$$

Note that when $(x-y)^2 > 0 \rightarrow$ there's no Lorentz transform that takes $y-x \rightarrow x-y$

\rightarrow so this is why possible because $(x-y)^2 < 0$ (spacelike).

$$\text{So } D(y-x) = D(x-y)$$

$$\text{So, } [\phi(x), \phi(y)] = 0 \quad \checkmark$$

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④ The Klein-Gordon Propagator