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 Course: **8.421 - AMO I**  
 Problem set: **#8**  
 Due: Friday, April 8, 2022.

**1. Optical Traps and Scattering.** What are the proper power and wavelength needed to trap an ultracold atomic gas? Consider an alkali atom with resonance frequency  $\omega_0$  on the principal  $nS \rightarrow nP$  transition. A sample of atoms in the ground state  $nS$  are exposed to monochromatic radiation of intensity  $I$  and frequency  $\omega_L < \omega_0$ . Using the fact that essentially all of the oscillator strength out of the ground state comes from the  $nS \rightarrow nP$  transition, we have

$$\alpha(\omega_L) \approx \frac{2e^2}{\hbar} |\langle nP | z | nS \rangle|^2 \frac{\omega_0}{\omega_0^2 - \omega_L^2} \implies |\langle nP | z | nS \rangle|^2 = \frac{\hbar \alpha(\omega_L)}{2e^2} \frac{\omega_0^2 - \omega_L^2}{\omega_0}.$$

(a) AC Stark shift:

(i) From lecture, the AC Stark shift  $U_i$  from time-dependent perturbation theory is given by

$$U_i = -\frac{1}{4} \alpha(\omega_L) \mathcal{E}^2 = -\frac{2I\alpha(\omega_L)}{4c\epsilon_0} = -\frac{I\alpha(\omega_L)}{2c\epsilon_0}.$$

(ii) Now, we want to use the rotating wave approximation to obtain the AC Stark shift. This can be done by first writing down the true (symmetrized) Hamiltonian:

$$\mathcal{H} = \frac{\hbar}{2} \begin{pmatrix} -\omega_0 & \omega_R e^{i\omega_L t} \\ \omega_R^* e^{-i\omega_L t} & \omega_0 \end{pmatrix}.$$

By going into the rotating frame, plus making the rotating wave approximation, we find that

$$\mathcal{H}_{\text{rot}}^{\text{RWA}} = \frac{\hbar}{2} \begin{pmatrix} -\delta & \omega_R \\ \omega_R & \delta \end{pmatrix}$$

where  $\delta = \omega_0 - \omega_L$ . The energy shifts can be obtained from the eigenvalues:

$$\Delta E = \pm \frac{\hbar}{2} \sqrt{\omega_R^2 + \delta^2} = \pm \frac{\hbar}{2} \sqrt{\omega_R^2 + (\omega_0 - \omega_L)^2} \approx \pm \frac{\hbar(\omega_0 - \omega_L)}{2} \pm \frac{\hbar\omega_R^2}{4(\omega_0 - \omega_L)}$$

where we have used the fact that  $\omega_R \ll |\omega_0 - \omega_L|$ . From here, we find that the shift is

$$U_{ii} = -\frac{\hbar\omega_R^2}{4(\omega_0 - \omega_L)}$$

In particular, the energy of the lower state gets shifted down while the energy of the higher state gets shifted up (since we're red-detuning).

(iii) From the previous two parts, we find that

$$\frac{U_i}{U_{ii}} = \frac{I\alpha(\omega_L)}{2c\epsilon_0} \frac{4(\omega_0 - \omega_L)}{\hbar\omega_R^2} = \frac{I}{c\epsilon_0\omega_R^2} \frac{4e^2}{\hbar^2} |\langle nP | z | nS \rangle|^2 \frac{\omega_0}{\omega_0 + \omega_L}.$$

To simplify this, we must write the Rabi frequency in terms of the intensity:

$$\omega_R = \frac{e\mathcal{E} |\langle nS | z | nP \rangle|}{\hbar} \implies \omega_R^2 = \frac{e^2 |\langle nS | z | nP \rangle|^2}{\hbar^2} \frac{2I}{c\epsilon_0}.$$

With this, we have

$$\boxed{\frac{U_i}{U_{ii}} = \frac{2\omega_0}{\omega_0 + \omega_L}}$$

When  $\omega_L \approx 0$ , we have

$$\frac{U_i}{U_{ii}} \approx 2.$$

When  $\omega_L \approx \omega_0$ , we may write  $\omega_L + \omega_0 = 2\omega_0$ , so that

$$\frac{U_i}{U_{ii}} \approx 1.$$

We see that if the intensity has spatial structure, with the appropriate detuning, the AC Stark shift can have energy minima where the atoms can be trapped.

(b) From time-dependent perturbation theory, the amplitude of the excited state is

$$c_2(t) = \frac{e\mathcal{E}}{2\hbar} \langle nS | z | nP \rangle \left[ \frac{e^{i(\omega_0 + \omega_L)t} - 1}{\omega_0 - \omega_L} + \frac{e^{i(\omega_0 - \omega_L)t} - 1}{\omega_0 + \omega_L} \right].$$

Ignoring the  $-1$  terms which are associated with transients, we have

$$P_e(t) = \frac{e^2 \mathcal{E}^2}{4\hbar^2} |\langle nS | z | nP \rangle|^2 \frac{2[\omega_0^2 + \omega_L^2 + (\omega_0^2 - \omega_L^2) \cos(2\omega_L t)]}{(\omega_0^2 - \omega_L^2)^2}.$$

After time-averaging, this quantity is

$$P_{e,i} = \boxed{\frac{\omega_R^2}{2} \frac{\omega_0^2 + \omega_L^2}{(\omega_0^2 - \omega_L^2)^2}}$$

In the RWA picture, we know that

$$P_{e,ii}(t) = \frac{\omega_R^2}{\omega_R^2 + \delta^2} \sin^2 \left( \frac{\sqrt{\omega_R^2 + (\omega_0 - \omega_L)^2} t}{2} \right).$$

After time-averaging this is

$$P_{e,ii} = \frac{\omega_R^2}{2(\omega_R^2 + \delta^2)} \approx \boxed{\frac{\omega_R^2}{2(\omega_0 - \omega_L)^2}}$$

where we have used the approximation that the Rabi frequency is much less than the detuning.

(c) Calculate the photon scattering rate:

(i) Starting with

$$P = \frac{ck^4 |d|^2}{3} = \frac{\omega_L^4}{3c^3} |d|^2,$$

if we say  $d = \alpha(\omega_L)\mathcal{E}$  then we have

$$R_{sc} = \frac{P}{\hbar\omega_L} = \frac{\omega_L^3}{3\hbar c^3} |\alpha(\omega_L)|^2 \mathcal{E}^2 = \frac{\omega_L^3}{3\hbar c^3} |\alpha(\omega_L)|^2 \frac{8\pi I}{c}$$

where we have converted the intensity into CGS units. Now we recall from perturbation theory that

$$\alpha(\omega_L) = \frac{2e^2}{\hbar} |\langle nS | z | nP \rangle|^2 \frac{\omega_0}{\omega_0^2 - \omega_L^2} = \frac{e^2}{\hbar} |\langle nS | z | nP \rangle|^2 \left( \frac{1}{\omega_0 - \omega_L} + \frac{1}{\omega_0 + \omega_L} \right).$$

From here we find that

$$|\alpha(\omega_L)|^2 = \frac{e^4}{\hbar^2} |\langle nS | z | nP \rangle|^4 \left( \frac{1}{\omega_0 - \omega_L} + \frac{1}{\omega_0 + \omega_L} \right)^2.$$

Putting everything together, we find

$$\begin{aligned} R_{sc,i} &= \frac{\omega_L^3}{3\hbar c^3} \frac{8\pi I}{c} \frac{e^4}{\hbar^2} |\langle nS | z | nP \rangle|^4 \left( \frac{1}{\omega_0 - \omega_L} + \frac{1}{\omega_0 + \omega_L} \right)^2 \\ &= \frac{8\pi I \omega_L^3 e^4}{3c^4 \hbar^3} |\langle nS | z | nP \rangle|^4 \left( \frac{1}{\omega_0 - \omega_L} + \frac{1}{\omega_0 + \omega_L} \right)^2. \end{aligned}$$

Under RWA, we simply drop the counter-rotating term to find

$$R_{sc,ii} = \frac{8\pi I \omega_L^3 e^4}{3c^4 \hbar^3} |\langle nS | z | nP \rangle|^4 \left( \frac{1}{\omega_0 - \omega_L} \right)^2.$$

(ii) Let us write  $R_{sc,i}$  and  $R_{sc,ii}$  in terms of the spontaneous emission rate  $\Gamma = 4e^2 \omega_0^3 |\langle nS | z | nP \rangle|^2 / 3\hbar c^3$ :

$$\begin{aligned} R_{sc,i} &= \frac{8\pi I \omega_L^3 e^4}{3c^4 \hbar^3} |\langle nS | z | nP \rangle|^4 \left( \frac{1}{\omega_0 - \omega_L} + \frac{1}{\omega_0 + \omega_L} \right)^2 \\ &= \frac{8\pi I \omega_L^3}{3c^4 \hbar^3} \frac{9\Gamma^2 \hbar^2 c^6}{16\omega_0^6} \left( \frac{1}{\omega_0 - \omega_L} + \frac{1}{\omega_0 + \omega_L} \right)^2 \\ &= \frac{3\pi c^2}{2\hbar \omega_0^3} \left( \frac{\omega_L}{\omega_0} \right)^3 \left( \frac{\Gamma}{\omega_0 - \omega_L} + \frac{\Gamma}{\omega_0 + \omega_L} \right)^2 I, \end{aligned}$$

which is a well-known result given in many textbooks. Making the RWA, we find

$$R_{sc,ii} = \frac{3\pi c^2}{2\hbar \omega_0^3} \left( \frac{\omega_L}{\omega_0} \right)^3 \left( \frac{\Gamma}{\omega_0 - \omega_L} \right)^2 I.$$

From Part (b), we have that

$$P_{e,ii} = \frac{\omega_R^2}{2(\omega_0 - \omega_L)^2},$$

which gives

$$\begin{aligned} R_{sc,ii} &= \frac{3\pi c^2}{2\hbar \omega_0^3} \left( \frac{\omega_L}{\omega_0} \right)^3 \left( \frac{\Gamma}{\omega_0 - \omega_L} \right)^2 I \\ &= \frac{3\pi c^2}{2\hbar \omega_0^3} \left( \frac{\omega_L}{\omega_0} \right)^3 \frac{\Gamma^2}{\omega_R^2} 2P_{e,ii} \\ &= \frac{3\pi c^2}{\hbar \omega_0^3} \left( \frac{\omega_L}{\omega_0} \right)^3 \Gamma^2 P_{e,ii} \frac{\hbar^2}{e^2} \frac{c}{8\pi} \frac{4e^2 \omega_0^3}{3\hbar c^3 \Gamma} \\ &= \boxed{\frac{1}{2} \left( \frac{\omega_L}{\omega_0} \right)^3 \Gamma P_{e,ii}} \end{aligned}$$

- (iii) (Optional) We describe scattering as spontaneous emission from a virtual energy level. The energy diagram for this process is:

The lifetime of this virtual state is:

*blah*

- (d) (i) The  $D_{1,2}$  lines of Na have  $\omega_0 \approx 2\pi \cdot 508$  THz. Now the infrared laser at 985 nm has  $\omega_L = 2\pi \cdot 304$  THz, which corresponds to a detuning of  $2\pi \cdot 204$  THz, which is much larger than the detuning of the yellow laser (a few GHz). This means that the RWA expression is more suitable for the yellow laser.
- (ii) Here we want to calculate the required power and scattering rate for each of the two types of lasers. Using results from time-dependent perturbation theory, we have two equations:

$$U_i = \frac{I(0)\alpha(\omega_L)}{2c\epsilon_0} = \frac{3\pi c^2}{2\omega_0^3} \left( \frac{\Gamma}{\omega_0 - \omega_L} + \frac{\Gamma}{\omega_0 + \omega_L} \right) I(0) = k_B \cdot 10 \mu\text{K}$$

$$R_{\text{sc},i} = \frac{3\pi c^2}{2\hbar\omega_0^3} \left( \frac{\omega_L}{\omega_0} \right)^3 \left( \frac{\Gamma}{\omega_0 - \omega_L} + \frac{\Gamma}{\omega_0 + \omega_L} \right)^2 I(0).$$

where we have converted to CGS units in the first line. To solve for  $P, R_{\text{sc},ii}$ , we will use  $I(0) = 2P/\pi w^2$  and Mathematica. For each laser, we will have:

Yellow laser:  $P = 0.102 \mu\text{W}$ ;  $R_{\text{sc},ii} = 7700 \text{ Hz}$

Infrared laser:  $P = 9.871 \text{ mW}$ ;  $R_{\text{sc},ii} = 0.017 \text{ Hz}$

Mathematica calculations:

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In[1]:= Solve[(2*P/(Pi*w^2))*3*Pi*c^2*
Gam*(1/(omega0 - omegal) + 1/(omega0 + omegal))/(2*omega0^3) ==
kB*T, P]

Out[1]= {{P -> (
kB omega0^2 (omega0 - omegal) (omega0 + omegal) T w^2)/(6 c^2 Gam)}}

In[15]:= numbers = {kB -> 1.3806*10^(-23),
omega0 -> 2*Pi*3*10^8/(589*10^-9),
omegal -> 2*Pi*3*10^8/(985*10^-9),
T -> 10*10^(-6),
w -> 6*10^(-6),
Gam -> 2*Pi*10*10^6,
c -> 3*10^8,
hbar -> 1.0545*10^(-34)};

In[16]:= (kB omega0^2 (omega0 - omegal) (omega0 + omegal) T w^2)/(
6 c^2 Gam) /. numbers

Out[16]= 0.00987114

In[17]:= R = ((3*Pi*c^2)/(2*hbar*omega0^3))*(omegal/
omega0)^3*(Gam/(omega0 - omegal) + Gam/(omega0 + omegal))^2*2*
P/(Pi*w^2);

In[18]:= R /. numbers /. {P -> 0.009871144461932576'}

Out[18]= 0.0171102
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**2. Magic Wavelength Optical Trap.** Here we have a system with lower state  $|S\rangle$ , upper state  $|P\rangle$  with bare energy separation  $\hbar\omega_{PS}$ . The dipole moment is  $d_{PS}$ .

- (a) (i) The AC Stark shift for the state  $S$  is given by

$$\Delta E_S = -\frac{1}{4}\alpha(\omega_L)\mathcal{E}^2 = -\frac{d_{PS}^2\mathcal{E}^2}{4\hbar} \left( \frac{1}{\omega_{PS} - \omega_L} + \frac{1}{\omega_{PS} + \omega_L} \right)$$

Nothing too surprising here.

- (ii) Assuming that state  $|P\rangle$  couples almost exclusively to state  $|S\rangle$  only, then the polarizability of state  $|P\rangle$  is simply the additive reciprocal of the polarizability of state  $|S\rangle$ . This means that the energy shift of state  $|P\rangle$  is

$$\Delta E_S = -\Delta E_P.$$

- (iii) Since the energy levels shift either *away from* or *toward* each other as a function of the detuning, it is not possible for the relative energy shift to be zero with the current setup. For that to happen, we will need a third level.
- (b) Consider a third, higher energy state  $|D\rangle$  and assume that the states  $|S\rangle$  and  $|D\rangle$  do not couple directly. With this information, we know that the energy shift in  $|S\rangle$  remains the same.
- (i) Since  $|P\rangle$  couples with  $|D\rangle$ , the energy shift is modified.

$$\begin{aligned}\Delta E_P &= -\Delta E_S - \frac{d_{PD}^2 \mathcal{E}^2}{4\hbar} \left( \frac{1}{\omega_{DP} - \omega_L} + \frac{1}{\omega_{DP} + \omega_L} \right) \\ &= \frac{d_{PS}^2 \mathcal{E}^2}{4\hbar} \left( \frac{1}{\omega_{PS} - \omega_L} + \frac{1}{\omega_{PS} + \omega_L} \right) - \frac{d_{PD}^2 \mathcal{E}^2}{4\hbar} \left( \frac{1}{\omega_{DP} - \omega_L} + \frac{1}{\omega_{DP} + \omega_L} \right) \\ &= \frac{n^2 d_{DP}^2 \mathcal{E}^2}{4\hbar} \left( \frac{1}{f\omega_{DP} - \omega_L} + \frac{1}{f\omega_{DP} + \omega_L} \right) - \frac{d_{PD}^2 \mathcal{E}^2}{4\hbar} \left( \frac{1}{\omega_{DP} - \omega_L} + \frac{1}{\omega_{DP} + \omega_L} \right)\end{aligned}$$

- (ii) We would like to set the  $|S\rangle \rightarrow |P\rangle$  transition frequency so that it is independent of the trap laser power. This is achieved exactly when the shifts in  $|S\rangle$  and in  $|P\rangle$  are the same:

$$\begin{aligned}2 \frac{n^2 d_{DP}^2 \mathcal{E}^2}{4\hbar} \left( \frac{1}{f\omega_{DP} - \omega_L} + \frac{1}{f\omega_{DP} + \omega_L} \right) &= \frac{d_{PD}^2 \mathcal{E}^2}{4\hbar} \left( \frac{1}{\omega_{DP} - \omega_L} + \frac{1}{\omega_{DP} + \omega_L} \right) \\ \iff \boxed{\omega_L = \omega_{DP} \sqrt{\frac{f(f - 2n^2)}{1 - 2fn^2}}}\end{aligned}$$

### 3. Species-Dependent and Spin-Dependent AC Stark shift

- (a) Here we consider a linearly polarized dipole trap laser with electric field polarization along the  $x$  direction propagating along the  $z$  direction.
- (i) For this part we want to find the laser wavelength for which the AC Stark shift in state  $|S\rangle_{1/2}$  is zero. The AC Stark shift is given by

$$\Delta E_S = -\frac{1}{4\hbar} \sum_i |C_{1,i}|^2 \left[ \frac{1}{\omega_1 - \omega} + \frac{1}{\omega_1 + \omega} \right] - \frac{1}{4\hbar} \sum_i |C_{2,i}|^2 \left[ \frac{1}{\omega_2 - \omega} + \frac{1}{\omega_2 + \omega} \right]$$

where  $C_{1,i}, C_{2,i}$  are the dipole matrix elements. For the  $D_1$  line, we have that

$$\begin{aligned}\sum_i |C_{1,i}|^2 &= \frac{1}{2} |e \langle S_{1/2}, +1/2 | \hat{\sigma}_- \cdot \mathbf{r} | \rangle|^2 + \frac{1}{2} |e \langle | \rangle|^2 \\ \sum_i |C_{2,i}|^2 &= \frac{1}{2} |e \langle | \rangle|^2 + \frac{1}{2} |e \langle | \rangle|^2\end{aligned}$$

where we have used

$$\langle a | \hat{x} \cdot \mathbf{r} | b \rangle = \frac{1}{\sqrt{2}} \langle a | \hat{\sigma}_+ \cdot \mathbf{r} | b \rangle + \frac{1}{\sqrt{2}} \langle a | \hat{\sigma}_- \cdot \mathbf{r} | b \rangle$$

and fact that the inference terms between the different polarizations vanish for this problem.

- (ii)
- (b)
  - (i)
  - (ii)