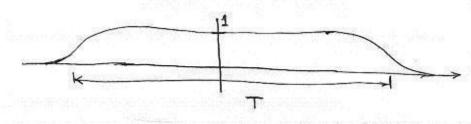
on the interactions, and then slowly switching off the interactions at some late time.

In other words H_t+) -> H_T+) ft+)
where f++) looks like



So $\int_{-\infty}^{\infty} f(t) dt = T$, $\int_{-\infty}^{\infty} (f(t))^2 dt = T$

Let S = Texp {-i \int_odt H_zurfier}

We define the S-matrix as

< final |S| initial >

where limitial > is a free particle state with

momenta \vec{k}_i^T + energies E_i^T and I-final > is a free particle state with momenta \vec{k}_i^T + energies E_i^T .

Let us concern ourselves with the nontrivial part of the 5-matrix,

<final | S-1 | initial >.

As $T \rightarrow \infty$, $V \rightarrow \infty$ we can write this amplitude as

 $\angle \text{final} \mid S-1 \mid \text{initial} >$ $= i \cdot \mathfrak{M} \cdot (2\pi)^4 S(E_{tot}^F - E_{tot}^I) S^{(3)}(\hat{k}_{tot}^F - \hat{k}_{tot}^I)$ (i is part of standard) $\text{definition for } \mathfrak{M}$

where M is a function of the momenta

For finite T and finite V we instead have

$$$$

$$= i M \int_{-\infty}^{\infty} f(t) e^{i (E_{tot}^{F} - E_{tot}^{T}) t} S_{\vec{k}_{tot}, \vec{k}_{tot}} \cdot V$$

$$So | | | | | | | | | | | | | | | |^{2}$$

=
$$|\mathcal{M}|^2 \cdot S_{\hat{k}_{ht}}, \hat{k}_{ht}^{I}, V^2 \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(t) f(t') e^{i(E_{ht}^F - E_{ht}^I)(t'-t)} dt' dt$$

as $T \rightarrow \infty$ this clearly gives some constact times $\delta(E_{tot}^F - E_{tot}^T)$.

What is this constant? To get let us integrate with respect to E_{tot}^F .

$$=2\pi\int_{-\infty}^{\infty}f_{tt}^{2}) dt = 2\pi T$$

So the constant is 2TT.T and

(<first) S-1) initial>|2 = |911|2 (211) S(EFT-EIN) SEFT, KINT V2T

$$|\langle final | S-1 | initial \rangle|^2$$

$$= |\mathcal{M}|^2 (2\pi) \delta(E_{bt}^F - E_{bt}^{\perp}) \delta_{\vec{K}_{bt}^F, \vec{K}_{bt}^{\perp}} V^2 T$$

We have been using relativistic normalizations for our states

$$\langle initial | initial \rangle = Tr(2E_i^T \cdot V)$$
 becomes $\langle \pi \rangle^3 \delta^{(3)}(0)$ as $V \to \infty$ $\langle final | final \rangle = Tr(2E_i^T \cdot V)$

To get the transition probability per unit time

$$\frac{\text{probability}}{\text{time}} = \frac{1}{T} \frac{\left| \left\langle \text{final} \right| \text{S-1} \right| \text{initial} > \right|^{2}}{\left\langle \text{final} \right| \text{final} \left| \left\langle \text{shitial} \right| \text{initial} > \right|}$$

$$= \frac{\left| \left| M \right|^{2} (2\pi) \, \delta \left(E_{\text{tot}}^{F} - E_{\text{tot}}^{F} \right) \, \delta_{\vec{K}_{\text{fot}}, \vec{K}_{\text{tot}}}^{F} \, V^{2}}{TT \left(2E_{i}^{F} \cdot V \right) \, TT \left(2E_{i}^{F} \cdot V \right)}$$

As $V \rightarrow \infty$, $\S_{\vec{K}_{bt}}^F, K_{bt}^{I} V \rightarrow (2\pi)^3 \S^{(5)}(\vec{K}_{tot}^F - \vec{K}_{tot}^{I})$

If we sum over final states in some window then we have

As
$$V \rightarrow \infty$$
, $\frac{d^{3}\vec{K}_{1}^{F}}{(2\pi)^{3}2E_{1}^{F}} = \frac{d^{3}\vec{K}_{1}^{F}}{(2\pi)^{3}2E_{1}^{F}} = \frac{d^{3}\vec{K}_{1}^{F}}{(2\pi)^{3}2E_{1}^{F}} = \# \text{ of final particles}$

Let's consider a single particle decay $(n_{\rm I}=1)$. The total decay rate is $\Gamma'=\int d\Gamma$ where

$$d(\Gamma^{7} = \frac{1}{2 E^{\pm}} \left(\frac{\prod_{i=1}^{n_{F}} d_{i}^{3} \vec{K}_{i}^{F}}{(2\pi)^{3} 2 E_{i}^{F}} \right) |\mathcal{M}|^{2} (2\pi)^{4} \delta^{(4)} (K_{bt}^{F} - K_{tot}^{\pm})$$

For a two-particle initial state, the cross-section is given by

6 = Probability time of lax density

The flux density is the relative velocity between the beam and target times density of incoming beam in the laboratory frame. We have normalized our probability for one incoming beam particle, so density = $\frac{1}{V}$, and

flux =
$$\frac{|\vec{V}_A - \vec{V}_B|}{V}$$
 \vec{V}_A, \vec{V}_B velocities of particles in the laboratory frame

So
$$d_6 = \left(\frac{\sqrt{r_E}}{\sqrt{2\pi}} \frac{d^3 \vec{k}_{\perp} \vec{F}}{(2\pi)^3 2 \vec{E}_{\perp} \vec{F}}\right) (2\pi)^4 \delta^{(4)} (K_{tot}^F - K_{tot}^I) \frac{100(1)^2}{2 \vec{E}_A 2 \vec{E}_B |\vec{V}_A - \vec{V}_B|}$$

$$\equiv d \pi_{n_F}$$

Let's consider a special case... two final particles $(M_F=2)$ in the center of mass frame.

$$\int d\Pi_{z} = \int \frac{d\rho_{1} \rho_{1}^{2} d\Omega}{(2\pi)^{3} 2E_{1} 2E_{2}} (2\pi) \int \frac{d\rho_{1} \rho_{2}^{2} d\Omega}{(2\pi)^{3} 2E_{1} 2E_{2}} (2\pi) \int \frac{d\rho_{2} \rho_{1}^{2} d\Omega}{(2\pi)^{3} 2E_{1}^{2} 2E_{2}^{2}} (2\pi) \int \frac{d\rho_{2} \rho_{1}^{2} d\Omega}{(2\pi)^{3} 2E_{1}^{2} 2E_{2}^{2}} (2\pi) \int \frac{d\rho_{2} \rho_{1}^{2} d\Omega}{(2\pi)^{3} 2E_$$

There the final particle energies are $E_1 = \sqrt{\vec{p}_1^2 + m_1^2}$ $E_2 = \sqrt{\vec{p}_2^2 + m_2^2}$

[recall S(fix)] = \frac{S(x-x_0)}{(x_0 is simple)} [f(x_0)]

$$\frac{dE_1}{d\rho_1} = \frac{d\int_{\ell_1^2 + m_1^2}^2}{d\ell_1} = \frac{1}{2} \frac{2\rho_1}{\sqrt{\rho_1^2 + m_1^2}} = \frac{\rho_1}{E_1}$$

$$\frac{dE_2}{d\rho_1} = \frac{\rho_1}{E_2}$$

So
$$\int d\Pi_2 = \int \frac{d\Omega}{16\pi^2} \frac{P^2}{E_1 E_2 (\frac{P_1}{E_1} + \frac{P_2}{E_2})} \Big|_{P_1 \text{ thosen so}}$$

$$= \int \frac{d\Omega}{16\pi^2} \frac{P_1}{E_1 + E_2} = \int \frac{d\Omega}{16\pi^2} \frac{P_1}{E_{cm}}$$

So for two particles -> two particles,

$$\left(\frac{d6}{d\Omega}\right)_{cm} = \frac{\left|\vec{p}^{finel}\right| \left|\vec{M}\right|^2}{2E_A 2E_B \left|\vec{V}_A \vec{V}_B\right| \left|6\Pi^2\right|} \left(E_{cm} = E_A + E_B\right)$$

It is conventional to define the T-matrix... S=1+iT

() zim:

= (fine Pi, ... Pi | Texp (-i] dt HI(t) | PA, PB > free)

A = connected diagrams only + "amputated" diagrams only

"Amputated" means that the diagram can't be broken into disconnected pieces by cutting one internal line. This is also called one-particle irreducible (1PI).

Not Amoutated

Can cut here

can cut here

Ampetated

 $\overline{}$

X

The claim won't be proven until later in the course, but the idea is similar to before...

and we would like something similar ...

But this is tricky. More on this in a couple of chapters ahead. For now we just take the claim as time pending later verification, though we will find some subtleties and fix them at that time.

Note that

relativistic normalization

$$\phi_{I}^{\dagger}(x) |\vec{p}\rangle_{ful} = \int_{(2\pi)^{3}}^{d^{3}\vec{k}} a_{\vec{k}} e^{-i\vec{k}\cdot x} \int_{2\vec{k}_{\vec{p}}}^{2\vec{k}_{\vec{p}}} a_{\vec{p}}^{\dagger} |o\rangle$$

$$= e^{-i\vec{p}\cdot x} |o\rangle$$

We can think of taking the commutator of $\phi_{T}^{+}(x)$ with the $a_{\overline{p}}^{+}$ from $|\overline{p}|_{free}$. This suggests the notation

$$\phi_{\mathbf{T}}^{\mathbf{r}}(\mathbf{x}) | \vec{p} \rangle_{\text{free}} = e^{-i \vec{p} \cdot \mathbf{x}}$$

We now drop the "free" subscript. Similarly

and we define

$$\langle \vec{p} | \vec{\phi}_{\Gamma}(x) = e^{+i\vec{p}\cdot x}$$

Feynman Rules in position space with external lines

For each propagator & J Dp (x-y)

For each internal vertex XZ (-ix) Sd42

Divide by symmetry factor S

Feynman Rules in momentum space with external lines

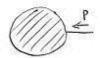
For each propagator =

For each internal vertex



-i) and momentum conservation

For each external line



no extra factor (i.e, 1)

Integrale over all unconstrained momenta and divide by symmetry factor S.

Example

<pi, Pz | iT | pa, PB > at lowert order

Feynman amplitude $iM = -i\lambda$

So
$$\left(\frac{d6}{d\Omega}\right)_{CM} = \frac{|\vec{p}_{final}| |m|^2}{2E_A 2E_B |\vec{v}_A - \vec{v}_B| |\vec{b}\vec{q}^2 E_{GM}|}$$

Let $p = |\vec{p}|^2 = |\vec{p}| = |\vec{p}|$ all some since masses are all the same

$$E_{cm} = 2E_A = 2E_B = 2\sqrt{p^2 + m^2}$$

$$|\vec{\nabla}_A - \vec{\nabla}_B| = 2|\vec{\nabla}_A| = \frac{2|\vec{P}_A|}{E_A} = \frac{2p}{E_A}$$