

# Lecture 5 - Quantum States of Light

## Review from last lecture

- Analogy with harmonic oscillator:

One e.m. field mode	Correspondence	Harmonic Oscillator
$\dot{\mathcal{A}}_i = -\mathcal{E}_i$	$\mathcal{A}_i \hat{=} x$	$\dot{x} = \frac{p}{m}$
$\dot{\mathcal{E}}_i = \omega_i^2 \mathcal{A}_i$	$\mathcal{E}_i \hat{=} -\frac{p}{m}$	$\frac{\dot{p}}{m} = -\omega^2 x$
$H_i = \frac{\epsilon_0}{2} \frac{1}{V} ( \mathcal{E}_i ^2 + \omega_i^2  \mathcal{A}_i ^2)$	$\epsilon_0 \frac{1}{V} \hat{=} m$	$H = \frac{m}{2} \left( \left( \frac{p}{m} \right)^2 + \omega^2 x^2 \right)$
$\alpha_i = \mathcal{N}_i \left( \mathcal{A}_i - \frac{i}{\omega_i} \mathcal{E}_i \right)$		$\alpha = \mathcal{N} \left( x + i \frac{p}{m\omega} \right)$
$\frac{d\alpha_i}{dt} = -i\omega_i \alpha_i$		$\frac{d\alpha}{dt} = -i\omega \alpha$

- Quantization and Commutation relations:

One e.m. field mode	Harmonic Oscillator
$\mathcal{A}_i \rightarrow \hat{\mathcal{A}}_i$	$x \rightarrow \hat{x}$
$\mathcal{E}_i \rightarrow \hat{\mathcal{E}}_i$	$p \rightarrow \hat{p}$
$[\hat{\mathcal{A}}_i, \hat{\mathcal{E}}_i] = -\frac{V}{\epsilon_0} i\hbar$	$[\hat{x}, \hat{p}] = i\hbar$
$\hat{a}_i$ annihilation operator associated to $\alpha_i$	$\hat{a}$ annihilation operator associated to $\alpha$
$[\hat{a}_i, \hat{a}_i^\dagger] = 1$ for	$[\hat{a}, \hat{a}^\dagger] = 1$ for
$\mathcal{N} = \sqrt{\frac{\epsilon_0 \omega_i}{2\hbar}} \frac{1}{V}$	$\mathcal{N} = \sqrt{\frac{m\omega}{2\hbar}}$

- Fields:

$$\vec{E}_\perp(\vec{r}) = i \sum_j \mathcal{E}_j \vec{\epsilon}_j \left( a_j e^{i\vec{k}_j \cdot \vec{r}} - a_j^\dagger e^{-i\vec{k}_j \cdot \vec{r}} \right)$$

$$\vec{B}(\vec{r}) = i \sum_j \mathcal{E}_j \frac{\vec{k}_j \times \vec{\epsilon}_j}{\omega_j} \left( a_j e^{i\vec{k}_j \cdot \vec{r}} - a_j^\dagger e^{-i\vec{k}_j \cdot \vec{r}} \right)$$

$$\vec{A}_\perp(\vec{r}) = \sum_j \frac{\mathcal{E}_j}{\omega_j} \vec{\epsilon}_j \left( a_j e^{i\vec{k}_j \cdot \vec{r}} + a_j^\dagger e^{-i\vec{k}_j \cdot \vec{r}} \right)$$

where  $\mathcal{E}_j = \sqrt{\frac{\hbar \omega_j}{2\epsilon_0 V}}$  (careful: This pure number should not to be confused with the complete Fourier transform of the classical field mode  $\vec{\mathcal{E}}_\perp$ , which is a function of  $\vec{k}$ .)

- Uncertainty relation:

Just like  $\Delta x \Delta p \geq \frac{\hbar}{2}$  we have, using our correspondence  $\mathcal{A}_j \equiv x$  and  $\mathcal{E}_j \equiv -p/m$  as well as  $\epsilon_0/V \equiv m$ :

$$\Delta \mathcal{A}_{\perp,j} \Delta \mathcal{E}_{\perp,j} \geq \frac{\hbar}{2} \frac{V}{\epsilon_0}$$

Quick reminder on units:  $\vec{\mathcal{E}}_\perp(\vec{k}_j) = \int d^3r \vec{E}_\perp(\vec{r}) e^{-i\vec{k}_j \cdot \vec{r}}$  has units of  $[EV]$ . Also,  $\vec{A}$  has units of an electric field, so  $[\mathcal{A}] = [EVt]$ . So  $[\epsilon_0 \mathcal{A} \mathcal{E}] = [\epsilon_0 E^2 V^2 t]$ .  $\epsilon_0 E^2$  is an energy density, so  $[\epsilon_0 \mathcal{A} \mathcal{E}] = [\text{Energy} \cdot \text{time} \cdot \text{Volume}] = [\hbar V]$ .

- Vacuum properties:

$$\text{We have } \langle 0 | \vec{E}_\perp(\vec{r}) | 0 \rangle = \langle 0 | \vec{B}(\vec{r}) | 0 \rangle = \langle 0 | \vec{A}(\vec{r}) | 0 \rangle = 0$$

but since  $\langle 0 | a_j a_k^\dagger | 0 \rangle = \delta_{jk}$  we find the

- Vacuum fluctuations:

$$\left( \Delta \vec{E}_\perp \right)^2 = \langle 0 | (\vec{E}_\perp(\vec{r}))^2 | 0 \rangle = \sum_j \mathcal{E}_j^2 = \sum_j \frac{\hbar \omega_j}{2\epsilon_0 V}$$

$$\left( \Delta \vec{B}_\perp \right)^2 = \langle 0 | (\vec{B}(\vec{r}))^2 | 0 \rangle = \sum_j \frac{1}{c^2} \mathcal{E}_j^2 = \sum_j \frac{\hbar \omega_j}{2\epsilon_0 c^2 V}$$

$$\left(\Delta \vec{A}_\perp\right)^2 = \langle 0 | (\vec{A}_\perp(\vec{r}))^2 | 0 \rangle = \sum_j \frac{1}{\omega_j^2} \mathcal{E}_j^2 = \sum_j \frac{\hbar}{2\epsilon_0 V \omega_j}$$

- Field values for number states:

$$|n_1 = 0, \dots, n_{j-1} = 0, n_j, n_{j+1} = 0, \dots\rangle \equiv |n_j\rangle$$

$$\langle n_j | \vec{E}_\perp(\vec{r}) | n_j \rangle = \langle n_j | \vec{B}(\vec{r}) | n_j \rangle = 0$$

$$c^2 \Delta B^2 = \Delta E_\perp^2 = \langle n_j | \vec{E}_\perp(\vec{r})^2 | n_j \rangle = (2n_j + 1) \mathcal{E}_j^2$$

So this is  $(2n_j + 1)$  times the  $\Delta E_\perp^2$  in vacuum.

## Coherent States (Quasi-classical states, Glauber states)

I want a state in which the E-field and B-field are as close to a classical state as possible.

$$\vec{E}_{\text{classical}} = \sum_j \mathcal{E}_j \vec{\epsilon} \left( \alpha_j e^{i\vec{k}_j \cdot \vec{r}} + \text{c.c.} \right)$$

$$\vec{E}_{\text{quantum}} = \sum_j \mathcal{E}_j \vec{\epsilon} \left( a_j e^{i\vec{k}_j \cdot \vec{r}} + \text{c.c.} \right)$$

This leads us to define the coherent state  $|\alpha\rangle$  as an eigenstate of the annihilation operator with eigenvalue  $\alpha_j$ :

$$\boxed{a_j |\alpha_j\rangle = \alpha_j |\alpha_j\rangle}$$

(for one mode, say mode  $j$  - let's drop the index  $j$  in the following to unclutter).

Expand over number states:

$$|\alpha\rangle = \sum_{n=0}^{\infty} c_n |n\rangle$$

$$\begin{aligned}
|\alpha\rangle &= \sum_{n=1}^{\infty} c_n \sqrt{n} |n-1\rangle = \sum_{n=0}^{\infty} c_{n+1} \sqrt{n+1} |n\rangle \\
&\stackrel{!}{=} \alpha \sum_{n=0}^{\infty} c_n |n\rangle
\end{aligned}$$

$$\Rightarrow c_{n+1} \sqrt{n+1} = \alpha c_n \quad \text{or} \quad c_n = \frac{\alpha}{\sqrt{n}} c_{n-1}$$

$$\Rightarrow c_n = \frac{\alpha^n}{\sqrt{n!}} c_0$$

$$\Rightarrow |\alpha\rangle = c_0 \sum_{n=0}^{\infty} \frac{\alpha^n}{\sqrt{n!}} |n\rangle$$

Normalization:

$$\langle \alpha | \alpha \rangle = 1 = |c_0|^2 \sum_{\substack{m=0 \\ n=0}}^{\infty} \frac{\alpha^n \alpha^{*m}}{\sqrt{n!m!}} \langle m | n \rangle = |c_0|^2 \sum_{n=0}^{\infty} \frac{|\alpha|^{2n}}{n!} = |c_0|^2 e^{|\alpha|^2}$$

$$\Rightarrow c_0 = e^{-\frac{|\alpha|^2}{2}}$$

up to a phase factor. So we finally have found

$$\boxed{|\alpha\rangle = e^{-\frac{|\alpha|^2}{2}} \sum_{n=0}^{\infty} \frac{\alpha^n}{\sqrt{n!}} |n\rangle}$$

We may immediately note that  $|\alpha\rangle$  is *not* an eigenstate of the creation operator:

$$a^\dagger |\alpha\rangle = e^{-\frac{|\alpha|^2}{2}} \sum_{n=0}^{\infty} \frac{\alpha^n}{\sqrt{n!}} \sqrt{n+1} |n+1\rangle = e^{-\frac{|\alpha|^2}{2}} \sum_{n=1}^{\infty} \frac{n \alpha^{n-1}}{\sqrt{n!}} |n\rangle$$

which is clearly not proportional to  $|\alpha\rangle$ . In fact, one sometimes works with an alternative normalization where  $||\alpha\rangle = e^{\frac{|\alpha|^2}{2}} |\alpha\rangle$ , which allows one to right simply

$$a^\dagger |\alpha\rangle = \frac{\partial}{\partial \alpha} |\alpha\rangle$$

$$(\text{simply from } \frac{\partial}{\partial \alpha} \alpha^n = n\alpha^{n-1})$$

- Time evolution:

Let's start at  $t = 0$  with  $|\alpha\rangle$ . At time  $t$  we find:

$$\begin{aligned} |\psi(t)\rangle &= e^{-iHt/\hbar} |\alpha\rangle \\ &= e^{-\frac{|\alpha|^2}{2}} \sum_{n=0}^{\infty} \frac{\alpha^n}{\sqrt{n!}} e^{-(n+\frac{1}{2})\omega t} |n\rangle \\ &= e^{-i\frac{\omega t}{2}} e^{-\frac{|\alpha|^2}{2}} \sum_{n=0}^{\infty} \frac{(\alpha e^{-i\omega t})^n}{\sqrt{n!}} |n\rangle \\ &= e^{-i\frac{\omega t}{2}} |\alpha e^{-i\omega t}\rangle \end{aligned}$$

Probability of finding the value  $(n + \frac{1}{2})\hbar\omega$  for the energy (or the value  $n$  for the photon number) is time independent:

$$P(n) = |c_n|^2 = e^{-|\alpha|^2} \frac{|\alpha|^{2n}}{n!}$$

$$\text{Mean photon number: } \bar{n} = \langle \hat{N} \rangle = \langle \alpha | a^\dagger a | \alpha \rangle = \alpha^* \alpha = |\alpha|^2$$

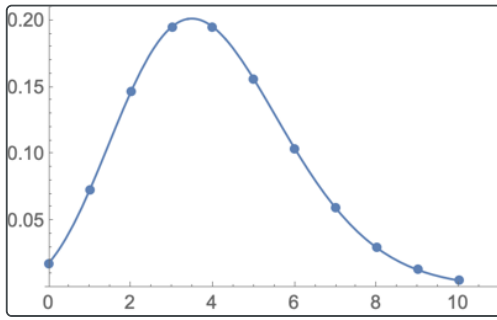
Equivalently we can calculate:

$$\langle \hat{N} \rangle = \sum_n n P(n) = \sum_n n e^{-|\alpha|^2} \frac{|\alpha|^{2n}}{n!} = |\alpha|^2 \sum_n e^{-|\alpha|^2} \frac{|\alpha|^{2n}}{n!} = |\alpha|^2$$

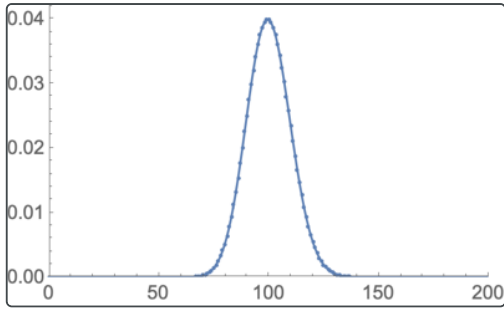
$$\text{So } \bar{n} = |\alpha|^2 \text{ and}$$

$$\boxed{P(n) = e^{-\bar{n}} \frac{\bar{n}^n}{n!}} \quad \text{Poisson Distribution}$$

$P(n)$  for  $\bar{n} = 4$ :



$P(n)$  for  $\bar{n} = 100$



- Standard deviation:

$$\begin{aligned}
 \Delta n^2 &= (\Delta N)^2 = \langle N^2 \rangle - \langle N \rangle^2 = \langle \alpha | a^\dagger a a^\dagger a | \alpha \rangle - |\alpha|^4 \\
 &= \langle \alpha | a^\dagger (a^\dagger a + 1) a | \alpha \rangle - |\alpha|^4 \\
 &= |\alpha|^4 + \langle N \rangle - |\alpha|^4 \\
 &= \langle N \rangle = \bar{n}
 \end{aligned}$$

So  $\boxed{\Delta n^2 = \bar{n}}$  as it should be for the Poisson distribution.

- Electric field: Use  $|\psi(t)\rangle = e^{-i\frac{\omega t}{2}} |\alpha e^{-i\omega t}\rangle$  and find (again, we omit the index  $j$ ):

$$\langle \psi(t) | \vec{E}_\perp(\vec{r}) | \psi(t) \rangle = i\mathcal{E}\vec{\epsilon} \left\{ \alpha e^{i(\vec{k}\cdot\vec{r} - \omega t)} - \alpha^* e^{-i(\vec{k}\cdot\vec{r} - \omega t)} \right\}$$

(which is just the classical field)

$$\begin{aligned}
 \langle \psi(t) | \vec{E}_\perp(\vec{r})^2 | \psi(t) \rangle &= \langle \psi(t) | (\dots a \dots - a^\dagger \dots)(\dots a \dots - \dots a^\dagger \dots) | \psi(t) \rangle \\
 &= \left( \langle \psi(t) | \vec{E}_\perp(\vec{r}) | \psi(t) \rangle \right)^2 + \mathcal{E}^2
 \end{aligned}$$

Therefore

$$\Rightarrow \boxed{\Delta \vec{E}_\perp^2 = \mathcal{E}^2}$$

which is the same fluctuation as in the vacuum state!

- Are coherent states orthonormal? Let's find out:

$$\begin{aligned}
 |\langle \beta | \alpha \rangle|^2 &= \left| e^{-\frac{|\alpha|^2}{2}} e^{-\frac{|\beta|^2}{2}} \sum_{m,n} \frac{\alpha^n \beta^{*m}}{\sqrt{n!} \sqrt{m!}} \langle m | n \rangle \right|^2 \\
 &= e^{-|\alpha|^2} e^{-|\beta|^2} \left| e^{\alpha \beta^*} \right| \\
 &= e^{-|\alpha - \beta|^2}
 \end{aligned}$$

They are not orthogonal!

But they do obey the closure relation:

$$\frac{1}{\pi} \int d(\text{Re } \alpha) d(\text{Im } \alpha) |\alpha\rangle \langle \alpha| = 1$$

Equivalently we can write

$$\frac{1}{\pi} \int d^2 \alpha |\alpha\rangle \langle \alpha| = 1$$

Proof:

$$\frac{1}{\pi} \int d^2 \alpha |\alpha\rangle \langle \alpha| = \frac{1}{\pi} \sum_{m,n} \frac{|m\rangle \langle n|}{\sqrt{m!n!}} \int d^2 \alpha e^{-|\alpha|^2} \alpha^m (\alpha^*)^n$$

$$\begin{aligned}
 \int d^2 \alpha e^{-|\alpha|^2} \alpha^m (\alpha^*)^n &= \int_0^\infty dr e^{-r^2} r^{m+n+1} \int_0^{2\pi} e^{i(m-n)\theta} d\theta \\
 &= 2\pi \int_0^\infty dr e^{-r^2} r^{m+n+1} \delta_{mn} \\
 &= 2\pi \delta_{mn} \int_0^\infty dr e^{-r^2} r^{2n+1} \\
 &= 2\pi \delta_{mn} \frac{1}{2} \int_0^\infty du e^{-u} u^n \\
 &= \pi n! \delta_{mn}
 \end{aligned}$$

So

$$\frac{1}{\pi} \int d^2\alpha |\alpha\rangle \langle\alpha| = \frac{1}{\pi} \sum_{m,n} \frac{|m\rangle \langle n|}{\sqrt{m!n!}} \pi n! \delta_{mn} = \sum_n |n\rangle \langle n| = 1$$

Application to number state:

$$|n\rangle = \frac{1}{\pi} \int d^2\alpha |\alpha\rangle \langle\alpha|n\rangle = \frac{1}{\pi} \int d^2\alpha e^{-\frac{|\alpha|^2}{2}} \frac{(\alpha^*)^n}{\sqrt{n!}} |\alpha\rangle$$

The weight of the state  $|\alpha\rangle$  can be seen to be peaked at  $|\alpha| = \sqrt{n}$ , which is independent of the argument of  $\alpha$ .

Application to a coherent state:

$$|\beta\rangle = \frac{1}{\pi} \int d^2\alpha |\alpha\rangle \langle\alpha|\beta\rangle = \frac{1}{\pi} \int d^2\alpha e^{-\frac{1}{2}(|\alpha|^2 + |\beta|^2) + \alpha^* \beta} |\alpha\rangle$$

So a coherent state can be written as a superposition of coherent states! The basis is said to be overcomplete. Still, we have

$$\langle\beta|\beta\rangle = \frac{1}{\pi} \int d^2\alpha e^{-|\alpha-\beta|^2} = 1$$

as it should be.

$$\text{Decomposition: } |\psi\rangle = \int d^2\alpha c_\alpha |\alpha\rangle \text{ with } c_\alpha = \frac{1}{\pi} \langle\alpha|\psi\rangle$$

This decomposition is not unique as

$$\langle\alpha|\alpha'\rangle \neq 0, \text{ but } |\langle\alpha|\alpha'\rangle|^2 = e^{-|\alpha-\alpha'|^2} \rightarrow 0$$

if  $\alpha$  and  $\alpha'$  are widely separated.

- We say that the basis of coherent states  $|\alpha\rangle$  is overcomplete.

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