Consider spinning top: R-center fran  $\mathcal{T} = \mathcal{R} \times \mathcal{F}$   $= -Mg \, \mathcal{R} \times \hat{z}$ total F=0! = MgR L× 2  $\vec{\tau} = \vec{\Gamma} = \vec{\Omega} \times \vec{L}$ with I = Mg R 2. Does precession depend on tipping angle? No! Look from top: dl = Lshrdp de = Long j  $= \Omega L \text{ sust}$   $= M_g R \text{ sust}$   $= M_g R \text{ independent of } v^2.$ Energy: E = Mg R cord MgR / . 2

Magnetic uranents.

force F = - TE = O for wriform torque: T = \$\vec{\pi} \times B\$ non-zero!

For classical charge distribution: 1/1/2

$$\int_{\mathbf{I}} \mathbf{I} \cdot \mathbf{A} \hat{\mathbf{z}}$$

$$T = \frac{e}{T} = \frac{eV}{2\pi r}$$

$$A = \pi r^{2}$$

$$A = \pi r^{2}$$

$$= \frac{ev}{2\pi r} \pi r^{2}$$

$$= 1 ev r$$

 $= \frac{1}{2} \frac{eh}{m} L$   $= \frac{1}{2} \frac{eh}{m} L$ 

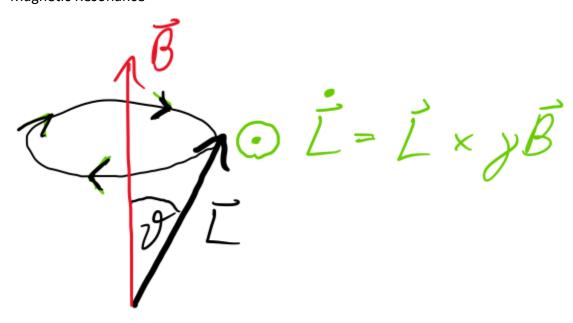
Lorents force produces dorgen  $\vec{t} = \vec{L} = \vec{n} \times \vec{B}$ 

We saw above that in ~ I for a classical charge distribution. Let's assume that's true for quantum objects as well:

Write  $\vec{p} = \vec{y} \perp$ Where  $\vec{y} - g \vec{y} romagnetic ratio$ 

So den \( \bar{\bar{L}} = \bar{L} \times \bar{B} \)

=> Precession of  $\vec{L}$  (and thus  $\vec{h}$ ) or bonf Magnetic Field with augular frequency  $\Omega_L = -\gamma B$ , He Larmor Frequency.



Tipping angle it is constant. For electron spin angular momentum, we have Ye = 211. 2.8 M/2 For protons, we have  $y_p = 2\pi \cdot 4.2 \frac{httr}{6}$ . For a classical charge distribution with mass me, and orbital angular momentum with L=1, we have  $\gamma = \frac{eh}{2m} = 1.4 \frac{4k}{6} = 1$ Why the factor of 2 for e sph? If turns out that yes, the magnetic moment of one electron is + pg in state 192 (spin up)
and - pg in state 182. So then the
energy difference between 192 and 162 in a magnetic field is 2 pg = 2.8 MHz. B. The precession frequency is the frequency at

superponna

US CHANCE , 50 :

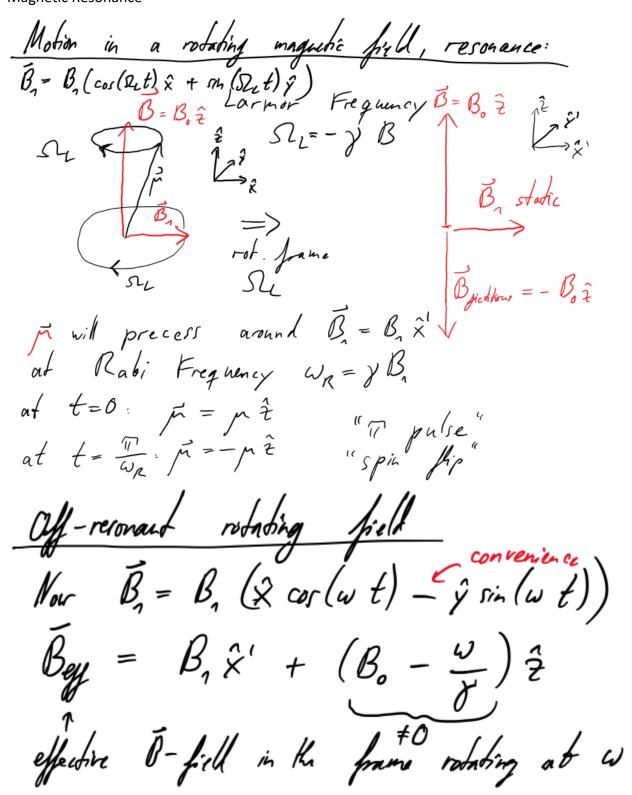
"classical" charge dishibution

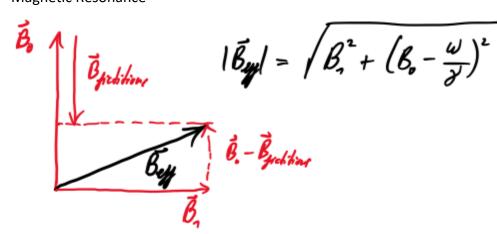
For the classical charge distribution (i.e. standard vectors, i.e. L=1), there are three star with  $m_L=\pm 1$  and 0. The energy of  $|\pm\rangle$  1.4 M/s. B, but state low stays at zero.

At one bies de bilt le magnetic moment away from my the 1-12 state (printing down), the superposition 21-1>+Blox oscillates at 1.4 Afte B, so that's the precession frequency Robating coordinate system: y a vector robuses with \$\overline{\Omega}, then  $A = \Omega \times A$ To relate an arbitrary rate of change in the inertial frame to the rate of charge in a robating frame robating at 52,  $(\widetilde{A})_{in} = (\widetilde{A})_{ref} + \widetilde{\Omega} \times \widetilde{A}_{in}$ 

(y. A is combact in the relating frame then  $\tilde{A}_{in} = \tilde{\Omega} \times \tilde{A}_{in}$ . If  $\tilde{\Omega} = 0$  then  $\tilde{A}_{in} = \tilde{A}_{in}$ . General relationship must be linear.) So ve have the operator equation:  $\left(\frac{d}{dt}\right)_{lat} = \left(\frac{d}{dt}\right)_{la} - \vec{\Omega} \times$ => Inot = In - Qx I  $= \mathcal{L} \times (\gamma \vec{B} + \vec{\Omega})$ = y L x (B + 2)

Choose  $\Omega - \overline{\Omega}_{L} = -\gamma \overline{B}$ The :  $\Rightarrow L$  is constant in the roboting frame.

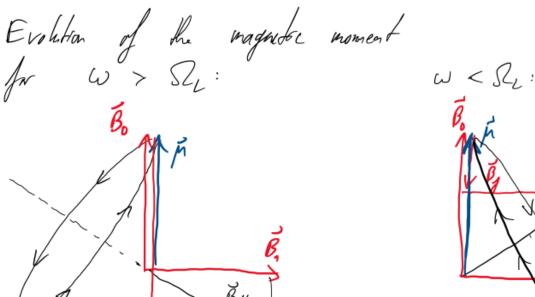


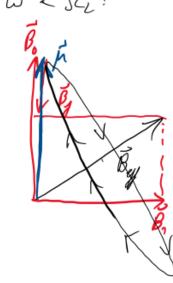


Precession Frequency
$$\Omega_{R} = \gamma B_{gy} = \sqrt{\omega_{R}^{2} + (\Omega_{L} - \omega)^{2}}$$

$$\Omega_{R} = \sqrt{\omega_{R}^{2} + \sigma^{2}}$$

$$\Omega_{R} = \sqrt{\omega_{R}^{2} + \sigma^{2}}$$

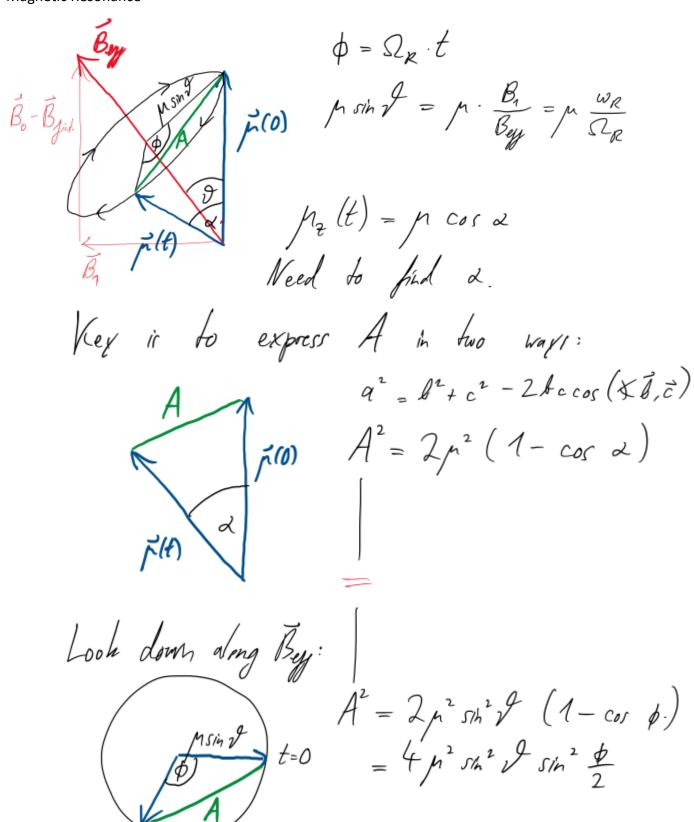




· In both case, precession frequency  $\Omega_R > \omega_R$ · In both case, predoes never fully invert.

Time evalution: If t=0,  $\vec{F}=\vec{h}\cdot\hat{z}$ 

Bpicktion



$$=) \cos z = 1 - 2 \sin^2 \theta \sin^2 \theta$$

$$=) p_{2}(t) = p\left(1 - 2\frac{\omega_{R}^{2}}{\Omega_{R}^{2}}\sin^{2}\left(\frac{\Omega_{R}t}{2}\right)\right)$$

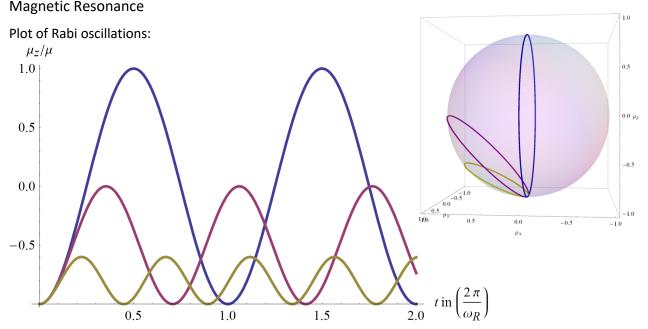
· exact result

· same as for expectation value in QM.

Full inversion only for  $W = \Omega_L$  (then  $w_R = \Omega_R$ !)

Off-resonant case: precession is faster

inversion incomplete



In all these examples, we start with the magnetic moment initially pointing down, in the -z direction (then the curves also look identical to the later probability to excite an atom from the ground-state in a two-level system, with the y-axis now going from 0 to 1).

The blue curve is for a resonant drive, with  $\delta=0$ . We see that after half a Rabi period,  $\frac{T}{2}=\frac{\pi}{\omega_R}$ , the magnetic moment got inverted (so-called  $\pi$ -pulse) and points up. After one Rabi period,  $T=\frac{2\pi}{\omega_R}$ , the magnetic moment is back to pointing in the original direction, down.

The purple curve shows the situation when the detuning is  $\delta=1~\omega_R$ . We see that the Rabi oscillations are now **incomplete** and are **faster**. We find in this special case the maximum value of the z-component of the magnetic moment to be

$$\frac{\mu_Z}{\mu} = -1 + 2\frac{\omega_R^2}{\omega_R^2 + \delta^2} = -1 + 2\frac{1}{1+1} = 0.$$

The generalized Rabi frequency is here  $\Omega_R=\sqrt{\omega_R^2+\delta^2}=\sqrt{2}~\omega_R$ , so the magnetic moment will be back pointing perfectly down after only  $T'=\frac{2\pi}{\Omega_R}=\frac{1}{\sqrt{2}}\frac{2\pi}{\omega_R}\approx 0.71~T$ 

The yellow curve is for an even larger detuning of  $\delta=2~\omega_R$ . The oscillations are even faster, and the contrast even smaller. The maximum  $\mu_Z$  is now

$$\frac{\mu_Z}{\mu} = -1 + 2\frac{\omega_R^2}{\omega_R^2 + \delta^2} = -1 + 2\frac{1}{1+4} = -\frac{3}{5} = -0.6.$$

The generalized Rabi frequency is here  $\Omega_R=\sqrt{\omega_R^2+\delta^2}=\sqrt{5}~\omega_R$ , so the magnetic moment will be back pointing down after only  $T'=\frac{2\pi}{\Omega_R}=\frac{1}{\sqrt{5}}\frac{2\pi}{\omega_R}\approx 0.45~T$ 

**Note:** The **initial** behavior of all the curves is identical!  $\mu_z$  grows initially quadratically with time, with a curvature that's independent of detuning! Indeed, we have

$$\frac{\mu_z}{\mu} = -1 + 2\frac{\omega_R^2}{\Omega_R^2} \sin^2\left(\frac{\Omega_R t}{2}\right) \approx -1 + \frac{1}{2}\omega_R^2 t^2$$

This is valid for times such that  $\Omega_R t \ll 1$ , which for large detuning means  $t \ll \frac{1}{\delta}$ . This we can also see directly by considering the Fourier uncertainty relation between frequency and time: At times  $t \ll \frac{1}{\delta}$ , the system did not even "have enough time to realize" that it was detuned from resonance! 3D View on Bloch sphere:

Rapid adiabati passage Technique for invending spin by (slowly) sweeping a cross the resonance "slow" compared to the Larmor frequency y By "rapid" compared to relaxation processes Physical produce: A magnetic moment in is initially aligned with a static field  $B_0 = B_0 \hat{\tau}$ , and there is a weak magnetic field B that rotates at frequency w in the x-y plane. start at a detuning  $\sqrt{\ll -\omega_R} = y B_1$ In the rotating frame (rotating at  $\omega$  about the z-axis), the effective magnetic field points almost in the z-direction (ar  $|B_n| \ll |B_n - \frac{\omega}{y}| = \left|\frac{\sigma}{y}\right|$ The magnetic moment in lightly precesses about le effective magnetic field, so is almost arigned in ? Now let us slawly (we will see what that means) sweep the frequency of the rotating field B. from  $S \ll - \omega_R$  through S = 0 and up to  $S \gg \omega_R$ . The situation in the rotating frame (rotating at the momentary frequency  $\omega(t)$  of B) locks like this:

B.- 3) î

d <<- W/R

E. F.

J= 0

ST B

 $\int \gg \omega_R$ 

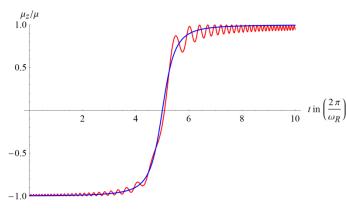
The magnetic moment always precesser tighty around the -story changing - effective field. At the land of the sweep, the magnetic moment finds itself inverted!

Adiabability requires that the magnetic moment always precesses tightly around By, that me ins: Willia one precession period 27 Starmer, the angle of Bey must not have advanced More than a few degrees, i.e.  $\Delta \vartheta \ll 2\pi$ .  $\Delta \vartheta = \vartheta \cdot \Delta t = \vartheta \frac{2\pi}{\Omega_{12000}} \ll 2\pi$ or just Shamm = > By > of. Now the smallest Larmor frightny, and (for a linear sweep d(+) = 2t, say) the largest rate of change of  $\mathcal{I}$  occur around resonance, at  $\mathcal{I} = 0$ , or  $\mathcal{I} = \frac{\pi}{2}$ . Here,  $\vartheta = \frac{T}{Z} - \frac{B_{t,y}}{B_t}$   $\dot{\vartheta} = \frac{|\dot{\omega}|}{\gamma B_t} = \frac{|\dot{\omega}|}{\omega_R}$   $\mathcal{J} = \frac{|\dot{\omega}|}{\beta} = \frac{|\dot{\omega}|}{\omega_R}$ S, (S=0) = JBn = WR  $\Rightarrow \dot{\vartheta} = \frac{|\dot{\omega}|}{\omega_R} \ll \omega_R \quad \text{or} \quad |\dot{\omega}| \ll \omega_R^2$ i.e. The change of w in one Rabi period should be much smaller than the Rabi figurescy. This implies that this inversion will be slower than on on-resonance Tr-puble!

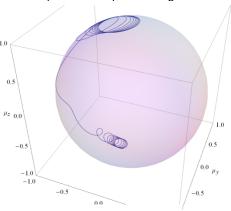
Rapid adiabatic passage with various parameters:

Blue:  $\cos\theta = \frac{\delta(t)}{\sqrt{\delta(t)^2 + \omega_R^2}}$  Cosine of B<sub>eff</sub> with z; Red: z-component of the magnetic moment

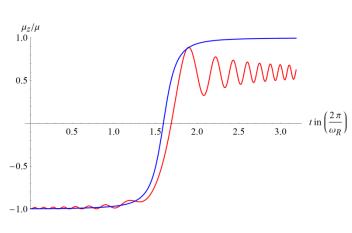
1. Start of sweep:  $\delta=-10~\omega_R$ , End of sweep:  $\delta=+10~\omega_R$ , Sweep rate:  $\dot{\omega}=\frac{20~\omega_R}{10~\frac{2\pi}{\omega_R}}=\frac{1}{\pi}~\omega_R^2$ 

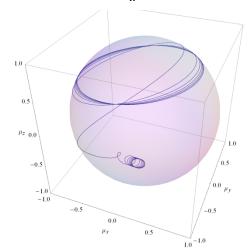


Parametric plots of the tip of the magnetic moment:



2. Start of sweep:  $\delta=-10~\omega_R$ , End of sweep:  $\delta=+10~\omega_R$ , Sweep rate:  $\dot{\omega}=\pi\frac{20~\omega_R}{10~\frac{2\pi}{\omega_R}}=\omega_R^2$ 





3. Start of sweep:  $\delta=-10~\omega_R$ , End of sweep:  $\delta=+10~\omega_R$ , Sweep rate:  $\dot{\omega}=\pi^2\frac{20~\omega_R}{10~\frac{2\pi}{\omega_R}}=\pi\omega_R^2$ 

