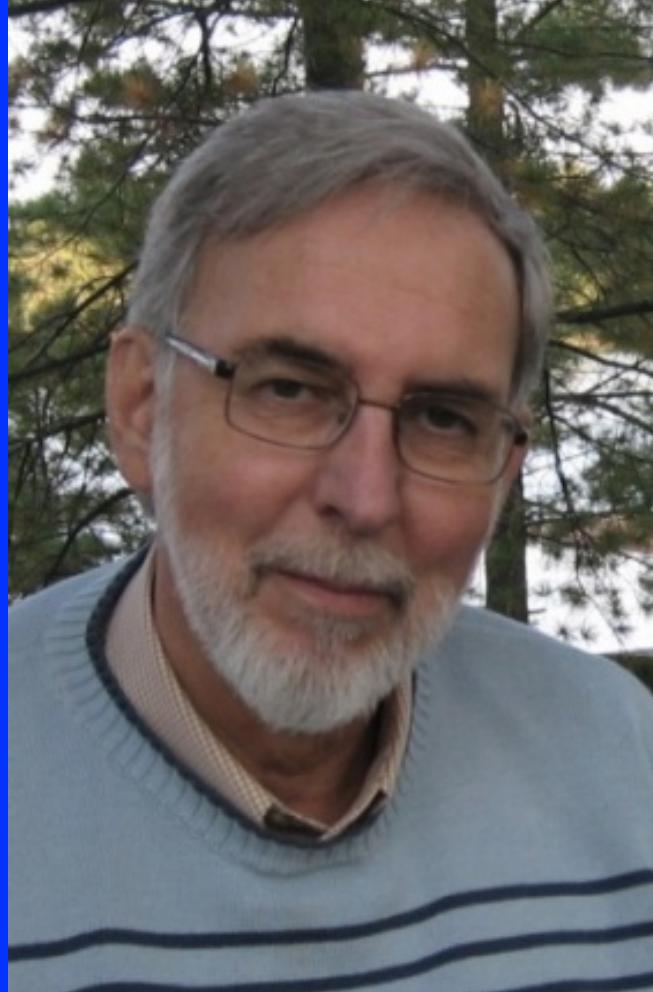


The superfluid mass density



**Gordon Baym
University of Illinois, Urbana**

**Finite Temperature Non-Equilibrium
Superfluid Systems
Heidelberg, 19 September 2011**

In fondest memory,
Allan Griffin, d. May 19, 2011



Landau Two-Fluid Model

Can picture superfluid ^4He as two interpenetrating fluids:

Normal: density $\rho_n(T)$, velocity v_n

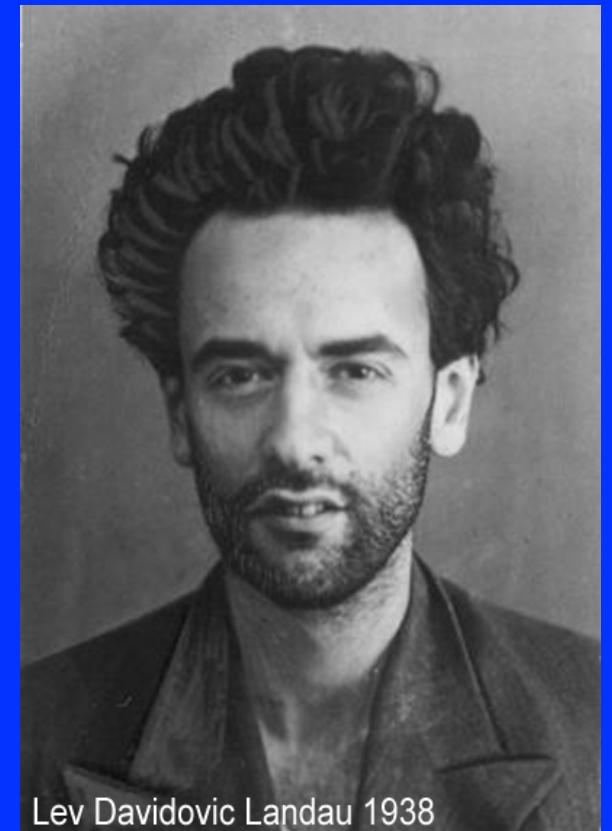
Superfluid: density $\rho_s(T)$, velocity v_s

$$\rho = \rho_n(T) + \rho_s(T)$$

$$\text{Mass current} = \rho_s v_s + \rho_n v_n$$

$$\text{Entropy current} = s v_n$$

:carried by normal fluid only



Lev Davidovic Landau 1938

Second sound (collective mode) =
counter-oscillating normal and superfluids

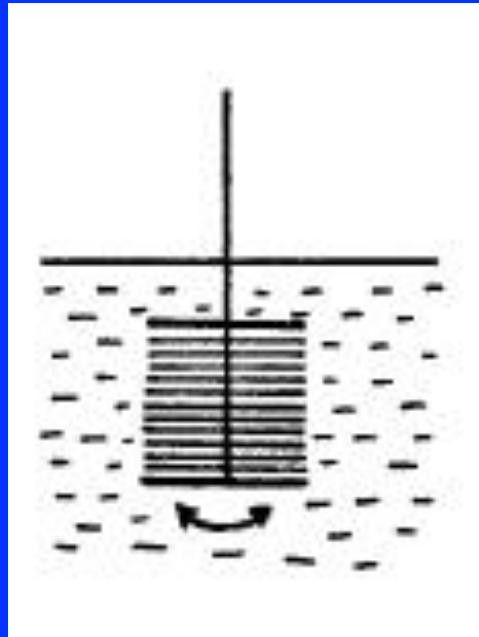
The discovery of superfluidity: a chronology of events in 1935-1938

Allan Griffin
Department of Physics,
University of Toronto
Toronto, Ontario,
Canada M5S 1A7
griffin@physics.utoronto.ca

(September 21, 2006)

The dramatic announcement of superfluidity of liquid ^4He in 1938 is one of the defining moments in modern physics. The two short notes which were published back to back in the Jan. 8 issue of *Nature* (by Kapitza [1] working in Moscow and Allen and Misener [2] working in Cambridge University) immediately caught the attention of the physics community. This stimulated feverish activity in the period leading up to World War II, and in the 1950s developed into a major research area called “quantum fluids.”

Try to rotate helium slowly. Normal fluid component rotates, but superfluid component stays put.



Androniskashvili experiment – with stack of closely spaced disks oscillating back and forth – measure how much fluid rotates

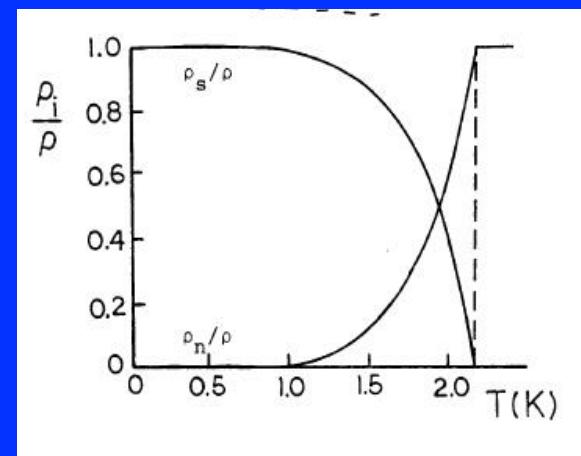
Moment of inertia

$$I = I_{\text{disk}} + I_{\text{fluid}}$$

E.L. Andronikashvili,
J. Physics, USSR, 1946

Measure resonant frequency,
and deduce I_{fluid} from

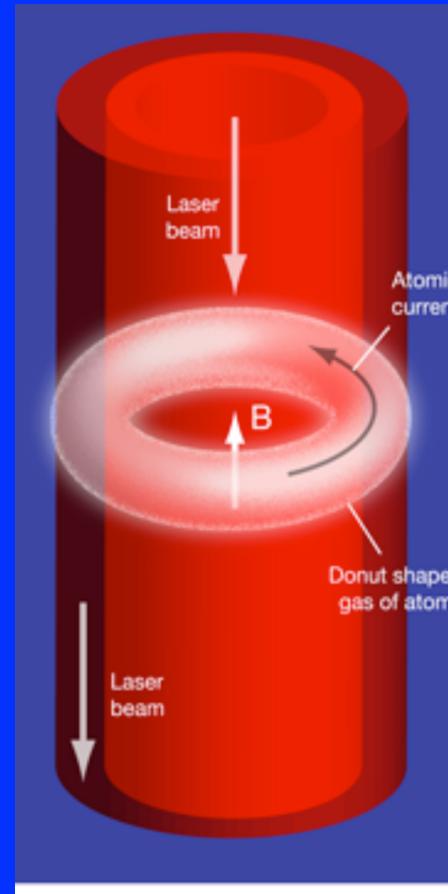
$$I \frac{d^2\theta}{dt^2} = -k\theta$$



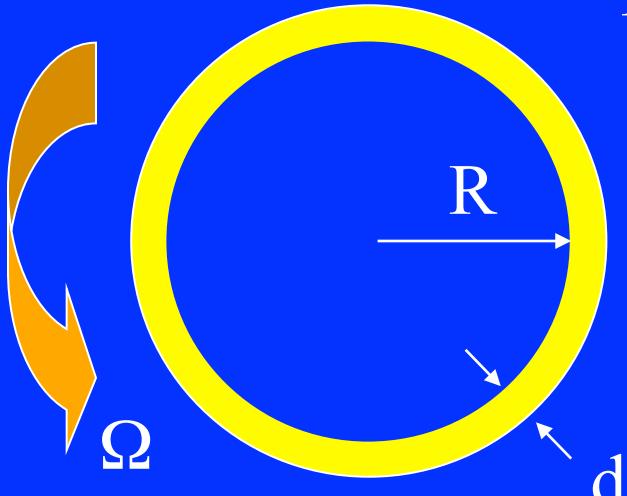
Andronikashvili experiment in cold atoms to measure superfluid mass density

N.R. Cooper and Z. Hadzibabic, PRL 104, 030401 (2010)

Two-photon Raman coupling via beams with different orbital angular momentum difference, to simulate uniform rotation



Hess-Fairbank experiment (*Phys. Rev. Lett.* 19, 216 (1967))



Rotate thin ($d \ll R$) annulus of liquid ⁴He at Ω

- 1) Rotate slowly at $T > T_\lambda$: $\Omega < \Omega_c \sim 1/mR^2$
liquid rotates classically with angular momentum $L = I_{\text{classical}}\Omega$.

$$I_{\text{classical}} = NmR^2$$

- 2) Cool to $T < T_\lambda$: liquid rotates with reduced moment of inertia
 $I(T) < I_{\text{classical}}$. $I(T=0) = 0$.

Only the normal fluid rotates. $I(T) = (\rho_n/\rho)I_{\text{classical}}$
The superfluid component remains stationary in the lab.

Reduction of moment of inertia is an equilibrium phenomenon.

Moment of inertia of superfluid

Reduction of moment of inertia due condensation
= analog of Meissner effect.

$$I = \frac{\rho_n}{\rho} I_{classical}$$

Rotational spectra of nuclei:
 $E = J^2/2I$, indicate moment of inertia, I , reduced from rigid body value, I_{cl} . Migdal (1959). BCS pairing.

Element	β [7]	x_p	x_n	$\left(\frac{I}{I_0}\right)_{rect.}$	$\left(\frac{I}{I_0}\right)_{osc.}$	$\left(\frac{I}{I_0}\right)$ [7] exper.
Nd ¹⁵⁰	0.26	0.54	0.94	0.15	0.38	0.35
Sm ¹⁵²	0.24	0.65	1.02	0.17	0.43	0.38
Gd ¹⁵⁴	0.26	0.52	0.88	0.13	0.35	0.36
Gd ¹⁵⁶	0.33	0.87	1.37	0.22	0.57	0.48
Gd ¹⁵⁷	0.29	0.93	1.60	0.22	0.64	0.60
Dy ¹⁶²	0.30	0.84	1.43	0.23	0.57	0.50
Hf ¹⁷⁹	0.20	0.99	1.75	0.27	0.66	0.52
Os ¹⁸⁶	0.18	0.44	0.69	0.09	0.26	0.28
Th ²³⁰	0.22	0.63	0.95	0.15	0.40	0.43
Th ²³²	0.22	0.84	1.42	0.24	0.60	0.44
U ²³⁸	0.24	0.83	1.29	0.22	0.54	0.43

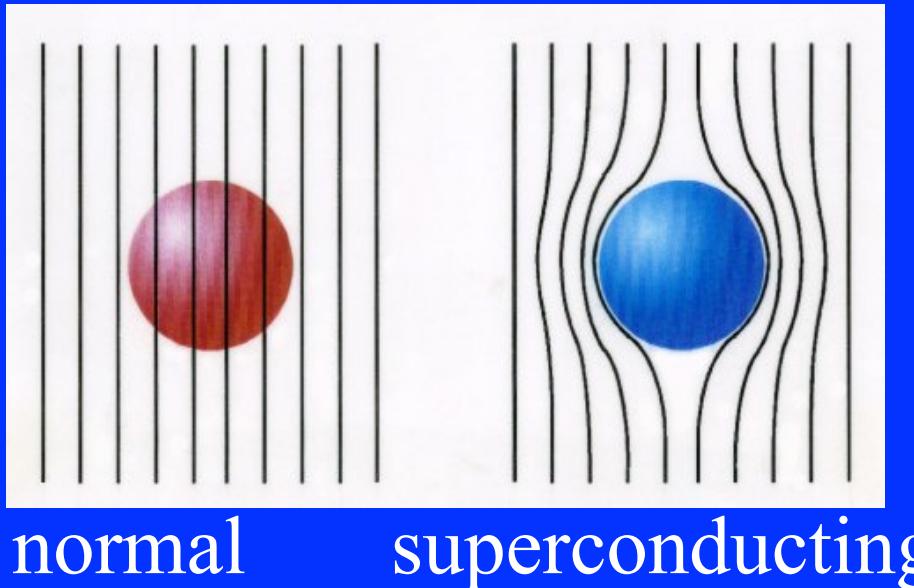
The Meissner effect

W. Meissner and R. Ochsenfeld, Berlin 1933

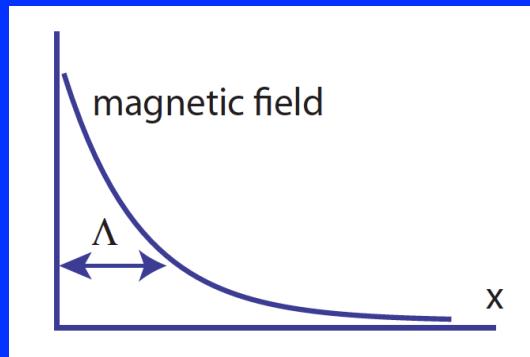


Walther Meissner
(1882-1974)

Superconductors expel magnetic fields
(below critical field):



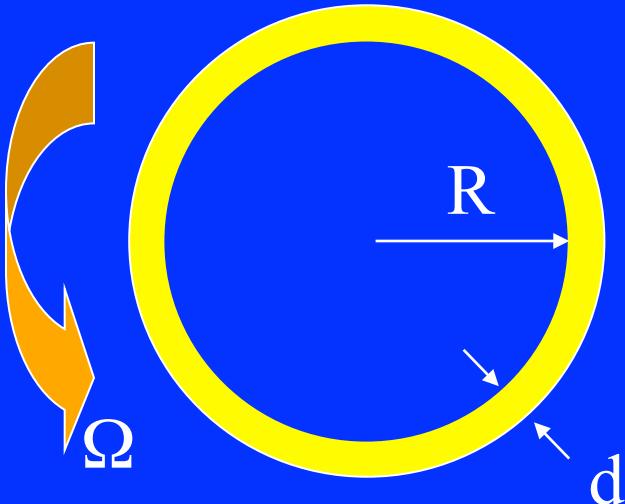
Screening of magnetic field
within penetration depth Λ



Fundamental property of superconductors:
perfect diamagnets -- not perfect
conductors! Equivalent to reduced
moment of inertia in neutral superfluids

$$\frac{1}{\Lambda^2} = \frac{4\pi n e^2 \rho_s}{mc^2 \rho}$$

Superfluid flow



- 1) Rotate rapidly at $T > T_\lambda$: $\Omega > \Omega_c$
liquid rotates classically with angular momentum $L = I_{\text{classical}}\Omega$.
- 2) Continue rotating, cool to $T < T_\lambda$:
liquid rotates classically
- 3) Stop rotation of annulus. Liquid keeps rotating with $L = I_s\Omega$,
where $I_s = (\rho_s/\rho) I_{\text{classical}}$

Only the superfluid rotates. The normal component is stationary.

Superfluid flow is metastable (albeit with huge lifetime in macroscopic system)

Order parameter of condensate

$$\Psi(\vec{r}) = |\psi| e^{i\phi(\vec{r})}$$

wave function of mode into which particles condense

Defined more rigorously by eigenfunction cf.

largest eigenvalue of density matrix

$$\langle \psi(\vec{r}) \psi^\dagger(\vec{r}') \rangle \rightarrow \Psi(\vec{r}) \Psi(\vec{r}')^*$$

Superfluid velocity:

$$\vec{v}_s(\vec{r}) = \frac{\hbar}{m} \nabla \phi$$

Chemical potential:

$$\mu = \partial \phi / \partial t$$

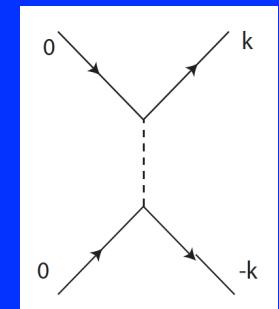
-

Superfluid acceleration eqn.: $\frac{\partial \vec{v}_s}{\partial t} + \nabla \mu = 0$

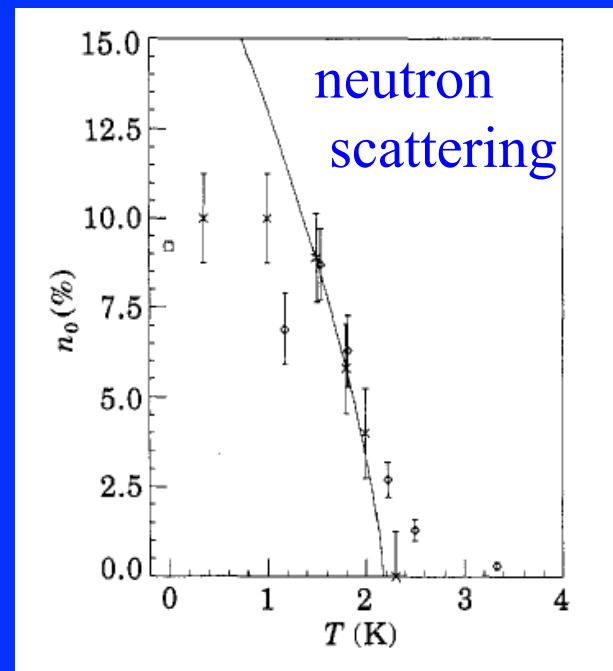
Condensate density is NOT superfluid density

$$\rho_s \neq m|\psi|^2$$

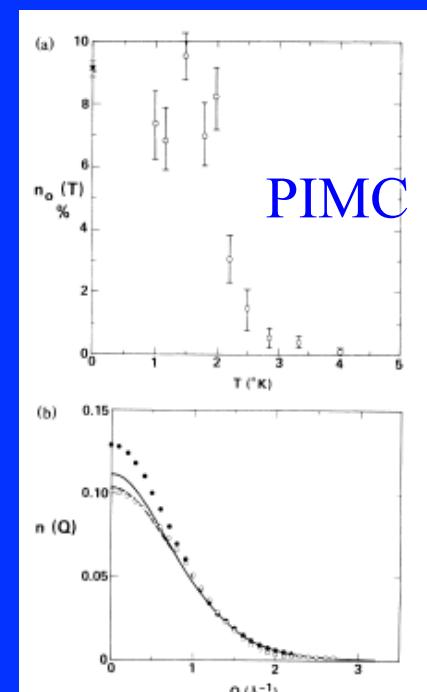
In ground state, interactions drive particles into non-zero momentum single particle states:



In ${}^4\text{He}$ at $T=0$,
 $\rho_s/\rho = 1$, while
 $<10\%$ of particles
are in condensate



*Snow, Wang, & Sokol,
Europhys. Lett. 19 (1992)*



*Ceperley & Pollock,
PRL 56 (1986)*

Order parameter of BCS paired fermions

Paired seen in amplitude to remove a pair of fermions ($\uparrow\downarrow$)
then add pair back, and come back to same state:

$$\langle \psi_{\uparrow}^{\dagger}(1)\psi_{\downarrow}^{\dagger}(2)\psi_{\downarrow}(3)\psi_{\uparrow}(4) \rangle \simeq \langle \psi_{\uparrow}^{\dagger}(1)\psi_{\downarrow}^{\dagger}(2) \rangle \langle \psi_{\downarrow}(3)\psi_{\uparrow}(4) \rangle$$

[Cf., $\langle \psi(\vec{r})\psi^{\dagger}(\vec{r}') \rangle \rightarrow \Psi(\vec{r})\Psi(\vec{r}')^*$ in Bose system]

Order parameter $\langle \psi_{\downarrow}(r)\psi_{\uparrow}(r) \rangle \rightarrow \Psi(r)$, as in Bose system

Similar physics as in Bose system

$$\Psi(\vec{r}) = |\psi| e^{i\phi(\vec{r})}$$

Supercurrent velocity:

$$\vec{v}_s(\vec{r}) = \frac{\hbar}{m} \nabla \phi$$

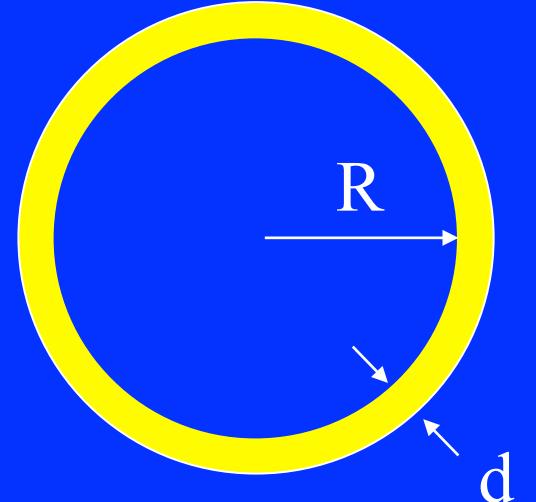
Chemical potential:

$$\mu = -\partial\phi/\partial t$$

(Meta)stability of superfluid flow

Bosons of density n in annulus, $T=0$

$$H = \frac{\hbar^2}{2m} \nabla \psi^* \cdot \nabla \psi + \frac{g}{2} \psi^* \psi^* \psi \psi$$



Condensates: $\psi_0 = \sqrt{n}$ at rest

$\psi_1 = e^{i\phi} \sqrt{n}$ single vortex

Can one slip continuously from single vortex state to rest state, via condensate $\psi = a\psi_0 + b\psi_1$ with $|a|^2 + |b|^2 = 1$?

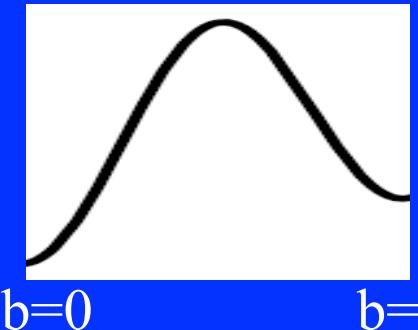
Energy density: $E/V = \frac{\hbar^2}{2mR^2} n|b|^2 + \frac{g}{2} n^2 (|a|^4 + |b|^4 + 4|a|^2|b|^2)$

$$= \frac{\hbar^2}{2mR^2} n|b|^2 + \frac{g}{2} n^2 + 2gn^2|b|^2(1 - |b|^2)$$

Energy density:

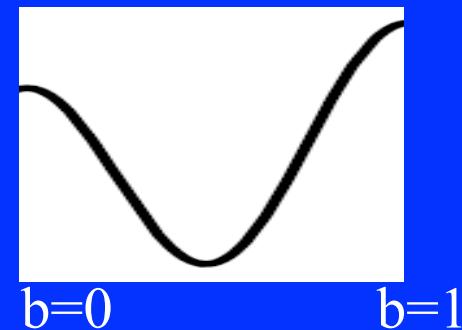
$$= \frac{\hbar^2}{2mR^2} n|b|^2 + \frac{g}{2} n^2 + 2gn^2|b|^2(1 - |b|^2)$$

$g > 0$: have barrier of height $\sim gnN$
for $0 < |b|^2 < 1$



Superfluid flow state (vortex) with $b=1$ is metastable

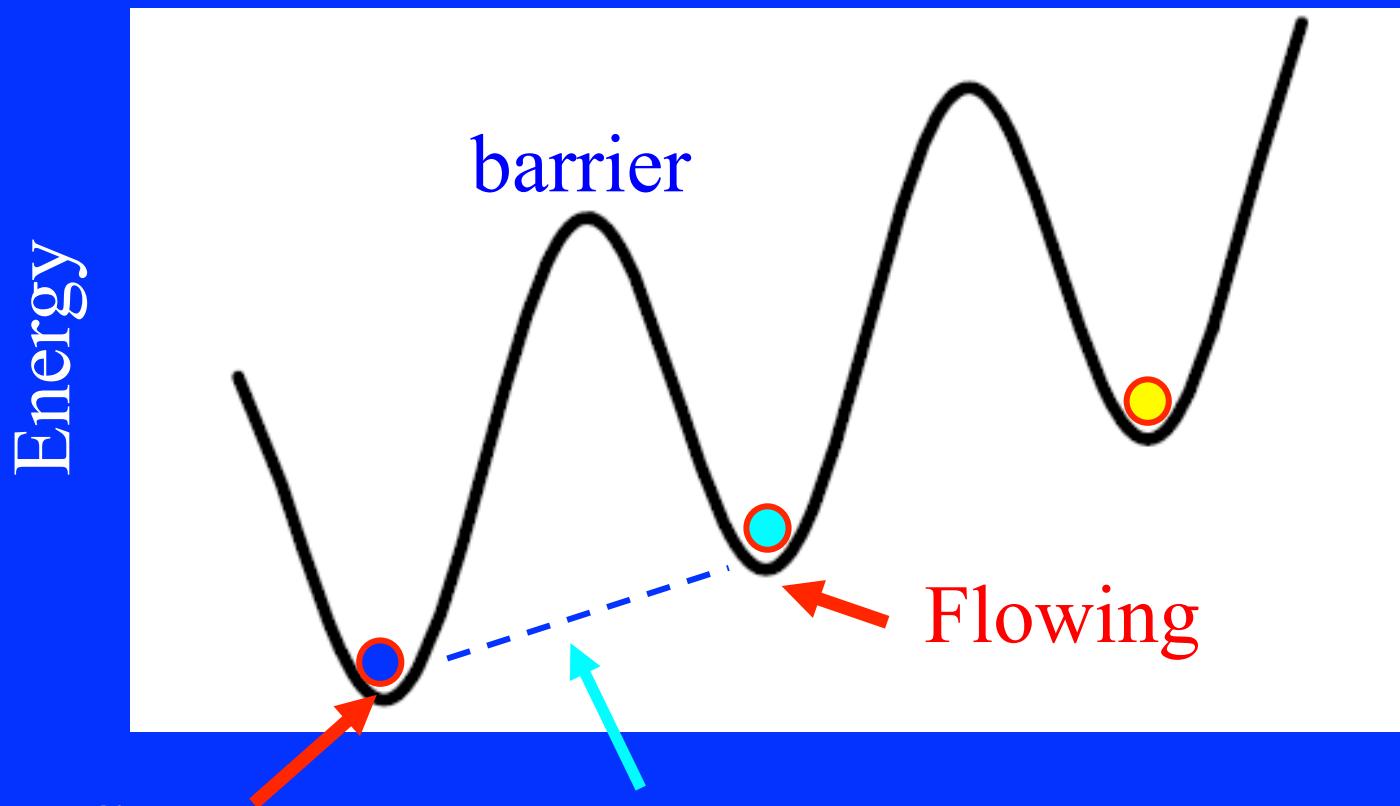
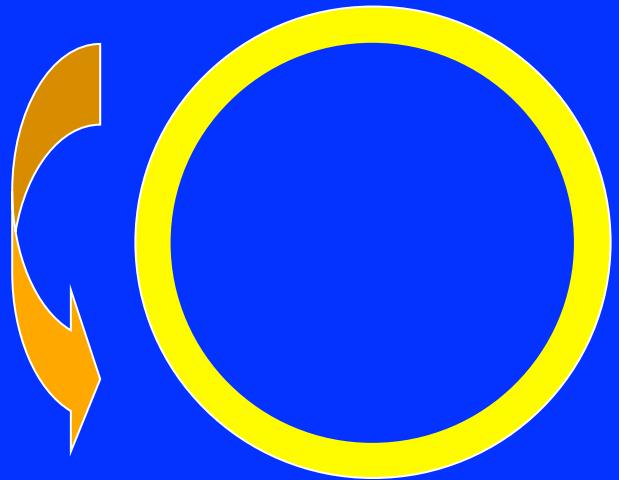
But for $g < 0$ have minimum



Flow is unstable!

Cf. H atom decaying from 2p to 1s state, emitting energy

Superfluid flow difficult to stop
because of enormous energy
barrier, a hill:



No flow Normally (not super), roll down with no barrier,
 from flowing state to resting

Condensate density and superfluid mass density

M. Holzmann & GB, Phys. Rev. B 76, 092502 (2007)

For a superfluid flowing down a pipe (at rest) with superfluid velocity v_s in z direction

$$F(v_s, T, \mu) = F(0, T, \mu) + \frac{1}{2} \rho_s v_s^2$$

$$\partial F / \partial v_s = - \langle P_z - M v_s \rangle / V$$

where P_z is the total momentum.

$$\partial^2 F / \partial v_s^2 = \rho - \beta \langle P_z^2 \rangle / V,$$

so that

$$\rho_n = \beta \langle P_z^2 \rangle / V.$$

In normal state, total momentum is Gaussianly distributed:

$$\propto \exp(-\beta P_z^2 / 2M) \quad \text{so that} \quad \beta \langle P_z^2 \rangle / V = M / V \quad \text{and} \quad \rho_n = \rho$$

In superfluid phase, the total momentum and v_s are entangled and total momentum distribution is not classical.

Exact relation between ρ_s and the condensate density, via the single particle Green's function

P.C. Hohenberg & P.C. Martin, *PRL* 22 (1963); B.D. B.D. Josephson, *PL* 21 (1966); GB, St. Andrews lectures (1967), A. Griffin *PR* B30 (1984)

$$G(k, z) = -i\langle T(\psi^\dagger \psi)\rangle(k, z) \quad z = \text{complex frequency}$$

$$\rho_s = -\lim_{k \rightarrow 0} \frac{n_0 m^2}{k^2 G(k, 0)}$$

Ex. in Bogoliubov mean field ($n_0 = n$),

$$G(k, z) = \frac{z + gn + k^2/2m}{z^2 - gnk^2/m + k^4/4m^2} \Rightarrow \rho_s = nm$$

Valid in 2D as well as 3D:

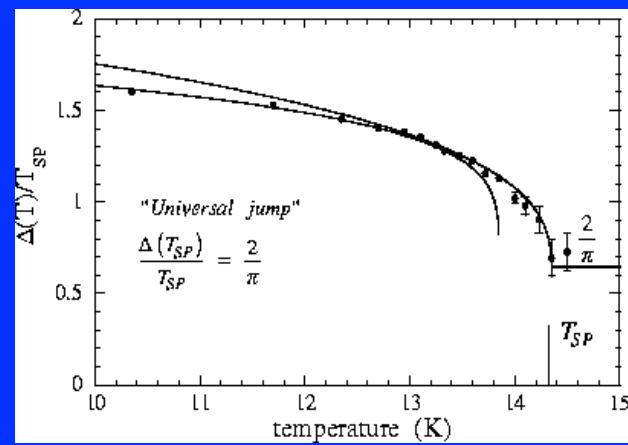
*M. Holzmann & GB, PR B 76 (2007);
 M. Holzmann, GB, J-P Blaizot,
 & F Laloë, PNAS 104 (2007)*

In 2D finite size Berestetskii-Kosterlitz-Thouless system,

$$n_0 \sim 1/(\text{size})^{2-\eta}$$

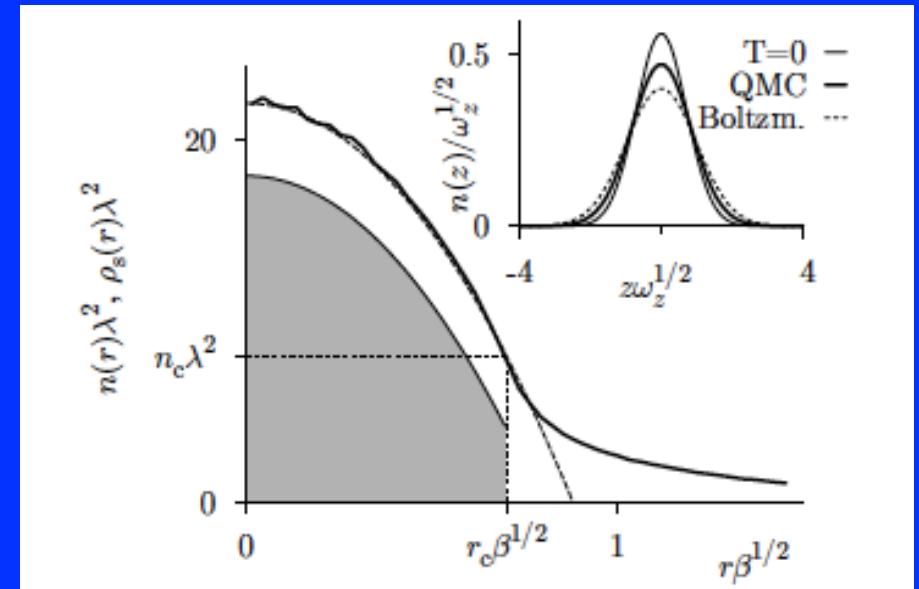
$$G(k, 0) \sim 1/k^{2-\eta}$$

At T_c $\rho_s = \frac{m^2 T}{2\pi\eta}$ $\eta = 1/4$



Order parameter in CuGeO₃

Lorenzo et al., EPL 45 (1999)



Density profile in 2D trap.
 Shaded region $\Leftrightarrow \rho_s$

Holtzmann & Krauth, EPL 82 (2008)

Exact (and equivalent) definition of ρ_s in terms of current-current correlation functions

$$Y_{ij}(\mathbf{r}, \mathbf{r}', \omega) = \int dt e^{i\omega(t-t')} \langle [j_i(\mathbf{r}, t), j_j(\mathbf{r}', t')] \rangle$$

$$\mathbf{j}(\mathbf{r}) = \frac{1}{2im} [\psi^\dagger(\mathbf{r}) \nabla \psi(\mathbf{r}) - \nabla \psi^\dagger(\mathbf{r}) \psi(\mathbf{r})]$$

Decompose into longitudinal and transverse components:

$$Y_{ij}(\mathbf{k}, \omega) = \frac{k_i k_j}{k^2} Y_L(k, \omega) + \left(\delta_{ij} - \frac{k_i k_j}{k^2} \right) Y_T(k, \omega).$$

f-sum rule =>

$$\rho = \lim_{k \rightarrow 0} m^2 \int \frac{d\omega}{2\pi} \frac{Y_L(k, \omega)}{\omega}$$

motion in tube with
closed ends

Define normal mass density

$$\rho_n = \lim_{k \rightarrow 0} m^2 \int \frac{d\omega}{2\pi} \frac{Y_T(k, \omega)}{\omega}$$

motion in tube with open ends



$$\rho_s = \rho - \rho_n$$

How does it work?

$$\text{At } T=0, \quad \int \frac{d\omega}{2\pi} \frac{\Upsilon(k, \omega)}{\omega} = 2 \sum_{a \neq 0} \frac{|\langle a | j_q | 0 \rangle|^2}{E_a - E_0}$$

In general, for low-lying states,

$$\langle a | j_k | 0 \rangle_L \sim k^{1/2} \quad E_a - E_0 \sim k$$

and $\rho_n = \rho$. Same for transverse in normal .

In BEC, $\langle \text{phonon} | \vec{j}_k | 0 \rangle \sim k^{1/2} \hat{k}$.

Thus $\langle \text{phonon} | \vec{j}_k | 0 \rangle_T = 0$ and $\rho_n = 0$.

In superconductor with gap, matrix elements vanish in long wavelength limit, while denominators in T integral remain finite and $\rho_n = 0$.

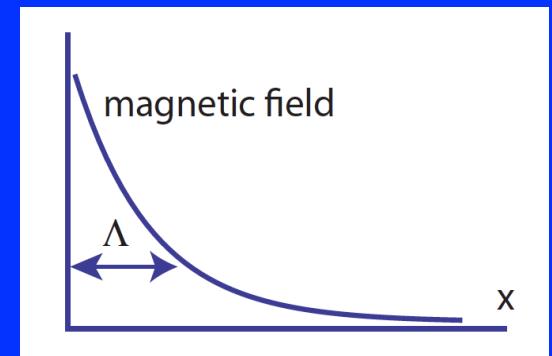
Meissner effect

The $\langle jj \rangle_T$ xx correlation gives fundamental description of the Meissner effect as well:

Penetration depth Λ in superconductor
 \Leftrightarrow screening of magnetic field

$$\frac{1}{\Lambda^2} = \frac{4\pi n e^2 \rho_s}{mc^2 \rho}$$

$$A_{\text{tot}}(\mathbf{k}) = \frac{k^2}{k^2 + (1/\Lambda^2)} A(\mathbf{k}) \quad \Rightarrow$$



Moment of inertia of superfluid

$$I = \frac{\rho_n}{\rho} I_{\text{classical}}$$

In terms of $\langle jj \rangle$:

$$\mathcal{J}_{ij} = \int \frac{d\mathbf{k}}{(2\pi)^3} \int d\mathbf{r} \int d\mathbf{r}' e^{i\mathbf{k} \cdot (\mathbf{r} - \mathbf{r}')} \epsilon_{isl} \epsilon_{jmn} r_s r'_m \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \frac{Y_{ln}(k\omega)}{\omega}$$

Landau calculation of ρ_s for system w. quasiparticles

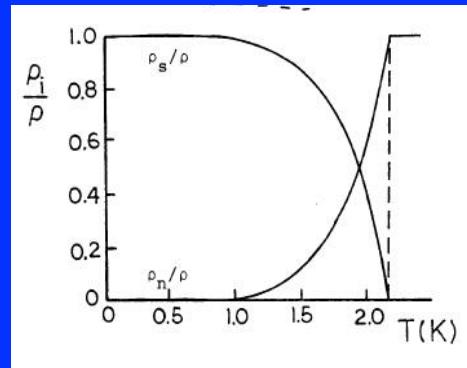
$$\langle \mathbf{P} \rangle = \sum \mathbf{p} \langle N_{\mathbf{p}} \rangle \quad \langle N_{\mathbf{p}} \rangle = [e^{\beta \{\epsilon_{\mathbf{p}} + \mathbf{p} \cdot (\mathbf{v}_s - \mathbf{v}_n)\}} - 1]^{-1}$$

$$\langle \mathbf{P} \rangle = - \sum_{\mathbf{p}} \mathbf{p} (\mathbf{p} \cdot \mathbf{v}_n) \frac{\partial}{\partial \epsilon_{\mathbf{p}}} \frac{1}{e^{\beta \epsilon_{\mathbf{p}}} - 1}$$

$$\rho_n = - \int \frac{d\mathbf{p}}{(2\pi)^3} \frac{p^2}{3} \frac{\partial}{\partial \epsilon_{\mathbf{p}}} \frac{1}{e^{\beta \epsilon_{\mathbf{p}}} - 1}$$

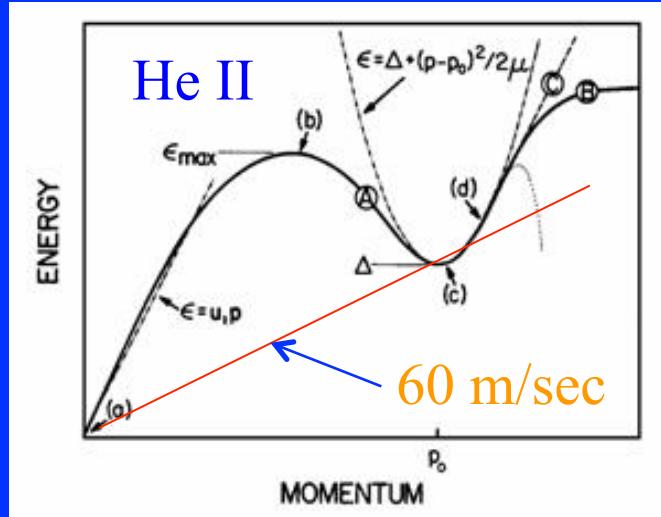
for phonons:

$$\rho_n(T) = \frac{2\pi^2}{45\hbar^3 s^3} T^4$$



The Landau criterion for superfluidity

Superfluid with elementary excitation spectrum $\varepsilon(q)$



Fluid flowing in pipe in x direction, velocity v with respect to walls. In wall frame excitation energy is

$$\varepsilon_v(q) = \varepsilon(q) + vq_x$$

According to Landau:

For $v < \varepsilon(q)/q$ cannot make spontaneous excitations (which would decay superflow) and *flow is superfluid.*

For v opposite to q_x and $v > \varepsilon(q)/q$ have $\varepsilon_v(q) < 0$
Can then make excitations spontaneously, and
superfluidity ceases. $v_{\text{crit}} = 60 \text{ m/sec}$ in superfluid He.

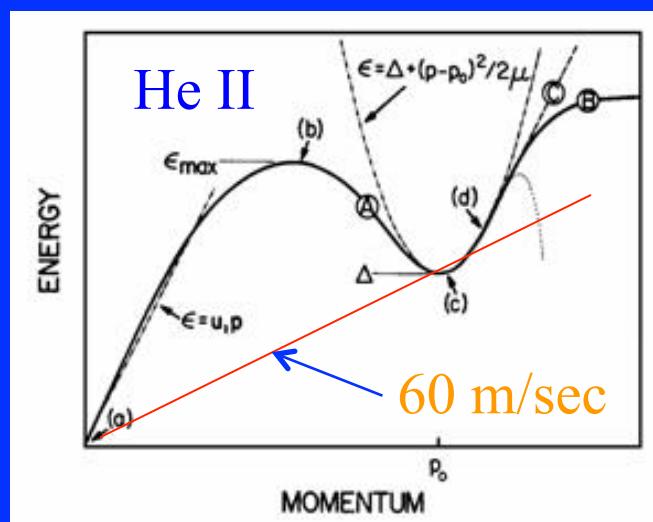
In this way we see that neither phonons nor rotons can be excited if the velocity of flow in helium II is not too large. This means that the flow of the liquid does not slow down, i.e. helium II discloses the phenomenon of superfluidity†.

SUFFICIENT

It must be remarked that already the reasons given above are enough to make the superfluidity vanish at sufficiently large velocities. We leave aside the question as to whether superfluidity disappears at smaller velocities for some other reason (the velocity limit obtained from (4.2) is large—the velocity of sound in helium equals 250 m/sec; (4.4) gives a value only several times lower).

NECESSARY

L.D. Landau, J. Phys. USSR 5, 71 (1941)



At Landau critical velocity, group and phase velocity of excitations are equal:

$$\frac{\partial \epsilon}{\partial q} = \frac{\epsilon}{q}$$

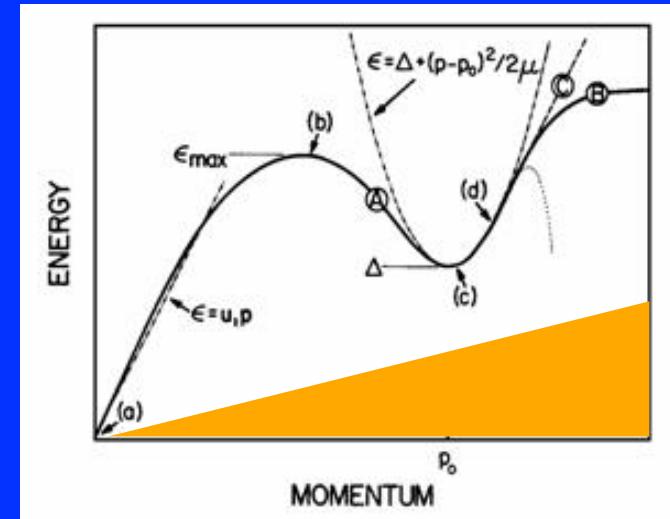
The Landau criterion is neither necessary nor sufficient

Superfluid systems with no “gap”:

- 1) Dilute solutions of degenerate ^3He in superfluid ^4He :
Particle-hole spectrum

$$\omega = (\vec{p} + \vec{q})^2 / 2m - \vec{p}^2 / 2m$$

reaches down to $\omega = 0$ at $\mathbf{q} \neq 0$.

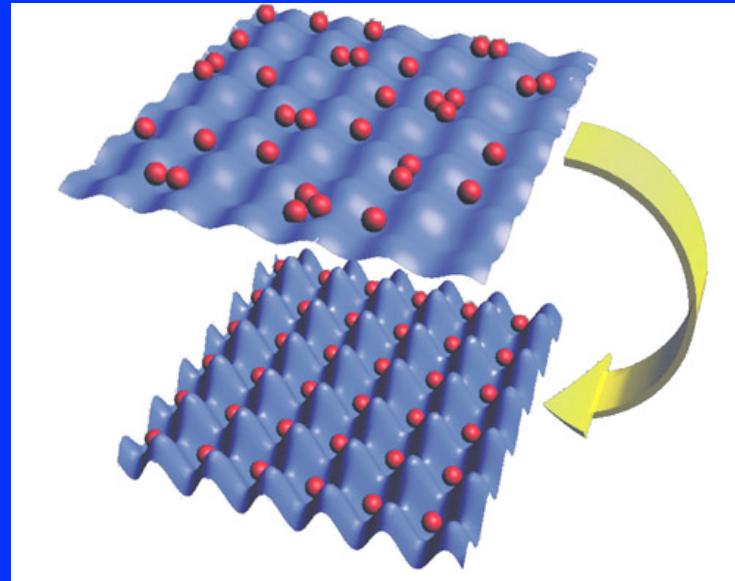
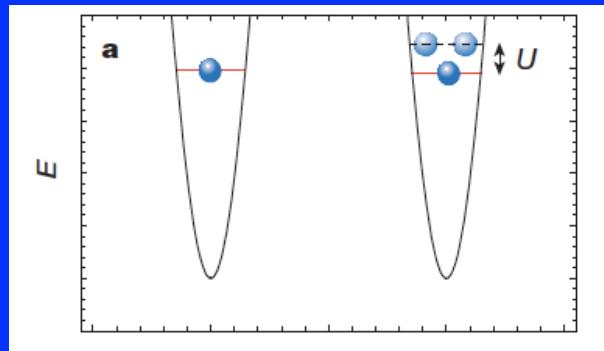


Landau critical velocity vanishes, but system is superfluid.

- 2) Superfluid ^4He at non-zero temperature: Can scatter a phonon of momentum \mathbf{k} to $-\mathbf{k}$ with zero energy change.
Again Landau critical velocity vanishes, but system is a perfectly good superfluid.

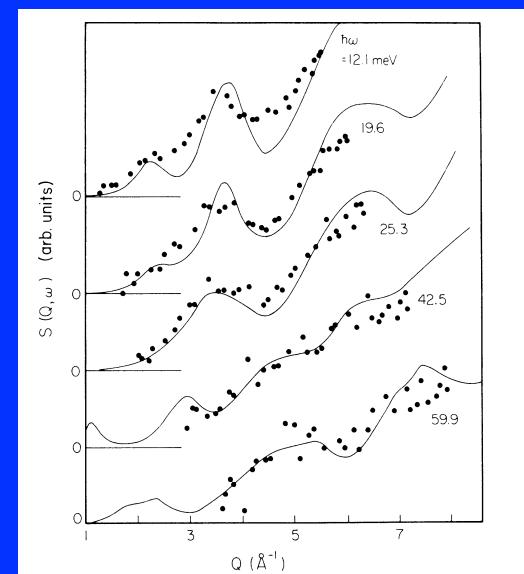
Gap also not sufficient to guarantee superfluidity:

ex. bosons in optical lattice:



superfluid
↓
Mott insulator

Amorphous solids, e.g., Si doped with H, not superfluid.



What happens when the Landau criterion is violated?

The superfluid mass density becomes less than the total mass density. It does not necessarily vanish!

In dilute solutions of ^3He in superfluid ^4He ,

$$\rho_s = \rho - (m^* - m_3)n_3$$

m^* = ^3He effective mass, m_3 = bare mass.

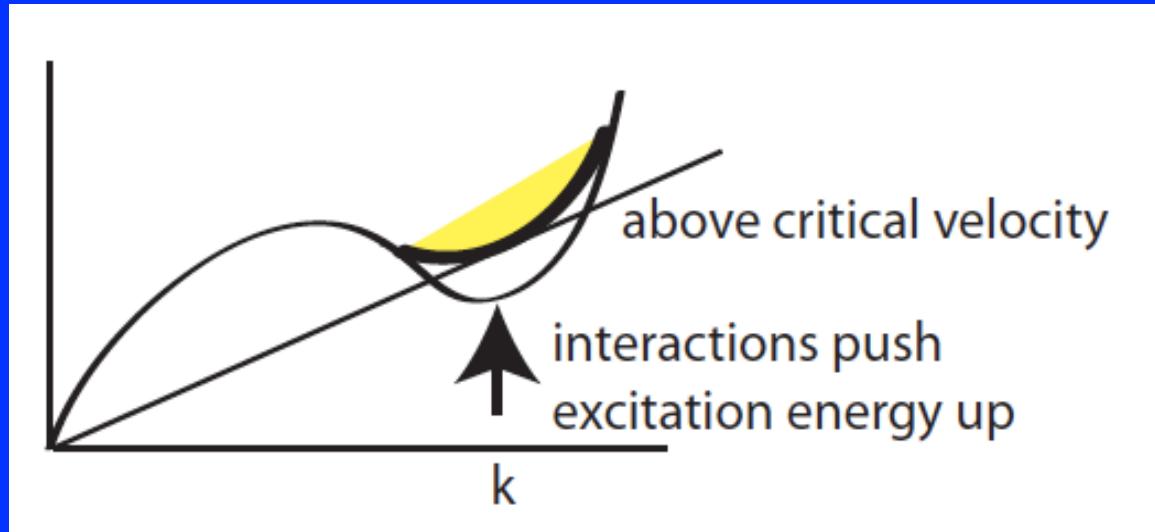
In superfluid ^4He at nonzero temperature,

$$\rho_s = \rho - aT^4 - \dots$$

Formation of non-uniform states

Formation of non-uniform states

L. Pitaevskii, JLTP 87, 127 (1992), GB & CJ Pethick



Beyond the critical velocity spontaneously form excitations of finite momentum k .

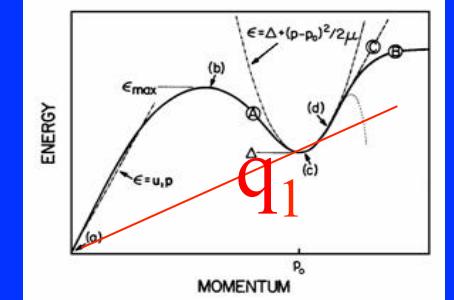
Interactions of these excitations, when repulsive, raises energy of unstable mode k , to make velocity just critical.

Mixing in of modes of momentum k causes condensate to become non-uniform.

Simple model when Landau critical velocity is exceeded

(GB and CJ Pethick)

Weakly interacting Bose gas with finite range interaction $g(r)$, and thus $g(q) (> 0)$, to produce v_{crit} at non-zero $q = q_1$



In Bogoliubov approx:

$$\varepsilon(q) = \left[\frac{ng(q)q^2}{m} + \left(\frac{q^2}{2m} \right)^2 \right]^{1/2}$$

Critical point: where group velocity = phase velocity,

$$\frac{d(gn)}{d(q^2/2m)} = -1,$$

Study stability of uniform condensate
for $q > q_1$

$$\psi(x) = \sqrt{n} e^{iqx}$$

Since excitations at q_1 can form spontaneously, generate new condensate of form

$$\psi(x) = e^{iqx} [\sqrt{n_0} + ue^{-iq_1 x} + ve^{iq_1 x}],$$

Let $u = \zeta \cosh(\phi/2)$, $v = \zeta \sinh(\phi/2)$

Stable solution $\zeta^2 = (v - v_{crit}) \frac{q_1}{2G}$ above critical velocity

$$G = \frac{g_1}{4\epsilon_1^2} \left[g_1 g_2 n^2 - \frac{q_1^2}{m} \left(\frac{q_1^2}{m} + 2g_1 n \right) \right] = \text{effective repulsion of excitations near } q_1$$

$$g_\kappa \equiv g(\kappa q_1)$$

Non uniform density: $n(x) = n + 2\sqrt{n_0} \zeta \cosh \phi \cos(q_1 x)$

Reduction of superfluid mass density

$$\rho_s = mn - mq_1 \frac{\partial \zeta^2}{\partial q} = mn - \frac{q_1^2}{2G}.$$

ΤΗΛΩΝ ΥΘΥ