

Homework #7 solutions

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1. Rayleigh and Thomson scattering using two different interaction Hamiltonians.

(a) Transition matrix calculation (8 points)

The quantized fields at $\mathbf{R}=\mathbf{0}$ are given by

$$\mathbf{A}(\mathbf{0}) = \int d^3k \sum_{\hat{\epsilon}} \frac{\mathcal{E}_{\omega}}{\omega} \left[a_{\hat{\epsilon}}(\mathbf{k}) + a_{\hat{\epsilon}}^{\dagger}(\mathbf{k}) \right] \hat{\epsilon}$$

$$\mathbf{E}_{\perp}(\mathbf{0}) = \int d^3k \sum_{\hat{\epsilon}} i\mathcal{E}_{\omega} \left[a_{\hat{\epsilon}}(\mathbf{k}) - a_{\hat{\epsilon}}^{\dagger}(\mathbf{k}) \right] \hat{\epsilon}$$

where $\mathcal{E}_{\omega} = \sqrt{\frac{\hbar\omega}{2\varepsilon_0 L^3}}$

The Born expansion of the transition matrix \mathcal{T}_{fi} is

$$\mathcal{T}_{fi} = \langle \psi_f | H_I | \psi_i \rangle + \lim_{\epsilon \rightarrow 0^+} \sum_j \frac{\langle \psi_f | H_I | \psi_j \rangle \langle \psi_j | H_I | \psi_i \rangle}{E_i - E_j + i\epsilon} + \dots = \mathcal{T}_{fi}(1) + \mathcal{T}_{fi}(2) + \dots$$

The Dipole Hamiltonian

The interaction Hamiltonian H'_I does not couple $|a; \mathbf{k}\hat{\epsilon}\rangle$ to $|a; \mathbf{k}'\hat{\epsilon}'\rangle$. Hence, the intermediate state of the scattering process is either $|b; 0\rangle$ or $|b; \mathbf{k}\hat{\epsilon}, \mathbf{k}'\hat{\epsilon}'\rangle$. The relevant matrix elements are the following:

$$\begin{aligned} \langle b; 0 | H'_I | a; \mathbf{k}\hat{\epsilon} \rangle &= -i\mathcal{E}_{\omega} \langle b | \mathbf{d} \cdot \hat{\epsilon} | a \rangle \\ \langle b; \mathbf{k}\hat{\epsilon}, \mathbf{k}'\hat{\epsilon}' | H'_I | a; \mathbf{k}\hat{\epsilon} \rangle &= i\mathcal{E}_{\omega} \langle b | \mathbf{d} \cdot \hat{\epsilon}' | a \rangle \end{aligned}$$

The transition matrix element to lowest order comes in at second order, where we sum over the intermediate states b .

$$\mathcal{T}'_{fi}(1) = 0$$

$$\mathcal{T}'_{fi}(2) = \frac{q^2 \hbar}{2\varepsilon_0 L^3} \frac{1}{\sqrt{\omega\omega'}} \sum_b (\omega\omega') \left[\frac{\langle a | \mathbf{r} \cdot \hat{\epsilon}' | b \rangle \langle b | \mathbf{r} \cdot \hat{\epsilon} | a \rangle}{\hbar(\omega - \omega_{ba})} - \frac{\langle a | \mathbf{r} \cdot \hat{\epsilon} | b \rangle \langle b | \mathbf{r} \cdot \hat{\epsilon}' | a \rangle}{\hbar(\omega' + \omega_{ba})} \right]$$

where $\omega_{ba} = (E_b - E_a)/\hbar$

The scattering is elastic, hence $\omega = \omega'$ and it follows that,

$$\mathcal{T}'_{fi} = \frac{q^2 \hbar}{2\varepsilon_0 L^3} \omega \sum_b \left[\frac{\langle a | \mathbf{r} \cdot \hat{\epsilon}' | b \rangle \langle b | \mathbf{r} \cdot \hat{\epsilon} | a \rangle}{\hbar(\omega - \omega_{ba})} - \frac{\langle a | \mathbf{r} \cdot \hat{\epsilon} | b \rangle \langle b | \mathbf{r} \cdot \hat{\epsilon}' | a \rangle}{\hbar(\omega + \omega_{ba})} \right]$$

(i) **Rayleigh scattering**

For Rayleigh scattering $\hbar\omega \ll E_b - E_a$

$$\mathcal{T}'_{fi} = -\frac{q^2\hbar}{2\varepsilon_0 L^3} \omega \sum_b \left[\frac{\langle a|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}'|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}|a\rangle + \langle a|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}'|a\rangle}{\hbar\omega_{ba}} \right]$$

(ii) **Thomson scattering**

For Thomson scattering $\hbar\omega \gg E_b - E_a$, we must expand \mathcal{T}'_{fi} to lowest non-vanishing order in $x \equiv (E_b - E_a)/\hbar\omega$

$$\begin{aligned} T'_{fi} &= \frac{q^2\hbar}{2\varepsilon_0 L^3} \sum_b [\langle a|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}'|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}|a\rangle (1+x)^{-1} - \langle a|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}'|a\rangle (1-x)^{-1}] \\ &= \frac{q^2}{2\varepsilon_0 L^3} \frac{1}{\hbar\omega} \sum_b [\langle a|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}'|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}|a\rangle + \langle a|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}'|a\rangle] (E_b - E_a) \end{aligned}$$

Using the commutation relation, $[\mathbf{r}, H] = \frac{i\hbar}{m}\mathbf{p}$, we can write the following identity:

$$\frac{\langle b|\mathbf{p} \cdot \hat{\boldsymbol{\varepsilon}}|a\rangle}{(E_b - E_a)} = i\frac{m}{\hbar} \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}|a\rangle$$

Plugging in this identity and using the closure relation of the states $|b\rangle$, we obtain,

$$\begin{aligned} T'_{fi} &= -i\frac{1}{m} \frac{q^2}{2\varepsilon_0 L^3} \frac{1}{\omega} \sum_b [\langle a|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}'|b\rangle \langle b|\mathbf{p} \cdot \hat{\boldsymbol{\varepsilon}}|a\rangle - \langle a|\mathbf{p} \cdot \hat{\boldsymbol{\varepsilon}}|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}'|a\rangle] \\ &= -i\frac{1}{m} \frac{q^2}{2\varepsilon_0 L^3} \frac{1}{\omega} \langle a|[\mathbf{r} \cdot \hat{\boldsymbol{\varepsilon}}', \mathbf{p} \cdot \hat{\boldsymbol{\varepsilon}}]|a\rangle \\ &= \frac{q^2\hbar}{2\varepsilon_0 L^3} \frac{1}{\omega} \frac{1}{m} \hat{\boldsymbol{\varepsilon}} \cdot \hat{\boldsymbol{\varepsilon}}' \end{aligned}$$

The Coulomb-gauge Hamiltonian

For the Coulomb-gauge Hamiltonian $H_I = -\frac{q}{m}\mathbf{p} \cdot \mathbf{A}(\mathbf{0}) + \frac{q^2}{2m}\mathbf{A}^2(\mathbf{0})$

Again we use the Born expansion, but this time the first order term is non-vanishing, because $\mathbf{A}^2(\mathbf{0})$ can couple $|i\rangle$ and $|f\rangle$ directly. There are two kind of terms, one corresponding to $\langle f|\mathbf{A}^2(\mathbf{0})|i\rangle$ and the other to $\langle f|\mathbf{p} \cdot \mathbf{A}(\mathbf{0})|i\rangle$. The relevant matrix elements are given by:

$$\begin{aligned}\langle b; 0|H_I|a; \mathbf{k}\hat{\epsilon}\rangle &= \frac{\mathcal{E}_\omega}{\omega}\langle b|\mathbf{p} \cdot \hat{\epsilon}|a\rangle \\ \langle b; \mathbf{k}\hat{\epsilon}, \mathbf{k}'\hat{\epsilon}'|H_I|a; \mathbf{k}\hat{\epsilon}\rangle &= \frac{\mathcal{E}_\omega}{\omega}\langle b|\mathbf{p} \cdot \hat{\epsilon}'|a\rangle\end{aligned}$$

The contribution due to the first and second terms are thus,

$$\begin{aligned}\mathcal{T}_{fi}(1) &= \frac{q^2\hbar}{2\varepsilon_0 L^3} \frac{1}{\sqrt{\omega\omega'}} \frac{1}{m} \hat{\epsilon} \cdot \hat{\epsilon}' \\ \mathcal{T}_{fi}(2) &= \frac{q^2\hbar}{2\varepsilon_0 L^3} \frac{1}{\sqrt{\omega\omega'}} \frac{1}{m^2} \sum_b \left[\frac{\langle a|\mathbf{p} \cdot \hat{\epsilon}'|b\rangle\langle b|\mathbf{p} \cdot \hat{\epsilon}|a\rangle}{\hbar(\omega - \omega_{ba})} - \frac{\langle a|\mathbf{p} \cdot \hat{\epsilon}|b\rangle\langle b|\mathbf{p} \cdot \hat{\epsilon}'|a\rangle}{\hbar(\omega' + \omega_{ba})} \right]\end{aligned}$$

(i) Rayleigh scattering

For Rayleigh scattering and assuming $\omega = \omega'$, we expand \mathcal{T}_{fi} in powers of $x \equiv \frac{\hbar\omega}{E_b - E_a}$.

$$\mathcal{T}_{fi}(2) = -\frac{q^2}{m^2} \frac{\hbar}{2\varepsilon_0 L^3} \frac{1}{\omega} \sum_b \left(\frac{\langle a|\mathbf{p} \cdot \boldsymbol{\epsilon}'|b\rangle\langle b|\mathbf{p} \cdot \boldsymbol{\epsilon}|a\rangle}{E_b - E_a} (1-x)^{-1} + \frac{\langle a|\mathbf{p} \cdot \boldsymbol{\epsilon}|b\rangle\langle b|\mathbf{p} \cdot \boldsymbol{\epsilon}'|a\rangle}{E_b - E_a} (1+x)^{-1} \right)$$

Expanding \mathcal{T}_{fi} up to second order, we obtain:

$$\begin{aligned}\mathcal{T}_{fi}(2) &= -\frac{q^2}{m^2} \frac{\hbar}{2\varepsilon_0 L^3} \frac{1}{\omega} \frac{m}{i\hbar} \sum_b \left(\langle a|\mathbf{p} \cdot \boldsymbol{\epsilon}'|b\rangle\langle b|\mathbf{r} \cdot \boldsymbol{\epsilon}|a\rangle (1-x)^{-1} - \langle a|\mathbf{r} \cdot \boldsymbol{\epsilon}|b\rangle\langle b|\mathbf{p} \cdot \boldsymbol{\epsilon}'|a\rangle (1+x)^{-1} \right) \\ &= -i \frac{q^2}{m\hbar} \frac{\hbar}{2\varepsilon_0 L^3} \frac{1}{\omega} \langle a|[\mathbf{p} \cdot \boldsymbol{\epsilon}', \mathbf{r} \cdot \boldsymbol{\epsilon}]|a\rangle - i \frac{q^2}{m\hbar} \frac{\hbar}{2\varepsilon_0 L^3} \frac{1}{\omega} (\hbar\omega) \sum_b \left(\frac{\langle a|\mathbf{p} \cdot \boldsymbol{\epsilon}'|b\rangle\langle b|\mathbf{r} \cdot \boldsymbol{\epsilon}|a\rangle}{E_b - E_a} + \frac{\langle a|\mathbf{r} \cdot \boldsymbol{\epsilon}|b\rangle\langle b|\mathbf{p} \cdot \boldsymbol{\epsilon}'|a\rangle}{E_b - E_a} \right) \\ &\quad - i \frac{q^2}{m\hbar} \frac{\hbar}{2\varepsilon_0 L^3} \frac{1}{\omega} (\hbar\omega)^2 \sum_b \left(\frac{\langle a|\mathbf{p} \cdot \boldsymbol{\epsilon}'|b\rangle\langle b|\mathbf{r} \cdot \boldsymbol{\epsilon}|a\rangle}{(E_b - E_a)^2} - \frac{\langle a|\mathbf{r} \cdot \boldsymbol{\epsilon}|b\rangle\langle b|\mathbf{p} \cdot \boldsymbol{\epsilon}'|a\rangle}{(E_b - E_a)^2} \right) + O(x^3)\end{aligned}$$

Using the same identity on $\frac{\langle b|\mathbf{p} \cdot \hat{\epsilon}|a\rangle}{(E_b - E_a)}$ and using the canonical commutation relation, we get

$$\mathcal{T}_{fi}(2) = -\frac{q^2}{m} \frac{\hbar}{2\varepsilon_0 L^3} \frac{1}{\omega} \boldsymbol{\epsilon} \cdot \boldsymbol{\epsilon}'$$

$$\begin{aligned}
& -\frac{q^2}{\hbar^2} \frac{\hbar}{2\epsilon_0 L^3} \frac{1}{\omega} \cdot (\hbar\omega) \sum_b (\langle a|\mathbf{r} \cdot \boldsymbol{\epsilon}'|b\rangle \langle b|\mathbf{r} \cdot \boldsymbol{\epsilon}|a\rangle - \langle a|\mathbf{r} \cdot \boldsymbol{\epsilon}|b\rangle \langle b|\mathbf{r} \cdot \boldsymbol{\epsilon}'|a\rangle) \\
& -\frac{q^2}{\hbar^2} \frac{\hbar}{2\epsilon_0 L^3} \frac{1}{\omega} \cdot (\hbar\omega)^2 \sum_b \left(\frac{\langle a|\mathbf{r} \cdot \boldsymbol{\epsilon}'|b\rangle \langle b|\mathbf{r} \cdot \boldsymbol{\epsilon}|a\rangle + \langle a|\mathbf{r} \cdot \boldsymbol{\epsilon}|b\rangle \langle b|\mathbf{r} \cdot \boldsymbol{\epsilon}'|a\rangle}{E_b - E_a} \right) + O(x^3)
\end{aligned}$$

The first term cancels the contribution from $\mathcal{T}_{fi}(1)$, the middle term is zero; therefore, we are left with

$$\mathcal{T}_{fi} = \frac{q^2 \hbar}{2\epsilon_0 L^3} \omega \sum_b \left[\frac{\langle a|\mathbf{r} \cdot \hat{\boldsymbol{\epsilon}}'|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\epsilon}}|a\rangle + \langle a|\mathbf{r} \cdot \hat{\boldsymbol{\epsilon}}|b\rangle \langle b|\mathbf{r} \cdot \hat{\boldsymbol{\epsilon}}'|a\rangle}{\hbar\omega_{ab}} \right] + O(x^3)$$

which is the same as the one we found with the dipole Hamiltonian.

(ii) **Thomson scattering**

For Thomson scattering $\hbar\omega \gg E_b - E_a$, we can ignore $E_b - E_a$ compared to $\hbar\omega$. One can expand \mathcal{T}_{fi}^2 in powers of $x \equiv \frac{E_b - E_a}{\hbar\omega}$, as has been done before for the dipole interaction. We find that the zeroth order term vanishes:

$$\mathcal{T}_{fi}(2) \approx -\frac{q^2}{m^2} \frac{\hbar}{2\epsilon_0 L^3} \frac{1}{\omega} \sum_b (\langle a|\mathbf{p} \cdot \hat{\boldsymbol{\epsilon}}'|b\rangle \langle b|\mathbf{p} \cdot \hat{\boldsymbol{\epsilon}}|a\rangle + \langle a|\mathbf{p} \cdot \hat{\boldsymbol{\epsilon}}|b\rangle \langle b|\mathbf{p} \cdot \hat{\boldsymbol{\epsilon}}'|a\rangle) \frac{(E_b - E_a)}{(\hbar\omega)^2}$$

By replacing $(E_b - E_a)$ by E_I in the numerator and by then using the closure relation over b we obtain

$$\mathcal{T}_{fi}(2) \approx -\frac{q^2}{m^2} \frac{\hbar}{2\epsilon_0 L^3} \frac{1}{\omega} \frac{E_I}{\hbar\omega} \frac{\langle a|(p^2/m)|a\rangle}{\hbar\omega}$$

Now, twice the average kinetic energy is the same order of magnitude as E_I . Thus, we prove that

$$\mathcal{T}_{fi}(2) \sim \mathcal{T}_{fi}(1) \left(\frac{E_I}{\hbar\omega} \right)^2$$

$\mathcal{T}_{fi}(1)$ is thus larger than $\mathcal{T}_{fi}(2)$ and we obtain the same result as with the dipole Hamiltonian upon substitution of $\langle b|\mathbf{r} \cdot \hat{\boldsymbol{\epsilon}}|a\rangle$ and using the commutation relation, namely

$$\mathcal{T}_{fi} = \frac{q^2 \hbar}{2\epsilon_0 L^3} \frac{1}{\omega} \frac{1}{m} \hat{\boldsymbol{\epsilon}} \cdot \hat{\boldsymbol{\epsilon}}'$$

One can also show from general arguments that the transition amplitudes calculated from both Hamiltonians are exactly the same (see Photon and Atoms by C.Cohen-Tannoudji)

- (b) Thomson scattering cross section calculations (2 points)

The transition probability per unit time and per unit solid angle is

$$\begin{aligned}\frac{\delta w_{fi}}{\delta \Omega'} &= \frac{2\pi}{\hbar} \left(\frac{q^2}{2m \varepsilon_0 L^3 \omega} \right)^2 (\hat{\varepsilon} \cdot \hat{\varepsilon}')^2 \frac{L^3}{8\pi^3} \frac{(\hbar c k)^2}{\hbar^3 c^3} \\ &= \frac{e^4}{m^2 c^3 L^3} (\hat{\varepsilon} \cdot \hat{\varepsilon}')^2\end{aligned}$$

where $e^2 = \frac{q^2}{4\pi\varepsilon_0}$. If we divide by the photon flux, which is equal to c/L^3 , the differential cross section is

$$\frac{d\sigma}{d\Omega'} = r_0^2 (\hat{\varepsilon} \cdot \hat{\varepsilon}')^2$$

To find the total cross section, sum $(\hat{\varepsilon} \cdot \hat{\varepsilon}')^2$ over the two polarizations orthogonal to \mathbf{k}' and to make the angular average.

Thus the total cross section for Thomson scattering is $\frac{8\pi}{3} r_0^2$, where r_0 is the classical electron radius.

2. Long-range (Van der Waals) interaction between ground-state atoms

- (a) (2 Points) What is the first non-vanishing term in the series for the perturbed ground state energy of the system?

Solution: Using standard perturbation theory, the series for the perturbed ground state energy of the system is

$$\Delta E_{gg} = V_{gg} + \sum_{i \neq g} \frac{|V_{ig}|^2}{E_g^0 - E_i^0} + \text{higher-order terms}$$

Take the z -axis to be along the internuclear separation vector \vec{R} . The perturbation Hamiltonian becomes

$$H_{E1} = e^2 \frac{(x_a x_b + y_a y_b - 2z_a z_b)}{R^3}$$

and

$$\begin{aligned}\Delta E_g &= e^2 \frac{\langle g_a g_b | x_a x_b + y_a y_b - 2z_a z_b | g_a g_b \rangle}{R^3} + e^4 \sum_{i_a i_b} \frac{|\langle i_a i_b | \frac{x_a x_b + y_a y_b - 2z_a z_b}{R^3} | g_a g_b \rangle|^2}{(E_g^a + E_g^b) - (E_i^a + E_i^b)} \\ &\quad + \text{higher-order terms}\end{aligned}$$

The first-order term in this perturbation series vanishes regardless of the parity of the ground state because H_{E1} is an odd-parity operator [1].

Thus, the first non-vanishing term in the perturbation series is the second-order term which varies like $\frac{1}{R^6}$.

(b) (5 Points) Express the result from (a) in terms of oscillator strengths.

Solution: We need to evaluate the second-order term:

$$\begin{aligned}\Delta E_{g_a g_b} &= \sum_{n_a, l_a, n_b, l_b} \sum_{m_a, m_b} e^4 \frac{|\langle i_a i_b | \frac{x_a x_b + y_a y_b - 2z_a z_b}{R^3} | g_a g_b \rangle|^2}{(E_g^a + E_g^b) - (E_i^a + E_i^b)} \\ &= e^4 \sum_{n_a, l_a, n_b, l_b} \frac{\frac{1}{R^6} [\sum_{m_a m_b} |\langle i_a i_b | x_a x_b + y_a y_b - 2z_a z_b | g_a g_b \rangle|^2]}{(E_g^a + E_g^b) - (E_i^a + E_i^b)}\end{aligned}$$

Notice that the sum over the upper states has been explicitly separated into sums over quantum numbers n and l and sums over the orientational quantum number m .

Next the sum over the upper state orientational quantum numbers m_a and m_b is calculated

$$\begin{aligned}\sum_{m_a m_b} |\langle i_a i_b | x_a x_b + y_a y_b - 2z_a z_b | g_a g_b \rangle|^2 &= \sum_{m_a m_b} S_a^{11} S_b^{11} + S_a^{12} S_b^{12} - 2S_a^{13} S_b^{13} + \\ &S_a^{21} S_b^{21} + S_a^{22} S_b^{22} - 2S_a^{23} S_b^{23} - 2S_a^{31} S_b^{31} - 2S_a^{32} S_b^{32} + 4S_a^{33} S_b^{33}\end{aligned}$$

To save space above, I have used some shorthand notation:

$$S_a^{ij} \equiv \langle n_a, l_a, m_a | x_a^i | g_a \rangle \langle g_a | x_a^j | n_a, l_a, m_a \rangle$$

$$S_b^{ij} \equiv \langle n_b, l_b, m_b | x_b^i | g_b \rangle \langle g_b | x_b^j | n_b, l_b, m_b \rangle$$

Where $(x^1, x^2, x^3) = (x, y, z)$. For simple cases (such as the hydrogen atom), we expect that the “off-diagonal” terms in the above sum will “average” to zero and that $S^{ij} = S^{ij} \delta_{ij}$. We can simplify the above sum to be

$$\begin{aligned}\sum_{m_a m_b} |\langle i_a i_b | x_a x_b + y_a y_b - 2z_a z_b | g_a g_b \rangle|^2 &= \sum_{m_a m_b} S_a^{11} S_b^{11} + 0 - 0 + 0 + \\ &S_a^{22} S_b^{22} - 0 - 0 - 0 + 4S_a^{33} S_b^{33}\end{aligned}$$

Let us calculate this explicitly for the case of upper states $|i\rangle$ which have $l = 1$ (i.e. $|P\rangle$ states) and lower states $|g\rangle$ which have $l = 0$ (i.e. $|S\rangle$ states). To do so we express the dipole operator and the angular dependences of the $|P\rangle$ and $|S\rangle$ states in terms of spherical harmonics

$$\begin{aligned}x &= r \sqrt{\frac{4\pi}{3}} \frac{1}{\sqrt{2}} (Y_1^{-1} - Y_1^1) \\ y &= r \sqrt{\frac{4\pi}{3}} \frac{i}{\sqrt{2}} (Y_1^{-1} + Y_1^1) \\ z &= r \sqrt{\frac{4\pi}{3}} Y_1^0\end{aligned}$$

Two facts:

- i. The angular dependence of a state $|l, m\rangle$ is given by Y_l^m .

ii. The orthogonality condition for spherical harmonics, $\int (Y_{l_1}^{m_1})^* Y_{l_2}^{m_2} d\Omega = \delta_{l_1 l_2} \delta_{m_1 m_2}$.

Using these facts we obtain the following results for our matrix elements (Of course, the calculation is made simpler by the fact that $|S\rangle$ states have no angular dependence because $Y_0^0 = \frac{1}{\sqrt{4\pi}}$. Another way: I chose the $|S\rangle$ states for a reason!)

$$\begin{aligned} \langle l=1, m=+1|x|l=0, m=0\rangle &= \frac{-\langle r\rangle}{\sqrt{6}} & \langle l=1, m=+1|y|l=0, m=0\rangle &= \frac{i\langle r\rangle}{\sqrt{6}} \\ \langle l=1, m=0|x|l=0, m=0\rangle &= 0 & \langle l=1, m=0|y|l=0, m=0\rangle &= 0 \\ \langle l=1, m=-1|x|l=0, m=0\rangle &= \frac{\langle r\rangle}{\sqrt{6}} & \langle l=1, m=-1|y|l=0, m=0\rangle &= \frac{i\langle r\rangle}{\sqrt{6}} \\ \langle l=1, m=+1|z|l=0, m=0\rangle &= 0 \\ \langle l=1, m=0|z|l=0, m=0\rangle &= \frac{\langle r\rangle}{\sqrt{3}} \\ \langle l=1, m=-1|z|l=0, m=0\rangle &= 0 \end{aligned}$$

Our sum over orientations becomes

$$\begin{aligned} \sum_{m_a m_b} |\langle i_a i_b | x_a x_b + y_a y_b - 2z_a z_b | g_a g_b \rangle|^2 &= \left(\frac{\langle r_a \rangle^2}{3} \right) \left(\frac{\langle r_b \rangle^2}{3} \right) + \left(\frac{\langle r_a \rangle^2}{3} \right) \left(\frac{\langle r_b \rangle^2}{3} \right) + \\ &4 \left(\frac{\langle r_a \rangle^2}{3} \right) \left(\frac{\langle r_b \rangle^2}{3} \right) = 6 \left(\frac{\langle r_a \rangle^2}{3} \right) \left(\frac{\langle r_b \rangle^2}{3} \right) \end{aligned}$$

There are several other ways to verify this result: (1) by writing the Hamiltonian in terms of $Y_{1(a,b)}^m$, to find that only terms $Y_{1(a)}^m Y_{1(b)}^{m'}$ satisfying the conservation of angular momentum $m = -m'$ remain or (2) by choosing the orthogonal p_x, p_y, p_z basis for $l=1$ where selection rules for x, y, z are simpler.

Thus we have the result

$$\Delta E_{g_a g_b} = -\frac{e^4}{R^6} \sum_{n_a, l_a, n_b, l_b} \frac{6 \left(\frac{\langle r_a \rangle^2}{3} \frac{\langle r_b \rangle^2}{3} \right)}{(E_i^a - E_g^a) + (E_i^b - E_g^b)}$$

Now we introduce oscillator strengths: $f_{ig} \equiv \frac{2m_e \omega_{ig}}{\hbar} |\langle m_i | x | m_g \rangle|^2$. Note that the mass, m_e , is the electron mass. After all, it is the oscillator.

The choice of x, y , or z in our definition of oscillator strength represents an artificial choice of axis; we should obtain the same f_{ig} if we sum over upper m states and average over lower m states.

Using the dipole operator matrix elements derived above for $|P\rangle \rightarrow |S\rangle$ transitions, we obtain (In this case there is no need to average over lower m states because there is only one.)

$$\sum_{m_i} |\langle m_i | x | m_g \rangle|^2 = \sum_{m_i} |\langle m_i | y | m_g \rangle|^2 = \sum_{m_i} |\langle m_i | z | m_g \rangle|^2 = \frac{\langle r \rangle^2}{3}$$

Therefore,

$$f_{ig} \equiv \frac{2m_e \omega_{ig}}{\hbar} \frac{\langle r \rangle^2}{3} \Rightarrow \frac{\langle r \rangle^2}{3} = \frac{\hbar}{2m_e} \frac{f_{ig}}{\omega_{ig}}$$

Finally,

$$\begin{aligned} \Delta E_{g_a g_b} &= -\frac{3}{2} \frac{e^4}{R^6} \frac{\hbar^2}{m_e^2} \sum_{n_a, l_a, n_b, l_b} \frac{\frac{f_{ig}^a}{\omega_{ig}^a} \frac{f_{ig}^b}{\omega_{ig}^b}}{(E_i^a - E_g^a) + (E_i^b - E_g^b)} \\ &= -\frac{C_6}{R^6} \end{aligned}$$

where

$$C_6 = \frac{3e^4 \hbar}{2m_e^2} \sum_{n_a, l_a, n_b, l_b} \frac{f_{ig}^a f_{ig}^b}{\omega_{ig}^a \omega_{ig}^b (\omega_{ig}^a + \omega_{ig}^b)}$$

- (c) (3 Points) Express your result for C_6 from (b) in terms of polarizabilities $\alpha_g^{(a)}$ and $\alpha_g^{(b)}$. We will make the dominant-level approximation: assume $f_{ig}^a \approx 1$, $f_{ig}^b \approx 1$ for only one transition and drop the sums over excited states.

$$\Delta E_{g_a g_b} = -\frac{3e^4 \hbar}{2m_e^2} \frac{1}{R^6} \frac{1}{\omega_{ig}^a \omega_{ig}^b (\omega_{ig}^a + \omega_{ig}^b)}$$

Recall that the static polarizability of the ground state is defined by

$$\alpha_g \equiv 2e^2 \sum_i \frac{|\langle i|z|g \rangle|^2}{E_i - E_g} = \frac{e^2}{m_e} \sum_i \frac{f_{ig}}{\omega_{ig}^2}$$

Using the dominant level approximation $\Rightarrow \alpha_g \approx \frac{e^2}{m_e \omega_{ig}^2}$ or $\frac{1}{\omega_{ig}} \approx \frac{m_e \omega_{ig} \alpha_g}{e^2}$. As a result we obtain

$$C_6 = \frac{3}{2} \frac{\alpha_g^a \alpha_g^b E_{ig}^a E_{ig}^b}{E_{ig}^a + E_{ig}^b} = \frac{3}{2} \hbar \frac{\alpha_g^a \alpha_g^b \omega_{ig}^a \omega_{ig}^b}{\omega_{ig}^a + \omega_{ig}^b}$$

Note that this result has the correct units of [Energy] \times [Length]⁶ because polarizability has units of [Volume]= [Length]³.

[1] There is an exception: if both “atoms” possess space-fixed electric dipole or permanent electric quadrupole moments. This can occur if the atoms are in a strong external field, if the “atoms” are a special type of polyatomic molecule, or if the atoms are in $l \geq 1$ ground states.