spin by a potential  $V(\vec{x}) = -\langle \vec{n} \rangle \cdot \vec{B}(\vec{x})$  where  $\langle \vec{n} \rangle$  is the magnetic moment.

So we conclude that  $\langle \vec{n} \rangle = \frac{e}{m} \left[ F_1(0) + F_2(0) \right] \xi^{\prime \dagger} \bar{\xi} \xi$ 

It is conventional to define

where q is called the Landé g-factor.

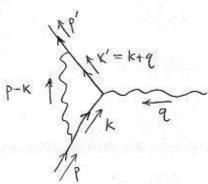
We find that  $g=2+\Theta(x)$ . The 2 was what Dirac was able to explain from his Dirac equation, and made people notice what he was doing seemed correct... hard to get g=2 from a classical picture of rotating charged object. The OO) is what Feynman, Schwinger, and Tomonaga was able to explain from QED, and made people notice that QED seemed correct despite the awful infinities.

We note that for an intrinsically composite

"bound-stude" particle like the proton, g is not close to 2 (not very close like for the electron). This is because the magnetic field "sees" the spins + masses of the constituent particles (quarks + gluons).

## One-loop magnetic moment

We calculate the one-loop correction to the magnetic moment



Let \(\tau = \gamma^m + \delta \gamma^m\)

Then  $\overline{u}(p') \delta \Gamma''(p',p) u(p')$   $= \int \frac{d^4k}{(2\pi)^4} \frac{-i g_{LB}}{(k-p)^2 + i \xi} \overline{u}(p') \left(-i e \delta''\right) \frac{i k' + m}{k'^2 - m^2 + i \xi} \delta'' \frac{i (k+m)}{k^2 - m^2 + i \xi}$ 

Using  $\gamma^{\mu}\gamma^{\nu}\gamma^{\nu}=-2\gamma^{\mu}$ ,  $\gamma^{\mu}\gamma^{\mu}\gamma^{\nu}=+6\gamma^{\mu}$ , and  $\gamma^{\mu}\gamma^{\mu}\gamma^{\nu}=-2\gamma^{\mu}\gamma^{\nu}\gamma^{\nu}$ ,

we have

$$2ie^{2}\int \frac{d^{4}k}{(2\pi)^{4}} \frac{\overline{u}(p')\left[k'y''k'+m^{2}y'''-2m(k+k')''\right]u(p)}{((k-p)^{2}+i\epsilon)(k'^{2}-m^{2}+i\epsilon)(k^{2}-m^{2}+i\epsilon)}$$

This integral looks hard. We need some integration tricks.

Identity: (Feynman's trick for combining denominators)

$$\frac{1}{A \cdot B} = \int_0^1 dx \frac{1}{[xA + (1+x)B]^2}$$

Check: For  $A \neq B$ , integral gives  $-\frac{1}{(xA+(1-x)B)(A-B)}\Big|_{0}^{1}$ =  $-\frac{1}{A-B}\Big(\frac{1}{A}-\frac{1}{B}\Big)=\frac{1}{A-B}$ 

For A = B, integral gives  $\int_0^1 dx \frac{1}{A^2} = \frac{1}{A^2} = \frac{1}{A \cdot B}$ 

## Generalizations:

$$\frac{1}{A_1A_2\cdots A_n} = \int_0^1 dx_1 \cdots dx_n \frac{S(Z_ix_i-1)\cdot (n-i)!}{[x_1A_1+x_2A_2+\cdots x_nA_n]^n}$$

By differentiating this repeatedly we get

$$\frac{1}{A_{i}^{m_{1}}A_{i}^{m_{2}}-A_{i}^{m_{n}}} = \int_{0}^{1} dx_{1}-dx_{n} \frac{\delta(z_{i}x_{i}-1) \cdot T_{i}x_{i}^{m_{i}-1} \Gamma(m_{1}+\cdots+m_{n})}{\sum_{i} x_{i}A_{i}^{2} \Gamma(m_{i})\Gamma(m_{2})-\Gamma(m_{n})}$$

This is time even for non-indeper mi's

In our case we have

$$\frac{1}{((K-p)^{2}+i\epsilon)(K^{2}-m^{2}+i\epsilon)(K^{2}-m^{2}+i\epsilon)} = \int_{0}^{1} \frac{dx\,dy\,dz}{dx\,dy\,dz} \frac{\delta(x+y+z-1)\cdot\Gamma(3)}{D^{3}}$$

where 
$$D = \chi(k^2-m^2) + y((k+q)^2-m^2) + z(k-p)^2 + iz$$
  
 $= \cdot k^2 + 2k \cdot (yq-zp) + yq^2 + zp^2 - (x+y)m^2 + iz$ 

We now complete the square ...

$$D = l^2 - \Delta + iz$$

$$\frac{z^2m^2 - 2m^2}{\sqrt{2}}$$
where  $l = k + yq - 2p$  and  $\Delta = + y^2q^2 - yq^2 + z^2p^2 - zp^2$ 

$$-zy \neq q \cdot p + (x+y)m^2$$

Since 
$$p^{2} = (q+p)^{2} = q^{2}+p^{2}+2q\cdot p$$
,  $m^{2} = q^{2}+m^{2}+2q\cdot p$   
and so  $2q\cdot p = -q^{2}$ .  
We can write  $\Delta = (q^{2}-y+yz)q^{2}+(z^{2}-z+x+y)m^{2}$   
 $= -xyq^{2}+(1-z)^{2}m^{2}$ 

Since 
$$q^2 = (p'-p)^2 = m^2 + m^2 - 2p'p$$
  
=  $2m^2 - 2EE + 2p'p$   
 $\leq 2m^2 - 2\sqrt{p^2 + m^2}\sqrt{p^2 + m^2} + 2\sqrt{p}\sqrt{p}$   
 $\leq 0$ 

· proof: let  $V_1 = {m \choose |\vec{p}|}$   $V_2 = {m \choose |\vec{p}|}$ then  $V_1 \cdot V_2 = m^2 + |\vec{p}'||\vec{p}|$ , and  $|V_1||V_2| = |m^2 + \vec{p}'^2||m^2 + \vec{p}|^2$ Therefore  $q^2 \le 0$ 

Since  $q^2 \le 0$ ,  $\Delta = -xy q^2 + (1-z)^2 m^2 > 0$ Let us work on the numerator of the integral. Since D is even in l, and only a function of  $l^2$ ,

$$\int \frac{d^4 l}{(2\pi)^4} \frac{l^7}{D^3} = 0$$

$$\int \frac{d^4 l}{(2\pi)^4} \frac{l^m l^n}{D^3} = C \int \frac{d^4 l}{(2\pi)^4} \frac{g^{mr} l^2}{D^3}$$

Contracting  $\mu$  and  $\nu$ , we find a factor of  $\frac{1}{4}$ ,  $\int \frac{d^4l}{l^2l^3} \frac{l^4l^4}{l^3} = \frac{1}{4} \int \frac{d^4l}{(2\pi)^4} \frac{g^{\mu\nu}l^2}{D^3}$ 

Our numerator was

$$\begin{array}{ll} \overline{u(p')} \left[ \frac{1}{K} \frac{3^m K' + m^2 }{3^m} - 2m \left( \frac{1}{K+K'} \right)^m \right) u(p) \\ \overline{u(p')} \left[ \frac{1}{K} \frac{3^m K' + m^2 }{3^m} \frac{3^m K' + (1-y) 9 \left( + 2p \right) + m^2 }{3^m} \frac{3^m}{2m \left( (1-y) 9 \left( + 2p \right) \right)} \right] u(p) \\ \overline{u(p')} \left[ \frac{1}{K} \frac{3^m K' + m^2 }{3^m} \frac{3^m}{2m \left( (1-y) 9 \left( + 2p \right) \right)} \right] u(p) \\ \overline{u(p')} \left[ \frac{1}{K} \frac{3^m K' + m^2 }{3^m} \frac{3^m}{2m \left( (1-y) 9 \left( + 2p \right) \right)} \right] u(p) \\ \overline{u(p')} \left[ \frac{1}{K} \frac{3^m K' + m^2 }{3^m} \frac{3^m K$$

Take the term ...

$$= \overline{U(p')} \left[ + y \times m^2 \, \chi''' + \frac{4y \times m^2 \, \chi''' (1-y)}{2m} + \frac{4y \times m^2 \, \chi'''' (1-y)}{2m} + \frac{4y \times m^2 \, \chi''''' (1-y)}{2m} + \frac{4y \times m^2 \, \chi''''' (1-y)}{2m} + \frac{4y \times m^2 \, \chi''''' (1-y)}$$

a little more work gives

Notice that the  $g(^m m(z-z)(x-y))$  term vanishes since the integral is symmetric under  $x \Longrightarrow y$  exchange

We can use the Gordon identity  $\overline{u(p')}8^mu(p) = \overline{u(p')}\left(\frac{p'''+p''}{2m} + \frac{i\sigma''''q''}{2m}\right)u(p)$ to put the p'''+p''' term in terms of  $8^m$  and  $i\sigma'''q''$ So we now have

$$= 2ie^{2} \int \frac{d^{4}l}{(2\pi)^{4}} \int_{0}^{1} \frac{dx \, dy \, dz}{D^{3}} \frac{S(x+y+z-1)}{\sum_{i=1}^{2} + (i-x)(i-y)q^{2}} + (1-4z+z^{2})m^{2}$$

$$+ i\frac{g^{2}q}{2m} (2m^{2}\overline{z}(1-\overline{z})) \int u(p)$$

$$\overline{u}(p') SP''(p',p) u(p)$$

$$= 2ie^2 \int \frac{d^4l}{(2\pi)^4} \int_0^1 dxdydx \frac{S(x+y+2-1)\cdot 2\cdot N}{D^3}$$

$$D = l^2 - \Delta + ix$$
where we showed  $\Delta > 0$ 

The l'integral has poles at  $\pm (\sqrt{12+\Delta} - i\delta)$ 

Let  $l^0 = i l_E^0$ ,  $l^{1,2,3}_{E} l_E^{1,2,3}$  E stands for Euclidean

We integrate  $l_{\rm E}^{\circ}$  from  $-\infty$  to  $\infty$ . Then

$$\int \frac{d^4 l}{(2\pi)^4} \frac{1}{[l^2 - \Delta + i\epsilon]^m} = \int \frac{id l_e^0 d^3 l_e}{(2\pi)^4} \frac{1}{[-l_e^2 - \Delta)^m}$$

$$= \frac{i(-1)^m}{(2\pi)^4} \int d\Omega_4 \int_0^\infty dl_e \frac{l_e^3}{[l_e^2 + \Delta]^m}$$

What is 
$$\int d\Omega_4 = ?$$
  
We know  $\int d\theta = 2\pi = \int d\Omega_2$   
 $\int d\phi \int_1^1 d\cos\theta = 4\pi = \int d\Omega_3$ 

Consider the Gaussian integral  $\int_{-\infty}^{\infty} - \int_{-\infty}^{\infty} dl_{\varepsilon}^{2} - dl_{\varepsilon}^{2} = \left(l_{\varepsilon}^{2}\right)^{2} = \left(l_{\varepsilon}^{2}\right)^{2} + \left(l_{\varepsilon}^{2}\right)^{2} = \int_{-\infty}^{\infty} d\Omega_{+} \int_{-\infty}^{\infty} d\Omega_$ 

S.  $\int dx_4 = 2\pi^2$ .

Now we think about the dle integration. In the denominator we have  $(l_{\epsilon}^2 + \Delta)^3$ . In the numerator we have  $\int_{\epsilon}^{3} l_{\epsilon}^3 dl_{\epsilon}$ . (something  $l_{\epsilon}^2 + \cdots$ ).

Notice that this integral diverges as  $l \in \to \infty$ . It is a logarithmic divergence since if we cutoff the integral at M, we get  $\sim \log M$  as  $M \to \infty$ .

We will hope that this divergence can be absorbed into an unknown constant (which is divergent), but then everything comes out convergent thereafter.

First we need to regulate the ultraviolet divergence by some means. The following is one type of regularization method called Pauli-Villars.

Idea: Modify the photon propagator ...

$$\frac{-i g_{\mu\nu}}{(K-p)^2 + i\epsilon} - \frac{-i g_{\mu\nu}}{(K-p)^2 - \Lambda^2 + i\epsilon}$$
 we then take the limit  $\Lambda \to \infty$ 

So our  $\triangle$  in the denominator becomes, for the massive photon,  $\triangle_{\Lambda} = - \times y q^2 + (1-z)^2 m^2 + z \Lambda^2$ Now we can do the integral.

$$\int \frac{d^4l}{(2\pi)^4} \left[ \frac{l^2}{(l^2 - \Delta)^3} - \frac{l^2}{(l^2 - \Delta)^3} \right] \sim \log \left( \frac{A^2}{\Delta} \right) \quad \text{as } \Lambda \to \infty$$

In any case the  $F_2(q^2)$  part is finite.  $\int \frac{d^4\ell}{(2\pi)^4} \frac{1}{(\ell^2 \Delta + i\epsilon)^3} = \frac{i \times (-1)}{(4\pi)^2} \cdot \frac{1}{2\Delta}$ 

Tu(p) SPM up) = \( \frac{\pi}{2\pi} \int \frac{1}{2} \dxdyd\(\frac{2}{2} \sum \frac{1}{2} \dxdyd\(\frac{2}{2} \sum \frac{1}{2} \dxdyd\(\frac{2}{2} \dxdyd\(\frac{2} \dxdyd\(\frac{2}{2} \dxdyd\(\frac{2}{2} \dxdyd\(\frac{2}{2} \d

This gives  $F_{2}(0) = \frac{1}{2\pi} \int_{0}^{1} dx dy dz S(x+y+7-1) \frac{z}{1-z}$ =  $\frac{1}{\pi} \int_{0}^{1} dz \int_{0}^{1-z} dy \frac{z}{1-z} = \frac{x}{2\pi}$ 

 $S_0$   $9^{-2}_2 = F_2(0) = \frac{\alpha}{2\pi} \approx 0.0011614$ 

first computed by Schninger in 1948

Experimentally F2(0) = 0.0011597