Adding interactions

To preserve causality, we consider only local interactions. For example,

Hint =
$$\int d^3x' \mathcal{H}(\phi(x)) = -\int d^3x' \mathcal{L}_{ind}(\phi(x))$$

function of
fields at the
same point

Common example in particle + condensed matter physics,

$$L_{\rm int} = -\frac{\lambda}{4!} \phi^{\dagger}$$

So $Z = \frac{1}{2} (\partial_{\mu} \phi) (\partial^{\mu} \phi) - \frac{1}{2} m^{2} \phi^{2} - \frac{\lambda}{4!} \phi^{4}$

Note that TICK is still 20 pcx since there are no interactions involving 20 p.

In general adding interactions will produce divergences.

These are ultraviolet (high momentum) divergences which signal that this is not the fundamental

theory at arbitrarily short distances. (could be string theory or something else).

However no matter what the true physics looks like at high momenta or short distances, the low momentum or long distance physics is well approximated by an "effective" field theory with only "renormatizable" interactions.

These are interactions where the coupling constant has dimensions [Mass] where $d \ge 0$.

Example: In 3 space + 1 time dimensions (3+1) it turns out that $\phi(x)$ has dimensions (Mass)!. If I compare $-\frac{1}{2}m^2\phi^2$ and $-\frac{\lambda}{4!}\phi^4$... both must have the same mass dimension. So

7 ~ [Mass),

and this is renormalizable.

On the other hand, something like $-\frac{\lambda_6}{6!}\phi^6$ would give

which is not renormalizable.

Perturbation Expansion

Let $H = H_0 + H_{int} \iff \text{for example,}$ $H_{int} = \int d^3x \frac{2}{4!} \phi^4(x)$ free Klein-Gordon, which we have be studying

We will generate a power series in I. At any fixed time to, we can write

 $\phi(to,\vec{x}) = \int \frac{d^3\vec{p}}{(2\pi)^3} \sqrt{2\vec{e}_{\vec{p}}} \left(a_{\vec{p}} e^{i\vec{p}\cdot\vec{x}} + a_{\vec{p}}^{\dagger} e^{-i\vec{p}\cdot\vec{x}} \right)$ we can absorb the $e^{-i\vec{e}_{\vec{p}}t_o} \text{ in the definition}$ of $a_{\vec{p}}$

The Heisenberg field $\phi(t,\vec{x})$ is then given by $\phi(t,\vec{x}) = e^{i \cdot H \cdot tt - t_0} \phi(t_0,\vec{x}) e^{-i \cdot H \cdot tt - t_0}$

If we shut off the interaction we have for the Hemiltonian —i Hott-to) $\phi(t_0, \hat{x}) \in H_0(t-t_0)$

= $\int \frac{d^3\vec{p}}{(2\pi)^3\sqrt{2}\vec{e}_{\vec{p}}} (a_{\vec{p}}e^{-i\vec{p}\cdot\vec{x}} + a_{\vec{p}}^{\dagger}e^{i\vec{p}\cdot\vec{x}}) \Big|_{X_{=}^{e}}$

... we define this to be $\phi(t,\vec{x})$, the interaction picture field.

This interaction picture field coincides with the Heisenberg field when $\lambda = 0$.

The Heisenberg field for $\lambda \neq 0$ is $\phi(t, \vec{x}) = e^{i \cdot H \cdot tt - t_0} e^{-i \cdot H \cdot tt - t_0}$ $= e^{i \cdot H \cdot tt - t_0} e^{-i \cdot H_0 \cdot tt - t_0} \phi_{I}(t, \vec{x}) e^{i \cdot H_0 \cdot (t - t_0)} e^{-i \cdot H_0 \cdot tt - t_0}$ $= U^{\dagger}(t, t_0) \phi_{I}(t, \vec{x}) U(t, t_0)$ where $U(t, t_0) = e^{i \cdot H_0 \cdot tt - t_0} e^{-i \cdot H_0 \cdot tt - t_0}$ we evolve the operators as $\phi_{I}(t, \vec{x})$ we evolve the states by $U(t, t_0)$...

Utt, to) lv>

Note:

$$i\frac{\partial}{\partial t}$$
 Utt, to) = $e^{iH_0(t-t_0)}$ (- H_0+H) $e^{-iH(t-t_0)}$
= $e^{iH_0(t-t_0)}$ Hint $e^{-iH_0(t-t_0)}$ $e^{iH_0(t-t_0)}$ $e^{-iH(t-t_0)}$
= $e^{iH_0(t-t_0)}$ Hint $e^{-iH_0(t-t_0)}$ $e^{iH_0(t-t_0)}$ $e^{-iH(t-t_0)}$
Hint ($\phi_{\mathbf{I}}(t,\mathbf{x})$) Utt, to)
= $H_{\mathbf{I}}(t,\mathbf{x})$ Utt, to)
 $H_{\mathbf{I}}(t,\mathbf{x})$ $H_{\mathbf{I}}(t,\mathbf{x})$ Dyson's formula
time ordering symbol

Why time ordering? Because $H_{\rm I}(t_i) + H_{\rm I}(t_z)$ don't commute for different times $t_i + t_z$.

The time ordering puts the latest operators on the left. That's why the $H_{\rm I}(t)$ is on the left. $i \stackrel{>}{\Rightarrow} U(t,t_0) = H_{\rm I}(t)U(t,t_0)$

As a power series in λ : $U(tt,t_0) = 1 + (-i) \int_{t_0}^{t} dt' H_{\mathbf{I}}(t') + \frac{(-i)^2}{2!} \int_{t_0}^{t} dt_1 dt_2 \{H_{\mathbf{I}}(t_1) H_{\mathbf{I}}(t_2)\} + \cdots$

Let us now generalize U...

Define Uttst's = Texp {-i [t', dt" H_I tt")}
for any t > t'

Then $i \frac{\partial}{\partial t} \mathcal{V}(t,t') = \mathcal{H}_{\mathbf{I}}(t) \mathcal{V}(t,t')$ $i \frac{\partial}{\partial t} \mathcal{V}(t,t') = -\mathcal{V}(t,t') \mathcal{H}_{\mathbf{I}}(t')$

will work ...

i = Ut,t') = e i Holt-to) (-Ho+H) e i Ht-to) -i Holt-to)

= H_Itt) Ut,t')

i = e i Holt-to) e -i H (t-t') (+H + Ho) e -i Holt-to)

= - Ult,t') H_T(t')

Note that U(t,t') is unitary, $U^{\dagger}(t,t') = U^{-1}(t,t')$

Also for tixtzxt3,

Ulti, t2) Ulti, t3) = Ulti, t3)

Let 10> be the ground state of Ho

Let 12> be the ground state of H.

Let In> label all the energy states of H (n=0 corresponds with U2>)

Let the corresponding energies be En.

Then e-iHT 10> =

= iE.T | 12><210> + = e -iEnT | 1N><10>

We will zero out Ho so that Holo> = 0.

Now let us consider the limit as T→∞.

Subtle point ... in stend of considering small positive T >+00, we will consider T > (1-ix).00 (what this means is that we use analytic condinuation to get the physical result)

Assuming there is a gap between $E_0 + other$ $E_{n's}$, then e^{-iE_nT} dies slowest for N=0.

As
$$T \rightarrow (1-i\epsilon) \cdot \infty$$
,
 $e^{-iHT}|_{0} \rightarrow e^{-iE \cdot T}|_{\Omega} \times \Omega|_{0}$.
So $12 > = \lim_{T \rightarrow (1-i\epsilon) \cdot \infty} (e^{-iE \cdot T} \times \Omega|_{0})^{T} e^{-iHT}|_{0} > 0$.
Also

12> = Lim (e = (T+to) < 2/0>) = i H(T+to) | 0> = lim (e-iE.(t.-(-T)) / e-iH(t.-(-T)) iH.(-T-t)) (0)

= lim (e (File(to-(-T))) (to,-T) 10> T>(1-ie)00

Also we have

$$\langle \Omega \rangle = \lim_{T \to (1-i\epsilon)\infty} (e^{-iE_0(T-t_0)} < 0|\Omega\rangle)^{-1} < 0|e^{-iH_0(T-t_0)}$$

 $= \lim_{T \to (1-i\epsilon)\infty} (e^{-iE_0(T-t_0)} < 0|\Omega\rangle)^{-1} < 0|e^{-iH_0(T-t_0)} = e^{-iH_0(T-t_0)}$

For xo>y°>to, we then have

<2/ dex) pry) 12>

= lim T=(1-is) (1<01,2>|2 e-iE.2T)" × <01 U(T, x°) \(\dag{\super} \) U(x°, y°) \(\dag{\super} \) U(y°, -T) \(\operatorname{\super} \)

We note that

$$\langle \Omega | \Omega \rangle = \lim_{T \to (1-i\epsilon)\infty} \left[(|\langle o|\Omega \rangle|^2 e^{-i\overline{E}_0 2T})^{-1} \right]$$

So it 12> normalized,

1<01 2>12 e-iE2T → <01 U(T,-T)10>

Therefore

< 21 0 cm pry 12>

= lm <0 | U(T, x°) \$\psi_{\text{L}}(x) U(x°, y°) \$\phi_{\text{L}}(y) U(y°, -T) /0> \tag{0 | U(T, -T) | 0>}

(x°>y°)

For x°< y° then roles of x+y on the right hand side are reversed. So

\[
 \left(\tau \phi \tau \right) \right) \right(\tau \right) \right) \right) \right(\tau \right) \right) \right) \right(\tau \right) \right