

March

14

Sunday
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2021/02/07

7:00 UNIVERSE ACROSS SCALES - ASSIGNMENT 2

10:00 $\text{Wavefunction}, \psi(n, 0) = \begin{cases} An/a, & 0 \leq n \leq a \\ A(b-n)/(b-a), & a \leq n \leq b \\ 0, & \text{otherwise} \end{cases}$

10:00 $(a) \text{ Normalising : } \int_{-\infty}^{\infty} |\psi(n, 0)|^2 dn = 1$

11:00 $\int_a^b |\psi(n, 0)|^2 dn + \int_a^b \left| \frac{A(b-n)}{(b-a)} \right|^2 dn = 1$

12:00 $\int_0^a \frac{A^2 n^2}{a^2} dn + \int_a^b \frac{A^2 (b-n)^2}{(b-a)^2} dn = 1$

13:00 $\frac{A^2}{a^2} \int_0^a n^2 dn + \frac{A^2}{(b-a)^2} \int_a^b (b-n)^2 dn = 1$

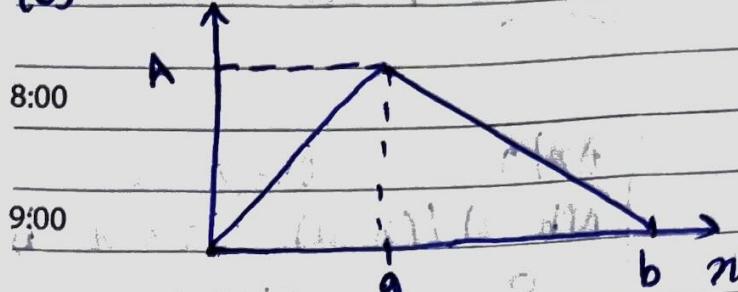
14:00 $\frac{A^2}{a^2} \left[\frac{n^3}{3} \right]_0^a + \frac{A^2}{(b-a)^2} \left[\frac{-(b-n)^3}{3} \right]_a^b = 1$

15:00 $\frac{A^2}{a^2} (a^3 - 0) + \frac{A^2}{(b-a)^2} \times \frac{(b-a)^3}{3} = 1$

16:00 $\frac{a^2 A^2}{3} + \frac{A^2 (b-a)}{3} = 1$

17:00 $\frac{b}{3} A^2 = 1 \Rightarrow A^2 = 3/b \Rightarrow A = \sqrt{3/b}$

789 $\psi(n, 0)$



(c) From the above graph, we can see that the particle is most likely to be found at $n=a$.

(d) Probability of finding particle b/w a and b is

$$\int_a^b |\psi(n, 0)|^2 dn$$

$$P(-\infty < n \leq a) = \int_{-\infty}^a |\psi(n, 0)|^2 dn + \int_0^a |\psi(n, 0)|^2 dn$$

$$= 0 + \int_0^a \frac{A^2 n^2}{a^2} dn$$

$$= \frac{2A^2}{a^2} \left[\frac{n^3}{3} \right]_0^a$$

$$= \frac{A^2}{3a^2} (a^3)^0$$

19:00

$$= \frac{a^2}{3} A^2$$

20:00

7:00 We know, $N = \sqrt{\frac{3}{b}}$

8:00 $P(-\infty < n \leq a) = \frac{a}{\sqrt{b}}$

9:00 $\frac{a^2}{b} \parallel (b \neq 1)$

10:00

CASE 1: $b = a$

11:00 $P(-\infty < n \leq a) = \frac{a}{a} = 1$

12:00

CASE 2: $b \neq 2a$

13:00 $P(-\infty < n \leq a) = \frac{a}{2a} = \frac{1}{2}$

14:00

(2) The Schrodinger equation is given by :-

15:00

$$\frac{\partial \psi}{\partial t} = i\hbar \frac{\partial^2 \psi}{\partial n^2} - \frac{i}{\hbar} v(n, t) \psi(n, t)$$

16:00

17:00 Let us now take the wave function $\psi(n, t)$

to have a potential energy of $v(n, t) + V_0$

18:00

$$\frac{\partial \psi}{\partial t} = i\hbar \frac{\partial^2 \psi}{\partial n^2} - \frac{i}{\hbar} [v(n, t) + V_0] \psi(n, t)$$

19:00

$$= i\hbar \frac{\partial^2 \psi}{\partial n^2} - \frac{i}{\hbar} v(n, t) \psi(n, t) - \frac{i}{\hbar} V_0 \psi(n, t)$$

20:00

$$\frac{\partial \psi}{\partial t} + \frac{i}{\hbar} V_0 \psi(n,t) = \frac{i\hbar}{\partial n} \frac{\partial^2 \psi}{\partial n^2} - \frac{i}{\hbar} V(n,t) \psi(n,t)$$

8:00

This is a differential equation, so the integrating factor

$$I = \exp \left(\int \frac{iV_0}{\hbar} dt \right)$$

10:00

$$= e^{iV_0 t / \hbar}$$

11:00

Multiplying both sides of the DE by I:-

$$I \frac{d\psi}{dt} + \frac{i}{\hbar} V_0 \psi(n,t) = \frac{i\hbar}{\partial n} \frac{\partial^2 \psi}{\partial n^2} - \frac{i}{\hbar} V(n,t) \psi(n,t)$$

13:00

$$\frac{d}{dt} \left(e^{iV_0 t / \hbar} \psi \right) + \frac{i}{\hbar} V_0 \psi(n,t) = e^{iV_0 t / \hbar} \left(\frac{i\hbar}{\partial n} \frac{\partial^2 \psi}{\partial n^2} - \frac{i}{\hbar} V(n,t) \psi(n,t) \right)$$

15:00

The LHS can be seen as a case of product rule for differentiation.

17:00

$$\therefore \frac{d}{dt} \left(e^{iV_0 t / \hbar} \psi \right) = e^{iV_0 t / \hbar} \frac{i\hbar}{\partial n} \frac{\partial^2 \psi}{\partial n^2} - e^{iV_0 t / \hbar} \frac{i}{\hbar} V(n,t) \psi(n,t)$$

18:00

$$\text{Bringing } e^{iV_0 t / \hbar} \text{ into the partial derivative,}$$

$$\frac{d}{dt} \left(e^{iV_0 t / \hbar} \psi \right) = \frac{i\hbar}{\partial n} \frac{\partial^2}{\partial n^2} \left(e^{iV_0 t / \hbar} \psi \right) - \frac{i}{\hbar} V(n,t) e^{iV_0 t / \hbar} \psi$$

February

March

7: Comparing this to Schrödinger's equation,
 $\phi = \psi(n,t) e^{iV_0 t/\hbar}$

8:00

$$\Rightarrow \psi(n,t) = \phi(n,t) e^{-iV_0 t/\hbar}$$

9:00

Hence, the wave function picked up a time
10: dependent $[e^{-iV_0 t/\hbar}]$ factor.

11:00

$$③ \frac{d}{dt} \int_{-\infty}^{\infty} \psi_1^* \psi_2 dn = 0 \quad [\text{TO BE PROVEN}]$$

12:00

$$\frac{d}{dt} \int_{-\infty}^{\infty} \psi_1^* \psi_2 dn = \int_{-\infty}^{\infty} \frac{d}{dt} (\psi_1^* \psi_2) dn$$

13:00

$$= \int_{-\infty}^{\infty} \left(\psi_1^* \frac{\partial \psi_2}{\partial t} + \psi_2 \frac{\partial \psi_1^*}{\partial t} \right) dn$$

14:00

$$= \int_{-\infty}^{\infty} \left(\psi_1^* \frac{\partial \psi_2}{\partial t} - \psi_2 \frac{\partial \psi_1^*}{\partial t} \right) dn$$

15:00

$$= \int_{-\infty}^{\infty} \left(\psi_1^* \frac{\partial \psi_2}{\partial t} - \psi_2 \frac{\partial \psi_1^*}{\partial t} \right) dn$$

16:00

$$\therefore \frac{\partial \psi_2}{\partial t} = \frac{i\hbar}{2m} \frac{\partial^2 \psi_2}{\partial n^2} - \frac{i}{\hbar} V \psi_2$$

17:00

$$\frac{\partial \psi_1^*}{\partial t} = -\frac{i\hbar}{2m} \frac{\partial^2 \psi_1^*}{\partial t^2} + \frac{i}{\hbar} V \psi_1^*$$

18:00

$$7:00 \int_{-\infty}^{\infty} \left[\Psi_1^* \frac{\partial \Psi_2}{\partial t} + \frac{\partial \Psi_1^*}{\partial t} \Psi_2 \right] dn =$$

$$8:00 \int_{-\infty}^{\infty} \Psi_1^* \left[i\hbar \frac{\partial^2}{\partial n^2} \Psi_2 - \frac{iV}{\hbar} \Psi_2 \right] + \Psi_2 \left[-i\hbar \frac{\partial^2}{\partial n^2} \Psi_1^* + \frac{iV}{\hbar} \Psi_1^* \right] dn$$

$$11:00 \int_{-\infty}^{\infty} \frac{i\hbar}{\alpha m} \left[\Psi_1^* \frac{\partial^2 \Psi_2}{\partial n^2} - \Psi_2 \frac{\partial^2 \Psi_1^*}{\partial n^2} \right] - \frac{iV \Psi_1^* \Psi_2}{\hbar} + \frac{iV \Psi_1^* \Psi_2}{\hbar}$$

$$13:00 = \int_{-\infty}^{\infty} \frac{i\hbar}{\alpha m} \left[\Psi_1^* \frac{\partial^2 \Psi_2}{\partial n^2} - \Psi_2 \frac{\partial^2 \Psi_1^*}{\partial n^2} \right]$$

$$14:00 = \int_{-\infty}^{\infty} \frac{i\hbar}{\alpha m} \left[\Psi_1^* \frac{\partial^2 \Psi_2}{\partial n^2} \Big|_{-\infty}^{\infty} - \frac{\partial \Psi_1^*}{\partial n} \frac{\partial \Psi_2}{\partial n} \right]$$

$$15:00 - \left[\Psi_2 \frac{\partial^2 \Psi_1^*}{\partial n^2} \Big|_{-\infty}^{\infty} + \frac{\partial \Psi_2}{\partial n} \frac{\partial \Psi_1^*}{\partial n} \right]$$

17:00 0

$$18:00 \text{ because } \left(\Psi_1^* \frac{\partial^2 \Psi_2}{\partial n^2} \Big|_{-\infty}^{\infty} \right) = \left(\Psi_2 \frac{\partial^2 \Psi_1^*}{\partial n^2} \Big|_{-\infty}^{\infty} \right) = 0$$

19:00 (Solutions are orthogonal to each other)

20:00

$$7:00 \quad \psi(n, 0) = \begin{cases} A(a^2 - n^2) & , -a \leq n \leq a \\ 0 & , \text{otherwise} \end{cases}$$

8:00

$$9:00 \quad \text{To normalize} \quad \int_{-\infty}^{\infty} |\psi(n, 0)|^2 dn = 1$$

10:00

$$A^2 (a^2 - n^2)^2 dn = 1$$

11:00

12:00

13:00

a

$$2A^2 \int_0^a (a^4 + n^4 - 2a^2 n^2) dn = 1 \quad [\text{changing limits from } (-a, a) \text{ to } (0, a)]$$

14:00

15:00

16:00

17:00

$$2A^2 \left[a^4 n + \frac{n^5}{5} - \frac{2a^2 n^3}{3} \right]_0^a = 1$$

$$2A^2 \left[a^5 + \frac{a^5}{5} - \frac{2a^5}{3} - 0 \right] = 1$$

18:00

19:00

20:00

21:00

22:00

23:00

7:00 $A^2 = \frac{15}{2} \times \frac{1}{895}$

8:00 $= \frac{15}{1695}$

9:00

$\Rightarrow A = \frac{1}{4} \sqrt{\frac{15}{895}}$

11:00

(b) ∞

12:00 $\sigma = \int_{-\infty}^{\infty} n |(\psi(n, 0)|^2 dn (c_1 + b - c_K + c_2))$

13:00 $-\infty$

14:00 $= \int_{-a}^a n |A(a^2 - n^2)|^2 dn$

15:00 $-a$

16:00 $= \int_{-a}^a A^2 n (a^2 - n^2)^2 dn$

17:00 $-a$

18:00 $= A^2 \int_{-a}^a n (a^2 - n^2)^2 dn$

now, $f(n) = n (a^2 - n^2)^2$

20:00 $f(-n) = -n (a^2 - n^2)^2 = -f(n)$

7:00 Thus, $f(n)$ is an odd function and the integration
8:00 of an odd integral over a symmetric interval is
zero.

$$9:00 \angle n > = 0$$

10:00

$$\int_a^b (x - b + a) dx$$

(c)

$$11:00 \angle p > = \angle M V >$$

$$= m \angle v >$$

12:00

$$= m \angle v > m \frac{d}{dt} \angle n > = 90$$

13:00

(d)

$$14:00 n^2 > = \int_{-a}^{\infty} n^2 |\psi(n, 0)|^2 dn$$

15:00

$$= A^2 \int_{-a}^a n^2 A^2 (a^2 - n^2)^2 dn$$

16:00

$$-a \quad a$$

$$17:00 = A^2 \int_{-a}^a n^2 (a^2 - n^2)^2 dn$$

18:00

$$= A^2 a \int_{-a}^a (a^4 n^2 + n^6 - 2a^2 n^4) dn$$

19:00

$$-a$$

$$20:00 = A^2 \left[\frac{a^4 n^3}{3} + \frac{n^7}{7} - \frac{2a^2 n^5}{5} \right]_{-a}^a$$

7:00

$$= \alpha A^2 \left[\frac{a^4 n^3}{3} + \frac{n^7}{7} - \frac{2a^2 n^5}{5} \right] \Big|_0^a$$

8:00

9:00

$$= 2 \times \frac{1}{16} \times \frac{15}{a^4} \left(\frac{a^7}{3} + \frac{a^7}{7} - \frac{2a^7}{5} \right)$$

10:00

$$= \frac{15}{8a^4} \left(\frac{35 + 15 - 42}{105} \right) a^7$$

11:00

12:00

$$= \frac{15}{8a^4} \times \frac{a^7}{105} \times 8$$

13:00

$$= \frac{1}{675} a^7 = \frac{a^7}{a^5 7}$$

14:00

$$= \underline{\underline{a^2/7}}$$

$$\text{pe) } \langle p^2 \rangle = \int_{-\infty}^{\infty} \Psi^* \left(-\hbar^2 \frac{d^2}{dn^2} \right) \Psi \, dn \quad [\text{using operators}]$$

8:00

$$= - \int_{-a}^a A(a^2 - n^2) \hbar^2 \frac{d^2}{dn^2} (A(a^2 - n^2)) \, dn$$

10:00

$$= - A^2 \hbar^2 \int_{-a}^a (a^2 - n^2) (-2) \, dn$$

12:00

$$= 2A^2 \hbar^2 \int_{-a}^a (a^2 - n^2) \, dn$$

13:00

$$= 2A^2 \hbar^2 \left[a^2 n - \frac{n^3}{3} \right]_{-a}^a$$

15:00

$$= 4A^2 \hbar^2 \left[a^2 n - \frac{n^3}{3} \right]_{-a}^a$$

16:00

$$= 4A^2 \hbar^2 \left(a^3 - \frac{a^3}{3} \right)$$

18:00

$$= \frac{2a^3}{3} \times 4A^2 \hbar^2$$

19:00

$$= \frac{2a^3}{3} \times \pi \hbar^2 \times \frac{1}{16\pi^2} \times \frac{15}{a^5} = \frac{5\hbar^2}{2a^2}$$

20:00

7) (P) The uncertainty in n is

$$\sigma_n = \sqrt{\langle n^2 \rangle - (\langle n \rangle)^2}$$

8:00

$$= \sqrt{\frac{a^2}{7} - 0}$$

9:00

$$= \frac{a}{\sqrt{7}} / 1$$

10:00

(g) The uncertainty in p is

$$\sigma_p = \sqrt{\langle p^2 \rangle - (\langle p \rangle)^2}$$

$$= \sqrt{\frac{5\hbar^2}{2} - 0}$$

14:00

$$= \sqrt{\frac{5}{2}} \cdot \frac{\hbar}{a}$$

(h) The uncertainty principle states : $\sigma_n \sigma_p \geq \frac{\hbar}{2}$

$$\sigma_p = \sqrt{\frac{5}{2}} \cdot \frac{\hbar}{a} \times \frac{a}{\sqrt{7}}$$

18:00

$$= \sqrt{\frac{5}{14}} \cdot \frac{\hbar}{2}$$

19:00

$$= \sqrt{\frac{5}{14}} \cdot \frac{\hbar}{2} \geq \frac{\hbar}{2}$$

2

Thus, the uncertainty principle is satisfied.

The time independent Schrodinger equation is

$$-\frac{\hbar^2}{2m} \frac{d^2 \psi(n)}{dx^2} = E \psi(n)$$

$$\frac{d^2 \psi(n)}{dx^2} = -\frac{2m}{\hbar^2} E \psi(n)$$

11:00

For infinite square well, $\psi(0) = \psi(a) = 0$

$$\therefore E=0 \Rightarrow \frac{d^2 \psi(n)}{dx^2} = 0 \quad [V=0 \text{ inside well}]$$

Thus, $\psi(n)$ will be a linear function of the form $An + B$.

$$\text{now, } \psi(0) = B = 0$$

$$\begin{aligned} \psi(a) &= Aa = 0 \\ \Rightarrow A &= 0 \end{aligned}$$

Thus, $\psi = 0$

There is no acceptable solution for the time-independent Schrodinger equation for $E=0$ for the infinite well.

20:00

14:00

⑥ $\psi(n, 0) = \begin{cases} A, & 0 \leq n \leq a/2 \\ 0, & a/2 \leq n \leq a \end{cases}$

To normalize, $\int_{-\infty}^{\infty} |\psi(n, 0)|^2 = 1$

15:00

a/2

16:00 $\int_0^{a/2} A^2 dn + \int_{a/2}^a 0 dn = 1$

17:00

$$A^2 \left[n \right]_0^{a/2} = 1$$

18:00

$$A^2 \left[\frac{a}{2} \right] = 1$$

19:00

$$A = \sqrt{\frac{2}{a}}$$

20:00

We know any wave function $\psi(n, 0) = \sum_{n=1}^{\infty} c_n \psi_n(x)$

8:00

Using Fourier Series formula,

$$c_n = \sqrt{\frac{2}{a}} \int_0^{a/2} \sin\left(\frac{n\pi n}{a}\right) \psi(n, 0) dn$$

$$= \sqrt{\frac{2}{a}} \int_0^{a/2} A \sin\left(\frac{n\pi n}{a}\right) dn$$

$$= A \sqrt{\frac{2}{a}} \int_0^{a/2} \sin\left(\frac{n\pi n}{a}\right) dn$$

$$= \sqrt{\frac{2}{a}} \cdot \sqrt{\frac{2}{a}} \left[-\cos\left(\frac{n\pi n}{a}\right) \right]_0^{a/2} \times \left(\frac{n\pi}{a}\right)^{-1}$$

16:00

$$\text{For } n=1, c_1 = \frac{2}{a} \times \left(\frac{\pi}{a}\right)^{-1} \times [-\cos(\pi/2) + \cos(0)]$$

$$= \frac{2}{\pi} (1)$$

$$= \frac{2}{\pi} / 1$$

The probability of measuring E_n is $|c_n|^2$.

$$7:00 \quad E_n = \frac{n^2 \pi^2 \hbar^2}{2ma^2}$$

8:00

$$\text{Hence, } E_i = 1C_1 l^2 = \frac{h}{\pi^2} //$$

9:00

7) The general solution for a particle in an infinite square well is

$$\Psi(n, t) = \sum_{n=1}^{\infty} c_n \sqrt{\frac{2}{a}} \exp\left(-i \frac{n^2 \pi^2 \hbar}{2ma^2} t\right) \sin\left(\frac{n\pi n}{a}\right)$$

10:00

$$TP: \Psi(n, 0) = \Psi(n, T)$$

11:00

$$\sum_{n=1}^{\infty} c_n \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi n}{a}\right) = \sum_{n=1}^{\infty} c_n \sqrt{\frac{2}{a}} \exp\left(-i \frac{n^2 \pi^2 \hbar}{2ma^2} T\right) \sin\left(\frac{n\pi n}{a}\right)$$

13:00

$$\sin\left(\frac{n\pi n}{a}\right)$$

14:00

$$\text{comparing both sides, } \exp\left(-i \frac{n^2 \pi^2 \hbar}{2ma^2} T\right) = 1$$

15:00

$$\cos\left(\frac{n^2 \pi^2 \hbar}{2ma^2} T\right) - i \sin\left(\frac{n^2 \pi^2 \hbar}{2ma^2} T\right) = 1$$

17:00

comparing real and imaginary parts,

18:00

$$\cos\left(\frac{n^2 \pi^2 \hbar}{2ma^2} T\right) = 1$$

19:00

$$\sin\left(\frac{n^2 \pi^2 \hbar}{2ma^2} T\right) = 0$$

7:00 Both sides of the equation are satisfied if

$$\frac{n^2 \pi^2 h}{2ma^2} T = n^2 2\pi$$

$$9:00 \Rightarrow T = \frac{2\pi \times 2ma^2}{\pi^2 h}$$

$$10:00 - \frac{4ma^2}{\pi h}$$

11:00

Hence, proved.

12:00

10:00

⑧ Let us assume that there is a unitary operator that copies two quantum states $| \psi \rangle$ and $| \phi \rangle$

$$| \psi \rangle \otimes | 0 \rangle \rightarrow | \psi \rangle \otimes | \psi \rangle$$

$$| \phi \rangle \otimes | 0 \rangle \rightarrow | \phi \rangle \otimes | \phi \rangle$$

13:00

Now let us consider the state $| \psi \rangle = \alpha | 0 \rangle + \beta | 1 \rangle$
which when cloned becomes :-

$$| \psi \rangle | 0 \rangle \rightarrow | \psi \rangle | \psi \rangle = (\alpha | 0 \rangle + \beta | 1 \rangle) (\alpha | 0 \rangle + \beta | 1 \rangle)$$

16:00

$$= \alpha^2 | 00 \rangle + \beta^2 | 11 \rangle + \beta \alpha | 01 \rangle + \alpha \beta | 10 \rangle$$

17:00

However, due to linearity of the cloning operator
which preserves inner product, we get.

$$(\alpha | 0 \rangle + \beta | 1 \rangle) | 0 \rangle \rightarrow \alpha | 0 \rangle | 0 \rangle + \beta | 1 \rangle | 0 \rangle$$

$$\text{Now, } \alpha | 0 \rangle | 0 \rangle \rightarrow \alpha^2 | 00 \rangle \quad \beta | 1 \rangle | 0 \rangle \rightarrow \beta^2 | 10 \rangle$$

This cloning misses the $|110\rangle$ and $|101\rangle$ terms.
 Thus, we have a contradiction and the operator
 cannot copy the linear superpositions of any two
 quantum states though it can copy them individually.
 This is because the operator is a unitary and
 linear operator.

10:00

(9) In the Stern Gerlach experiment, the probability of
 spin-up is $\cos^2(\theta/2)$ and that of spin-down
 is $\sin^2(\theta/2)$.
 where, $\theta \rightarrow$ angle between the two magnetic fields

11:00, $P(\text{up}) = 2 P(\text{down})$

$$\cos^2(\theta/2) = 2 \sin^2(\theta/2)$$

12:00 $\tan^2(\theta/2) = 1/2$

13:00 $\tan \theta/2 = \sqrt{1/2} \Rightarrow \frac{\theta}{2} = \tan^{-1}(\sqrt{1/2})$

14:00 $\theta = 2 \tan^{-1}\left(\frac{1}{\sqrt{2}}\right)$

15:00

According to the quantum entanglement theory,
 the particle follows a probabilistic curve until the
 moment of measurement.
 After this Alice measures the wave which forces
 the wave function to collapse and simultaneously

16:00

7 produces Bob's spin.

To ensure that there is no causal influence, we
8 need to make sure that light also doesn't reach

9:00 $c = 3 \times 10^8 \text{ m/s}$

$t = 1 \text{ ms} = 10^{-3} \text{ s}$

10:00 distance = $c \times t$

= $3 \times 10^5 \text{ m}$

11:00 = 300 km

Thus, they need to be atleast 300km for there to be no causal interference.

13:00

(11)

14:00 $m (\text{neutron}) = 1.674 \times 10^{-27} \text{ kg}$

$v = 10^6 \text{ m/s}$

15:00

$\Delta p = M \Delta v = 1.67 \times 10^{-27} \times 10^6$

16:00 = $1.67 \times 10^{-21} \text{ kgm/s}$

17 Uncertainty principle, $\Delta n \Delta p \geq \frac{\hbar}{2}$

18:00

$\Delta n \geq \frac{1.054 \times 10^{-34}}{2 \times 1.67 \times 10^{-21}}$

19:00

$\Delta n \geq 0.315 \times 10^{-13} \text{ m}$

20:00

$\Delta n \geq \underline{3.15 \times 10^{-14} \text{ m}}$

8:00 $m = 50\text{kg}$
 $v = 2\text{m/s}$

9:00 Uncertainty principle, $\Delta x \Delta p \geq \frac{\hbar}{2}$

10:00 $\Delta x m \Delta v \geq \frac{\hbar}{2}$

11:00 $\Delta x \times 50 \times 2 \geq \frac{1.05 \times 10^{-34}}{2}$

12:00 $\Delta x \geq \frac{1.05 \times 10^{-34}}{2 \times 50 \times 2}$

13:00 $\Delta x \geq 0.527 \times 10^{-34} \text{ m}$

14:00 $\Delta x \geq 5.27 \times 10^{-37} \text{ m}$

15:00 (20) The reason for why an electron cannot reside inside the nucleus can be explained using the uncertainty principle.

16:00 The size of the nucleus is of the order of 10^{-15} m .
Thus Δx should be less than 10^{-15} m for an e^- .
we also know, $m_e = 9.31 \times 10^{-31} \text{ kg}$

17:00 $\therefore \Delta x \Delta p \geq \frac{\hbar}{2}$ (Uncertainty principle)

7:00 $\Delta n \cdot m \Delta v \geq \frac{\hbar}{d}$

V...

8:00 $\Delta v \geq \frac{1.054 \times 10^{-34}}{2 \times 10^{-15} \times 9.1 \times 10^{-31}}$

A...

9:00

q...

$\Delta v \geq 0.057 \times 10^{12} \text{ m/s}$

10:00 $\Delta v \geq 5.7 \times 10^{10} \text{ m/s}$

I...

11:00 This implies that the e^- will need to have a speed much higher than the speed of

12:00 light [$c = 3 \times 10^8 \text{ m/s}$]; which is impossible.

Hence, the e^- cannot reside inside the nucleus.

13:00 IP...