

General Relativity (I)

homework for week 5

due: week 7

1. [the curvature tensor and related tensors; use 2D Riemann manifold as an example] 80%
The **Riemann curvature tensor** can be computed by:

$$\boxed{R^\alpha_{\beta\mu\nu} \equiv \Gamma^\alpha_{\beta\nu,\mu} - \Gamma^\alpha_{\beta\mu,\nu} + \Gamma^\alpha_{\sigma\mu} \Gamma^\sigma_{\beta\nu} - \Gamma^\alpha_{\sigma\nu} \Gamma^\sigma_{\beta\mu}} , \quad (1)$$

and a associated $\begin{pmatrix} 0 \\ 4 \end{pmatrix}$ tensor can be obtained by

$$\boxed{R_{\alpha\beta\mu\nu} \equiv g_{\alpha\kappa} R^\kappa_{\beta\mu\nu}} .$$

Consider a 2D sphere with coordinate (θ, ϕ) and radius a , the metric tensor is

$$g_{\alpha\beta} = \begin{pmatrix} a^2 & 0 \\ 0 & a^2 \sin^2 \theta \end{pmatrix} .$$

(a) In the class we have learned that the covariant derivative of a one-form p_α is

$$p_{\alpha;\beta} = p_{\alpha,\beta} - p_\mu \Gamma^\mu_{\alpha\beta} .$$

From this, show that

$$p_{\alpha;\beta\gamma} - p_{\alpha;\gamma\beta} = R^\mu_{\alpha\beta\gamma} p_\mu .$$

That is, unlike partial derivatives, the order of covariant derivatives matters (unless $R^\mu_{\alpha\beta\gamma} = 0$).

(b) As seen in problem set 1(f) of the week 4 homework, in a locally inertial frame at a point \mathcal{P} , we have find $\Gamma^\alpha_{\beta\mu}|_{\mathcal{P}} = 0$ and $\Gamma^\alpha_{\beta\mu,\nu}|_{\mathcal{P}} = 0$ (that is, second derivatives of $g_{\mu,\nu}$ cannot be zero in general). From eqn. (1) and

$$\Gamma^\mu_{\alpha\beta} = \frac{1}{2} g^{\mu\nu} (g_{\nu\alpha,\beta} + g_{\nu\beta,\alpha} - g_{\alpha\beta,\nu}) ,$$

we get

$$R^\alpha_{\beta\mu\nu} = \frac{1}{2} g^{\alpha\sigma} (g_{\sigma\nu,\beta\mu} - g_{\sigma\mu,\beta\nu} + g_{\beta\mu,\sigma\nu} - g_{\beta\nu,\sigma\mu}) . \quad (2)$$

Show the cyclic identity:

$$R^\alpha_{\beta\mu\nu} + R^\alpha_{\nu\beta\mu} + R^\alpha_{\mu\nu\beta} = 0 \quad (3)$$

or, alternatively

$$R_{\alpha\beta\mu\nu} + R_{\alpha\nu\beta\mu} + R_{\alpha\mu\nu\beta} = 0$$

Note that eqn. (2) is not a valid tensor equation since it involves partial derivative rather than covariant ones. However, eqn. (3) is a tensor equation since it is constructed by the (Riemann) tensors.

(c) According to the relation

$$R_{\alpha\beta\mu\nu} = -R_{\beta\alpha\mu\nu} ,$$

$$R_{\alpha\beta\mu\nu} = -R_{\alpha\beta\nu\mu} ,$$

$$R_{\alpha\beta\mu\nu} = R_{\mu\nu\alpha\beta} ,$$

argue that $R_{\alpha\mu\nu}^\alpha = 0$ and $R_{\beta\mu\alpha}^\alpha = -R_{\beta\alpha\mu}^\alpha$. Therefore, for the 2D sphere considered here, all the Riemann tensor are either zero or $\pm R_{\theta\phi\theta\phi}$.

(d) As a result of (c), the only non-zero contraction of the Riemann tensor is the **Ricci tensor**

$$R_{\alpha\beta} \equiv R_{\alpha\mu\beta}^\mu = R_{\beta\alpha} .$$

Show that $R_{\theta\theta} = -1$ and $R_{\phi\phi} = -\sin^2\theta$.

Note that the Riemann curvature tensor in a spherical surface considered here is NOT zero, as expected. In comparison, the Riemann curvature tensor vanishes for a *spherical coordinate* in a 3D space, since it is a *flat* space. This was verified in problem set 1(g) in the week 4 homework.

(e) The Ricci tensor is symmetric ($R_{\alpha\beta} = R_{\beta\alpha}$). Show this by contacting the cyclic identity, eqn. (3).

(f) The **Ricci scalar** is defined by

$$R \equiv g^{\mu\nu} R_{\mu\nu} .$$

Show that $R = -2/a^2$.

(g) We can further define an symmetric tensor, the **Einstein tensor**

$$G_{\mu\nu} \equiv R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R$$

Can you see there is only one divergence: $G^{\mu\nu}_{;\mu}$. That is, $G^{\mu\nu}_{;\mu} = G^{\mu\nu}_{;\nu}$.

Although you are not asked to show the covariant of the Einstein tensor is zero (the property is extremely important when constructing *Einstein's field equation*), it is good to know such property can be shown by using the Bianchi identity

$$R_{\beta\mu\nu}^\alpha + R_{\nu\beta\mu}^\alpha + R_{\mu\nu\beta}^\alpha = 0$$

2. [stress-energy tensor] 20%

The stress-energy tensor for a perfect fluid reads

$$T^{\alpha\beta} = (\rho + P)u^\alpha u^\beta + P g^{\alpha\beta} ,$$

where ρ and P are *rest-frame* energy density and pressure.

By perfect fluid we mean there is no viscosity (and therefore $T^{ij} = 0$) and no heat conduction (and therefore $T^{0i} = T^{i0} = 0$) in the MCRF (*Momentarily Comoving Reference Frame*).

(a) First, from $u^\alpha u_\alpha = -1$, proof that $u^\nu_{;\mu} u_\nu = 0$.

(b) *Dust* is a fluid without internal stress or pressure. By using the above relation, show that $T^{\mu\nu}_{;\nu} = 0$ implise that the dust particles follow geodesics.