

# The Non Equilibrium Steady State of Sparse Systems: Energy Absorption and Current

Daniel Hurowitz [1,2]

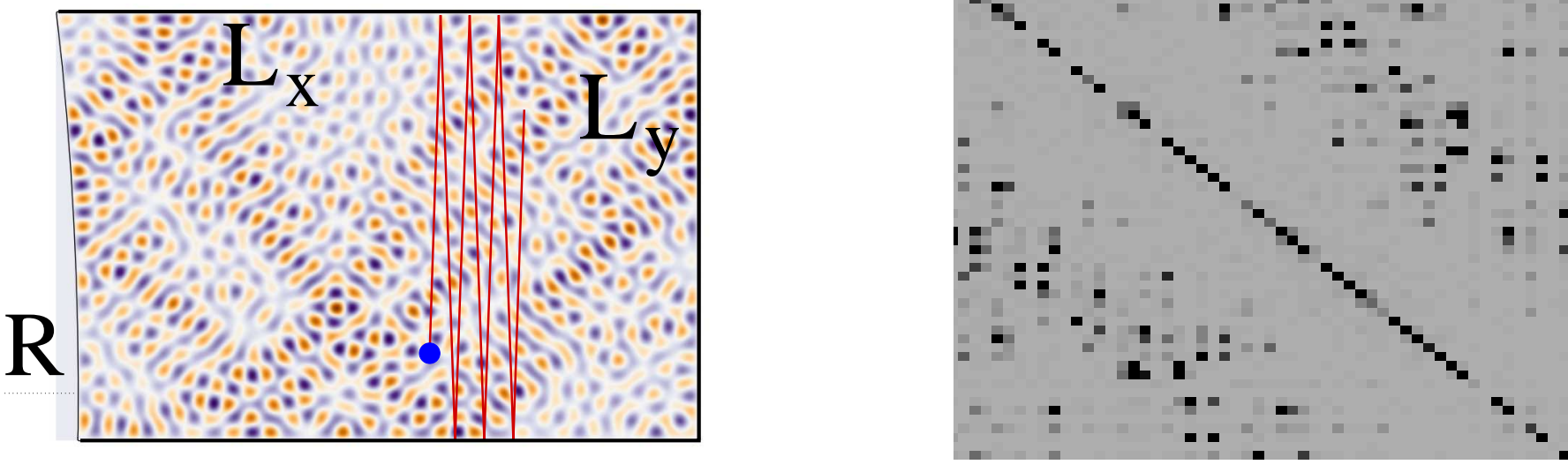
Ben Gurion University in the Negev, Beer Sheva, Israel



## What is a "sparse" matrix?

- ▶ The majority of matrix elements are small, yet some are very large.
- ▶ The elements have a log-wide distribution with a median that is much smaller compared to the average.

### Example: Weakly chaotic billiard [3]



## Why sparsity is interesting

- ▶ Energy absorption rate (EAR) calculation goes beyond linear response theory (requires resistor network calculation) [1].
- ▶ Leads to a novel NESS that has "glassy" nature [1].
- ▶ Novel quantum saturation effect [1].
- ▶ Current vs. driving goes beyond LRT (emergence of a Sinai regime) [2].

## Driven system + bath : A paradigm for NESS

The Hamiltonian:

$$\mathcal{H}_{\text{total}} = \mathcal{H}_0 - f(t)\{V_{nm}\} + F(t)\{W_{nm}\} + \mathcal{H}_{\text{Bath}}$$

Definitions:

$$\langle f(t)f(t') \rangle = \epsilon^2 \delta(t - t')$$

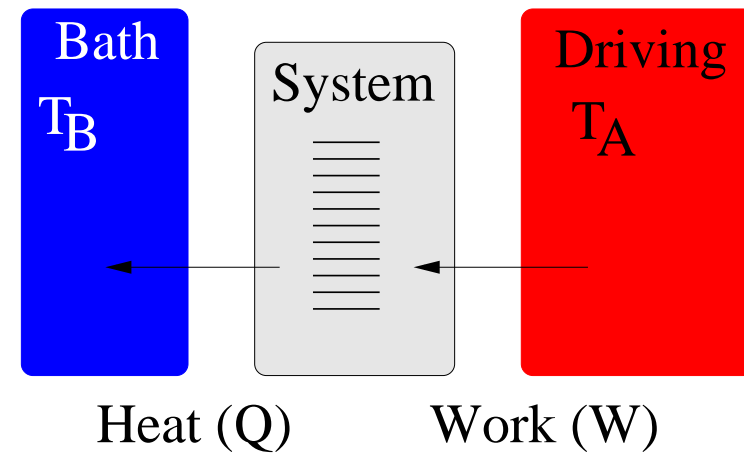
$T_A = \infty \equiv$  Temperature of A

$T_B \equiv$  Temperature of B

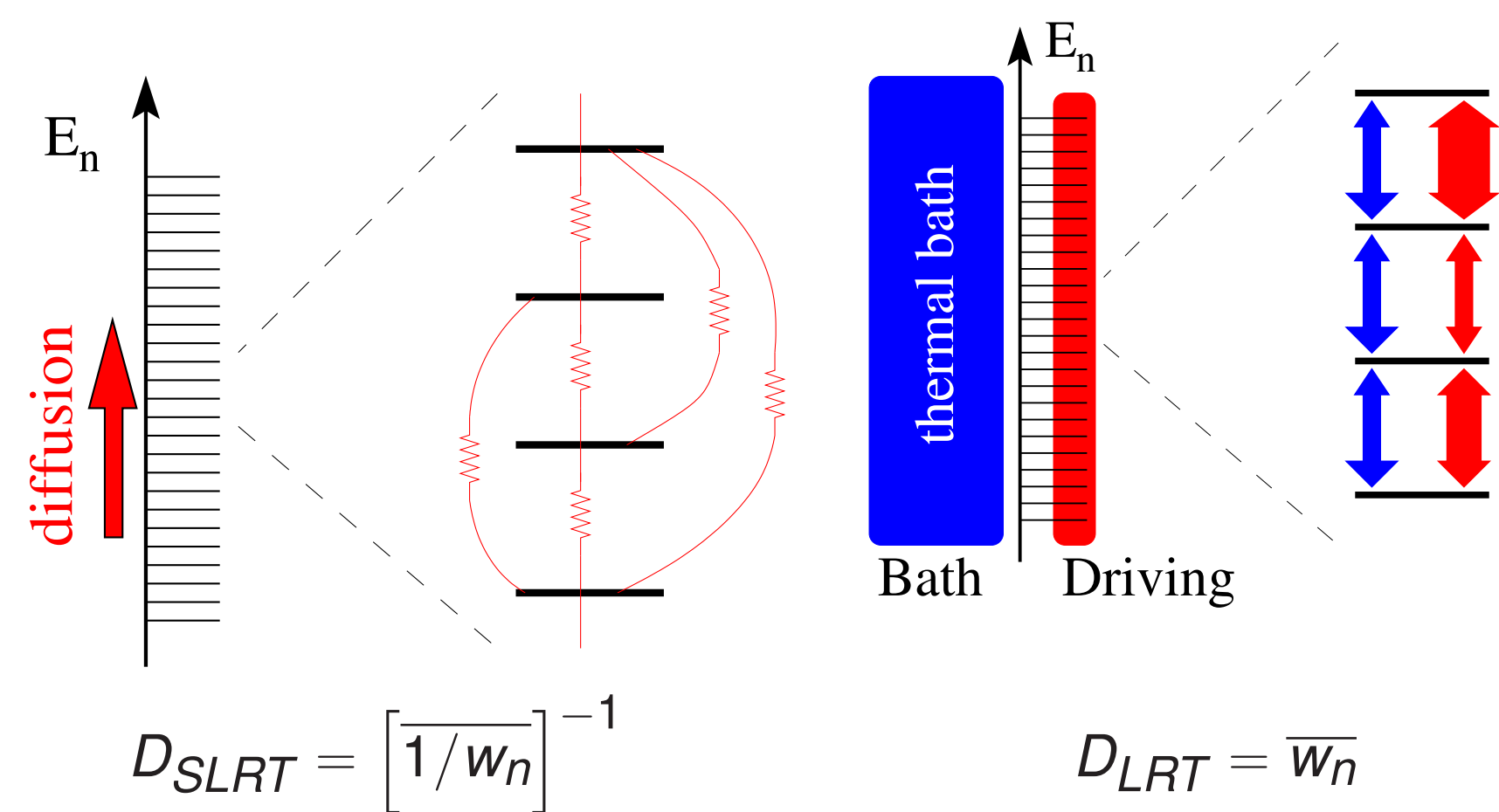
Steady state:

$$\dot{E} = \dot{W} - \dot{Q} = 0$$

$T_A \neq T_B$  : NESS is non-canonical



## Crossover from Linear response to semi-linear response



## Master equation description of dynamics

Quantum master equation for the reduced probability matrix:

$$\frac{d\rho}{dt} = -i[\mathcal{H}_0, \rho] - \frac{\epsilon^2}{2}[V, [V, \rho]] + \mathcal{W}^\beta \rho$$

Corresponding stochastic rate equation:

$$\frac{dp_n}{dt} = \sum_m w_{nm} p_m - w_{mn} p_n$$

Steady state equation:

$$\dot{\rho} = \mathcal{W}\rho = 0$$

The transition rates

$$\begin{aligned} w_{nm} &= w_{nm}^{\epsilon} + w_{nm}^{\beta} \\ w_{nm}^{\epsilon} &= w_{mn}^{\epsilon} \propto \epsilon^2 \\ w_{nm}^{\beta} &= \exp\left[-\frac{E_n - E_m}{T_B}\right] \\ w_{mn}^{\beta} & \end{aligned}$$

## The generalized fluctuation - dissipation phenomenology

$$\begin{aligned} \dot{W} &= \text{rate of heating} = \frac{D}{T_{\text{system}}} \\ \dot{Q} &= \text{rate of cooling} = \frac{D_B}{T_B} - \frac{D_B}{T_{\text{system}}} \end{aligned}$$

At the NESS:

$$\begin{aligned} T_{\text{system}} &= \left(1 + \frac{D(\epsilon)}{D_B}\right) T_B \\ \dot{Q} = \dot{W} &= \frac{1/T_B}{D_B^{-1} + D(\epsilon)^{-1}} \end{aligned}$$

Experimental determination of response:

$$D(\epsilon) = \frac{Q(\epsilon)}{Q(\infty) - Q(\epsilon)} D_B$$

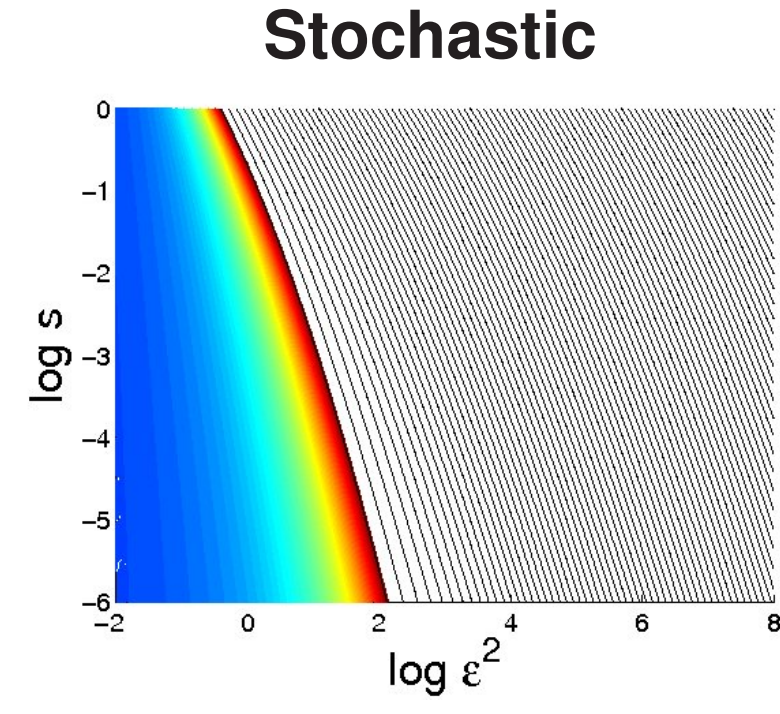
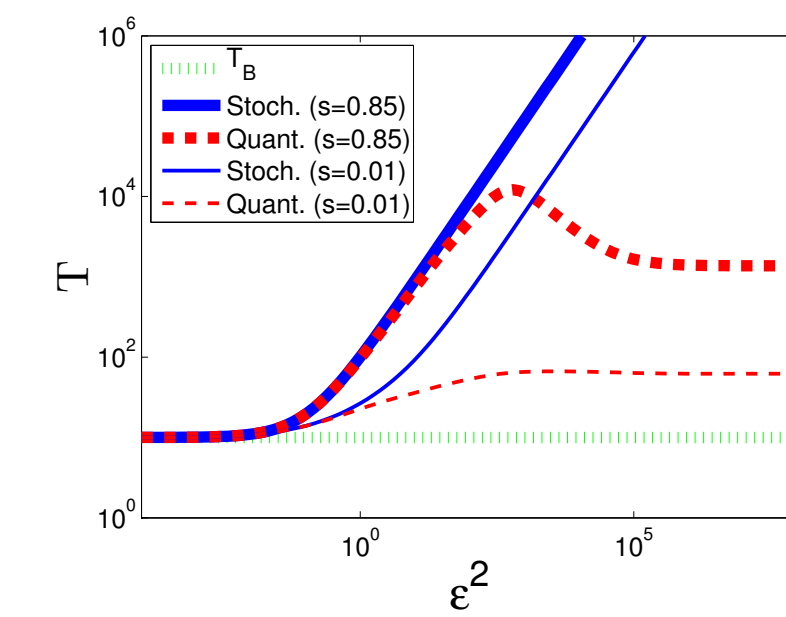
## The steady state temperature

### Microscopic and macroscopic Temperature

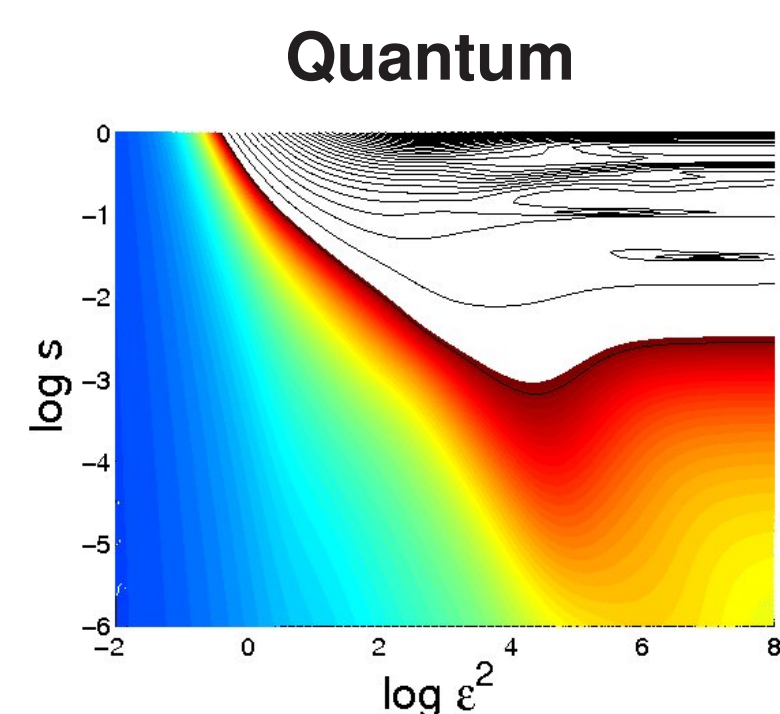
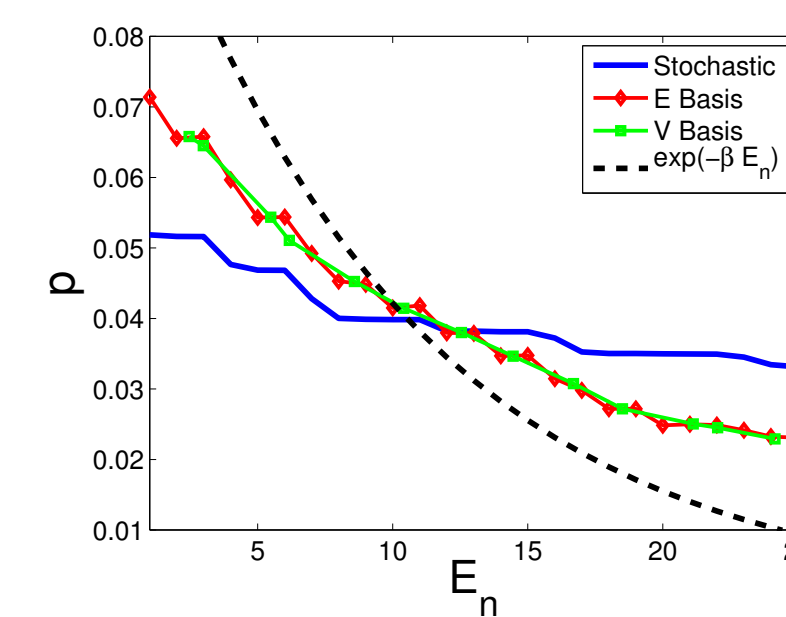
$$\frac{p_n}{p_m} = \exp\left(-\frac{E_n - E_m}{T_{nm}}\right), \quad T_{\text{system}} = \text{average}[T_{nm}]$$

This resembles a glassy state

### Temperature vs. Driving



### Occupation probability



### The quantum saturation effect

For very strong driving, the NESS is a mixture of V eigenstates:

$$p_r \sim \exp(-\langle E \rangle_r / T_{\text{mix}})$$

For non-sparse V, the eigenstates are extended in energy space -  $T_{\infty} \rightarrow \infty$ .

If V is sparse, the states are localized. As  $s \rightarrow 0$ ,  $T_{\infty} \sim T_{\text{mix}} \sim T_B$ .

$$T_B < T_{\infty} < \infty \quad [\text{depends on the sparsity}]$$

## Derivation of the cooling rate formula

cooling rate:

$$\dot{Q} = - \sum_{n,m} (E_n - E_m) w_{nm}^{\beta} p_m$$

occupation imbalance:

$$p_n - p_m = \left[2 \tanh\left(-\frac{E_n - E_m}{2T_{nm}}\right)\right] \bar{p}_{nm}$$

up/down transitions imbalance

$$w_{nm}^{\beta} - w_{mn}^{\beta} = \left[2 \tanh\left(-\frac{E_n - E_m}{2T_B}\right)\right] \bar{w}_{nm}^{\beta}$$

$$\begin{aligned} \dot{Q} &= \frac{1}{2} \sum_{n,m} (E_n - E_m)^2 \frac{\bar{w}_{nm}^{\beta}}{T_B} \bar{p}_{nm} - \frac{1}{2} \sum_{n,m} (E_n - E_m)^2 \frac{\bar{w}_{nm}^{\beta}}{T_{nm}} \bar{p}_{nm} \\ &= \frac{D_B}{T_B} - \frac{D_B}{T_{\text{system}}} \end{aligned}$$

definition of the diffusion coefficient:

$$D_B \equiv \left[ \frac{1}{2} \sum_n (E_n - E_m)^2 \bar{w}_{nm}^{\beta} \right]$$

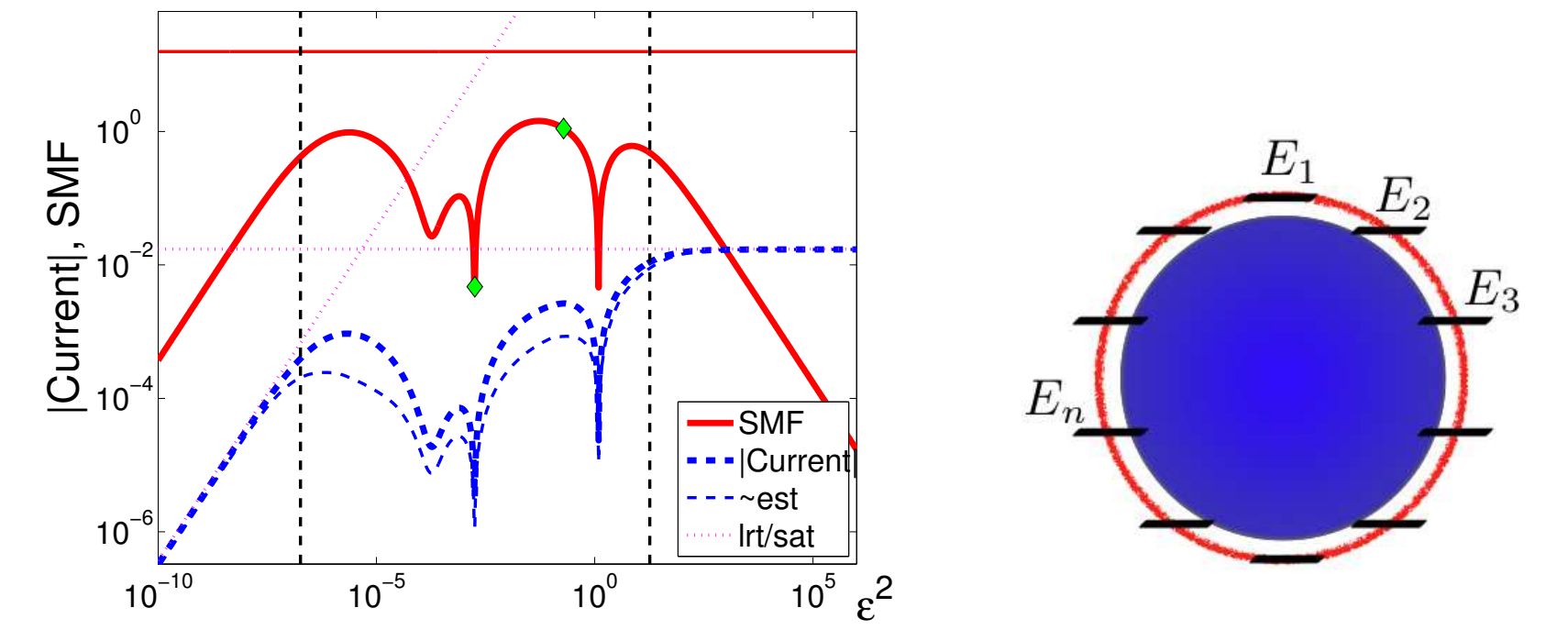
definition of effective system temperature:

$$\frac{1}{T_{\text{system}}} \equiv \left[ \frac{1}{T_{nm}} \right]$$

## Current vs. driving

Driving  $\leadsto$  Stochastic Motive Force  $\leadsto$  Current

Regimes: LRT regime, Sinai regime, Saturation regime



$$I = w_{\bar{n}} p_{n-1} - w_{\bar{n}} p_n, \quad \sum_n p_n = 1$$

$$I \sim \frac{1}{N} \bar{w} \exp\left[-\frac{\mathcal{E}_n}{2}\right] 2 \sinh\left(\frac{\mathcal{E}_{\odot}}{2}\right)$$

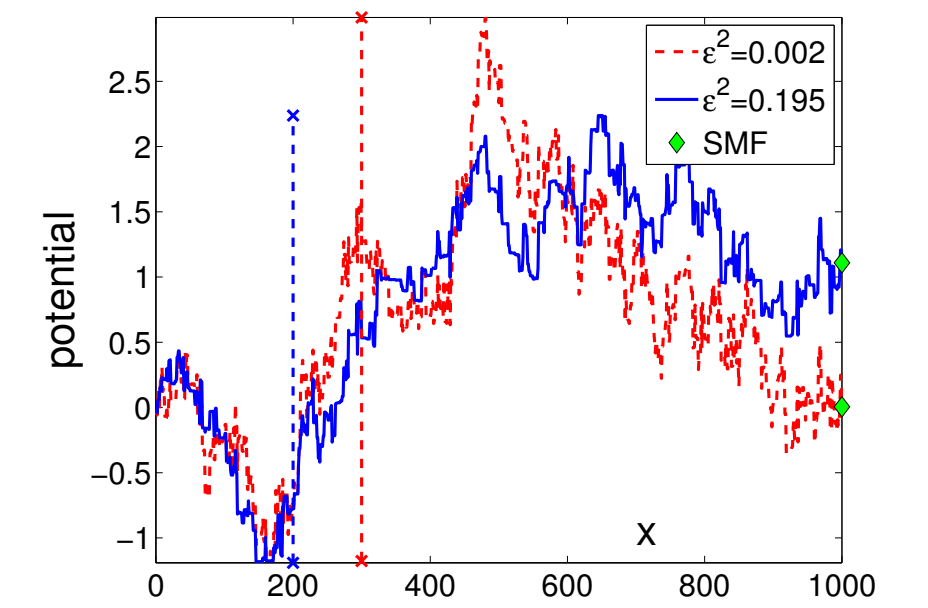
## The Stochastic Motive Force (SMF)

If we had only a bath

$$\frac{w_{nm}}{w_{mn}} = \exp\left[-\frac{E_n - E_m}{T_B}\right]$$

We define a "field"

$$\mathcal{E}(x) \equiv \ln\left[\frac{w_{nm}}{w_{mn}}\right]$$



The "potential" variation along a segment

$$\mathcal{E}(x_1 \leadsto x_2) = \sum_{x=x_1}^{x_2} \mathcal{E}(x) = \int_{x_1}^{x_2} \mathcal{E}(x) dx$$

$$\mathcal{E}_n \equiv \text{maximum}\left\{|\mathcal{E}(x_1 \leadsto x_2)|\right\}$$

$$\mathcal{E}_{\odot} \equiv \oint \mathcal{E}(x) dx \quad \text{if no driving} = 0$$

## Emergence of the "Sinai regime"

Sinai [1982]: Transport in a chain with random transition rates.

Assume transition rates are uncorrelated.

$\leadsto$  exponential build up of a potential barrier  $\mathcal{E}_n \propto \sqrt{N}$

$\leadsto$  exponentially small current.

But... we have telescopic correlations:

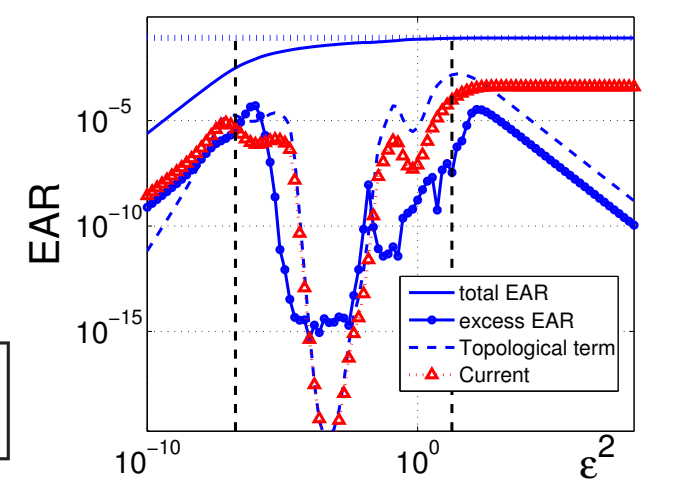
$$\mathcal{E}_{n,n+1} \sim \Delta_n \equiv (E_n - E_{n+1})$$

$$\mathcal{E}_{\odot} \approx - \sum_n \left[ \frac{1}{1 + g_n \epsilon^2} \right] \frac{\Delta_n}{T_B} \sim \frac{1}{T_B} \begin{cases} \epsilon^2, & \epsilon^2 < 1/g_{\text{max}} \\ 1/\epsilon^2, & \epsilon^2 > 1/g_{\text{min}} \\ [\pm] \sqrt{N} \Delta, & \text{otherwise} \end{cases}$$

Build up may occur if  $g_n$  are from a log-wide distribution.

## Beyond fluctuation dissipation phenomenology: Topological term in EAR formula

$$\begin{aligned} \dot{Q} &= \sum_n \left[ w_{\bar{n}}^{\beta} p_n - w_{\bar{n}}^{\beta} p_{n-1} \right] \Delta_n \\ &\approx \left[ \frac{D_B}{T_B} - \frac{D_B}{T(0)} \right] + T_B \mathcal{E}_{\odot} I \\ &\approx \frac{D_B}{T_B} \left[ (g_n \epsilon^2) - (g_n \epsilon^2)^2 + \text{Var}(g) \epsilon^4 \right] \end{aligned}$$



The EAR is correlated with the current.

## Conclusions

1. The stochastic NESS resembles a glassy phase (wide distribution of microscopic temperatures).
2. Definition of effective NESS temperature and extension of FDR phenomenology.
3. Prediction of LRT  $\rightarrow$  SLRT crossover.
4. For very strong driving, quantum saturation of effective NESS temperature.
5. Topological aspects: A wide distribution of rates is crucial for a Sinai regime.
6. Topological term in EAR is proportional to the entropy production, but sub-linear in driving intensity. The EAR is correlated with the current.

## References

- [1] D. Hurowitz, D. Cohen (EPL 2011)
- [2] D. Hurowitz, S. Rahav, D. Cohen (EPL 2012)
- [3] D. Cohen, T. Kottos, H. Schanz (JPA 2006)
- [4] M. Wilkinson, B. Mehlig, D. Cohen (EPL 2006)
- [5] A. Stotland, T. Kottos, D. Cohen (PRB 2010)
- [6] A. Stotland, D. Cohen, N. Davidson (EPL 2009)