

PH100: Mechanics & Thermodynamics
Tutorial 7

1. A furnace has walls of temperature 1600°C . What is the wavelength of maximum intensity emitted when a small door is opened?

From Wien's Displacement Law,

$$\lambda_m = \frac{b}{T}$$

$\lambda_m \rightarrow$ maximum intensity wavelength

$T \rightarrow$ temperature in K

$b \rightarrow$ Wien's Constant

$$b = 2.898 \times 10^{-3} \text{ m K}$$

$$\text{given } T = 1600^{\circ}\text{C} = 1873 \text{ K}$$

$$\therefore \lambda_m = \frac{2.898 \times 10^{-3}}{1873}$$

$$\lambda_m = 0.001547 \times 10^{-3} \text{ m}$$
$$= 1.547 \mu\text{m}$$

Hence, the wavelength corresponding to maximum intensity of radiation emitted is $1.547 \mu\text{m}$.

2. Calculate max for blackbody radiation for (a) liquid helium (4.2 K) (b) room temperature (293 K) (c) a steel furnace (2500 K) and, (d) a blue star (9000 K)

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Using Wien's displacement law

$$\lambda_m = \frac{b}{T} \text{, where } b = 2.898 \times 10^{-3} \text{ mK}$$

a) $T = 4.2 \text{ K}$

$$\begin{aligned}\therefore \lambda_m &= \frac{2.898 \times 10^{-3}}{4.2} \\ &= 0.69 \times 10^{-3} \text{ m} \\ &= 690 \text{ nm}\end{aligned}$$

Hence, max intensity for blackbody radiation at $T = 4.2 \text{ K}$ corresponds to wavelength of 690 nm.

b) $T = 293 \text{ K}$

$$\lambda_m = \frac{2.898 \times 10^{-3}}{293}$$

$$\begin{aligned}\lambda_m &= 0.00989 \times 10^{-3} \text{ m} \\ &= 9.89 \text{ nm}\end{aligned}$$

Hence, wavelength corresponding to max intensity is 9.89 nm.

c) $T = 2500 \text{ K}$

$$\begin{aligned}\lambda_m &= \frac{2.898 \times 10^{-3}}{2500} \\ &= 0.1159 \times 10^{-5}\end{aligned}$$

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$$\lambda_m = 1.16 \text{ } \mu\text{m}$$

Hence, wavelength corresponding to maximum ~~wavelength~~ intensity is $1.16 \text{ } \mu\text{m}$.

d) $T = 9000 \text{ K}$

$$\begin{aligned}\lambda_m &= \frac{2.898 \times 10^{-3}}{9000} \\ &= 0.322 \times 10^{-6} \\ &= 322 \text{ nm}\end{aligned}$$

Hence, wavelength corresponding to max intensity is 322 nm .

3. Calculate the temperature of a blackbody if the spectral distribution peaks at

(a) gamma rays, $\lambda = 1.5 \times 10^{-14} \text{ m}$

(b) X-rays, $\lambda = 1.5 \text{ nm}$

(c) red light, $\lambda = 640 \text{ nm}$

(d) broadcast TV waves, $\lambda = 1 \text{ m}$

and (e) AM radio waves, $\lambda = 204 \text{ m}$

Using Wien's displacement law,

$$\lambda_m = \frac{b}{T}$$

$$T = \frac{b}{\lambda_m} \quad \text{where } b = 2.898 \times 10^{-3} \text{ mK.}$$

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(a) gamma rays ($\lambda = 1.5 \times 10^{-14} \text{ m}$)

$$\therefore T = \frac{2.898 \times 10^3}{1.5 \times 10^{-14}}$$

$$T = 1.932 \times 10^{11} \text{ K}$$

$$T = 1.932 \times 10^{11} \text{ K}$$

Hence, temperature of blackbody is $1.932 \times 10^{11} \text{ K}$.

(b) X-rays ($\lambda = 1.5 \text{ nm}$)

$$\therefore T = \frac{2.898 \times 10^3}{1.5 \times 10^{-9}}$$

$$= 1.932 \times 10^6 \text{ K}$$

Hence, temperature of blackbody is $1.932 \times 10^6 \text{ K}$.

(c) red light ($\lambda = 640 \text{ nm}$)

$$\therefore T = \frac{2.898 \times 10^3}{640 \times 10^{-9}}$$

$$= 0.00453 \times 10^6$$

$$= 4530 \text{ K}$$

Hence, temperature of blackbody is 4530 K.

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d) broadcast TV waves ($\lambda = 1\text{m}$)

$$\therefore T = \frac{2.898 \times 10^{-3}}{1}$$

$$= \cancel{2.898} \cancel{\times 10^{-3}}$$

$$= 2.898 \times 10^{-3} \text{ K}$$

Hence, temperature of blackbody is $2.898 \times 10^{-3} \text{ K}$

e) AM radio waves ($\lambda = 204\text{m}$)

$$\therefore T = \frac{2.898 \times 10^{-3}}{204}$$

$$= 0.0142 \times 10^{-3}$$

$$= 14.2 \times 10^{-6} \text{ K}$$

Hence, the temperature of blackbody is $14.2 \times 10^{-6} \text{ K}$.

4. The wavelength of maximum intensity of sun's radiation is observed to be near 500 nm. Assume the sun to be a blackbody and calculate

a) the sun's surface temperature

b) the power per unit area emitted from the sun's surface, and

c) Energy received by Earth each day from sun's radiation.

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a) Using Wien's displacement law,

$$T = \frac{b}{\lambda m}$$

$$\begin{aligned} \therefore T &= \frac{2.898 \times 10^{-3}}{500 \times 10^{-9}} \\ &= 0.5796 \times 10^4 \\ &= 5796 \text{ K} \end{aligned}$$

Hence, the Sun's surface temperature is 5796 K.

b) From Stefan-Boltzmann Law,

$$E = e \sigma A T^4$$

E is power and e is emissivity, σ is Stefan's constant for blackbody $e=1$.

\therefore power per unit area (I)

$$I = \frac{E}{A} = \sigma T^4$$

$$\begin{aligned} I &= 5.67 \times 10^{-8} \times (5796)^4 \\ &= 6.399 \times 10^7 \text{ W/m}^2 \end{aligned}$$

Hence, the power per unit area emitted from Sun's surface is $6.399 \times 10^7 \text{ W/m}^2$.

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c) Let d be the distance between Sun and Earth and R be the radius of Earth.

: power received at Earth due to

$$\text{sun's radiation} = \frac{\sigma T^4 (4\pi R^2)}{(4\pi d^2)}$$

$$\therefore \frac{R}{d} = \frac{696340 \text{ (km)}}{1.496 \times 10^8 \text{ (cm)}} = 4.65 \times 10^{-3}$$

$$\begin{aligned}\text{: Power received at Earth} &= 5.67 \times 10^{-8} \times (5796)^4 \times (4.65 \times 10^{-3})^2 \\ &= 1383.62 \text{ W}\end{aligned}$$

: Energy received by Earth in one day

$$\begin{aligned}&= 1383.62 \times 24 \times 60 \times 60 \\ &= 1.195 \times 10^{10} \text{ J}\end{aligned}$$

Hence, Earth receives $1.195 \times 10^{10} \text{ J}$ of energy from Sun's radiation each day.

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5. (a) A blackbody's temperature is increased from 900 K to 2300 K. By what factor does the total power radiated per unit area increase?
- b) If the original temperature is again 900 K, what final temperature is required to double the power output?

(a) By Stefan's - Boltzmann law, power radiated (E) is given by,

$$E = \epsilon \sigma A T^4$$

∴ Power radiated per unit area (I),

$$I = \frac{E}{A} = \epsilon \sigma T^4$$

for blackbody $\epsilon = 1$,

$$\therefore I = \sigma T^4$$

Now, $T_1 = 900\text{ K}$ and $T_2 = 2300\text{ K}$.

$$\therefore \frac{I_2}{I_1} = \frac{\sigma T_2^4}{\sigma T_1^4}$$

$$\frac{I_2}{I_1} = \frac{(2300)^4}{(900)^4}$$

$$I_2 = \left(\frac{23}{9}\right)^4 I_1$$

$$I_2 = 42.652 I_1$$

Hence, total power radiated per unit area

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increased by a factor of 42.652 when temperature of blackbody is increased from 900 K to 2300 K.

(b) Let power per unit area be I_1 at 900 K. and power per unit area be $2I_1$ at T_2 K.

$$\therefore 2I_1 = \sigma T_2^4 \quad \text{--- (1)}$$

and, $I_1 = \sigma (900)^4 \quad \text{--- (2)}$

$$(1)/(2)$$

$$2 = \left(\frac{T_2}{900} \right)^4$$

$$\frac{T_2}{900} = 1.1892$$

$$T_2 = 1070.286 \text{ K}$$

Hence, the final temperature of 1070.286 K is required to double the power output.

6. a) At what wavelength will the human body radiate the maximum radiation? b) Estimate the total power radiated by a person of medium build (assume an area given by a cylinder of 175 cm height and 13 cm radius). c) Using your answer to (b), compare the energy radiated by a person in one day with energy intake

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of a 2000-kcal diet.

normal

a) The temperature of human body is 310 K .

Using Wien's displacement law,

$$\begin{aligned} \lambda_m &= \frac{b}{T} \\ &= \frac{2.898 \times 10^{-3}}{310} \\ &= 0.00935 \times 10^{-3} \\ &= 9.35 \times 10^{-6} \text{ m} \\ &= 9.35 \text{ um} \end{aligned}$$

Hence, at a wavelength of 9.35 um human body will radiate the maximum radiation.

b) Using Stefan's-Boltzmann law,

$$E = \epsilon \sigma A T^4$$

for blackbody $\epsilon = 1$,

Since, the body is considered to be a cylinder
 $A = 2\pi r(h+r)$

where h and r are height and radius of cylinder respectively.

$$\therefore E = \sigma \{2\pi r(r+h)\} T^4$$

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$$= 5.67 \times 10^{-8} \left\{ 2 \times 3.14 \times 13 \times 10^{-2} (188 \times 10^{-2}) \right\} (31)^4$$

$$= 5.67 \times 6.28 \times 13 \times 188 \times (31)^4 \times 10^{-8}$$

$$= 8.037 \times 10^{10} \times 10^{-8}$$

$$= 803.7 \text{ W}$$

Hence, the total power radiated is 803.7 W.

(c) Energy radiated by the person in one-day (u)

$$u = Et$$

$$= 803.7 \times 24 \times 60 \times 60$$

$$= 6.944 \times 10^7 \text{ J}$$

If person takes 2000 kcal per day,

$$\therefore 1 \text{ kcal} = 4184 \text{ J}$$

$$\therefore \text{energy taken per day} = 2000 \times 4184 \text{ J}$$

$$= 8.368 \times 10^6 \text{ J}$$

Hence, the person radiates more energy than his/her intake in one day.

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7. If we have waves in a one-dimensional box, such that the wave displacement $\psi(x, t) = 0$ for $x=0$ and $x=L$, where L is the length of the box, and

$$\frac{1}{c} \frac{\partial^2 \psi}{\partial t^2} - \frac{\partial^2 \psi}{\partial x^2} = 0 \quad (\text{wave equation})$$

Show that the solutions are of the form

$$\psi(x, t) = a(t) \sin\left(\frac{n\pi x}{L}\right) \quad n \in \{1, 2, 3, \dots\}$$

and $a(t)$ satisfies the (harmonic oscillator) equation

$$\frac{d^2 a(t)}{dt^2} + \omega_n^2 a(t) = 0$$

where $\omega_n = \frac{n\pi c}{L}$ is angular frequency $2\pi f$.

The given wave equation

$$\frac{1}{c} \frac{\partial^2 \psi}{\partial t^2} - \frac{\partial^2 \psi}{\partial x^2} = 0 \quad \rightarrow (1)$$

Let the solution be

$$\psi(x, t) = a(t) \sin(kx)$$

Given $\psi(x, t) = 0$ for $x=0$ and $x=L$
 for $x \leq L$

$$\therefore a(t) \sin(KL) = 0 \\ \sin(KL) = 0$$

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$$KL = n\pi$$

$$\therefore K = \frac{n\pi}{L}$$

Hence, solution is $\Psi(x, t) = a(t) \sin\left(\frac{n\pi x}{L}\right)$ 11

Substitute ⑪ in ①

$$\frac{1}{c^2} \frac{\partial^2 \Psi}{\partial t^2} = \frac{1}{c^2} \left(\frac{\partial^2 a(t)}{\partial t^2} \right) \sin\left(\frac{n\pi x}{L}\right)$$

$$\frac{\partial^2 \Psi}{\partial x^2} = \frac{\partial^2}{\partial x^2} \left[a(t) \sin\left(\frac{n\pi x}{L}\right) \right]$$

$$= a(t) \left(\frac{n\pi}{L} \right)^2 \sin\left(\frac{n\pi x}{L}\right)$$

① now changes to

$$\frac{1}{c^2} \left[\frac{\partial^2 a(t)}{\partial t^2} \right] \sin\left(\frac{n\pi x}{L}\right) + a(t) \left(\frac{n\pi}{L} \right)^2 \sin\left(\frac{n\pi x}{L}\right) = 0$$

$$\Rightarrow \frac{\partial^2 a(t)}{\partial t^2} + a(t) \left[\frac{n\pi c}{L} \right]^2 = 0$$

$$\frac{\partial^2 a(t)}{\partial t^2} + w_n^2 a(t) = 0$$

Hence, proved.