# Identifying Notable Objects from Spitzer Enhanced Imaging Astronomical Observations

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#### Goals

- Identify and locate notable objects (outliers) within the sky
- Define an area of interest around that object
- Narrow down search area for manual analysis

#### Approach

Excess infrared light could mean:

- Young Stellar Objects
- Active Galactic Nucleii
- Colliding Galaxies

## Hypotheses: Young Stellar Objects

- Notable objects can be identified as extrema in terms of infrared light
- Some notable objects are grouped into interesting structures

#### Data

- NASA/IPAC Infrared Science Archive:
  - Wide-field Infrared Survey Explorer (WISE)
    - Identifies objects, and readings on the energies they emit
  - 800m objects (records), 815 GB
  - Contains:
    - Location (right ascension, declination)
    - Movement
    - Colour
    - Readings across a number of bands

#### Unexpected Colour Interesting Object

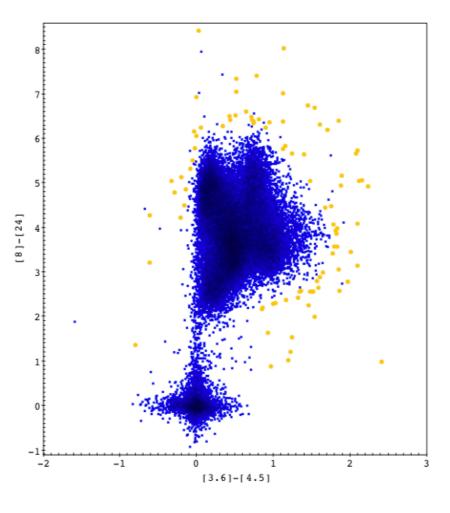


FIGURE 4: Color-color plot of SEIP sources, SNR $\geq$ 10, not in the galactic plane, and not in major survey areas (blue), vetted color outliers (yellow).

Source: Gorjian et al.

## Algorithm I

## Algorithm I, Step 1: Preprocessing with K-Means Classification

- Canned K-Means (Dataproc Pending)
- Split points into two groups, ~ N within each

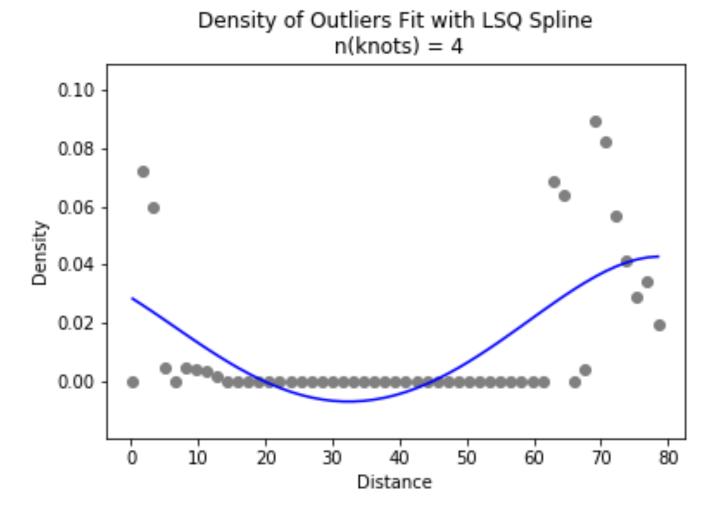
## Algorithm I, Step 2: Outliers as Nodes, Distances as Edges

- MRJob
- Within each group, take outliers:  $x_i$ :  $x_i \ge \mu \pm 2\sigma$
- Developed algorithm with parameters N, P, K.
- Complexity:
  - If number of cases < N(P): Does not run</li>
  - Otherwise, approximately:

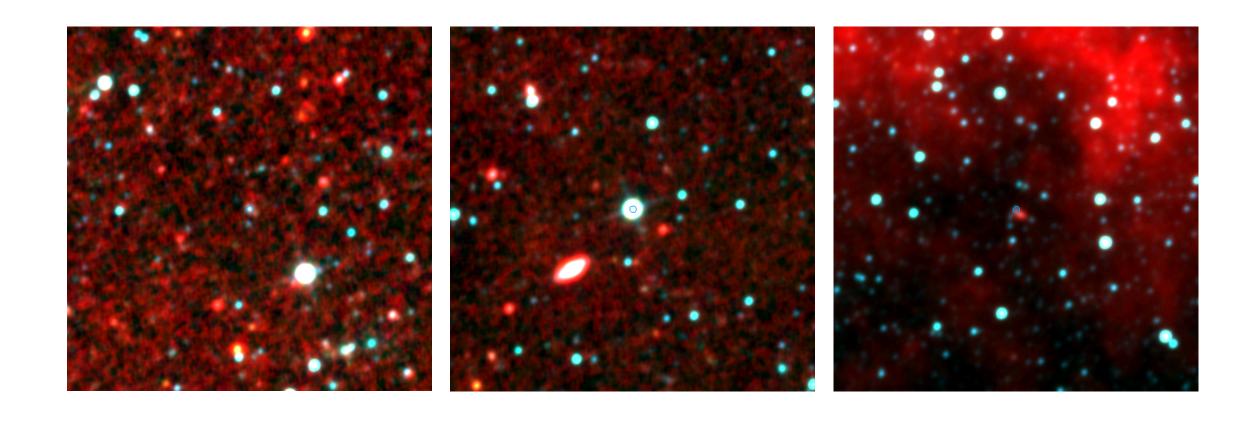
• 
$$(\sum_{i=1}^{(N-N*P)} (N*P+i)) * \frac{size - (N*P)}{(N*P)}$$

- Compare to:
  - $size^2$

## Algorithm I: Distance Density with Fitted Spline



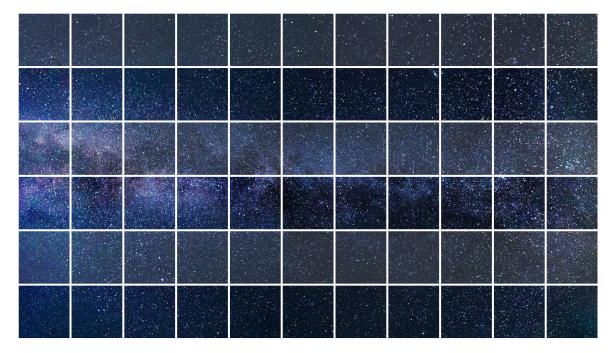
## Algorithm I: Results



## Algorithm II

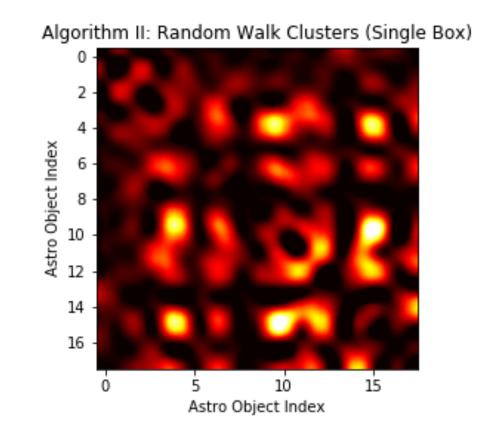
## Algorithm II (MapReduce)

- Step 1: Split Sky (Dataproc Implementation in Progress)
  - Create grids of the entire sky
  - Create graphs of each grid



### Algorithm II (MapReduce)

- Step 1: Split Sky (Dataproc Implementation in Progress)
  - Create grids of the entire sky
  - Create graphs of each grid
- Step 2: Random Walk
- (Dataproc in Progress)
- Step 3: Find Probabilistic Clusters (Coding In Progress)



#### Runtime Comparison

#### Algorithm I

- Local:
  - > 85 min. 51 sec.
  - 365,601 rows
- Dataproc:
  - 8 min. 44 sec.
    - 25 workers with n1-standard-4 specs
    - 358,169 rows

#### Algorithm II (RandomWalk)

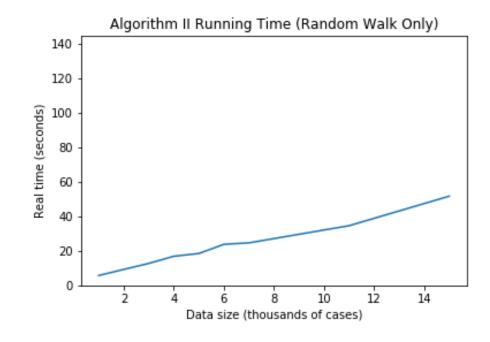
- Local:
  - 16 min. 9 sec.
  - 365,601 rows
- Dataproc:
  - In Progress

#### Runtime Comparison

#### Algorithm I



#### Algorithm II (RandomWalk)



### Challenges

#### Problems:

- Graph algorithms without all data in memory (sky is too large)
- Complexity too great for a fully connected graph

#### Solutions:

- Running random sample (Algorithm I)
- Grids and streaming processing (Algorithm II)