Some Thoughts on Filtering

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Outline

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- The Real Kalman
- 6 Some References

Apollo 8 launched on December 21, 1968

03:17:45:17 (Dec. 25, 1968, 6:36 a.m. UTC)

Jim Lovell (Commander Module Pilot): Roger. Do you wish me to reinitialize the W-matrix at this time?

03:17:45:26

Ken Mattingly (CAPCOM): Affirmative, Apollo 8

Setting

Samples

- I can sample from a normal distribution with the mean known only to me and with a publicly known variance.
- Based on your knowledge about me you have a (prior) belief about the mean of this distribution: it's centred at 0 and is probably between -3 and +3.
- At time 1, I give you some information: a sample from my distribution.
- At time 2, I give you more information: another sample from my distribution.
- And so on ...

Bayes' Theorem

$$\mathbb{P}(\theta \mid D) \triangleq \frac{\mathbb{P}(\theta \cap D)}{\mathbb{P}(D)}$$

Also

$$\mathbb{P}(D \mid \theta) \triangleq \frac{\mathbb{P}(\theta \cap D)}{\mathbb{P}(\theta)}$$

Thus

$$\mathbb{P}(\theta \mid D) \propto \mathbb{P}(D \mid \theta) \mathbb{P}(\theta)$$

I have told you that the data are drawn from

$$\mathbb{P}(x \mid \mu) \propto \exp{-\frac{(x-\mu)^2}{2\sigma^2}}$$

And, based on your knowledge about me, you have a prior

$$\mathbb{P}(\mu) \propto \exp{-rac{(\mu-\mu_0)^2}{2\sigma_0^2}}$$

This gives

$$\mathbb{P}(\mu \mid x) \propto \exp{-\frac{(x-\mu)^2}{2\sigma^2}} \times \exp{-\frac{(\mu-\mu_0)^2}{2\sigma_2^2}}$$

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$$= \exp{-\left[\frac{x^2 - 2\mu x + \mu^2}{2\sigma^2} + \frac{\mu^2 - 2\mu\mu_0 + \mu_0^2}{2\sigma_0^2}\right]}$$

$$\mathbb{P}(\mu \mid x) \propto \exp{-\frac{(x-\mu)^2}{2\sigma^2}} \times \exp{-\frac{(\mu-\mu_0)^2}{2\sigma_0^2}}$$

$$= \exp{-\left[\frac{x^2 - 2\mu x + \mu^2}{2\sigma^2} + \frac{\mu^2 - 2\mu\mu_0 + \mu_0^2}{2\sigma_0^2}\right]}$$

$$= \exp{-\left[\mu^2\left(\frac{1}{2\sigma^2} + \frac{1}{2\sigma_0^2}\right) - 2\mu\left(\frac{x}{2\sigma^2} + \frac{\mu_0}{2\sigma_0^2}\right) + \left(\frac{x^2}{2\sigma^2} + \frac{\mu_0^2}{2\sigma_0^2}\right)\right]}$$

$$\begin{split} \mathbb{P}(\mu \,|\, x) &\propto \exp{-\frac{(x-\mu)^2}{2\sigma^2}} \times \exp{-\frac{(\mu-\mu_0)^2}{2\sigma_0^2}} \\ &= \exp{-[\frac{x^2-2\mu x+\mu^2}{2\sigma^2} + \frac{\mu^2-2\mu\mu_0+\mu_0^2}{2\sigma_0^2}]} \\ &= \exp{-[\mu^2(\frac{1}{2\sigma^2} + \frac{1}{2\sigma_0^2}) - 2\mu(\frac{x}{2\sigma^2} + \frac{\mu_0}{2\sigma_0^2}) + (\frac{x^2}{2\sigma^2} + \frac{\mu_0^2}{2\sigma_0^2})]} \\ &\triangleq \exp{-[\frac{1}{2\sigma_1^2}(\mu^2 - 2\mu\mu_1 + \mu_1^2)]} \end{split}$$

$$\exp -\left[\mu^2 \left(\frac{1}{2\sigma^2} + \frac{1}{2\sigma_0^2}\right) - 2\mu \left(\frac{x}{2\sigma^2} + \frac{\mu_0}{2\sigma_0^2}\right) + \left(\frac{x^2}{2\sigma^2} + \frac{\mu_0^2}{2\sigma_0^2}\right)\right] \triangleq \exp -\left[\frac{1}{2\sigma_1^2} (\mu^2 - 2\mu\mu_1 + \mu_1^2)\right]$$

Collecting together terms

$$\exp -\left[\mu^2 \left(\frac{1}{2\sigma^2} + \frac{1}{2\sigma_0^2}\right) - 2\mu \left(\frac{x}{2\sigma^2} + \frac{\mu_0}{2\sigma_0^2}\right) + \left(\frac{x^2}{2\sigma^2} + \frac{\mu_0^2}{2\sigma_0^2}\right)\right] \triangleq \exp -\left[\frac{1}{2\sigma_1^2} (\mu^2 - 2\mu\mu_1 + \mu_1^2)\right]$$

Collecting together terms

$$\frac{1}{\sigma_1^2} = \frac{1}{\sigma_0^2} + \frac{1}{\sigma^2}$$

$$\exp -\left[\mu^2 \left(\frac{1}{2\sigma^2} + \frac{1}{2\sigma_0^2}\right) - 2\mu \left(\frac{x}{2\sigma^2} + \frac{\mu_0}{2\sigma_0^2}\right) + \left(\frac{x^2}{2\sigma^2} + \frac{\mu_0^2}{2\sigma_0^2}\right)\right] \triangleq \exp -\left[\frac{1}{2\sigma_1^2} (\mu^2 - 2\mu\mu_1 + \mu_1^2)\right]$$

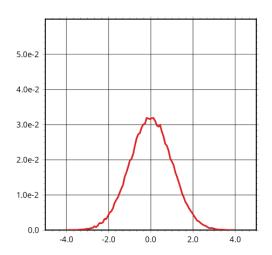
Collecting together terms

$$\begin{split} \frac{1}{\sigma_1^2} &= \frac{1}{\sigma_0^2} + \frac{1}{\sigma^2} \\ \frac{\mu_1}{\sigma_1^2} &= \frac{\mu_0}{\sigma_0^2} + \frac{x}{\sigma^2} \end{split}$$

Prior by Sampling

We can also represent the prior as a histogram

```
\mu_0, \sigma_0:: Double \mu_0 = 0.0 \sigma_0 = 1.00 priors :: Histogram BinD Double priors = hist $ runSampler (normal \mu_0 \sigma_0) 42 100000
```



Secret, Model and Data

But in reality the data come from a normal distribution with mean and (known) variance ${}^{\prime}$

```
\mu :: Double \mu = 1.00
```

 σ :: Double

 $\sigma = 0.9$

Secret, Model and Data

But in reality the data come from a normal distribution with mean and (known) variance

```
\mu :: Double \mu = 1.00 \sigma :: Double \sigma = 0.9
```

The function in Haskell for the likelihood $\mathbb{P}(x \mid \mu)$

```
likelihood :: Double \rightarrow Double \rightarrow Double likelihood \times \mu = n / d where n = \exp\left(-(x - \mu) \uparrow 2 / (2 * \sigma \uparrow 2)\right) d = \operatorname{sqrt}\left(2 * \pi * \sigma \uparrow 2\right)
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```

Finally I give you some noisy data

```
ds :: [Double]

ds = runSampler (normal \mu \sigma) 2 10
```



```
posteriorize :: Histogram BinD Double \rightarrow Double \rightarrow Histogram BinD Double posteriorize h \times = H.bmap bayes h where bayes :: Double \rightarrow Double \rightarrow Double bayes \theta p = p * likelihood <math>\times \theta
```

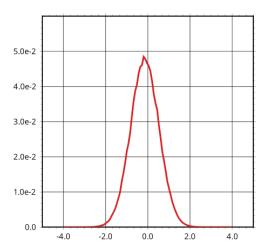
```
qs :: [Histogram BinD Double]
qs = scanl posteriorize priors ds
```

```
posteriorize :: Histogram BinD Double \rightarrow
                Double \rightarrow
                Histogram BinD Double
posteriorize h x = H.bmap bayes h
  where
     bayes :: Double \rightarrow Double \rightarrow Double
     bayes \theta p = p * likelihood \times \theta
qs :: [Histogram BinD Double]
gs = scanl posteriorize priors ds
ss :: [Double]
ss = map \ H.sum \ gs
```

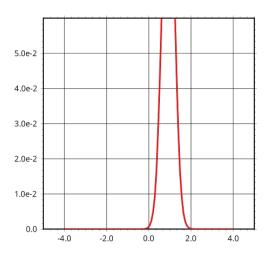
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posteriorize :: Histogram BinD Double \rightarrow
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posteriorize h x = H.bmap bayes h
  where
     bayes :: Double \rightarrow Double \rightarrow Double
     bayes \theta p = p * likelihood \times \theta
qs :: [Histogram BinD Double]
gs = scanl posteriorize priors ds
ss :: [Double]
ss = map H.sum qs
ns :: [Histogram BinD Double]
ns = zipWith (\lambda s \ q \rightarrow H.map (/s) \ q) \ ss \ qs
```



After 1 Observation



After 10 Observations



Observing Apollo (in 1D)

We want to track the lunar module

Newton's second law of motion

$$ma = F$$

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We want to track the lunar module

Newton's second law of motion

$$ma = F$$

As first order equation system

$$\frac{\mathrm{d}v}{\mathrm{d}t} = F$$

$$\frac{\mathrm{d}x}{\mathrm{d}t} = v$$

Mathematical Model

Let's assume there are no forces acting on the lunar module.

Discretizing

$$\frac{v_k - v_{k-1}}{\Delta t} = 0$$
$$\frac{x_k - x_{k-1}}{\Delta t} = v_k$$

Writing x_1 for x and x_2 for v

Mathematical Model

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$$\begin{bmatrix} x_1^{(k)} \\ x_2^{(k)} \end{bmatrix} = \begin{bmatrix} 1 & \Delta t \\ 0 & 1 \end{bmatrix} \begin{bmatrix} x_1^{(k-1)} \\ x_2^{(k-1)} \end{bmatrix}$$

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Writing x_1 for x and x_2 for v

$$\begin{bmatrix} x_1^{(k)} \\ x_2^{(k)} \end{bmatrix} = \begin{bmatrix} 1 & \Delta t \\ 0 & 1 \end{bmatrix} \begin{bmatrix} x_1^{(k-1)} \\ x_2^{(k-1)} \end{bmatrix} + \boldsymbol{Q}_k$$

Can only observe position

Observing

$$\begin{bmatrix} y_1^{(k)} \end{bmatrix} = \begin{bmatrix} 1 & 0 \end{bmatrix} \begin{bmatrix} x_1^{(k)} \\ x_2^{(k)} \end{bmatrix}$$

Can only observe position

Observing

$$\begin{bmatrix} y_1^{(k)} \end{bmatrix} = \begin{bmatrix} 1 & 0 \end{bmatrix} \begin{bmatrix} x_1^{(k)} \\ x_2^{(k)} \end{bmatrix} + \mathbf{R}_k$$

In vector notation

$$\mathbf{x}_k = \mathbf{A}\mathbf{x}_{k-1} + \mathbf{Q}_k$$

 $\mathbf{y}_k = \mathbf{H}\mathbf{x}_k + \mathbf{R}_k$

Sampling

I can sample a whole path from this model from which I will reveal a step at a time.

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Prior

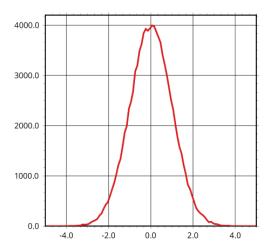
Based on your knowledge about Apollo you have a (prior) belief about the mean of this distribution: it's centred at (0,1) and both the position and velocity are probably ± 3 .

```
m{m}_0 :: M.Vector\ Double
m{m}_0 = M.fromList\ [0,1]
m{\Sigma}_0 :: M.Herm\ Double}
m{\Sigma}_0 :: M.Sym\ (2><2)\ [1.0,0.0,\ 0.0,1.0]
priorsApollo :: Histogram\ (Bin2D\ BinD\ BinD)\ Double
priorsApollo = hist2\ conv\ prePriors\ 2\ 100000
\mbox{where}\ prePriors\ seed\ n = \ evalState\ (replicateM\ n\ (sample\ G.Normal\ m{m}_0\ m{\Sigma}_0))
(pureMT\ (fromIntegral\ seed\ ))
```

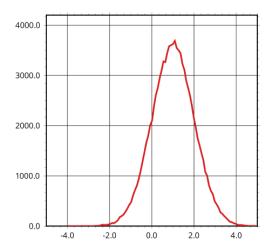
An Aside: Marginals

marginalX, marginalY :: Histogram BinD Double marginalX = H.reduceX H.sum priorsApollo marginalY = H.reduceY H.sum priorsApollo

Marginal Prior for Position



Marginal Prior for Velocity



Recall the state update

$$\boldsymbol{x}_k = \boldsymbol{A}\boldsymbol{x}_{k-1} + \boldsymbol{Q}_k$$

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$$\boldsymbol{x}_k = \boldsymbol{A}\boldsymbol{x}_{k-1} + \boldsymbol{Q}_k$$

In Haskell

```
newState :: MonadRandom m \Rightarrow

M.Vector\ Double \rightarrow m\ (M.Vector\ Double)

newState \mathbf{x}^{\flat} = sample \ rvar\ (G.Normal\ (\mathbf{A} \ \# > \mathbf{x}^{\flat})\ \mathbf{Q})
```

Recall the state update

$$\boldsymbol{x}_k = \boldsymbol{A}\boldsymbol{x}_{k-1} + \boldsymbol{Q}_k$$

In Haskell

newState :: MonadRandom $m \Rightarrow$ $M.Vector\ Double \rightarrow m\ (M.Vector\ Double)$ newState $\mathbf{x}^{\flat} = sample \ rvar\ (G.Normal\ (\mathbf{A} \ \# > \mathbf{x}^{\flat})\ \mathbf{Q})$

Recall the observations are given by

$$\mathbf{y}_k = \mathbf{H}\mathbf{x}_k + \mathbf{R}_k$$

Recall the state update

$$\boldsymbol{x}_k = \boldsymbol{A}\boldsymbol{x}_{k-1} + \boldsymbol{Q}_k$$

In Haskell

newState :: MonadRandom $m \Rightarrow$ $M.Vector\ Double \rightarrow m\ (M.Vector\ Double)$ $newState\ x^{\flat} = sample\ \ rvar\ (G.Normal\ (A\ \#> x^{\flat})\ Q)$

Recall the observations are given by

$$\mathbf{y}_k = \mathbf{H}\mathbf{x}_k + \mathbf{R}_k$$

Thus the likelihood of observation b given the state a is given by

likelihood $A :: M.Vector\ Double \rightarrow M.Vector\ Double \rightarrow Double$ likelihood $A \ a \ b = pdf\ (G.Normal\ (\textbf{\textit{H}} \# > a)\ \textbf{\textit{R}})\ b$

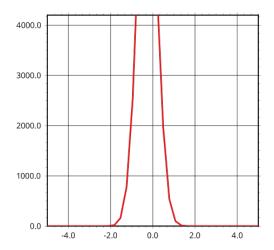


Apply Bayes'

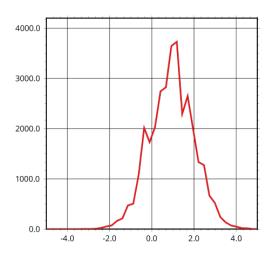
Now you can use Bayes' to track the path of the apollo.

```
posteriorizeA :: MonadRandom m \Rightarrow
  Histogram (Bin2D BinD) Double \rightarrow
  M.Vector\ Double \rightarrow
  m (Histogram (Bin2D BinD BinD) Double)
posteriorizeA q d = do
  let xfs = H.asList q
    xs = map pair2Vec $ map fst xfs
     fs = map \ snd \ xfs
  x \leftarrow mapM newState xs
  let newQ = hist2' \$ zip (map vec2Pair x) fs
  return $ H.bmap (\lambda(u, v) p \rightarrow p * likelihoodA (M.vector [u, v]) d) new Q
```

Position Estimate After 1 Observation



Velocity Estimate After 1 Observation



Marginals Again

```
*Test> take 1 apolloSamples
[([9.99586e-2,0.75767],[-0.36189])]
   marginal1X, marginal1Y :: Histogram BinD Double
   marginal1X = H.reduceX H.sum test
   marginal1Y = H.reduceY H.sum test
   m0X1, m1X1, muX1 :: Double
   m0X1 = H.sum\ marginal1X
   m1X1 = H.sum \ H.bmap \ (\lambda f \ v \rightarrow v * f) \ marginal 1X
   muY1 = m1Y1 / m0Y1
   muX1 = m1X1 / m0X1
*Test> muX1
```

```
*Test> muX
0.94108
```

*Test> muY1

-0.26897



Recall the Model

Wait a Minute ...

Compute power in 1968?

Recall the Model

Wait a Minute . . .

Compute power in 1968?

Recall our model:

$$\mathbf{x}_i = \mathbf{A}_{i-1}\mathbf{x}_{i-1} + \mathbf{Q}_{i-1}$$

 $\mathbf{y}_i = \mathbf{H}_i\mathbf{x}_i + \mathbf{R}_i$

where

$$Q_i \sim \mathcal{N}(0, \mathbf{\Sigma}_i^{(x)}), \quad R_i \sim \mathcal{N}(0, \mathbf{\Sigma}_i^{(y)})$$

The model is linear and the errors are Gaussian

Pólya

One way to solve a problem is to generalize it

$$x_k \sim p(x_k \mid x_{k-1})$$

 $y_k \sim p(y_k \mid x_k)$

- State update is Markovian
- The past is independent of the future given the present
- The measurement given the current state is independent of the past (quasi-Markovian)

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One way to solve a problem is to generalize it

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- State update is Markovian
- The past is independent of the future given the present
- The measurement given the current state is independent of the past (quasi-Markovian)

Re-writing our linear Gaussian model

$$p(x_k | x_{k-1}) = \mathcal{N}(x_k | A_{k-1} x_{k-1}, Q_k)$$

$$p(y_k | x_k) = \mathcal{N}(y_k | H_k x_k, R_k)$$

Bayesian Filtering Equations

What we'd like is

$$p(x_k \mid y_{1:k})$$

$$p(x_k, x_{k-1} | y_{1:k-1}) = p(x_k | x_{k-1}, y_{1:k-1}) p(x_{k-1} | y_{1:k-1})$$

= $p(x_k | x_{k-1}) p(x_{k-1} | y_{1:k-1})$

Marginalising

$$p(x_k \mid y_{1:k-1}) = \int p(x_k \mid x_{k-1}) p(x_{k-1} \mid y_{1:k-1}) dx_{k-1}$$

A Histogram Interpretation

We can think of our histograms as roughly being a distribution

$$G=\sum w_i\delta(x^{(i)})$$

where the x_i are the midpoints of the cells and the w_i are the (normalised) number of observations in the cell and δ is the Dirac delta "function".

$$p(x_k | y_{1:k-1}) = \int p(x_k | x_{k-1}) p(x_{k-1} | y_{1:k-1}) dx_{k-1}$$

$$= \sum p(x_k | x_{k-1}) w_i \delta(x_{k-1}^{(i)})$$

$$= \sum w_i \delta(x_k^{(i)})$$

where
$$x_{k}^{(i)} \sim p(x_{k} \,|\, x_{k}^{(i-1)})$$

Bayesian Filtering Equations

$$p(x_k \mid y_{1:k}) = \frac{1}{Z_k} p(y_k \mid x_k, y_{1:k-1}) p(x_k \mid y_{1:k-1})$$
$$= \frac{1}{Z_k} p(y_k \mid x_k) p(x_k \mid y_{1:k-1})$$

In histogram terms

$$p(x_k^{(i)} | y_{1:k}) = \frac{1}{Z_k} p(y_k | x_k^{(i-1)}) w_i \delta(x_{k-1}^{(i)})$$
$$= \frac{1}{Z_k} w_i' \delta(x_{k-1}^{(i)})$$

where $w_i' = \frac{1}{Z_k} p(y_k \mid x_k^{(i-1)}) w_i$ and $\frac{1}{Z_k} = \sum w_i'$.



Kalman Itself

A **lot** of algebraic manipulation gives the optimal solution.

Prediction Step

$$\hat{\pmb{x}}_i^{\flat} = \pmb{A}_{i-1}\hat{\pmb{x}}_{i-1}$$

$$\hat{\boldsymbol{\Sigma}}_{i}^{\flat} = \boldsymbol{A}_{i-1}\hat{\boldsymbol{\Sigma}}_{i-1}\boldsymbol{A}_{i-1}^{\top} + \boldsymbol{Q}_{i-1}$$

Correction Step

$$egin{aligned} oldsymbol{v}_i &= oldsymbol{y}_i - oldsymbol{H}_i \hat{oldsymbol{\chi}}_i^{eta} \ oldsymbol{S}_i &= oldsymbol{H}_i \hat{oldsymbol{\Sigma}}_i^{eta} oldsymbol{H}_i^{oldsymbol{ op}} oldsymbol{S}_i^{-1} \ \hat{oldsymbol{\chi}}_i &= \hat{oldsymbol{\chi}}_i^{eta} + oldsymbol{K}_i oldsymbol{v}_i \ \hat{oldsymbol{\Sigma}}_i &= \hat{oldsymbol{\Sigma}}_i^{oldsymbol{ op}} - oldsymbol{K}_i oldsymbol{S}_i oldsymbol{K}_i^{oldsymbol{ op}} \end{aligned}$$

Re-express Original

$$x_k = x_{k-1} + q_{k-1}, \quad q_{k-1} \sim \mathcal{N}(0, 0)$$

 $y_k = x_k + r_k, \quad r_k \sim \mathcal{N}(0, R)$

Prediction Step

$$\hat{x}_i^{\flat} = \hat{x}_{i-1} \quad \hat{\Sigma}_i^{\flat} = \hat{\Sigma}_{i-1}$$

Correction Step

$$v_i = y_i - \hat{x}_i^{\flat} \quad S_i = \hat{\Sigma}_i^{\flat} + R_i \quad K_i = \frac{\hat{\Sigma}_i^{\flat}}{S_i}$$
$$\hat{x}_i = \hat{x}_i^{\flat} + K_i v_i \quad \hat{\Sigma}_i = \hat{\Sigma}_i^{\flat} - K_i S_i K$$

Prediction Step

Introduction

$$\hat{x}_{i}^{\flat} = \hat{x}_{i-1} \quad \hat{\Sigma}_{i}^{\flat} = \hat{\Sigma}_{i-1} \\ \hat{x}_{1}^{\flat} = \hat{x}_{0} = \mu_{0} \quad \hat{\Sigma}_{1}^{\flat} = \hat{\Sigma}_{0} = \sigma_{0}^{2}$$

Correction Step

$$v_i = y_i - \hat{x}_i^{\flat} \quad S_i = \hat{\Sigma}_i^{\flat} + R_i \quad K_i = \frac{\hat{\Sigma}_i^{\flat}}{S_i}$$
$$\hat{x}^i = \hat{x}_i^{\flat} + K_i v_i \quad \sigma_1^2 = \hat{\Sigma}_i^{\flat} - K_i S_i K$$

Correction Step

$$v = y - \mu_0$$
 $S = \sigma_0^2 + \sigma^2$ $K = \frac{\sigma_0^2}{\sigma_0^2 + \sigma^2}$ $\mu_1 = \mu_0 + Kv$ $\sigma_1^2 = \sigma_0^2 - KSK$

Some References

Correction Step

$$v = y - \mu_0$$
 $S = \sigma_0^2 + \sigma^2$ $K = \frac{\sigma_0^2}{\sigma_0^2 + \sigma^2}$
 $\mu_1 = \mu_0 + Kv$ $\sigma_1^2 = \sigma_0^2 - KSK$

Correction Step

$$\mu_1 = \mu_0 + \frac{\sigma_0^2}{\sigma_0^2 + \sigma^2} (y - \mu_0) \quad \sigma_1^2 = \sigma_0^2 - \frac{(\sigma_0^2)^2}{\sigma_0^2 + \sigma^2}$$

Correction Step

$$v = y - \mu_0$$
 $S = \sigma_0^2 + \sigma^2$ $K = \frac{\sigma_0^2}{\sigma_0^2 + \sigma^2}$
 $\mu_1 = \mu_0 + Kv$ $\sigma_1^2 = \sigma_0^2 - KSK$

Correction Step

$$\mu_1 = \mu_0 + \frac{\sigma_0^2}{\sigma_0^2 + \sigma^2} (y - \mu_0)$$
 $\sigma_1^2 = \sigma_0^2 - \frac{(\sigma_0^2)^2}{\sigma_0^2 + \sigma^2}$

Correction Step

$$\frac{\mu_1}{\sigma_1^2} = \frac{\mu_0}{\sigma_0^2} + \frac{\mu}{\sigma^2}$$
 $\frac{1}{\sigma_1^2} = \frac{1}{\sigma^2} + \frac{1}{\sigma_0^2}$

Further Reading

- Simo Särkkä's book "Bayesian Filtering and Smoothing"
- Libbi: http://libbi.org
- Haskell package random-fu for random variables
- Haskell package histogram-fill for histograms
- Haskell package kalman for kalman, extended kalman, unscented kalman, particle filtering and smoothing
- Same in Python: https://filterpy.readthedocs.io/en/latest
- Haskell package hmatrix for statically typed matrices