

#### IT00CG19-3003

# **GPU Programming**

Project Assignment: Cosmic Dark Matter / 2-point Angular Correlation

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## Program Design and Implementation

- To pass this course you are asked to
  - design
  - implement
  - run and
  - report

the calculation of three histograms of the 2-point angular correlation function for two sets of galaxies

- D: measured set of 100 000 galaxies D
- R: synthetic random evenly distributed set of 100 000 galaxies



## Program Design and Implementation

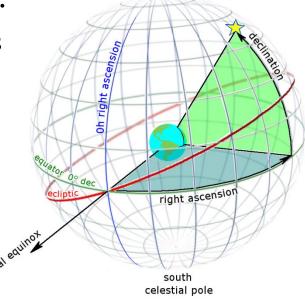
- Input data
- Figure out what needs to be calculated
- Design a sequential program on CPU
- Design your GPU threads and thread blocks
- Implementation on GPU (start with CPU?)
- Built in check points
- Output data





## Input Data

- Two lists of astronomical locations for N galaxies: real measured galaxies and synthetic evenly distributed random galaxies
- For each galaxy, real or synthetic, the list contains the galactic coordinates in this order:
  - right ascension  $\alpha$ , in arc minutes
  - declination  $\delta$ , in arc minutes
- Convert from arc minutes to radians by multiplying with  $1/60*\pi/180$
- Lists available on course page





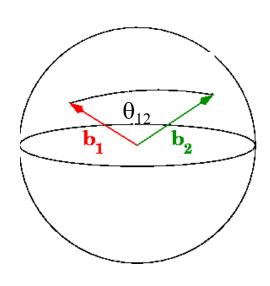
## Input Data in moodle

#### Real data

#### 100000 4646.98 3749.51 4644.35 3749.52 4643.24 3748.67 4646.98 3750.89 4641.13 3748.23 4646.75 3750.26 4643.29 3751.99 4640.78 3747 4638.95 3749.74 4647.29 3749.82 4651.65 3749.02 4649.65 3747.16 4646.66 3752.23 4649.38 3749.14 4648.22 3750.32 4639.27 3747.96 4637.3 3749.5 4649.07 3751.94 4640.39 3752.43

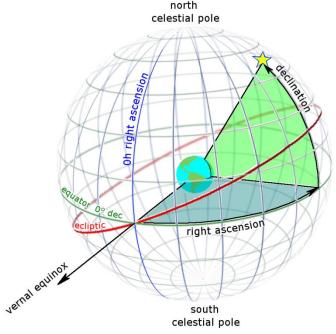
#### Synthetic data

```
100000
840.961426 387.991697
387.368692 2967.285746
2667.070581 3385.866638
3942.141923 3720.545649
5164.800068 3205.065003
1680.286209 653.214854
3553.889989 152.983872
1011.581615 76.716702
298.341944 3752.723427
4044.605401 2943.458158
1030.720073 2342.109887
54.257097 3769.420145
2478.545632 336.072740
970.044311 3556.987037
840.791266 2610.257310
5272.628047 2441.450374
380.633015 3771.599991
5049.735705 4627.545008
2849.250851 2945.861883
```





- Initialize the three histograms DD, DR and RR to zero, covering  $0 \rightarrow 180$  degrees, bin width = 0.25 degrees.
- For DD, find the angle between each pair of real-real galaxy (into which histogram bin?)
- For DR, find the angle between each pair of real-random galaxy
- For RR, find the angle between each pair of random-random galaxy
- Check point: how many entries in DR? In DD and RR?





- Given two points (galaxies) on the surface of a sphere, how do we calculate the angle between those two points as seen from the center of the sphere?
- Basic idea: represent each galaxy by a unit vector from the center to the surface of the sphere and calculate the dot product between these two vectors.
- The dot product between two 3-D vectors  $\mathbf{r}_1$  and  $\mathbf{r}_2$  is given by  $\mathbf{r}_1 \bullet \mathbf{r}_2 = |\mathbf{r}_1| |\mathbf{r}_2| \cos(\theta_{12})$  where  $|\mathbf{r}|$  is the length of the vector  $\mathbf{r}$ , and  $\theta_{12}$  is the angle between the two vectors.
- Here we have unit vectors, hence  $|\mathbf{r}_1| = |\mathbf{r}_2| = 1$



The formula we can use is now

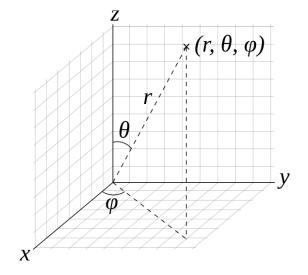
$$\theta_{12} = \arccos(\mathbf{r}_1 \bullet \mathbf{r}_2)$$

$$\mathbf{r} = \mathbf{r}_x \mathbf{e}_x + \mathbf{r}_y \mathbf{e}_y + \mathbf{r}_z \mathbf{e}_z$$

$$\mathbf{r}_1 \bullet \mathbf{r}_2 = \mathbf{r}_{1x} * \mathbf{r}_{2x} + \mathbf{r}_{1y} * \mathbf{r}_{2y} + \mathbf{r}_{1z} * \mathbf{r}_{2z}$$

• Using the spherical coordinates  $(r,\theta,\phi)$ , the Cartesian components of  ${\bf r}$  are given by

$$r_x = r*sin(\theta)*cos(\phi)$$
  
 $r_y = r*sin(\theta)*sin(\phi)$   
 $r_z = r*cos(\theta)$ 



Cartesian: x, y, z Spherical: r,  $\theta$ ,  $\phi$ 



- Finally, what is the connection between the spherical coordinates  $\theta$ ,  $\phi$  and the galactic coordinates right ascension  $\alpha$  and declination  $\delta$ ?
- Answer:  $\varphi = \alpha$  and  $\theta = 90 \delta$
- $\mathbf{r}_1 \bullet \mathbf{r}_2 = \sin(\theta_1) \cos(\phi_1) \sin(\theta_2) \cos(\phi_2) + \sin(\theta_1) \sin(\phi_1) \sin(\phi_2) \sin(\phi_2) + \cos(\theta_1) \cos(\theta_2)$ 
  - $= \cos(\delta_1) \cos(\alpha_1) \cos(\delta_2) \cos(\alpha_2) + \cos(\delta_1) \sin(\alpha_1) \cos(\delta_2) \sin(\alpha_2) + \sin(\delta_1) \sin(\delta_2)$
  - $= \cos(\delta_1) \cos(\delta_2) \left[\cos(\alpha_1) \cos(\alpha_2) + \sin(\alpha_1) \sin(\alpha_2)\right]$  $+ \sin(\delta_1) \sin(\delta_2)$
  - =  $\cos(\delta_1) \cos(\delta_2) \cos(\alpha_1 \alpha_2) + \sin(\delta_1) \sin(\delta_2)$



• Final result: the angle  $\theta_{12}$  between two galaxies  $(\alpha_1,\delta_1), (\alpha_2,\delta_2)$  is given by

$$\theta_{12} = \arccos(\sin(\delta_1) * \sin(\delta_2) + \cos(\delta_1) * \cos(\delta_2) * \cos(\alpha_1 - \alpha_2))$$

- Data type for  $\alpha_i, \delta_i$ : single or double precision floating point numbers? Will single precision suffice?
- Data layout in memory?





#### Threads and Thread Blocks

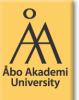
- Choose your threads and thread blocks
- If you have the time, play around with the size of your thread block and the amount of work one thread does.
- Do you need to synchronize your threads?
- Do you need to use atomic operations?





## Implementation on CPU and GPU

- Use the very basic program template from the lecture slides
  - read the data from the files, manipulate as needed
  - transfer data to the GPU (or use unified memory)
  - start the kernel(s)
  - transfer data back to the CPU
  - present your results: the first few omega values
- Start out with a small set of galaxies, perhaps with known relative angles?



## Output Data

- If you have the time, plot the histograms DD and RR to see if there are any visible differences.
- The scientific measure for the difference between the two distributions of galaxies R and D is

$$\omega_i(\theta) = (DD_i - 2*DR_i + RR_i)/RR_i$$
  
 $DD_i$ ,  $DR_i$ ,  $RR_i = value$  in histogram bin i

- If the  $\omega_i$  values are closer to zero than one, in the range [-0.5,0.5], then D has approximately the same distribution as R, and we have a random distribution of real galaxies
- If the  $\omega_i$  values are different from zero on the scale of one, then we have a non-random distribution of real galaxies



### Short Instructions for dione.abo.fi

Apply for an account and log on to the front node of dione:

```
ssh user name@dione.abo.fi
```

Load necessary modules:

```
module load GCC
```

Compile your program prog.cu nvcc -arch=sm\_70 prog.cu -o a.out -lm

Run your program a.out on the batch queue system

```
srun -p gpu -- mem=1G -t=1:00:00 -o prog.out -e prog.err
./a.out real_data sim_data
```